

Morgan Wascko Imperial College London April 12, 2007



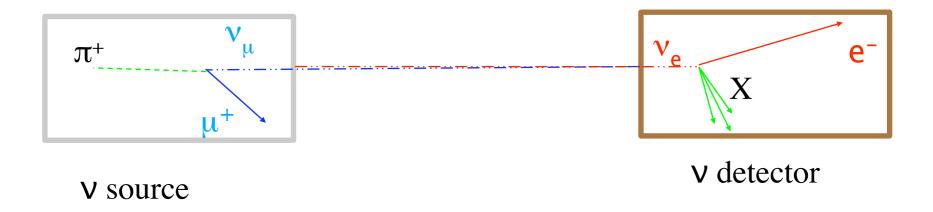
- 1. Motivation & Introduction
- 2. Description of the Experiment
- 3. Analysis Overview
- 4. Two Independent Oscillation Searches
- 5. First Results

### **Motivation: Neutrino Oscillations**

if neutrinos have mass, a neutrino that is produced as a  $\nu_{\mu}$  (e.g.  $\pi^+ \rightarrow \mu^+ \nu_{\mu}$ ) has a non-zero probability to oscillate and some time later be detected as a  $\nu_e$  (e.g.  $\nu_e$  n  $\rightarrow$  e<sup>-</sup>p)



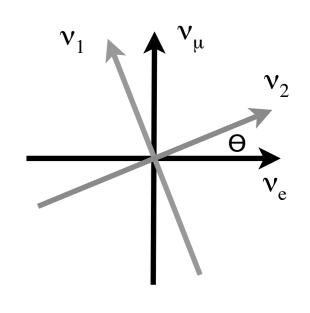
Pontecorvo, 1957



### **Motivation: Neutrino Oscillations**

In a world with 2 neutrinos, if the weak eigenstates  $(v_e, v_\mu)$  are different from the mass eigenstates  $(v_1, v_2)$ :

$$\begin{pmatrix} \mathbf{v}_e \\ \mathbf{v}_{\mu} \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \mathbf{v}_1 \\ \mathbf{v}_2 \end{pmatrix}$$



The weak states are mixtures of the mass states:

$$|\mathbf{v}_{\mu}\rangle = -\sin\theta |\mathbf{v}_{1}\rangle + \cos\theta |\mathbf{v}_{2}\rangle$$

$$|\mathbf{v}_{\mu}(t)\rangle = -\sin\theta (|\mathbf{v}_{1}\rangle e^{-iE_{1}t}) + \cos\theta (|\mathbf{v}_{2}\rangle e^{-iE_{2}t})$$

The probability to find a  $v_e$  when you started with a  $v_u$  is:

$$P_{oscillation}(\mathbf{v}_{\mu} \rightarrow \mathbf{v}_{e}) = |\langle \mathbf{v}_{e} | \mathbf{v}_{\mu}(t) \rangle|^{2}$$

### **Motivation: Neutrino Oscillations**

In units that experimentalists like:

$$P_{oscillation}(v_{\mu} \rightarrow v_{e}) = sin^{2}2\theta sin^{2} \left(\frac{1.27 \Delta m^{2}(eV^{2}) L(km)}{E_{v}(GeV)}\right)$$

Oscillation probability between 2 flavour states depends on:

1. fundamental parameters

 $\Delta m^2 = m_1^2 - m_2^2 = \text{mass squared difference between states}$  $\sin^2 2\theta = \text{mixing between } \nu \text{ flavours}$ 

2. experimental parameters

L = distance from v source to detector

 $E = \nu$  energy















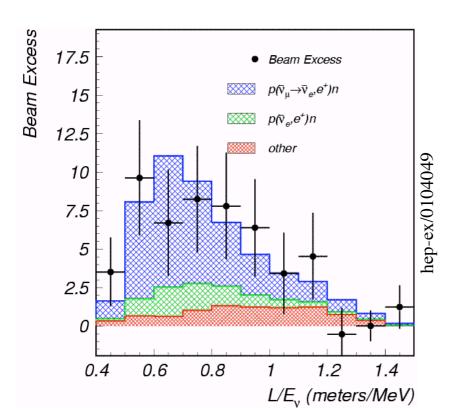


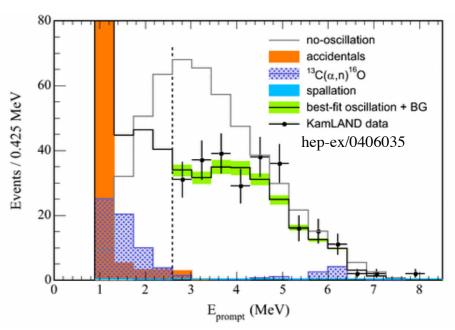
## **Motivation: Oscillation Signals**

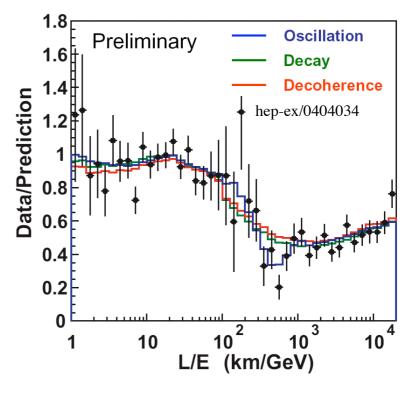
Solar v: measured by Homestake, ..., SNO confirmed by KamLAND

Atmospheric v: measured by K-II, ..., Super-K confirmed by K2K, MINOS

Accelerator v: measured by LSND unconfirmed



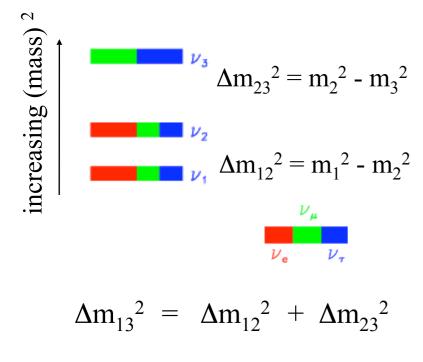


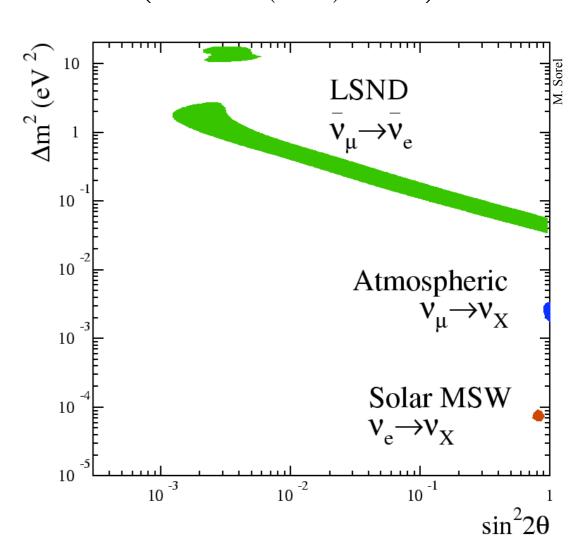


#### **Motivation: The Problem**

$$P_{oscillation}(v_{\mu} \rightarrow v_{e}) = sin^{2}2\theta sin^{2} \left(\frac{1.27 \Delta m^{2}(eV^{2}) L(km)}{E_{v}(GeV)}\right)$$

A standard 3 neutrino picture:



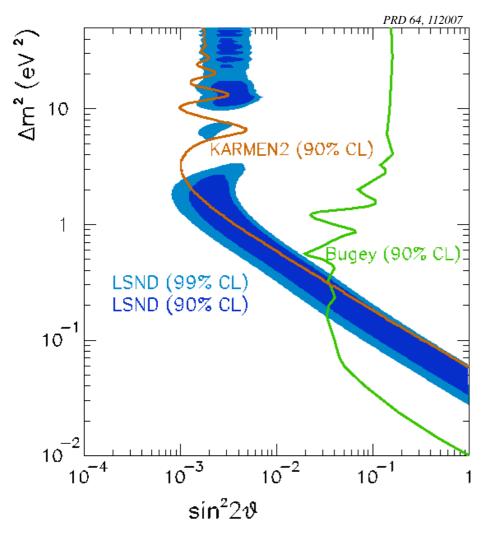


The oscillation signals cannot be reconciled without introducing physics beyond the Standard Model.

#### **Motivation: LSND**

MiniBooNE was proposed in 1997 to address the LSND result.

LSND observed a  $4\sigma$  excess of  $\overline{v}_e$  events in a  $\overline{v}_{\mu}$  beam:  $87.9 \pm 22.4 \pm 6.0$  interpreted as 2-neutrino oscillations,  $P(\overline{v}_{\mu} \rightarrow \overline{v}_e) = 0.26\%$ 



$$P = sin^2 2\theta sin^2 \left( \frac{1.27 \Delta m^2 (eV^2) L(km)}{E_{\nu}(GeV)} \right)$$

#### MiniBooNE strategy:

Keep  $(L/E_v)$  same as LSND but change systematics, including event signature:

- Order of magnitude higher  $E_v$  than LSND
- Order of magnitude longer baseline *L* than LSND
- Search for excess of  $v_e$  events above background

#### The MiniBooNE Collaboration

A. A. Aguilar-Arevalo<sup>5</sup>, A. O. Bazarko<sup>12</sup>, S. J. Brice<sup>7</sup>, B. C. Brown<sup>7</sup>, L. Bugel<sup>5</sup>, J. Cao<sup>11</sup>, L. Coney<sup>5</sup>,
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G. T. Garvey<sup>9</sup>, J. A. Green<sup>8,9</sup>, C. Green<sup>7,9</sup>, T. L. Hart<sup>4</sup>, E. Hawker<sup>15</sup>, R. Imlay<sup>10</sup>, R. A. Johnson<sup>3</sup>, P. Kasper<sup>7</sup>,
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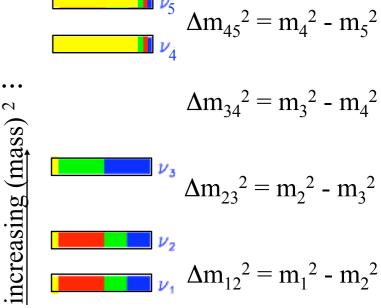
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### Motivation: MiniBooNE and LSND

If MiniBooNE observes LSND-type v oscillations...

The simplest explanation is to add more vs, to allow more independent  $\Delta m^2$  values.

The new vs would have to be **sterile**, otherwise they would have been seen already.



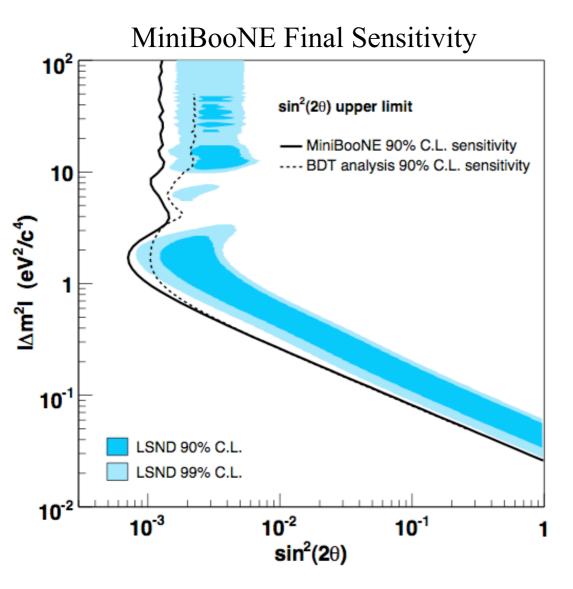
$$\nu_{\mu}$$
 $\nu_{s}$ 
 $\nu_{s}$ 

If MiniBooNE does not observe LSND-type oscillations...
The Standard Model wins again!

<u>Today</u>: MiniBooNE's initial results on testing the LSND anomaly

- A generic search for a  $v_e$  excess in our  $v_u$  beam,
- An analysis of the data within a  $\nu_{\mu} \rightarrow \nu_{e}$  appearance-only context

## **MiniBooNE Summary**



MiniBooNE performed a *blind analysis* for the  $v_u \rightarrow v_e$  appearance search

- Did not look at  $v_e$  events while developing reconstruction, particle identification algorithms
- Final cuts made with no knowledge of the number of  $v_e$  events in the box

Final sensitivity to  $v_e$  appearance shown for two independent analyses

- "Primary" analysis chosen based on slightly better sensitivity



We opened the box on March 26, 2007

**Primary Analysis** 

Cross-check Analysis

**Counting Experiment:** 

 $475 < E_v^{QE} < 1250 \text{ MeV}$ 

Counting Experiment: 300<E<sub>v</sub>QE<1500 MeV

expectation:  $358 \pm 19$  (stat)  $\pm 35$  (sys)

expectation:  $1070 \pm 33 \text{ (stat)} \pm 225 \text{ (sys)}$ 

data:

data:

significance:

significance:

Primary Analysis

Cross-check Analysis

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data: **380** 

data:

significance:  $0.55 \sigma$ 

significance:

**Primary Analysis** 

Cross-check Analysis

Counting Experiment: 475<E<sub>V</sub>QE<1250 MeV

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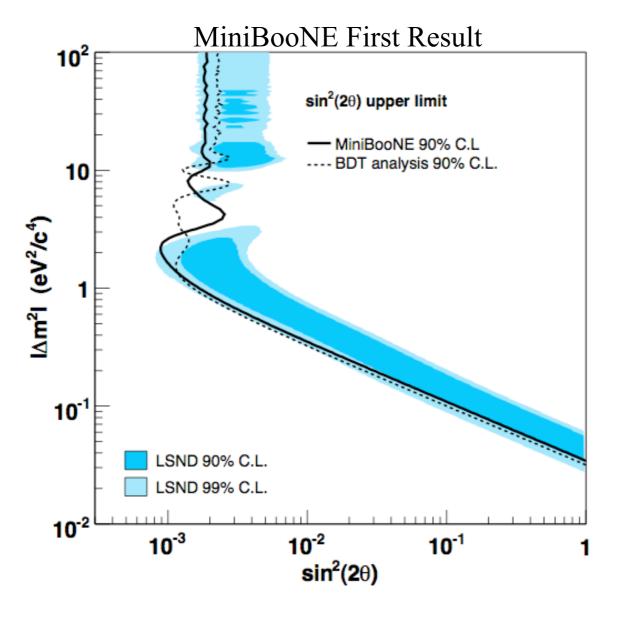
data: 380

data: **971** 

significance:  $0.55 \sigma$ 

significance: -0.38  $\sigma$ 

MiniBooNE observes no evidence for  $v_{\mu} \rightarrow v_{e}$  appearance-only oscillations.



The two independent oscillation analyses are in agreement!

The rest of this talk is a presentation of the experimental methods used to get here.



- 1. Motivation & Introduction
- 2. Description of the Experiment
  - -Beam
  - -Detector
- 3. Analysis Overview
- 4. Two Independent Oscillation Searches
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### MiniBooNE Overview: Beam and Detector

**Protons:**  $4 \times 10^{12}$  protons per 1.6 µs pulse, at 3 - 4 Hz from Fermilab Booster accelerator, with E<sub>proton</sub>=8.9 GeV. *First result uses*  $(5.58 \pm 0.12) \times 10^{20}$  protons on target.

**Mesons:** mostly  $\pi^+$ , some K<sup>+</sup>, produced in p-Be collisions, + signs focused into 50 m decay region.

**Neutrinos:** traverse 450 m soil berm before the detector hall. Intrinsic  $v_e$  flux  $\sim 0.5\%$  of  $v_u$  flux.

**Detector:** 6 m radius, 250,000 gallons of mineral oil (CH<sub>2</sub>), which emits Cherenkov and scintillation light. 1280 inner PMTs, 240 PMTs in outer veto region

MO Wascko, Imperial HEP Seminar

Absorber

#### **Booster Neutrino Beam: Neutrino Flux**

MiniBooNE is searching for an excess of  $v_e$  in a  $v_u$  beam

Modeled with a Geant4 Monte Carlo

"Intrinsic"  $v_e + \overline{v}_e$  content: 0.5%  $v_e$  Sources:

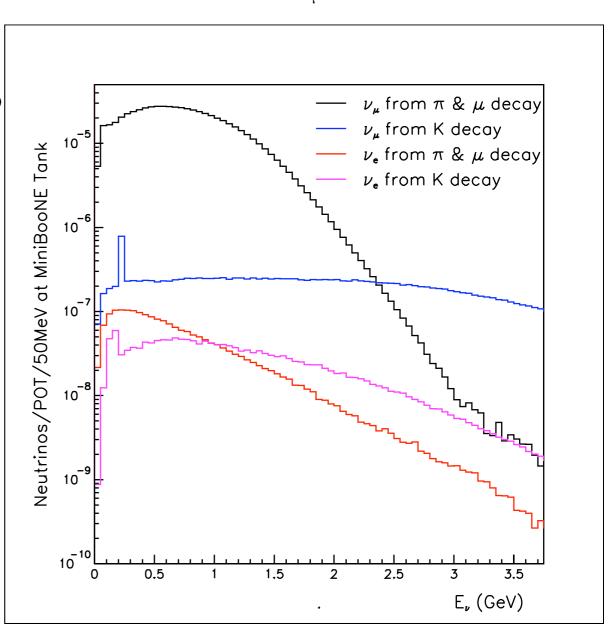
$$\mu^+ \rightarrow e^+ \overline{\nu}_{\mu} \nu_{e} \quad (42\%)$$

$$K^+ \to \pi^0 e^+ \nu_e (28\%)$$

$$K^0 \to \pi^+ e^- \nu_e \ (16\%)$$

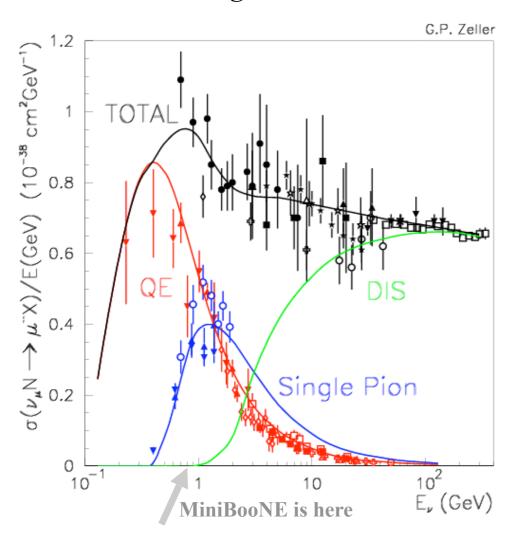
$$\pi^+ \rightarrow e^+ \nu_e \qquad (4\%)$$

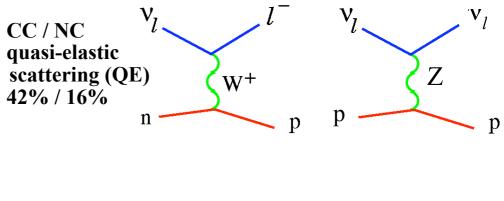
Antineutrino content: 6%

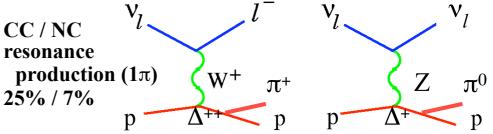


### MiniBooNE Detector: Neutrino Cross Sections

Modeling what the neutrinos do in the detector







Cross section predictions from NUANCE Monte Carlo

Use CCQE events for oscillation analysis signal channel:

$$E_{
m v}^{QE} = rac{1}{2} rac{2 M_p E_\ell - m_\ell^2}{M_p - E_\ell + \sqrt{(E_\ell^2 - m_\ell^2)} cos \Theta_\ell}$$

Only need lepton direction and angle to find v energy!

# **MiniBooNE Detector: Optics**

charged final state particles produce Ys

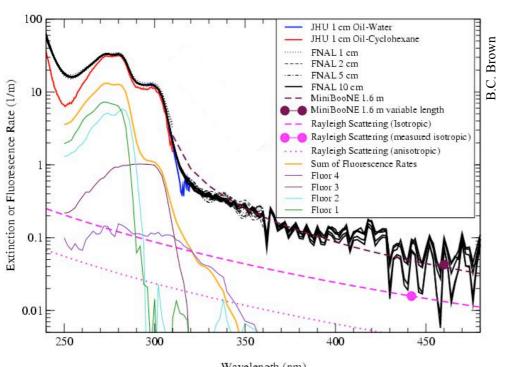
#### Cherenkov radiation

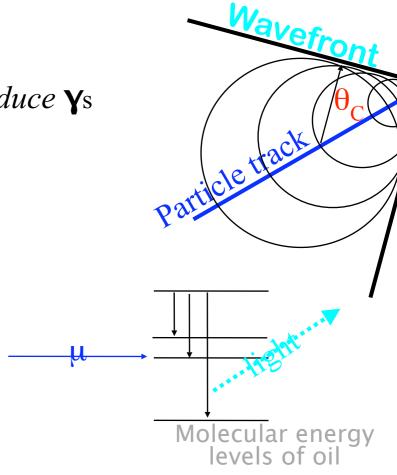
- Light emitted by oil if particle v > c/n
- forward and prompt in time

#### **Scintillation**

- Excited molecules emit de-excitation Ys
- isotropic and late in time

Extinction Rate for MiniBooNE Marcol 7 Mineral Oil





Ys are (possibly) detected by PMTs after undergoing absorption, reemission, scattering, fluorescence

"the optical model"

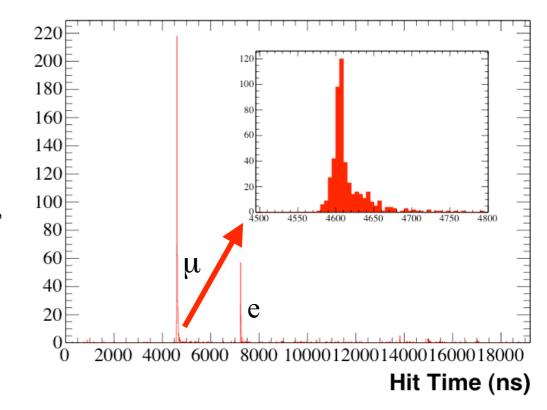
MC Wavelength (nm) April 12, 2007

### **MiniBooNE Detector: Hits**

First set of cuts based on simple hit clusters in time: "sub-events."

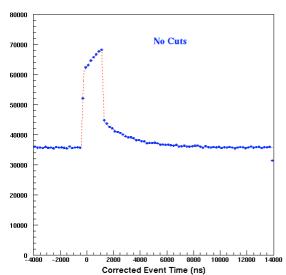
Most events are from  $\nu_{\mu}$  CC interactions, with characteristic two "sub-event" structure from stopped  $\mu$  decay.

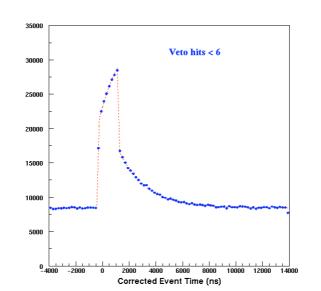
 $v_e$  CC interactions have 1 "sub-event".

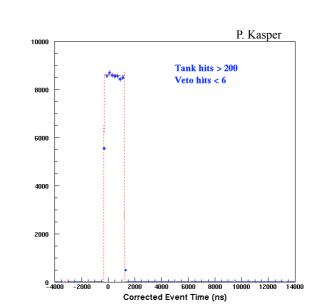


Simple cuts eliminate cosmic ray events:

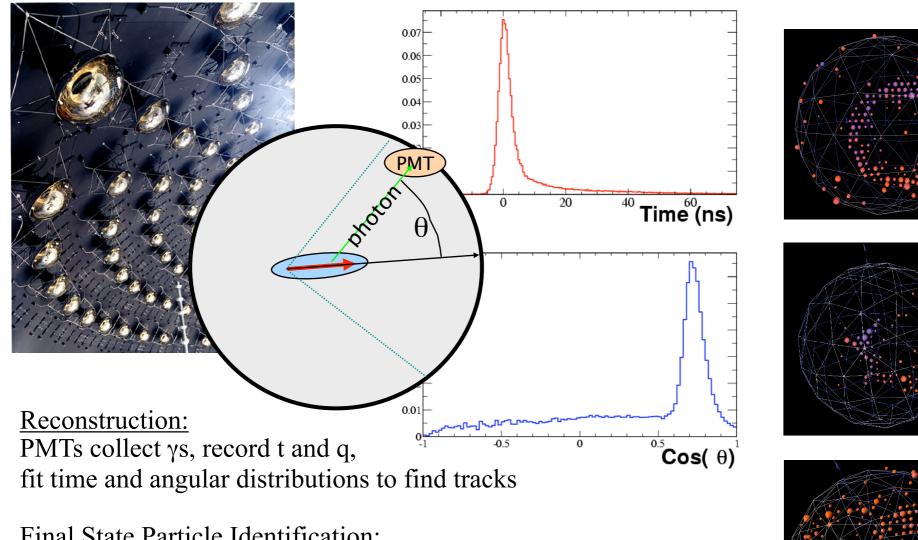
- 1. Require < 6 veto PMT hits,
- 2. Require > 200 tank PMT hits.







### MiniBooNE Detector: Reconstruction and Particle ID



#### **MiniBooNE Detector: PMT Calibration**



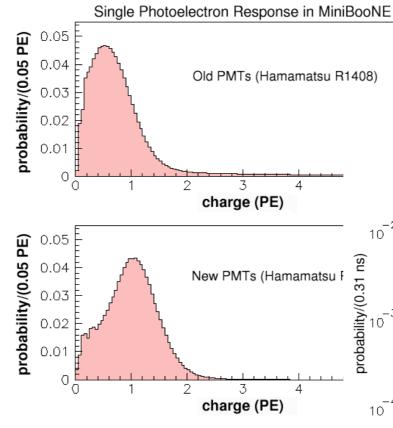
10% photo-cathode coverage

Two types of 8" Hamamatsu Tubes: R1408, R5912

Laser data are acquired at 3.3 Hz to continuously calibrate PMT gain and timing constants

PMTs are calibrated with a laser + 4 flask system

PMT Charge Resolution: 1.4 PE, 0.5 PE PMT Time Resolution: 1.7 ns, 1.1 ns



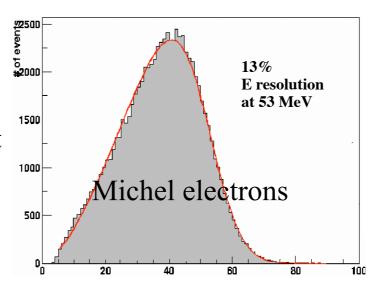
Timing Distribution for Laser Events (old tubes) prompt light late-pulsing reflections dark noise scattering (tail) pre-pulsing 10 80 corrected time (ns)

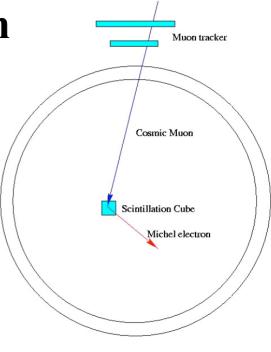
### MiniBooNE Detector: Cosmic Calibration

use cosmic muons and their decay electrons (Michels)

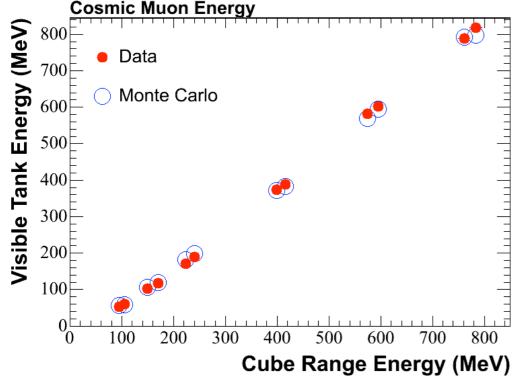
#### Michel electrons:

-set absolute energy scale and resolution at 53 MeV endpoint -optical model tuning





Muon tracker 7 scintillator cubes



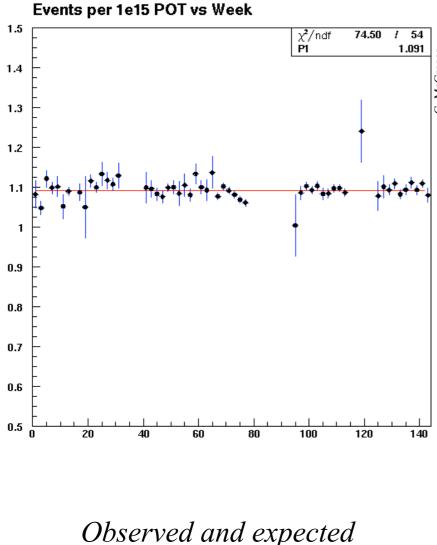
#### Cosmic muons which stop in cubes:

- -test energy scale extrapolation up to 800 MeV
- measure energy, angle resolution
- compare data and MC

Muon tracker + cube calibration data continuously acquired at 1 Hz

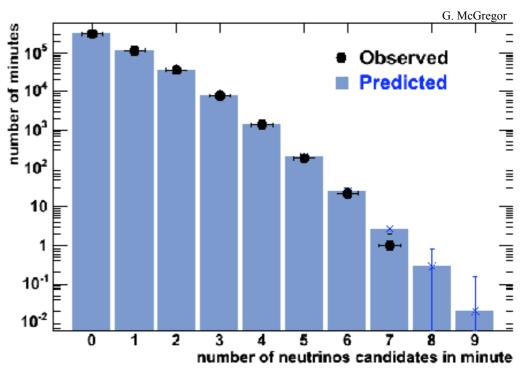
## MiniBooNE Beam & Detector: Stability

Neutrinos per proton on target throughout the neutrino run:



Observed and expected events per minute

MiniBooNE observes ~1 neutrino interaction per 1E15 protons.





- 1. Motivation & Introduction
- 2. Description of the Experiment
- 3. Analysis Overview
  - -Signal and Backgrounds
  - -Strategy
- 4. Two Independent Oscillation Searches
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## **Analysis Overview: Blind Analysis**

To avoid bias, MiniBooNE has done a blind analysis.

"Closed Box" Analysis

To study the data, we defined specific event sets with  $< 1\sigma v_e$  signal for analysis.

Initial	O.	pen	Boxes

all non-beam-trigger data

0.25% random sample

 $\nu_{\mu}$  CCQE

 $\nu_{\rm u} NC1\pi^0$ 

"dirt"

all events with  $E_v > 1.4 \text{ GeV}$ 

 $\nu_{\mu} CC1\pi^{+}$ 

 $\nu_{\mu}$ -e elastic

#### Second Step:

One closed signal box

#### <u>Use</u>

calibration and MC tuning

an unbiased data set

measure flux,  $E_{\nu}^{QE}$ , oscillation fit

measure rate for MC

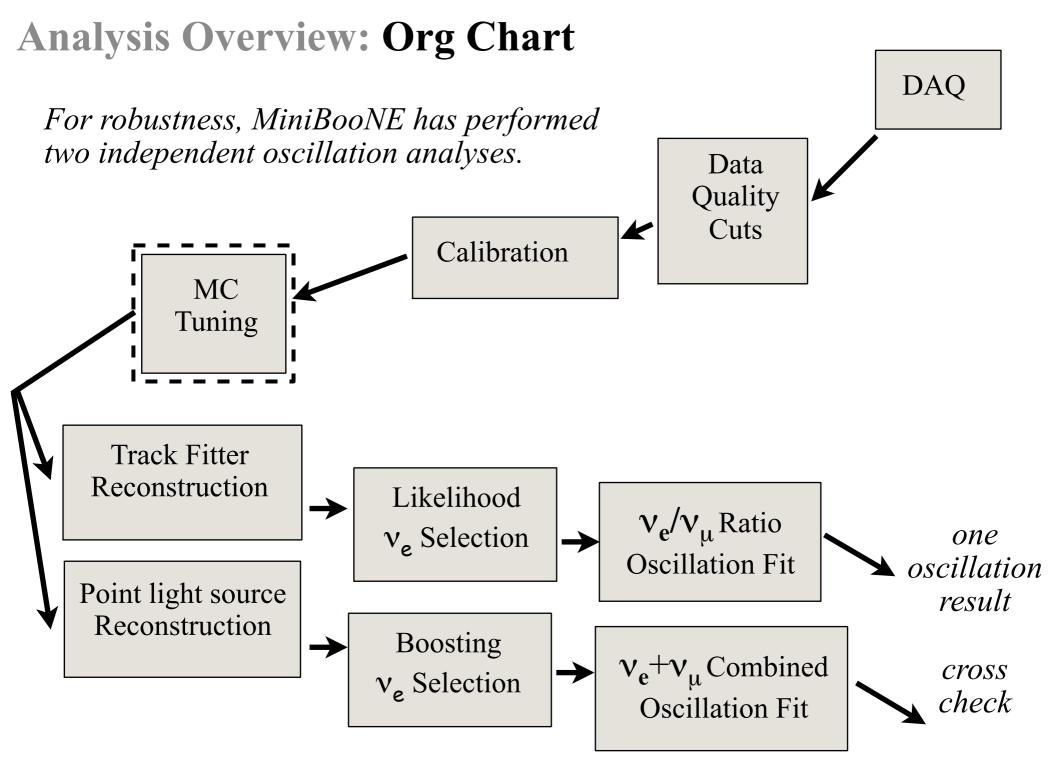
measure rate for MC

check MC rate

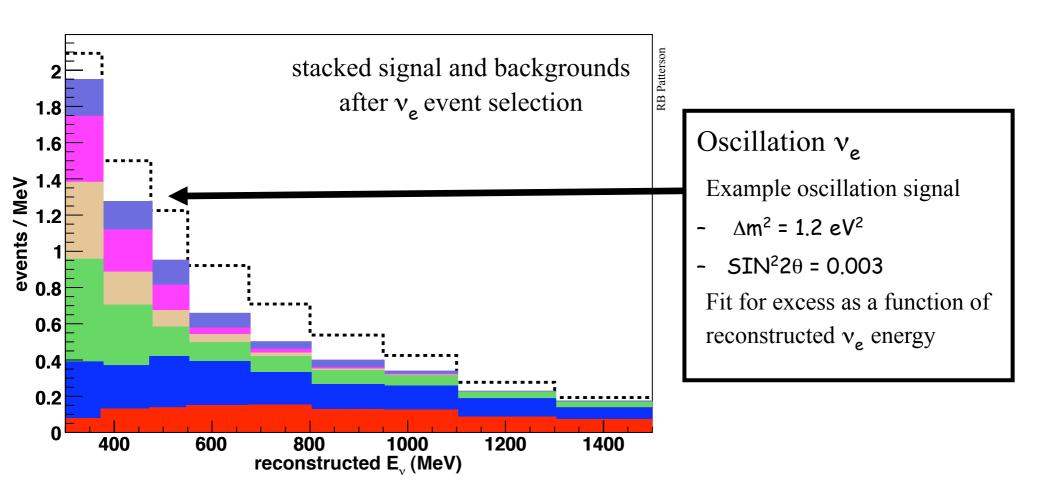
check MC rate

check MC rate

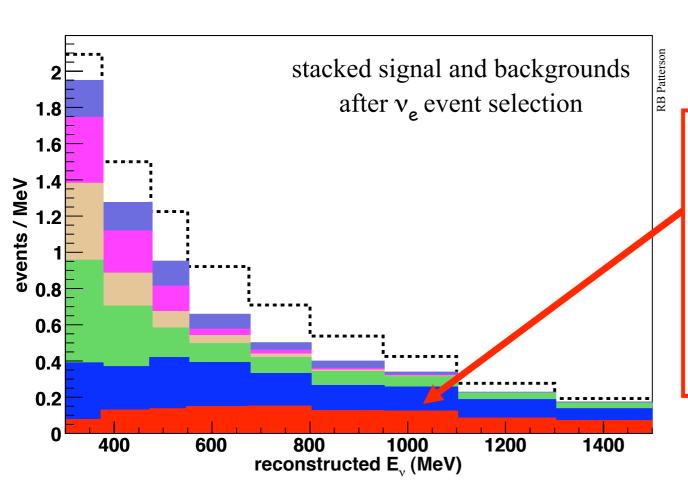
explicitly sequester the signal, 99% of data open



what we predict for the full  $\mathbf{v}$  data set (5.6E20 protons on target):



what we predict for the full  $\mathbf{v}$  data set (5.6E20 protons on target):

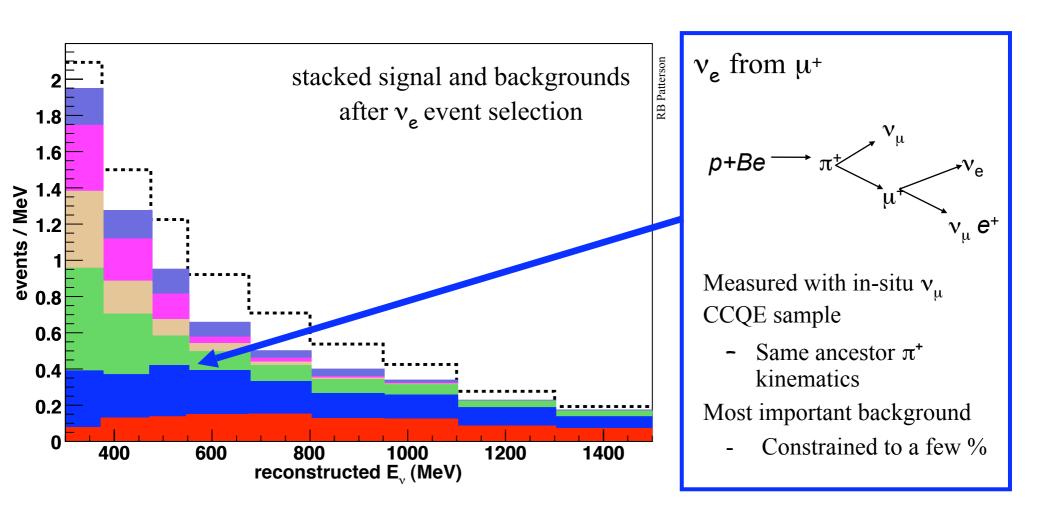


### $v_e$ from $K^+$ and $K^0$

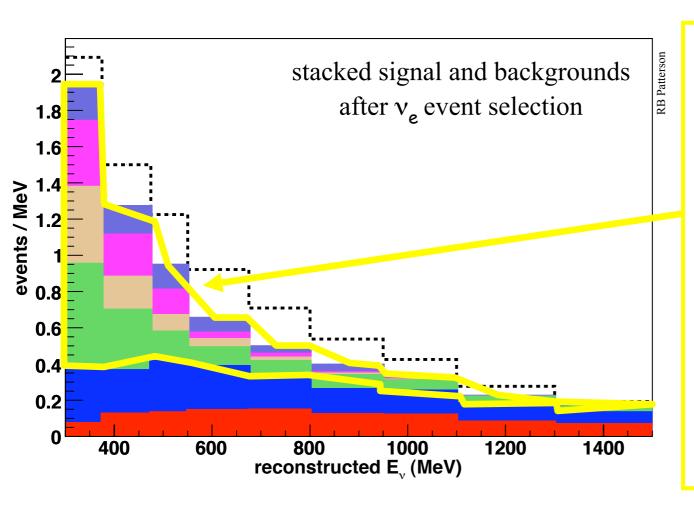
Use high energy  $v_e$  and  $v_\mu$  in-situ data for normalization cross-check

Use fit to kaon production data for shape

what we predict for the full  $\mathbf{v}$  data set (5.6E20 protons on target):



what we predict for the full  $\mathbf{v}$  data set (5.6E20 protons on target):



## MisID $\nu_{\mu}$

 $\sim 46\% \ \pi^{0}$ 

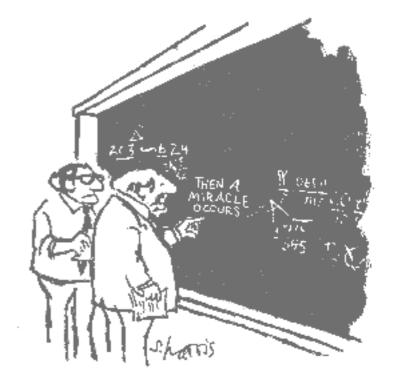
- Determined by clean  $\pi^0$  measurement
- ~14% "dirt"
  - Measure rate to normalize and use MC for shape
- ~16% Δ γ decay
  - $\pi^0$  measurement constrains
- ~24% other
  - Use  $v_{\mu}$  CCQE rate to normalize and MC for shape

# **Analysis Overview: Strategy**

recurring theme: good data/MC agreement

#### in-situ data are incorporated wherever possible...

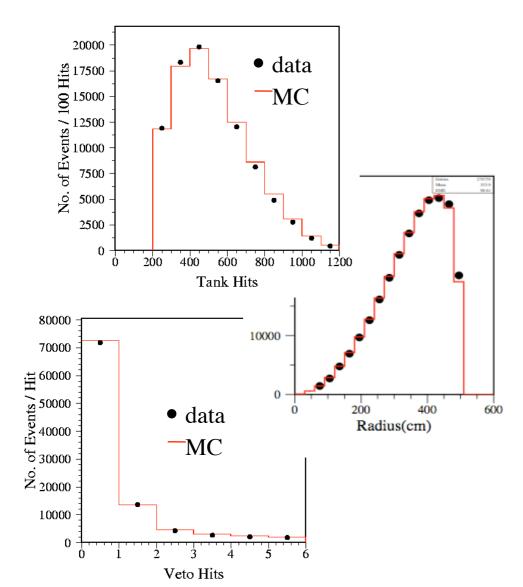
- (i) MC tuning with calibration data
  - energy scale
  - PMT response
  - optical model of light in the detector
- (ii) MC fine-tuning with neutrino data
  - cross section nuclear model parameters
  - $\pi^{\circ}$  rate constraint

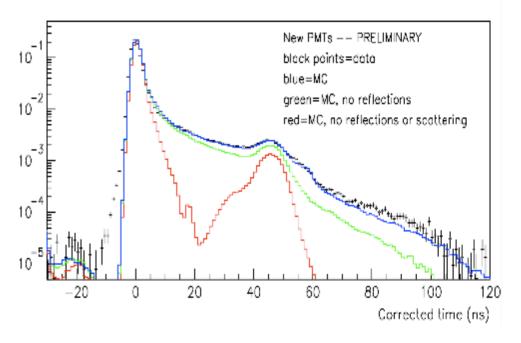


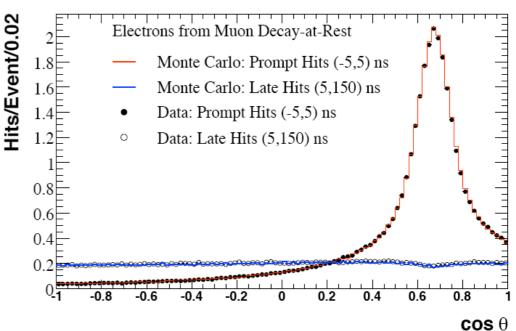
"I think you should be more explicit here in step two."

- (iii) constraining systematic errors with neutrino data
  - ratio method example:  $v_e$  from  $\mu$  decay background
  - combined oscillation fit to  $\nu_{\mu}$  and  $\nu_{e}$  data

#### MC tuning with calibration data



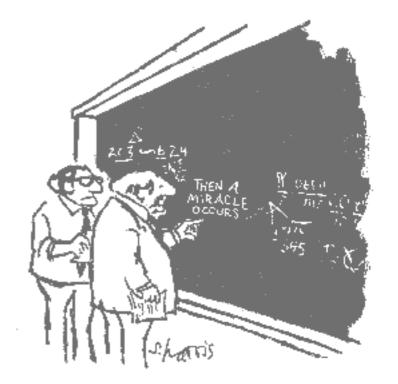




# **Analysis Overview: Strategy**

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  - PMT response
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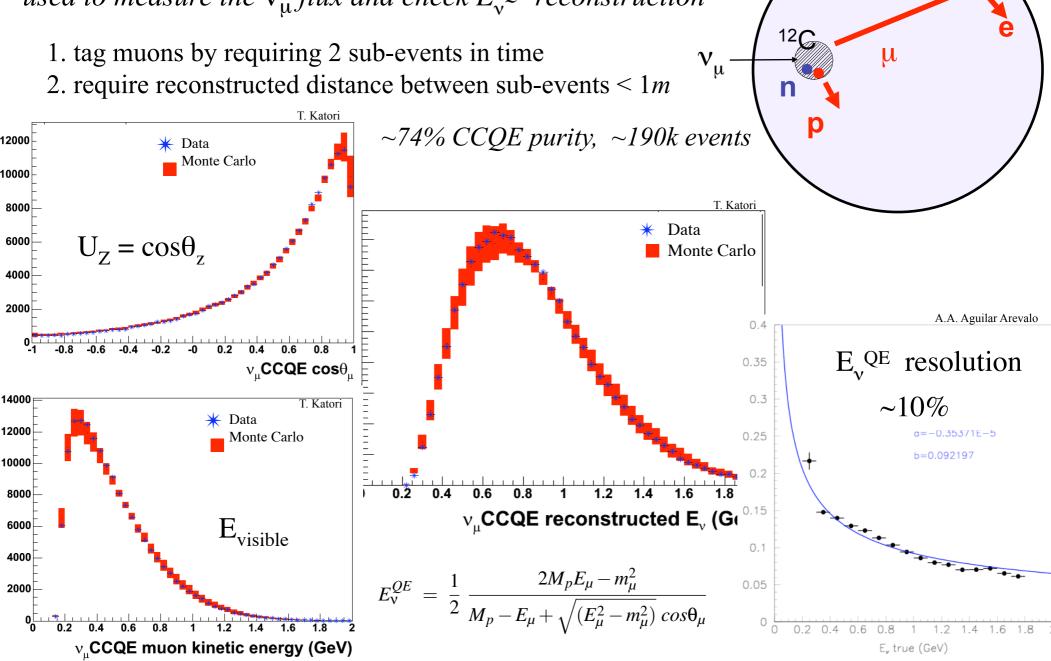


"I think you should be more explicit here in step two."

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# Analysis Strategy: $\nu_{\mu}$ CCQE Events

used to measure the  $v_{\mu}$  flux and check  $E_{\nu}^{QE}$  reconstruction

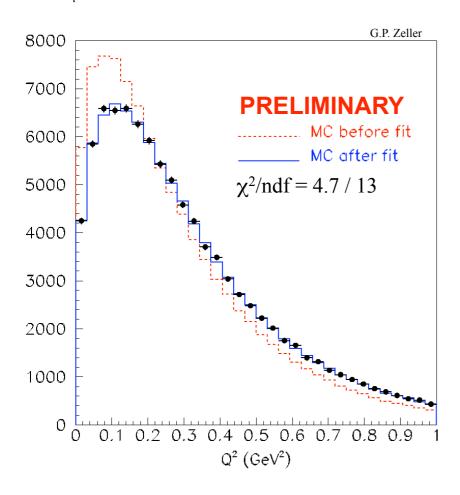


MO Wascko, Imperial HEP Seminar

April 12, 2007

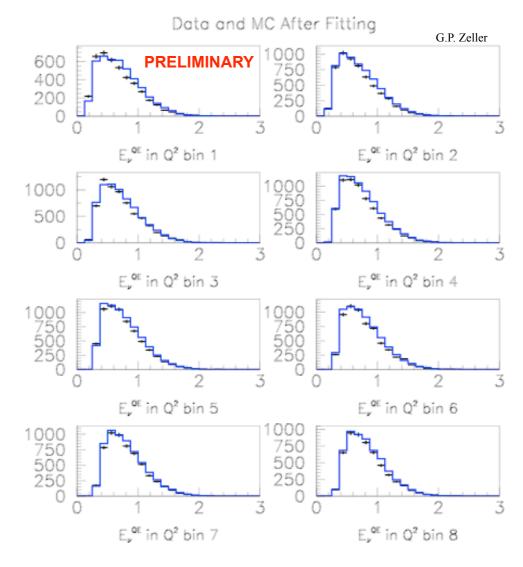
# Incorporating $v_{\mu}$ Data: CCQE Cross Section

The  $v_{\mu}$  CCQE data  $Q^2$  distribution is fit to tune empirical parameters of the nuclear model ( $^{12}$ C target)

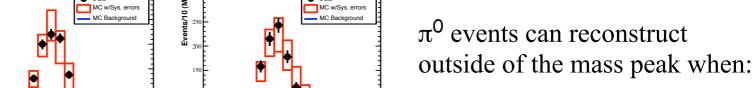


the tuned model is used for both  $\nu_{\mu}$  and  $\nu_{e}$  CCQE

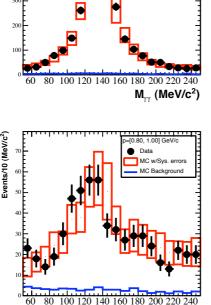
# this results in good data-MC agreement for variables **not** used in tuning



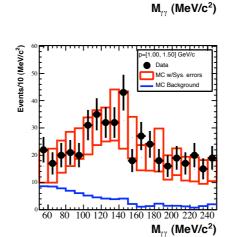
Analysis Strategy: π<sup>0</sup> Mis-ID Background clean  $\pi^0$  events are used to tune the MC rate vs.  $\pi^0$  momentum n(p) 100 120 140 160 M<sub>yy</sub> (MeV/c<sup>2</sup>) M<sub>yy</sub> (MeV/c<sup>2</sup>) M<sub>yy</sub> (MeV/c<sup>2</sup>) nts/10 (MeV/c<sup>2</sup> MC w/Sys. errors MC w/Sys. errors MC w/Sys. errors



- 1. asymmetric decays fake 1 ring
- 2. 1 of the 2 photons exits the detector
- 3. high momentum  $\pi^{\circ}$  decays produce overlapping rings



 $M_{yy}$  (MeV/c<sup>2</sup>)



60 80 100 120 140 160 180 200 220 240

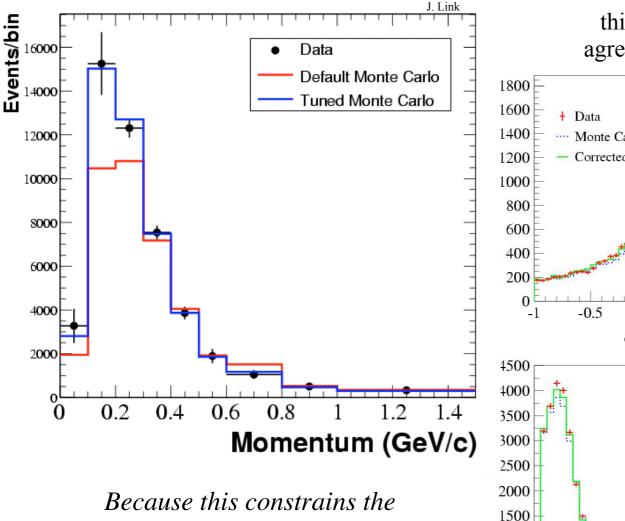
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80 100 120 140 160 180 200 220 240

 $M_{yy}$  (MeV/c<sup>2</sup>)

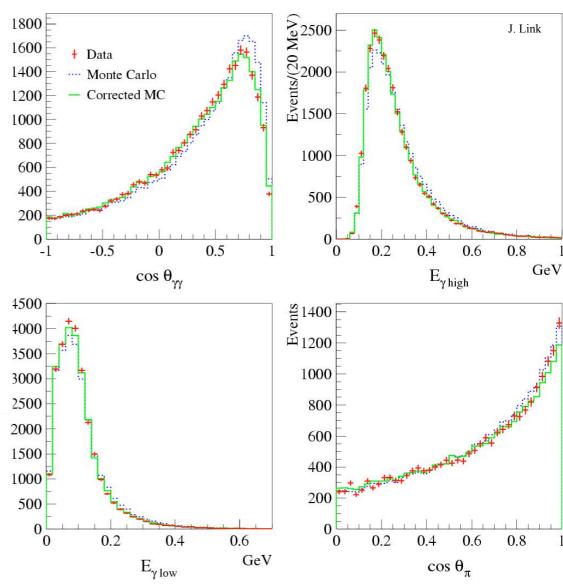
# **Analysis Strategy:** π<sup>0</sup> **Mis-ID Background**

The MC  $\pi^0$  rate (flux × xsec) is reweighted to match the measurement in  $p_{\pi}$  bins.



Because this constrains the  $\Delta$  resonance rate, it also constrains the rate of  $\Delta \rightarrow N\gamma$  in MiniBooNE

this procedure results in good data-MC agreement for variables **not** used in tuning



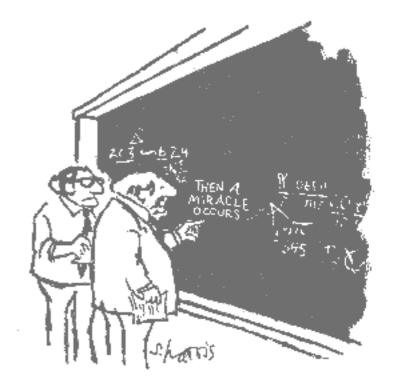
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# **Analysis Overview: Strategy**

#### in-situ data is incorporated wherever possible...

- (i) MC tuning with calibration data
  - energy scale
  - PMT response
  - optical model of light in the detector
- (ii) MC fine-tuning with neutrino data
  - cross section nuclear model parameters
  - $\pi^{\circ}$  rate constraint



"I think you should be more explicit here in step two."

- (iii) constraining systematic errors with neutrino data
  - ratio method example:  $v_e$  from  $\mu$  decay background
  - combined oscillation fit to  $\nu_{\mu}$  and  $\nu_{e}$  data

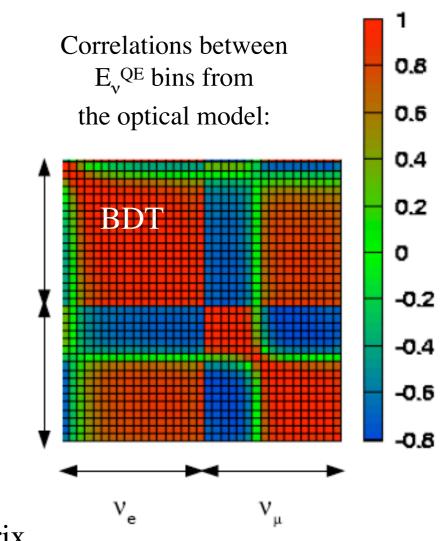
# **Analysis Strategy: Error Matrix**

$$E_{ij} \approx \frac{1}{M} \sum_{\alpha=1}^{M} \left( N_i^{\alpha} - N_i^{MC} \right) N_j^{\alpha} - N_j^{MC}$$

- N is number of events passing cuts
- •MC is standard Monte Carlo
- α represents a different MC draw (called a "multisim")
- M is the total number of MC draws
- i,j are E<sub>v</sub>QE bins

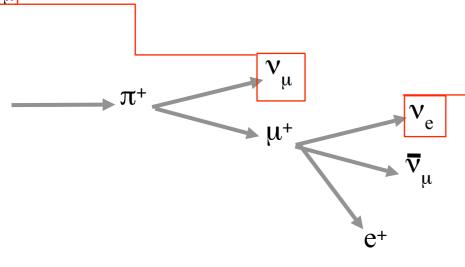
Total error matrix is sum from each source.

Primary (TB):  $v_e$ -only total error matrix Cross-check (BDT):  $v_u$ - $v_e$  total error matrix



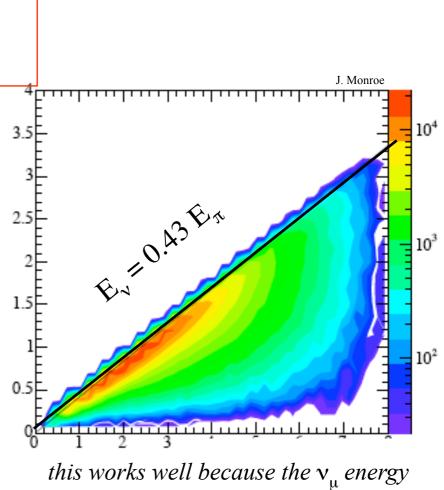
# Incorporating $\nu_{\mu}$ Data: $\mu^{+}$ -Decay $\nu_{e}$ Background

 $v_{\mu}$  CCQE events measure the  $\pi^+$  spectrum, this constrains the  $\mu^+$ -decay  $v_e$  flux



#### **Ratio Method Constraint:**

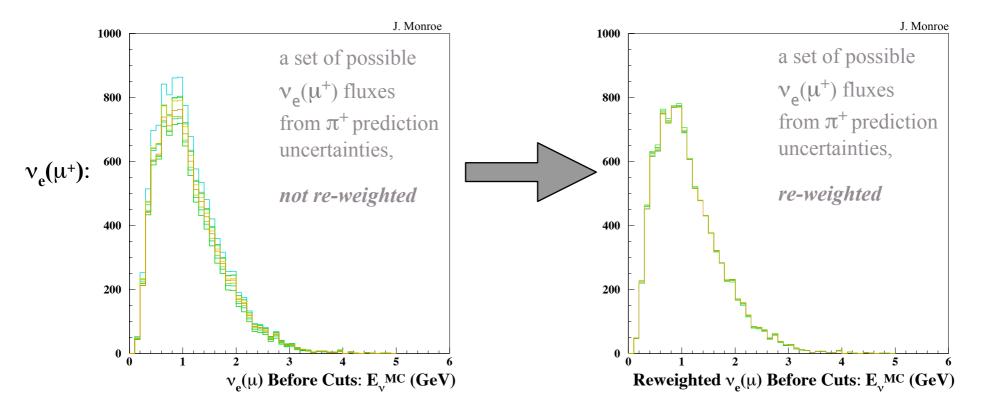
- 1. MC based on external data predicts a central value and a range of possible  $v_{\mu}(\pi)$  fluxes
- 2. make Data/MC ratio vs.  $E_{\nu}^{QE}$  for  $\nu_{\mu}CCQE$  data
- 3. re-weight each possible MC flux by the ratio (2) including the  $\nu_u$ , its parent  $\pi^+$ , sister  $\mu^+$ , and niece  $\nu_e$



this works well because the  $\nu_{\mu}$  energy is highly correlated with the  $\pi^+$  energy

## **Analysis Strategy 1: Ratio Method**

Impact of re-weighting the simulation using "fake data" (MC):



this reduction in the spread of possible fluxes translates directly into a reduction in the  $\mu^+$ -decay  $\nu_e$  background uncertainty

Can use ratio method to constrain most BG sources

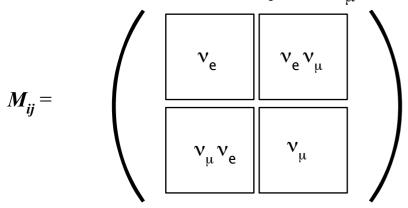
## **Analysis Strategy 2: Combined Fit**

Fit the  $E_{\nu}^{QE}$  distributions of  $\nu_e$  and  $\nu_{\mu}$  events for oscillations, together

Raster scan in  $\Delta m^2$ , and  $sin^2 2\theta_{\mu e}$  ( $sin^2 2\theta_{\mu x} == 0$ ), calculate  $\chi^2$  value over  $v_e$  and  $v_u$  bins

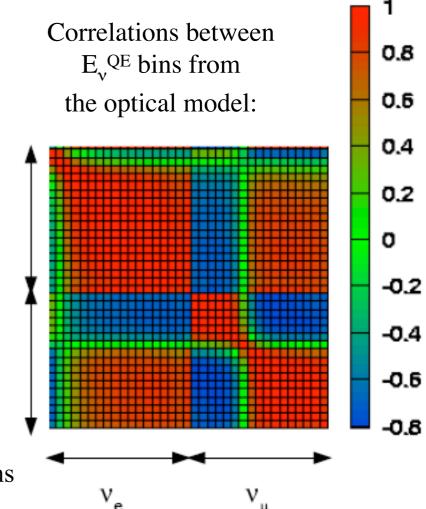
$$\chi^2 = \sum_{i=1}^{N_{bins}} \sum_{j=1}^{N_{bins}} (m_i - t_i) \, \mathcal{M}_{ij}^{-1} \, (m_j - t_j)$$

In this case, systematic error matrix  $M_{ij}$  includes predicted uncertainties for  $v_e$  and  $v_u$  bins



Left: example,  $m_i$  = "fake data" = MC with no oscillations

a combined fit constrains uncertainties in common



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## **Analysis Overview: Systematic Errors**

A long list of systematic uncertainties are estimated using Monte Carlo:

### neutrino flux predictions

- $\pi^+$ ,  $\pi^-$ ,  $K^+$ ,  $K^-$ ,  $K^0$ , n, and p total and differential cross sections
- secondary interactions of mesons
- focusing horn current
- target + horn system alignment

#### neutrino interaction cross section predictions

- nuclear model
- rates and kinematics for relevant exclusive processes
- resonance width and branching fractions

### detector modelling

- optical model of light propagation in oil (39 parameters!)
- PMT charge and time response
- electronics response
- neutrino interactions in dirt surrounding detector hall

**√Most** are constrained or checked using in-situ MiniBooNE data.

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- 1. Motivation & Introduction
- 2. Description of the Experiment
- 3. Analysis Overview
- 4. Two Independent Oscillation Searches
  - -Reconstruction and Event Selection
  - -Systematic Uncertainties
- 5. First Results

## Two Independent Oscillation Searches: Methods

Method 1: Track-Based Analysis

- Use careful reconstruction of particle tracks
- Identify particle type by likelihood ratio
- •Use ratio method to constrain backgrounds

### Strengths:

Relatively insensitive to optical model Simple cut-based approach with likelihoods

Method 2: Boosted Decision Trees

Independent cross-check

- Classify events using "boosted decision trees"
- Apply cuts on output variables to improve separation of event types
- •Use combined fit to constrain backgrounds

### Strengths:

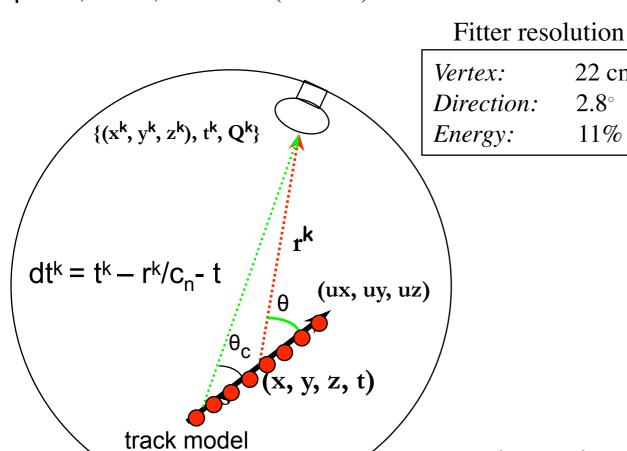
Combination of many weak variables form strong classifier Better constraints on background events

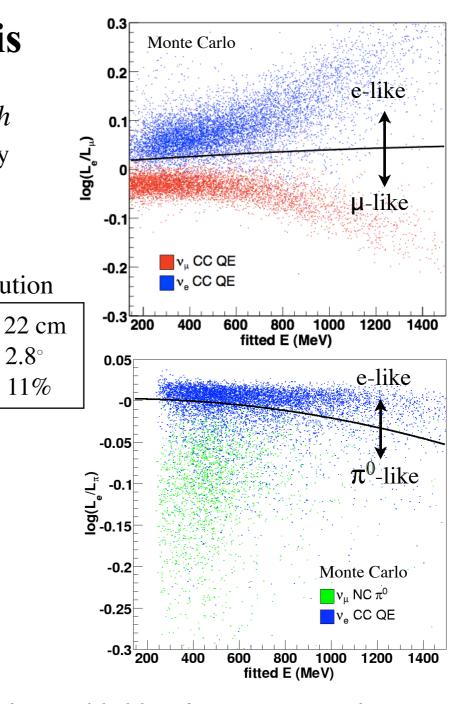
Primary analysis

# **Method 1: Track-Based Analysis**

Reconstruction fits an extended light source with 7 parameters: vertex, direction  $(\theta, \phi)$ , time, energy

Fit events under 3 possible hypotheses:  $\mu$ -like, e-like, two track ( $\pi^0$ -like)





Particle ID relies on likelihood ratio cuts to select  $v_e$ , cuts chosen to maximise sensitivity to  $v_{\mu} \rightarrow v_{e}$  oscillation

 $2.8^{\circ}$ 

## Track-Based Analysis: e/µ Likelihood

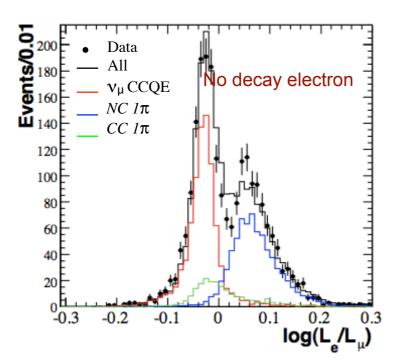
Test µ-e separation on data:

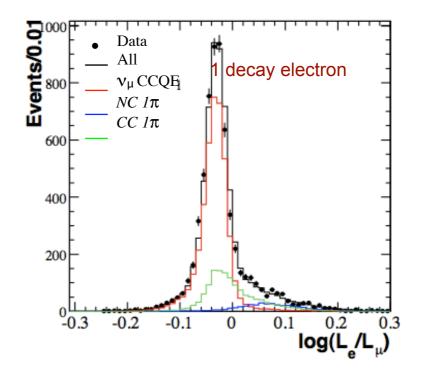
### $\nu_{\mu}$ CCQE data sample

Pre-selection cuts

Fiducial volume: (R < 500 cm)

2 subevents: muon + decay electron





### "All-but-signal" data sample

Pre-selection cuts

Fiducial volume: (R < 500 cm)

1 subevent: 8% of muons capture on <sup>12</sup>C

Events with  $\log(L_e/L_{\mu}) > 0$  (e-like) undergo additional fit with two-track hypothesis.

# Track-Based Analysis: $e/\pi^0$ Likelihood

Test  $e^{-\pi^0}$  separation on data: events/5 MeV/c2 2000 Monte Carlo Simulation "All-but-signal" data sample Data Monte Carlo  $\pi^0$  only 1500 Pre-selection cuts Fiducial volume cut (R < 500 cm) Signal 1000 1 subevent Region Invariant mass  $> 50 \text{ MeV/c}^2$ Events/0.01  $\log(L_e/L_\pi) < 0 \ (\pi\text{-like})$ **BLINDED REGION** REGION INDED Events/10.0 (MeV/c²) -0.3 -0.4BLINDED 50 100 150 200 250 Invariant Mass (MeV/c²) Data Tighter selection cuts: Monte Carlo Mass  $< 50 \text{ MeV/c}^2$ : Invariant mass < 200 MeV/c<sup>2</sup>  $10^{2}$  $\chi^2/\text{ndf} =$  $log(L_e/L_{\mu}) > 0$  (e-like) 20 40 60 80 100 120 140 160 180 200  $\log(L_e/L_\pi) < 0$  ( $\pi$ -like) Mass (MeV/c²)

### **Method 2: Boosted Decision Trees**

**Decision Trees:** A machine-learning technique which tries to recover signal events that would be eliminated in cut-based analyses.

#### *Training a decision tree:*

Step 1: For a set of N variables, determine the cut value for each variable that gives the best separation between signal and background.

Step 2: Choose the variable with the overall strongest separation; divide the events between two branches:

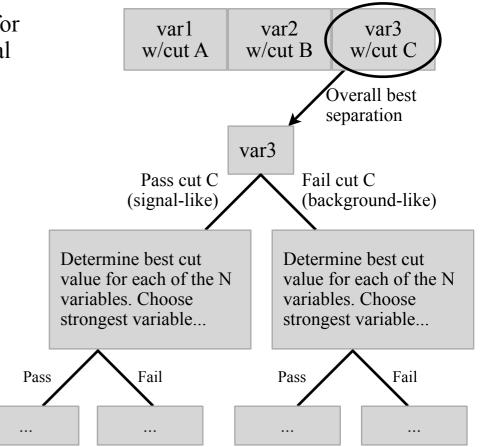
Events which passed the cut (classified as "signal") Events which failed the cut (classified as "background")

Steps 3-n: Repeat recursively for each node...

Stop when reach min. purity or max. iterations (Stopping points called "leaves" or "terminal nodes")

Final score: For each leaf,

- 1 for events on a "background-like" leaf
- +1 for events on a "signal-like" leaf

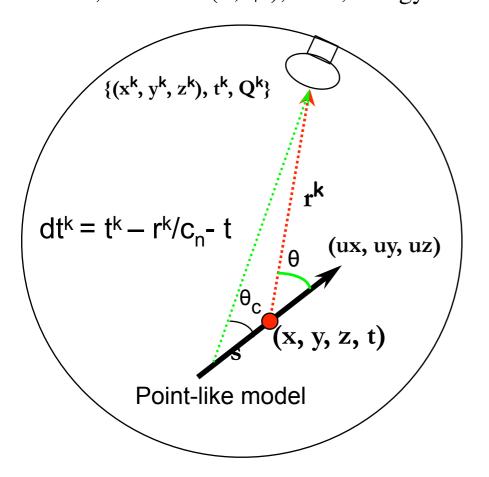


**Boosting:** Give a new weight to all events in leaves of decision tree; misclassified events receive a stronger weight. Re-training the decision tree with newly weighted events improves performance.

B.P. Roe, et al., NIM A543 (2005) 577. H. Yang, B.P. Roe, J. Zhu, NIM A555 (2005) 370

### **Boosted Decision Trees: Reconstruction and Particle ID**

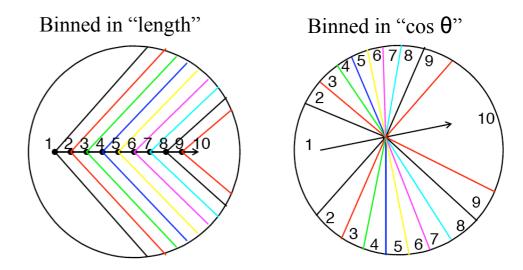
Reconstruction fits a point-like light source: vertex, direction  $(\theta, \phi)$ , time, energy



#### Fitter resolution

Vertex: 24 cm
Direction: 3.8°
Energy: 14%

Characterize topology of each event by dividing detector into "bins" relative to track:

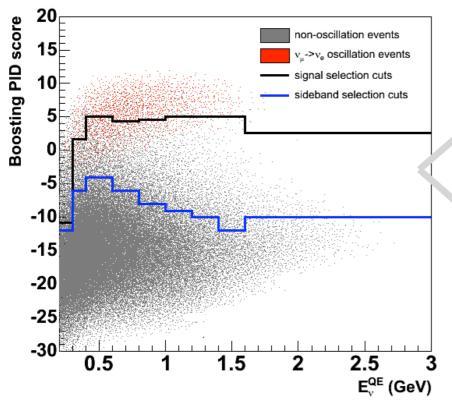


Particle ID "input variables" for the boosted decision trees are created from basic quantities in each bin: *e.g.*, charge, number of hits... *To select events, a particle ID cut is made on the Boosting output score*.

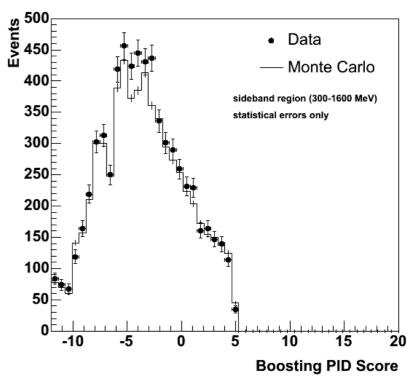
### **Boosted Decision Trees: Particle ID**

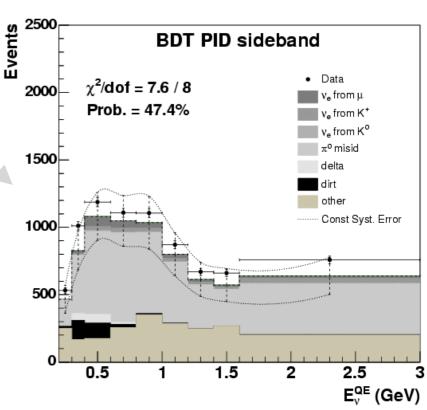
A sideband region is selected to validate MC in region near signal.

Sideband contains mostly misidentified  $\pi^0$  background events.



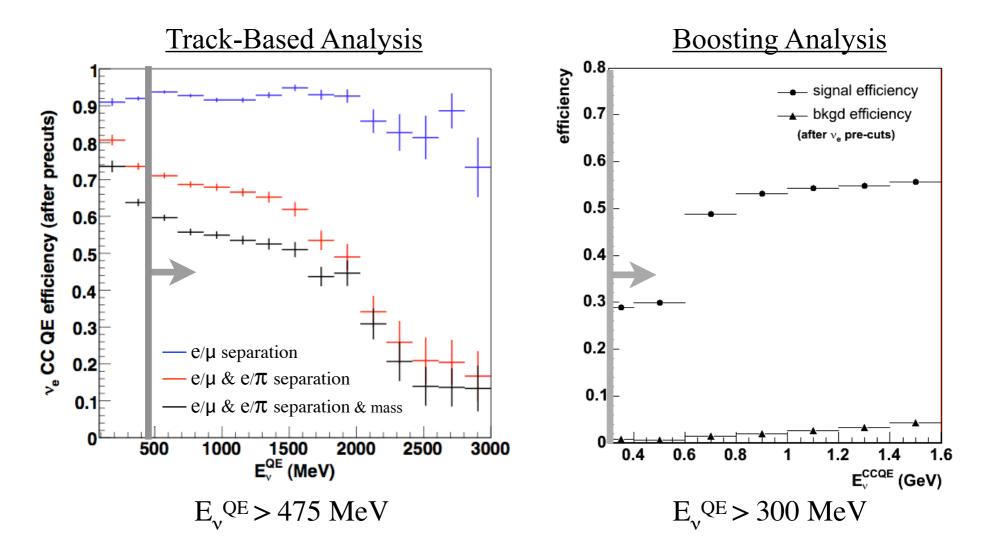
A  $\chi^2$  is calculated using the full systematic error matrix, data and MC are consistent.





## **Comparison: Efficiencies**

The two analyses have different event selection efficiency vs. energy trends,



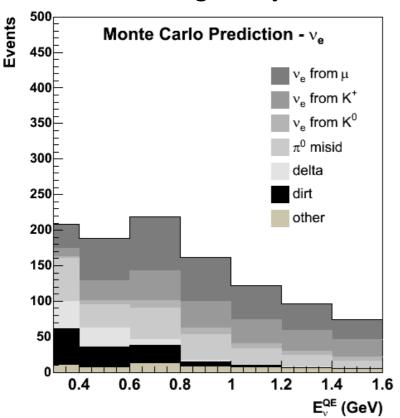
and different reconstructed  $E_{\gamma}$  regions for the oscillation analyses.

## **Comparison: Backgrounds**

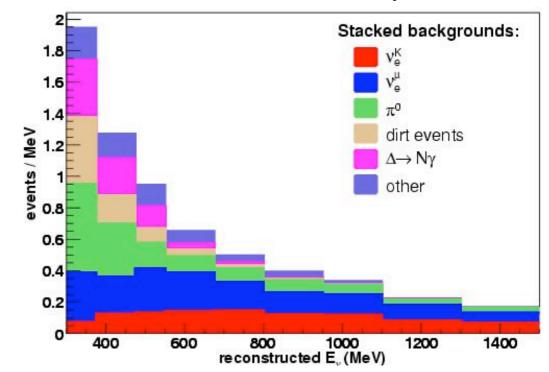
The two analyses have somewhat different background compositions.

Source	T-B	В
$v_e$ from $\mu$ decay	0.37	0.32
$v_e$ from $K$ decay	0.26	0.24
$\pi^0$ mis – ID	0.17	0.21
$\Delta  o N \gamma$	0.06	0.07
Dirt	0.05	0.11
Other	0.09	0.05

#### **Boosting Analysis**



### **Track-Based Analysis**



## **Comparison: Systematic Errors**

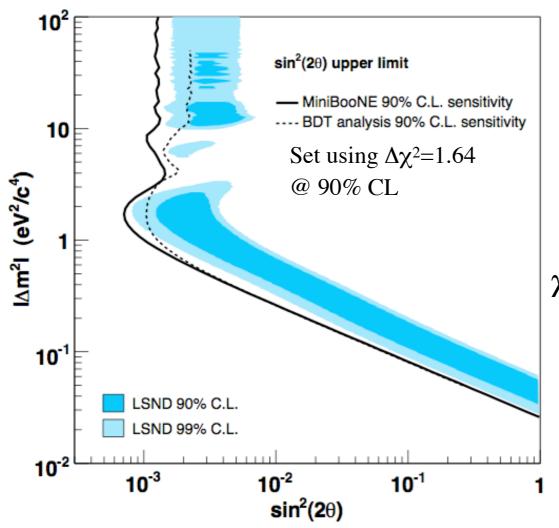
Both analyses construct error matrices for the oscillation fit, binned in  $E_{\nu}$ , to estimate the uncertainty on the expected number of  $\nu_e$  background events.

	source	track-based (%)	boosting (%)
$\sqrt{}$	Flux from π+/μ+ decay	6.2	4.3
$\sqrt{}$	Flux from K+ decay	3.3	1.0
$\sqrt{}$	Flux from K <sup>0</sup> decay	1.5	0.4
	Target and beam models	2.8	1.3
$\sqrt{}$	V-cross section	12.3	10.5
	NC π <sup>0</sup> yield	1.8	1.5
	External interactions	0.8	3.4
$\sqrt{}$	Optical model	6.1	10.5
	DAQ electronics model	7.5	10.8
	constrained total	9.6	14.5

Note:

"total" is **not**the quadrature
sum-- errors are
further reduced
by fitting with  $v_{\mu}$  data  $\sqrt{\phantom{a}}$ 

# **Comparison: Sensitivity**



Fit the Monte Carlo  $E_{\nu}^{QE}$  event distributions for oscillations

Raster scan in  $\Delta m^2$ , and  $\sin^2 2\theta_{\mu e}$ (assume  $\sin^2 2\theta_{\mu x} == 0$ ), calculate  $\chi^2$  value over  $E_v$  bins

$$\chi^{2} = \sum_{i=1}^{N_{bins}} \sum_{j=1}^{N_{bins}} (m_{i} - t_{i}) \, \mathcal{M}_{ij}^{-1} \, (m_{j} - t_{j})$$

 $m_i$  = Number of measured data events in bin i

 $t_i$  = Number of predicted events in bin i

( $t_i$  events are a function of  $\Delta m^2$ ,  $\sin^2 2\theta$ ,

 $M_{ij}^{-1}$  = Inverse of the covariance matrix

Since the track-based analysis achieved better sensitivity than the boosted decision tree analysis, we decided (before opening the box) that it would be used for the primary result.



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# **Results: Opening the Box**

### After applying all analysis cuts:

- Step 1: Fit sequestered data to an oscillation hypothesis
   Fit does not return fit parameters
   Unreported fit parameters applied to MC; diagnostic variables compared to data
   Return only the χ² of the data/MC comparisons (for diagnostic variables only)
- Step 2: Open plots from Step 1 (Monte Carlo has unreported signal)
  Plots chosen to be useful diagnostics, without indicating if signal was added (reconstructed position, direction, visible energy...)
- Step 3: Report only the  $\chi^2$  for the fit to  $E_{\nu}^{QE}$ No fit parameters returned
- Step 4: Compare  $E_{\nu}^{QE}$  for data and Monte Carlo, Fit parameters **are** returned At this point, the box is open (March 26, 2007)

Step 5: Present results two weeks later (today!)

Track-Based

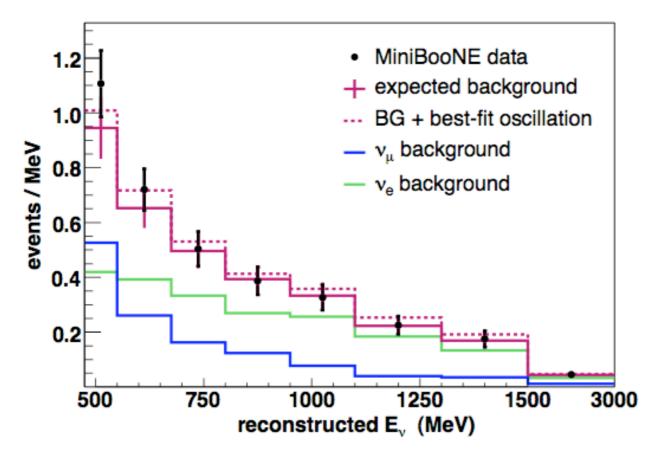
 $\chi^2$  Probability: 99%

**Boosting** 

 $\chi^2$  Probability: 62%



# Results: Track Based Analysis



We observe no significant evidence for an excess of  $v_e$  events in the energy range of the analysis.

Best Fit (dashed):  $(\sin^2 2\theta, \Delta m^2) = (0.001, 4 \text{ eV}^2)$ 

Counting Experiment: 475<E, QE<1250 MeV

data: 380

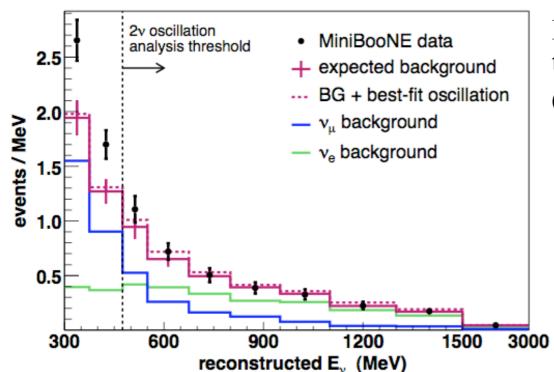
expectation:  $358 \pm 19 \text{ (stat)} \pm 35 \text{ (sys)}$ 

significance:

 $0.55 \sigma$ 

 $\chi^2$  probability of best-fit point: 99%  $\chi^2$  probability of null hypothesis: 93%

# Results: Track Based Analysis, Lower Energy Threshold

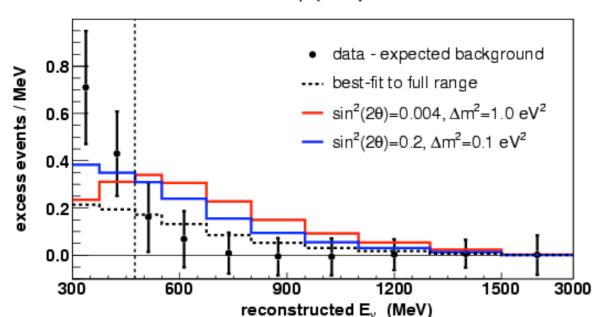


Extending down to energies below the analysis range:  $E_v^{QE} > 300 \text{ MeV}$  (we agreed to report this before box opening)

Data deviation for  $300 < E_v^{QE} < 475 \text{ MeV}: 3.7\sigma$ 

Oscillation fit to  $E_v^{QE} > 300 \text{ MeV}$ :

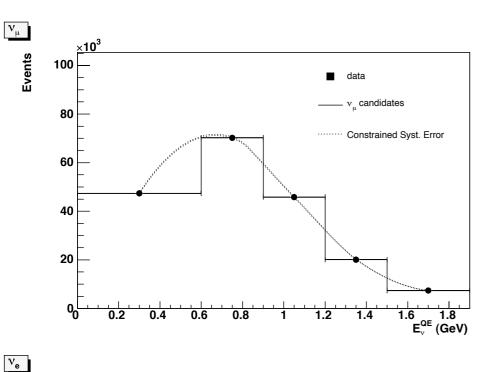
Best Fit  $(\sin^2 2\theta, \Delta m^2) = (1.0, 0.03 \text{ eV}^2)$ 



 $\chi^2$  probability at best-fit point: 18%

Fit is inconsistent with  $v_{\mu} \rightarrow v_{e}$  oscillations.

# Results: Boosted Decision Tree Analysis



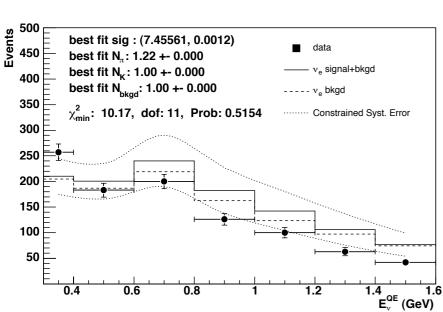
We observe no significant evidence for an excess of  $\nu_e$  events in the energy range of the analysis.

### **Counting Experiment:**

 $300 < E_v^{QE} < 1500 \text{ MeV}$ 

data: 971

expectation:  $1070 \pm 33 \text{ (stat)} \pm 225 \text{ (sys)}$ 



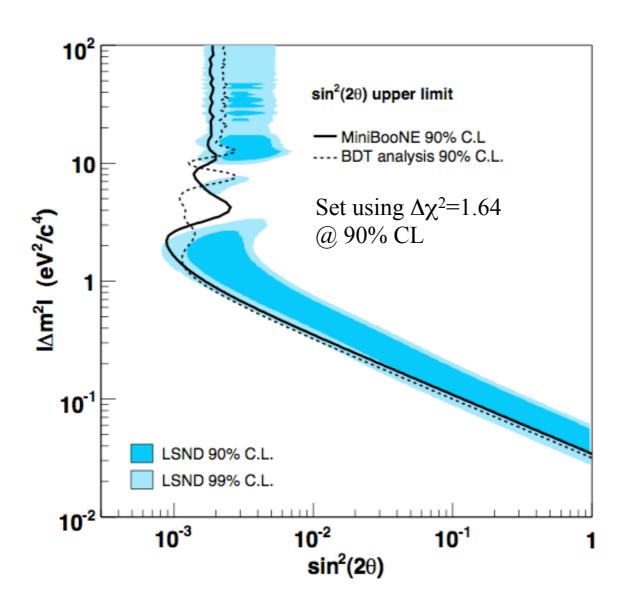
Best Fit Point (dashed):  $(\sin^2 2\theta, \Delta m^2) = (0.001, 7 \text{ eV}^2)$ 

 $\chi^2$  probability of best-fit point: 62%  $\chi^2$  probability of null hypothesis: 52%

significance: -0.38 σ

# **Results: Comparison**

MiniBooNE observes no evidence for  $v_{\mu} \rightarrow v_{e}$  appearance-only oscillations.



The two independent oscillation analyses are in agreement.

solid: track-based  $\Delta \chi^2 = \chi^2_{best fit} - \chi^2_{null}$ = 0.94

dashed: boosting

$$\Delta \chi^2 = \chi^2_{best fit} - \chi^2_{null}$$
$$= 0.71$$

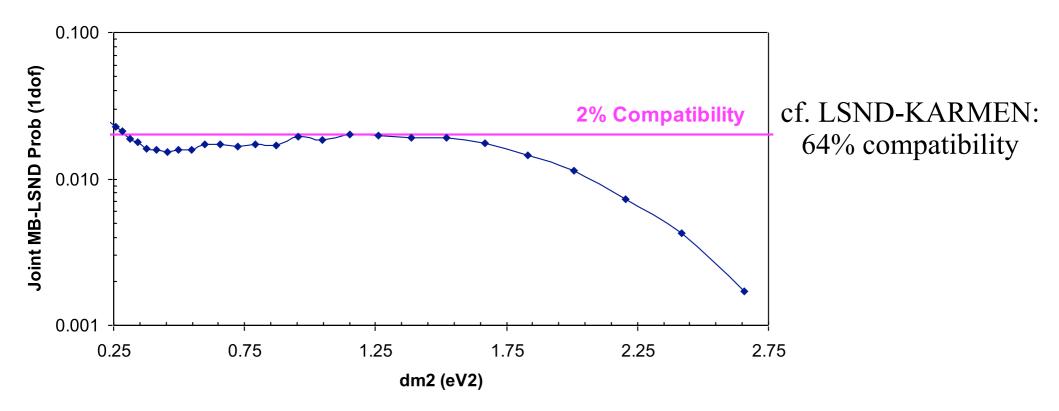
Therefore, we set a limit.

### **Results: Compatibility with LSND**

A MiniBooNE-LSND Compatibility Test:

$$\chi_0^2 = \frac{(z_{MB} - z_0)^2}{\sigma_{MB}^2} + \frac{(z_{LSND} - z_0)^2}{\sigma_{LSND}^2}$$

- For each  $\Delta m^2$ , form  $\chi^2$  between MB and LSND measurement
- Find  $z_0$  ( $sin^22\theta$ ) that minimises  $\chi^2$  (weighted average of 2 measurements), this gives  $\chi^2_{min}$
- Find probability of  $\chi^2_{min}$  for 1 dof = joint compatibility probability for this  $\Delta m^2$



MiniBooNE is incompatible with a  $\nu_{\mu} \rightarrow \nu_{e}$  appearance-only interpretation of LSND at **98% CL** 

### **Results: Plans**

A paper on this analysis is posted to the archive.

Many more papers supporting this analysis will follow, in the very near future:

 $\nu_{\mu}$  CCQE production  $\pi^0$  production

We are pursuing further analyses of the neutrino data, including: an analysis which combines TB and BDT, less simplistic models for the LSND effect.

MiniBooNE is presently taking data in antineutrino mode.

### SciBooNE will start taking data in June!

Will improve constraints on  $v_e$  backgrounds (intrinsic  $v_e$ s, improved  $\pi^0$  kinematics)

Will provide important constraints on "wrong-sign" BGs for antineutrino oscillation analysis

### **Conclusions**

- 1. Within the energy range of the analysis, MiniBooNE observes no statistically significant excess of  $v_{\rho}$  events above background.
- 2. In two independent oscillation analyses, the observed  $E_{\nu}$  distribution is inconsistent with a  $\nu_{\mu} \rightarrow \nu_{e}$  appearance-only model.
- 3. Therefore, we set a limit on  $v_{\mu} \rightarrow v_{e}$  oscillations at  $\Delta m^{2} \sim 1 \text{ eV}^{2}$ . The MiniBooNE LSND joint probability is 2%.



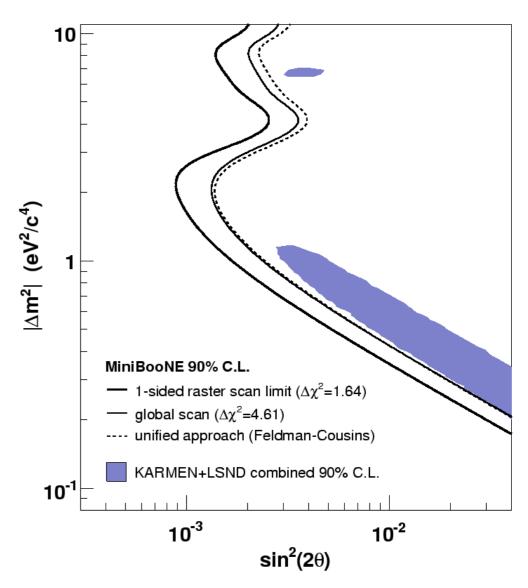
# **Results: Interpreting Our Limit**

There are various ways to present limits:

- Single sided raster scan (historically used, presented here)
- Global scan
- Unified approach (most recent method)

This result must be folded into an LSND-Karmen joint analysis.

Church, et al., PRD 66, 013001



We will present a full joint analysis soon.

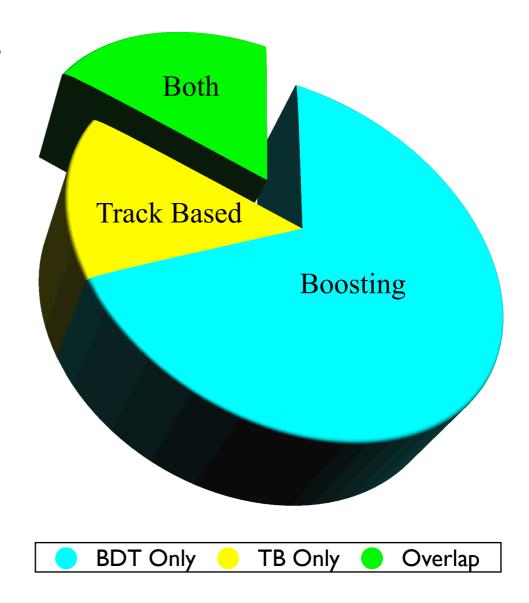
## **Results: Event Overlap**

#### Counting experiment numbers:

Track Based Algorithm finds 380 events Boosting Algorithm finds 971 events

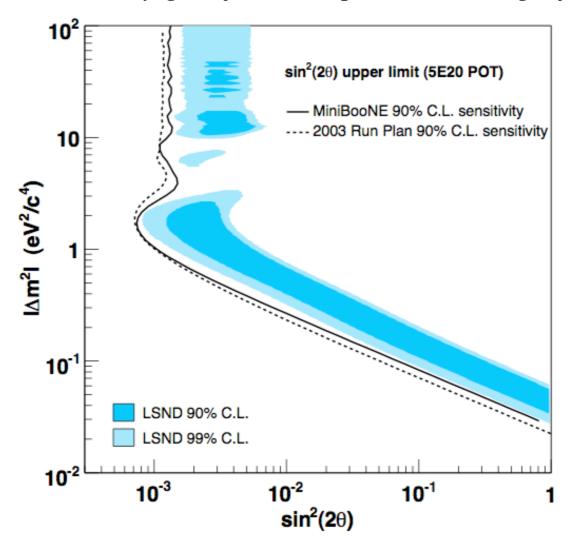
However, only 1131 events total, because 220 overlap

- chosen by both algorithms!



# **Results: Sensitivity Goal**

Compared to our sensitivity goal for 5E20 protons on target from 2003 Run Plan



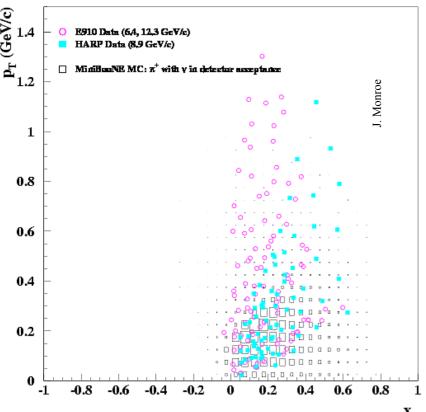
Set using  $\Delta \chi^2 = 1.64$  @ 90% CL

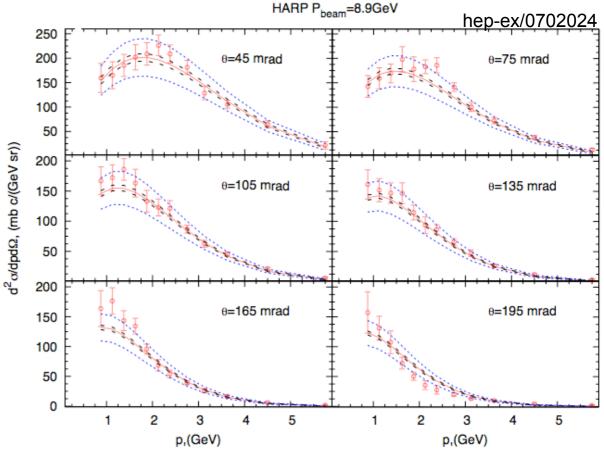
# **Booster Neutrino Beam: Modeling Meson Production**

**Prediction** from a fit to p Be  $\rightarrow \pi^+ X$  production data from E910 and HARP experiments ( $p_p = 6\text{-}12 \text{ GeV/c}$ ,  $\Theta_p = 0 \text{-} 330 \text{ mrad.}$ )

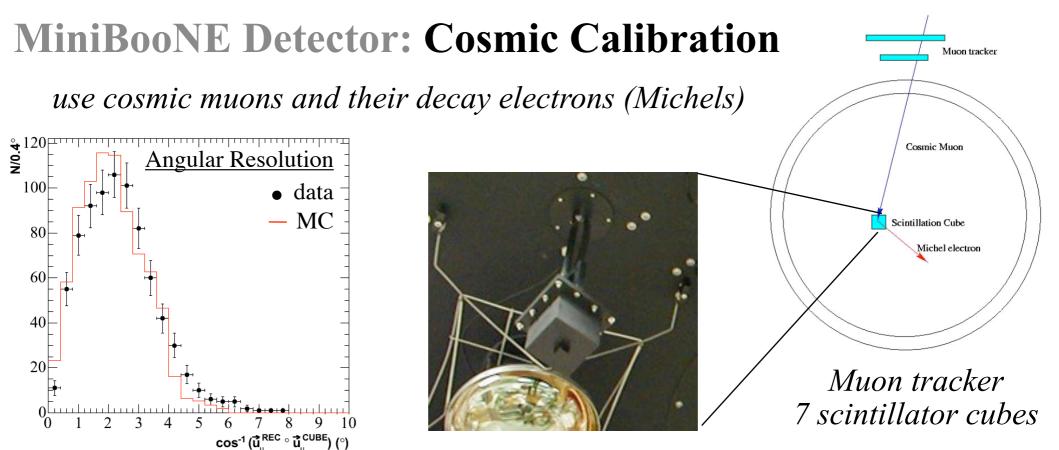
**Fit** (shown at right) uses Sanford-Wang parametrisation

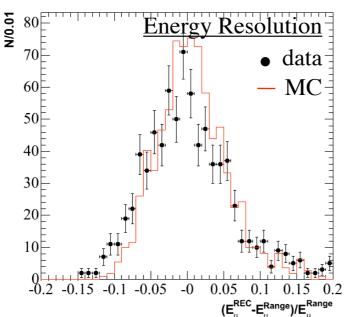
HARP has excellent phase space coverage for MiniBooNE





π<sup>-</sup> similarly parametrised
Kaons flux predictions use a Feynman Scaling
parametrisation (no HARP data)





### Cosmic muons which stop in cubes:

- -test energy scale extrapolation up to 800 MeV
- measure energy, angle resolution
- compare data and MC

Muon tracker + cube calibration data continuously acquired at 1 Hz

Analysis Strategy: Delta Background

 $\nu$  induced interactions that produce single  $\gamma$ s in the final state

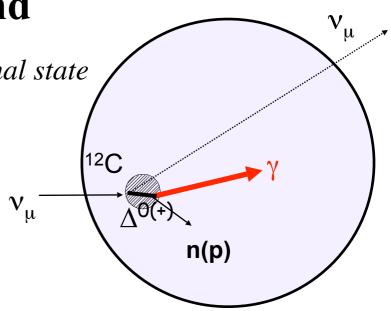
### Radiative Delta Decay (NC)

(i) Use  $\pi^0$  events to measure rate of NC  $\Delta$  production

- (ii) Use PDG branching ratio for radiative decay
  - 15% uncertainty on branching ratio

### Inner Bremsstrahlung (CC)

- (i) Hard photon released from neutrino interaction vertex
- (ii) Use events where the  $\mu$  is tagged by the decay e
  - study misidentification using BDT algorithm.

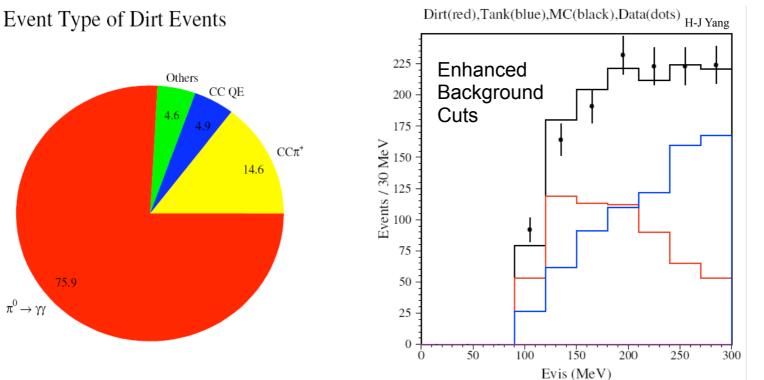


**Analysis Strategy: External Backgrounds** 

interactions outside the detector that deposit energy in the fiducial volume and pass the veto PMT hits cut

#### 1. "Dirt" Events

v interactions outside of the detector are measured in the "dirt box:"  $N_{data}/N_{MC} = 0.99 \pm 0.15$ 



2. Cosmic Ray Background Events

Measured from 126E6 strobe data triggers:  $2.1 \pm 0.5$  events.

MO Wascko, Imperial HEP Seminar April 12, 2007

n(p)

# **Analysis Overview: Background Summary**

Summary of predicted backgrounds for the primary MiniBooNE result (Track-Based Analysis):

Process	Number of Events
$\nu_{\mu}$ CCQE	10
$ u_{\mu}e ightarrow u_{\mu}e$	7
Miscellaneous $\nu_{\mu}$ Events	13
$NC \pi^0$	62
$NC \Delta \rightarrow N\gamma$	20
NC Coherent & Radiative $\gamma$	< 1
Dirt Events	17
$\nu_e$ from $\mu$ Decay	132
$\nu_e$ from $K^+$ Decay	71
$\nu_e$ from $K_L^0$ Decay	23
$\nu_e$ from $\pi$ Decay	3
Total Background	358
$0.26\% \nu_{\mu} \rightarrow \nu_{e}$	(example signal) <sup>163</sup>