MAGIC BASELINE BETA BEAM

Sandhya Choubey

Harish-Chandra Research Institute, Allahabad, India



ISS/IDS meeting, CERN, 29/30 March 2007

Sanjib Kumar Agarwalla, Sandhya Choubey, and Amitava Raychaudhuri, hep-ph/0610333 Sanjib Kumar Agarwalla, Amitava Raychaudhuri, and Abhijit Samanta, hep-ph/0505015

The "Golden Channel"
$$(\nu_e \rightarrow \nu_\mu)$$

 $P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$
 $\pm \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$
 $+ \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$
 $+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}$

The "Golden Channel"
$$(\nu_e \rightarrow \nu_\mu)$$

 $P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$
 $\pm \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$
 $+ \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$
 $+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}$
 $\bullet \alpha = \Delta m_{21}^2 / \Delta m_{31}^2$, $\hat{A} = A / \Delta m_{31}^2$
 $\bullet \Delta = \Delta m_{31}^2 L / 4E$, $A = \pm 2\sqrt{2}G_F n_e E$
 $\bullet A$ is positive for neutrinos
 $\bullet A$ is negative for antineutrinos Cervera *et al.*, hep-ph/0002108
Freund, Huber, Lindner, hep-ph/0105071

∇ SANDHYA CHOUBEY

MAGIC BASELINE BETA BEAM

ISS, CERN, 30.03.07 - p.1/14

The "Golden Channel"
$$(\nu_e \rightarrow \nu_\mu)$$

 $P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$
 $\pm \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$
 $+ \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$
 $+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}$

• Depends on θ_{13} , δ_{CP} and $sgn(\Delta m_{31}^2)$ and hence can measure them all simultaneously: **GOLDEN**

▽SANDHYA CHOUBEY

The "Golden Channel"
$$(\nu_e \rightarrow \nu_\mu)$$

 $P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$
 $\pm \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$
 $+ \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$
 $+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}$

• Depends on θ_{13} , δ_{CP} and $sgn(\Delta m_{31}^2)$ and hence can measure them all simultaneously: **GOLDEN**

This dependence however brings in the problem of PARAMETER DEGENERACIES

SANDHYA CHOUBEY

Degeneracies in the Golden Channel

$$P_{e\mu} \simeq \sin^{2} \theta_{23} \sin^{2} 2\theta_{13} \frac{\sin^{2}[(1-\hat{A})\Delta]}{(1-\hat{A})^{2}}$$

$$\pm \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$$

$$+ \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$$

$$+ \alpha^{2} \cos^{2} \theta_{23} \sin^{2} 2\theta_{12} \frac{\sin^{2}(\hat{A}\Delta)}{\hat{A}^{2}}$$
Octant of θ_{23} **Degeneracy**
Sgn(Δm_{31}^{2}) - δ_{CP} **Degeneracy**
Sgn(Δm_{31}^{2}) - δ_{CP} **Degeneracy**
Intrinsic (δ_{CP} , θ_{13}) **Degeneracy**
Burguet-Castell, Gavela, Gomez-Cadenas, Hernandez, Mena, hep-ph/0103258

Degeneracies in the Golden Channel

$$P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$$

$$\pm \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$$

$$+ \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$$

$$+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}$$
• Octant of θ_{23} Degeneracy
• $Sgn(\Delta m_{31}^2) - \delta_{CP}$ Degeneracy
• Intrinsic $(\delta_{CP}, \theta_{13})$ Degeneracy
Burguet-Castell, Gavela, Gomez-Cadenas, Hernandez, Mena, hep-ph/0103258
• Degeneracies create Clone Solutions

Killing the Clones at The Magic Baseline

$$P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$$

$$\pm \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$$

$$+ \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$$

$$+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}$$

∇ SANDHYA CHOUBEY

$$P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$$

$$\pm \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$$

$$+ \alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$$

$$+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}$$

 $\mathbf{Qlf}\sin(\hat{A}\Delta)\simeq 0$

SANDHYA CHOUBEY

$$P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$$

$$P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$$

• All δ_{CP} dependent terms drop out • (δ_{CP}, θ_{13}) and ($\delta_{CP}, sgn(\Delta m_{31}^2)$) degeneracies vanish • "Clean" measurement of θ_{13} and $sgn(\Delta m_{31}^2)$

$$P_{e\mu} \simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2[(1-\hat{A})\Delta]}{(1-\hat{A})^2}$$

• All δ_{CP} dependent terms drop out • (δ_{CP}, θ_{13}) and ($\delta_{CP}, sgn(\Delta m_{31}^2)$) degeneracies vanish • "Clean" measurement of θ_{13} and $sgn(\Delta m_{31}^2)$

$$\sin(\hat{A}\Delta) \simeq 0$$
$$\Rightarrow$$
$$L_{magic} \simeq 7690 \text{ km}$$

Barger, Marfatia, Whisnant, hep-ph/0112119

Huber, Winter, hep-ph/0301257

Smirnov, hep-ph/0610198

• Large Distance \Rightarrow Large Matter effects

- Large Distance \Rightarrow Large Matter effects
- Resonance energy

$$E_{res} = \frac{|\Delta m_{31}^2|\cos 2\theta_{13}}{2\sqrt{2}G_F N_e}$$

- Large Distance \Rightarrow Large Matter effects
- Resonance energy

$$E_{res} = \frac{|\Delta m_{31}^2|\cos 2\theta_{13}}{2\sqrt{2}G_F N_e}$$

• For $\Delta m_{31}^2 = 2.5 \times 10^{-3} \text{ eV}^2$, $\sin^2 2\theta_{13} = 0.1$ and the PREM profile $\rho_{av} = 4.13$ gm/cc, $E_{res} = 6.1$ GeV

- Large Distance \Rightarrow Large Matter effects
- Resonance energy

$$E_{res} = \frac{|\Delta m_{31}^2|\cos 2\theta_{13}}{2\sqrt{2}G_F N_e}$$

• For $\Delta m_{31}^2 = 2.5 \times 10^{-3} \text{ eV}^2$, $\sin^2 2\theta_{13} = 0.1$ and the PREM profile $\rho_{av} = 4.13$ gm/cc, $E_{res} = 6.1$ GeV

• At the magic baseline, largest matter effects will come when $E \simeq 6 \text{ GeV}$

The Probability



▽SANDHYA CHOUBEY

MAGIC BASELINE BETA BEAM

ISS, CERN, 30.03.07 - p.8/14



Agarwalla, S.C., Raychaudhuri, hep-ph/0610333

MAGIC BASELINE BETA BEAM

■ INO will come up at PUSHEP in Southern India

- INO will come up at PUSHEP in Southern India
- ICAL at INO will be a 50 kton magnetized iron calorimeter

- INO will come up at PUSHEP in Southern India
- ICAL at INO will be a 50 kton magnetized iron calorimeter
- Lt will have good energy and angle reconstruction

- INO will come up at PUSHEP in Southern India
- ICAL at INO will be a 50 kton magnetized iron calorimeter
- Lt will have good energy and angle reconstruction
- Lt will have charge identification capability

- INO will come up at PUSHEP in Southern India
- ICAL at INO will be a 50 kton magnetized iron calorimeter
- Lt will have good energy and angle reconstruction
- Lt will have charge identification capability
- It can be a far detector for a neutrino factory owing to its capacity for charge identification of muons

- INO will come up at PUSHEP in Southern India
- ICAL at INO will be a 50 kton magnetized iron calorimeter
- Lt will have good energy and angle reconstruction
- Lt will have charge identification capability
- It can be a far detector for a neutrino factory owing to its capacity for charge identification of muons
- It can also be a far detector for a Beta-Beam owing to its capacity to detect muons

- INO will come up at PUSHEP in Southern India
- ICAL at INO will be a 50 kton magnetized iron calorimeter
- Lt will have good energy and angle reconstruction
- Lt will have charge identification capability
- It can be a far detector for a neutrino factory owing to its capacity for charge identification of muons
- It can also be a far detector for a Beta-Beam owing to its capacity to detect muons
- However, this Beta-Beam far-detector is different its energy threshold is going to be about 1 GeV



- INO will come up at PUSHEP in Southern India
- ICAL at INO will be a 50 kton magnetized iron calorimeter
- Lt will have good energy and angle reconstruction
- Lt will have charge identification capability
- It can be a far detector for a neutrino factory owing to its capacity for charge identification of muons
- It can also be a far detector for a Beta-Beam owing to its capacity to detect muons
- However, this Beta-Beam far-detector is different its energy threshold is going to be about 1 GeV
- So the Beta-Beam itself must be different it must be a multi-GeV beam and thats good for matter effects!!

SANDHYA CHOUBEY

A Beta-Beam facility might come up at CERN

- A Beta-Beam facility might come up at CERN
- CERN-INO distance is equal to 7152 km



- A Beta-Beam facility might come up at CERN
- CERN-INO distance is equal to 7152 km
- This is tantalizingly close to the magic baseline

- A Beta-Beam facility might come up at CERN
- CERN-INO distance is equal to 7152 km
- This is tantalizingly close to the magic baseline
- Energy threshold of ICAL would be about 1 GeV

- A Beta-Beam facility might come up at CERN
- CERN-INO distance is equal to 7152 km
- This is tantalizingly close to the magic baseline
- Energy threshold of ICAL would be about 1 GeV
- The standard Beta-Beam ions ¹⁸Ne and ⁶He would require very large gamma

- A Beta-Beam facility might come up at CERN
- CERN-INO distance is equal to 7152 km
- This is tantalizingly close to the magic baseline
- Energy threshold of ICAL would be about 1 GeV
- The standard Beta-Beam ions ¹⁸Ne and ⁶He would require very large gamma
- Alternative ions ${}^{8}B$ and ${}^{8}Li$ have large end-point energy and hence "harder" spectra. Works!!



- A Beta-Beam facility might come up at CERN
- CERN-INO distance is equal to 7152 km
- This is tantalizingly close to the magic baseline
- Energy threshold of ICAL would be about 1 GeV
- The standard Beta-Beam ions ¹⁸Ne and ⁶He would require very large gamma
- Alternative ions ${}^{8}B$ and ${}^{8}Li$ have large end-point energy and hence "harder" spectra. Works!!
- For $\gamma = 350 500$ the flux peaks at $E \simeq 6$ GeV allows exploitation of near-resonant matter effects

SANDHYA CHOUBEY



Agarwalla, SC, Raychaudhuri, hep-ph/0610333

• Flux peaks at $E \simeq 6$ GeV for $\gamma = 350 - 500$ making this a near-magic baseline as well as near-resonant Beta-Beam experiment

▽SANDHYA CHOUBEY



Agarwalla, SC, Raychaudhuri, hep-ph/0610333

The rate shows a sharp dependence on the hierarchy and θ_{13}

▽SANDHYA CHOUBEY

ISS, CERN, 30.03.07 - p.12/14

• Sensitivity to θ_{13}



• At 3σ , $\sin^2 2\theta_{13} < 8.5 \times 10^{-4} (1.5 \times 10^{-3})$ with 80% detection efficiency and 10(5) years data

▽SANDHYA CHOUBEY

• Sensitivity to $sgn(\Delta m_{31}^2)$



• At 3σ , $\sin^2 2\theta_{13} < 8.5 \times 10^{-3}(9.8 \times 10^{-3})$ with 80% detection efficiency and 10(5) years data

SANDHYA CHOUBEY

ICAL at INO will be a 50 kton magnetized iron calorimeter.

- ICAL at INO will be a 50 kton magnetized iron calorimeter.
- CERN-INO baseline will be 7152 km tantalizingly close to the magic baseline and hence will be free of the clone solutions.

- ICAL at INO will be a 50 kton magnetized iron calorimeter.
- CERN-INO baseline will be 7152 km tantalizingly close to the magic baseline and hence will be free of the clone solutions.
- At this baseline we can get near-resonant matter effect for $E \approx 6 \text{ GeV} \Rightarrow {}^{8}B$ and ${}^{8}Li$ can do the job.

- ICAL at INO will be a 50 kton magnetized iron calorimeter.
- CERN-INO baseline will be 7152 km tantalizingly close to the magic baseline and hence will be free of the clone solutions.
- At this baseline we can get near-resonant matter effect for $E \approx 6 \text{ GeV} \Rightarrow {}^{8}B$ and ${}^{8}Li$ can do the job.
- Near-resonant matter effect gives largest possible $P_{e\mu} \Rightarrow$ enough statistics even though flux is smaller since its a Beta-Beam at 7152 km.



- ICAL at INO will be a 50 kton magnetized iron calorimeter.
- CERN-INO baseline will be 7152 km tantalizingly close to the magic baseline and hence will be free of the clone solutions.
- At this baseline we can get near-resonant matter effect for $E \approx 6 \text{ GeV} \Rightarrow {}^{8}B$ and ${}^{8}Li$ can do the job.
- Near-resonant matter effect gives largest possible $P_{e\mu} \Rightarrow$ enough statistics even though flux is smaller since its a Beta-Beam at 7152 km.
- The CERN-INO Beta-Beam experiment (Beta-INO) is expected to give sensitivity to θ_{13} and $sgn(\Delta m_{31}^2)$ better than all other rival proposals, apart from a high performance neutrino factory.

SANDHYA CHOUBEY