



Commissioning the CMS micro-strip tracker

R.Bainbridge



- Introduction
 - CMS and LHC
 - Requirements for tracker design
 - Micro-strip tracker readout system
- Commissioning and operations
 - Magnet Test Cosmic Challenge
 - Tracker Integration Facility
 - Tracker commissioning at P5
 - Global runs and cosmic data-taking



- LHC to probe physics at TeV scale
 - Requires high CoM energies
 - Requires large integrated luminosities
- QCD-dominated environment, so focus on:
 - Good muon id and momentum resolution
 - Good e/γ energy resolution, isolation, pi^0 rejection
 - Tracker: good charged particle momentum resolution and reconstruction efficiencies, vertexing, efficient tau id and b-tagging







LHC environment

- Compared with Tevatron
 - \sim 7x increase in CoM energy
 - \sim 5x higher BX frequency
 - \sim 20x increase in luminosity
- High multiplicity environment
 - 20 super-imposed events per 25 ns
 - 1000 charged particles per 25 ns
- Consequences for (inner) detector design
 - Require rad-hard sensors and front-end ASICs
 - Need highly granular detectors for low occupancies
 - Need fast signal processing for BX id
 - Precision vertexing to identify signal event and decays

beam parameter	p-p	p-p	Pb-Pb	units
	high	low	high	
luminosity	10^{34}	2×10^{33}	10^{27}	$\mathrm{cm}^{-2}\mathrm{s}^{-1}$
energy per nucleon	7	7	2.76	TeV
bunch spacing	25	25	99.8	\mathbf{ns}
particles per bunch	1.15×10^{11}	4×10^{10}	7.0×10^7	-
inelastic collisions per crossing	17.5	3.5	7	-







Design considerations for CMS tracker

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- Full silicon design for micro-strip
 - MSGC not sufficiently rad-hard
 - More than 200 m2 active surface area
 - 10 (11) layers in Barrel (End-Cap)
 - Coverage up to |eta| < 2.4
- Requirements (constraints)
 - Reliability (inaccessible for 10 yrs)
 - Radiation-hard sensors and ASICs
 - Many layers (robust pattern recognition)
 - Low noise (small signals in Si)
 - Low material budget (Bremsstrahlung, photon conversions, multiple scattering)
 - Low power (less power cables, cooling)
 - Fast signal processing (BX id)
 - High spatial precision (vertexing)
 - High granularity (low occupancy)



region	layers	radius	modules	thickness	mean pitch	resolution
						$r - \phi \times z$
		m		$\mu{ m m}$	$\mu\mathrm{m}$	$\mu m \times \mu m$
TIB	4	0.2 - 0.55	2724	320	81/118	$23-34 \times 230$
TOB	6	0.55 - 1.16	5208	500	81/183	$35-52 \times 530$
TID	3	0.55 - 1.20	816	320	97/128/143	$23-34 \times 230$
TEC	9	1.20 - 2.80	6400	320/500	96-183	$35-52 \times 530$

Pixel system

Resolution: 10 (20) µm for r-ø (z) Barrel: r=4.4, 7.3, 10.2 cm EndCap: r=34, 46.5 cm 100 x 150 µm pixels 66M channels







Control and readout (micro-strip tracker)



11

OM + ASP

7.1 x 8.1 mm

192-cell

pipeline

memory

- 128-channel design
 - IC and RAL design using 0.25 μm CMOS commercial process
 - BX id as signal constrained to single 25 ns time window
 - Data frame contains BX info and 128 analogue samples
- Intrinsically radiation-hard
 - Thin transistor gate oxides, enclosed transistor geometries
 - 10 Mrad dose: no significant effect on performances





- Front-End Driver "receiver card"
 - VME64 form factor
 - RAL and Imperial

• Functions

- Optical/electrical conversion
- A/D conv. (10bit @ 40 MHz)
- Zero-suppression
- Data formatting

Scale

- Input from 192 APV25 chips
- (25k detector strips)
- 440 FEDs for µstrip tracker
- (~700 for CMS)

• Data rates

- Input: 3.4 GB/s
- Output: ~200 MB/s









Production testing of FED

Optical

outputs

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OHS

System

& VME

FPGAs

- "FED tester" used for performance and production tests
 - Provides D/A and E/O conversions for known input
 - FED output can be checked against known input signals





- Several areas of study on the tracker readout system are not covered in this talk
 - Radiation hardness qualification on the APV25
 - Single event upset studies on the APV25
 - Effect of highly ionising events on the tracker readout system
 - APV25 emulator (prevents pipeline overflows in APV25)
 - Performance studies on the APV25, optical link system and FED
 - And more...
- Worthy of a future seminar...

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Magnet Test and Cosmic Challenge (MTCC)

- First CMS "slice test", during summer 2006
- Magnet test
 - Successful ramping of of magnet to 4 Tesla
 - Tests of service infrastructure and field mapping
- Cosmic challenge
 - Commission parts of all sub-dets (no pixels)
 - DAQ tests and global running







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14



- Successful operation with central DAQ
- Excellent and stable S/N performance
- No operational issues with B-field

B-field [T]	Runs	Events	"Pointing"
0.0	43	11M	240k
3.8	38	14M	250k
4.0	4	2M	31k



"3-week shift" for tracker DAQ experts!





Tracker Integration Facility

- Environmentally controlled clean-room
- Infrastructure to operate 1.6M channels
- Large cooling capacity (tests down to -15°C)
- Tracker Analysis Centre in close proximity
- Tracker assembly completed at TIF in March 2007







Tracker sub-systems at the TIF



- Extensive detector commissioning during spring
 - Commissioned ${\sim}15\%$ of the tracker
 - Services and infrastructure (power, cooling, racks)
 - Development of software tools and procedures
- Data-taking with cosmic triggers up to July 2007

Sub-system	Modules	Fraction [%]
TIB+	640	36
TOB+	720	28
TEC+	800	25





Large-scale tests using DAQ and DCS





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Cosmic triggers at the TIF



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Signal and noise performance





Installation within CMS at P5





- Aim: synchronous operation with CMS, maximize S/N, minimize occupancies
 - Procedures are usually done in "stand-alone" mode using local DAQ prior to CMS global runs
- Connectivity
 - Test communication with front-end devices via control system
 - Determine readout fibre connections b/w front-end and off-detector FEDs (and flag any problems)
 - Establish HV/LV power supply connections to front-end devices and sensors
- Tracker calibration
 - Internal synchronization (guarantee synchronous sampling at the front-end)
 - Tune optical link system (optimise dynamic range and gain, calibrate)
 - Tune configuration of front-end ASICs (operational modes, biasing, gain settings...)
 - Tune and calibrate APV25 pulse shaping (maximise S/N, minimise "fakes" in adjacent BX)
 - Determine FED calibration constants (pedestals and noise constants, flag dead and noisy channels)

• Commission with CMS

- Synchronize with global trigger (latency scans to "time in" tracker system)
- Other commissioning activities concerning services infrastructure and systems
 - Power Supplies, Detector Control and Safety System, VME electronics (trigger, control, readout)
 - DAQ software, databases, procedures







Synchronisation of APV25 chips





Tuning of readout optical link system

- Tune linear laser driver on front-end module
 - Four gain settings, target gain is 0.8 $\left[\text{V/V}\right]$
 - Tune bias setting to give max dynamic range
- Procedure allows to identify/measure:
 - Optimum gain, laser thresholds...
 - Best gain/bias settings







Summary of commissioning results

- Modules or fibres are disabled if:
 - LV problems
 - Unable to configure device(s)
 - Unable to lock onto clock
 - Faulty fibre ("low light" or broken)
 - Some can be recovered
- \bullet Further ${\sim}1\%$ modules exhibit noise problems
 - Largely due to power supply problems and tuning
 - Can be solved for most modules



Working system for CMS global runs, as defined by commissioning procedures									
Sub-system		Front-end modu	les	Fibre co	nnections	Active [%] (max)			
	Total	Responsive	Disabled	Total	Missing				
Outer Barrel	5208	5196	4	12043	19	99.5 (99.9)			
Inner Barrel	3540	3481	0	9012	14	98.2 (99.3)			
End-Cap +	3200	3183	6	7506	4	99.4 (99.9)			
End-Cap -	3200	3192	2	7535	9	99.6 (99.8)			

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Several flavours of global runs

- CRUZET (Cosmic RUn at Zero Telsa),
- CRAFT (Cosmic Run At Four Tesla)
- MWGR (Mid-Week Global Run)
- Private runs (eg, tracker/muon synchronisation)

• Tracker participation

- CRUZET3 (mid-July, ~week)
- CRUZET4 (late-August, ~week)
- 95% "uptime" for CRAFT (Oct/Nov, 3 weeks)
- Highest average L1 trigger rate: 550 Hz
- Tracks: few Hz with DT top or bottom



Global run type	B-f	field
	off	on
CRUZET1 (no TK)	40M	-
CRUZET2 (no TK)	32M	-
CRUZET3	68M	-
CRUZET4	61M	-
MWGR	40M	-
Commissioning	76M	-
CRAFT	39M	290M

Datasets and statistics for CRAFT						
Dataset	DBS Catalogue Name					
RAW	/Cosmics/Commissioning08-v1/RAW					
RECO	/Cosmics/Commissioning08-PromptReco-v2/RECO					
AICaReco 1	/Cosmics/Commissioning08_CRAFT_V2P_StreamALCARECOTkAlCosmics0T_v7/ALCARECO					
AlCaReco 2	/Cosmics/Commissioning08_CRAFT_V3P_StreamALCARECOTkAlCosmics0T_v11/ALCARECO					
Pointing	/Cosmics/Commissioning08_CRAFT_V1P_TrackerPointing_v3/RECO					
SuperPointing	/Cosmics/Commissioning08_CRAFT_V3P_SuperPointing_v4/RECO	664326				



- Tracking algorithm: 3 steps
 - Seed creation: first estimate of parameters
 - Pattern recognition: associates hits to a track
 - Fitting procedure: best estimate of track parameters
 - (Only fitting procedure is common to all algorithms)
- Presently using three tracking algorithms
 - Combinatorial Track Finder (CTF, default)
 - Road Search (RS)
 - Cosmic Tracker Finder (CoTF)
 - (CTF/RS designed for pp collisions, CoTF for cosmics)





First tracks during CRUZET and CRAFT





Detector coverage

- Distribution of "on-track" clusters within barrel
 - Example track shown opposite
 - Hit density maps shown for three different tracking algorithms
 - Data from CRUZET3 runs (no B-field)

S/N	for on-track clusters	
	Expected S/N	Measured S/N
TIB	24	15
TIB	28	29





CMS

Track statistics and momentum spectrum



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- Two algorithms used
 - Millipede (layer-level alignment followed by module-level)
 - HIP (moduel-level alignment with survey constraint)
- Each plot (TIB,TOB,TEC) shows four curves
 - Residuals based on survey position
 - Based on alignment parameters from CRUZET data
 - Based on CRAFT data for both algos







- Plots contain one entry per module
 - Red is using HIP on CRAFT data
 - Blue is using HIP on CRUZET4 data
- Significant improvement using CRAFT data
 - Improvement for pixel barrel too: 112 μm to 47 μm











- Tracker has extensive monitoring tools for inspecting event data and reco objects
 - Provides immediate on detector performance issues and is used to classify runs quality
 - Feedback often very useful for shifters and experts
 - Has already revealed several, subtle effects in during global runs some examples on next slides...





 Runs during the CRAFT data-taking period ClusterStoNCorr OnTrack landauPeak TIB тов TID 35 TEC 0.1 30 Mean number of 0.08 tracks per event 25 20 0.06 0.04 S/N for on-track clusters 0.02 5 0 66500 67000 67500 68000 68500 69000 69500 70000 70500 71000 Run number ______ 66500 67000 67500 68000 68500 69000 69500 70000 70500 71000 Run number NumberOfRecHitsPerTrack CKFTk_mean ClusterNoise OnTrack gaussMean TIB CosmicTk_mean тов RSTk mean 5 TID 14 TEC 4.5 12 10 3.5 8 3 6 Hits per tracks 2.5 Noise for on-track clusters 66500 67000 67500 68000 68500 69000 69500 70000 70500 71000 66500 67000 67500 68000 68500 69000 69500 70000 70500 71000 Run number Run number

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Front-end synchronization issues







- Number of digis per event is not constant
 - Digis are equivalent to FED zero-suppressed data
 - Average number is 1000 c.f. 10M channels (so normally 10⁻⁴ effect)
 - (Can ignore contribution from cosmic muons very small)





- Noise in few TEC modules at high trigger rates
 - Small fraction of modules ($\sim 10^{-3}$)
 - Low peak in S/N distribution for on-track clusters
 - Present only for low p_t tracks ("loopers")
 - Appear to be susceptible to fake tracks in TEC
 - Source of the noise being investigated
 - Studies on-going, protection being prepared









• CMS tracker installed and commissioned in tracks with Pt>1GeV

- Fraction of working channels is >99% (will improve further)
- Excellent detector performance (noise, signal, stability)
- Synchronous operation with CMS during global runs
- Good alignment already achieved based on large track sample
- Detailed studies of global run data ongoing
- Eagerly anticipating first collisions...

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(Some of) the Imperial CMS group





