

# LATEST NEUTRINO-ARGON SCATTERING RESULTS FROM MICROBOONE

Kirsty Duffy, Fermi National Accelerator Laboratory

HEP Seminar, Imperial College London

10th March 2021



55 cm

Run 3469 Eve

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Most neutrino experiments rely on **understanding neutrino interactions with nuclei**, to reconstruct neutrino flavour and energy

The success of future liquid argon experiments relies on a good understanding of these effects on **argon nuclei**

MicroBooNE is the world's **only running liquid-argon experiment taking data in a neutrino beam** — uniquely placed to answer these questions

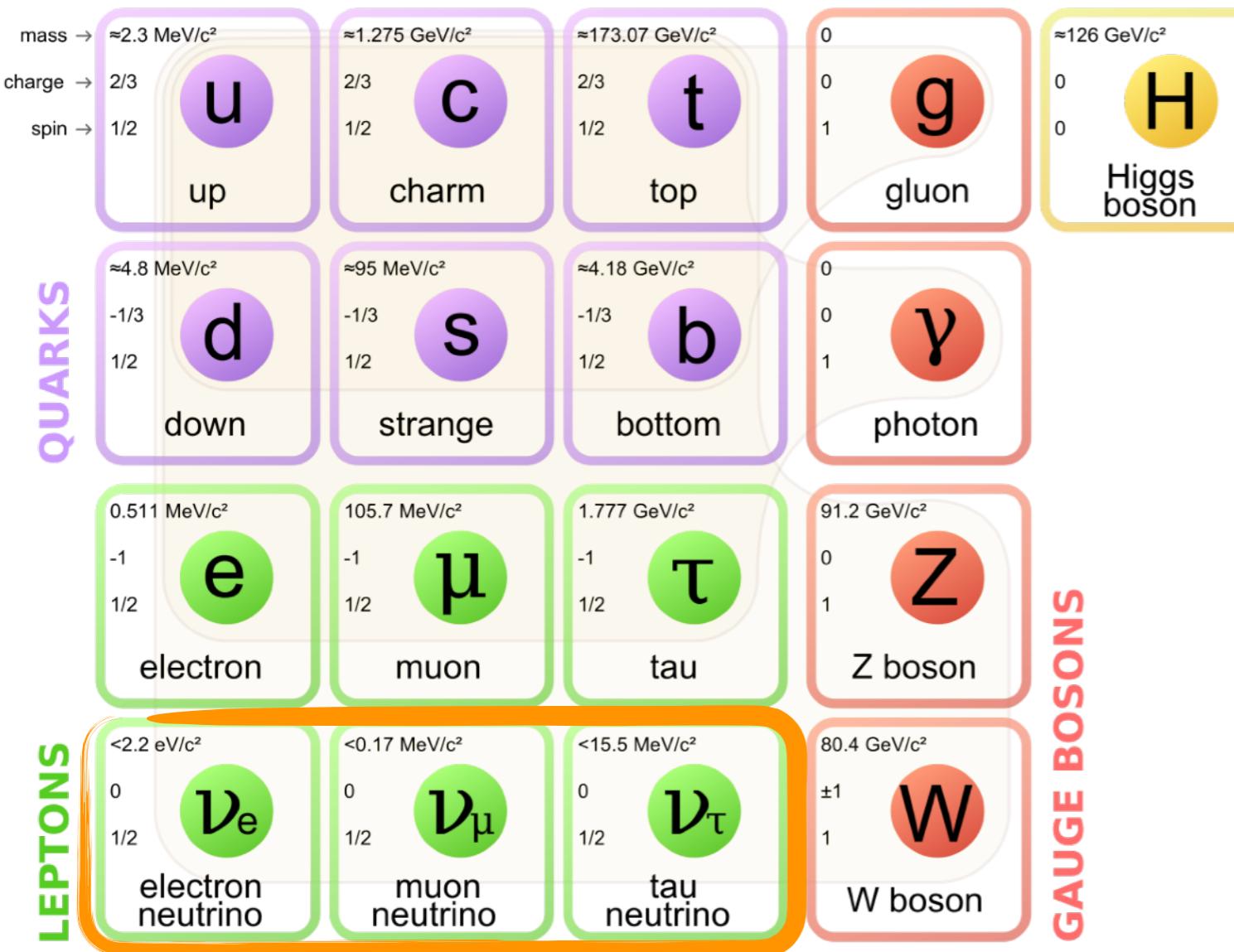
I will use “**neutrino interactions**” and “**neutrino cross sections**” interchangeably in this talk

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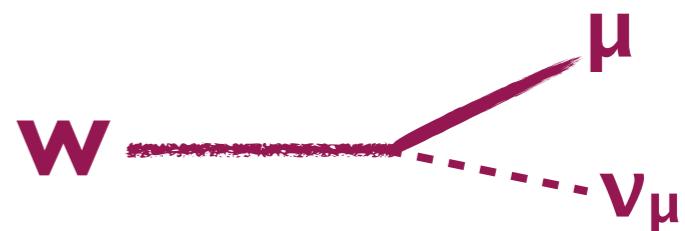
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# NEUTRINOS: WHAT WE KNOW

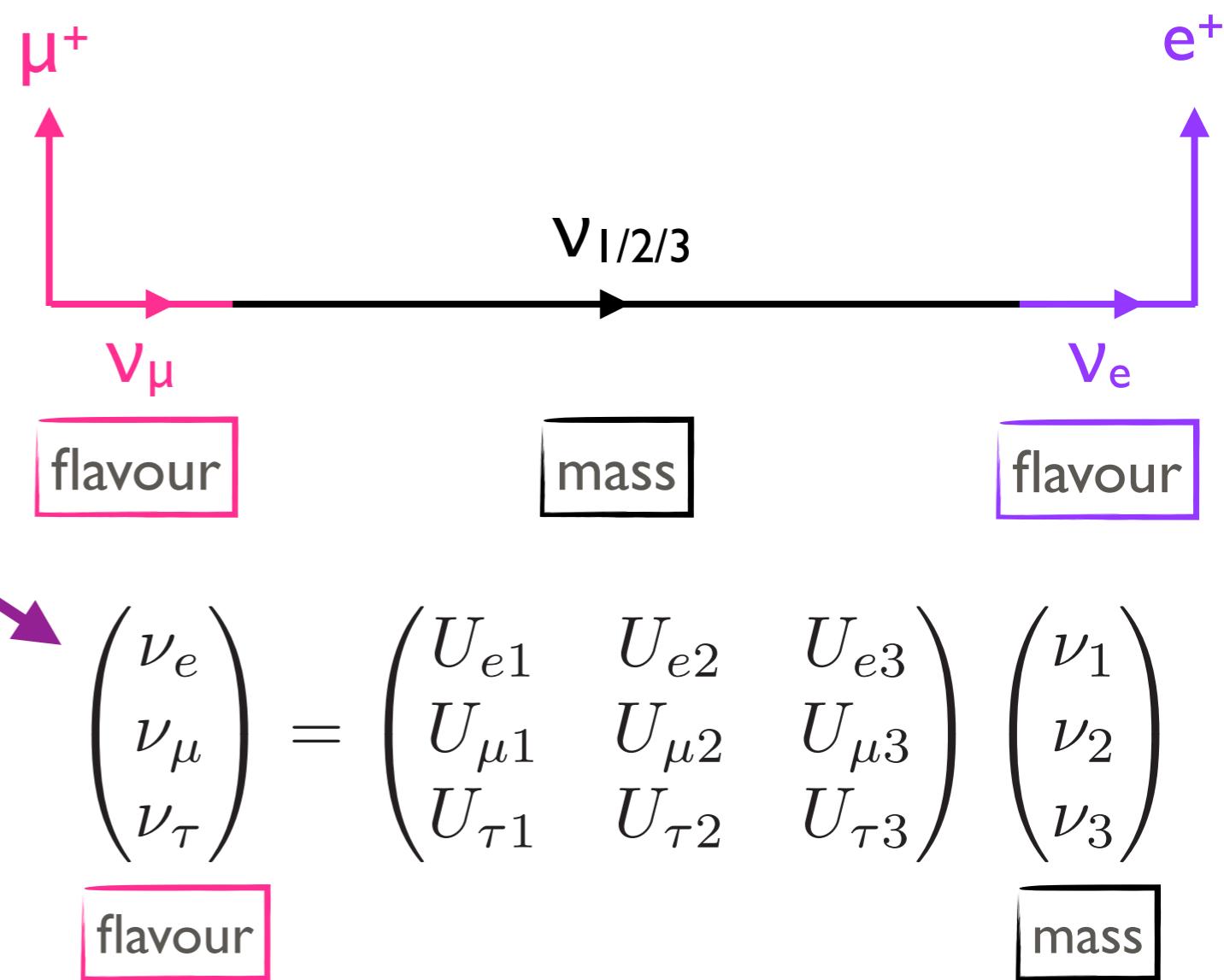


- Fundamental particles in the Standard Model
- Very small (non-zero) mass
- Interact via weak force
- “Paired” with charged leptons



# NEUTRINOS: WHAT WE KNOW

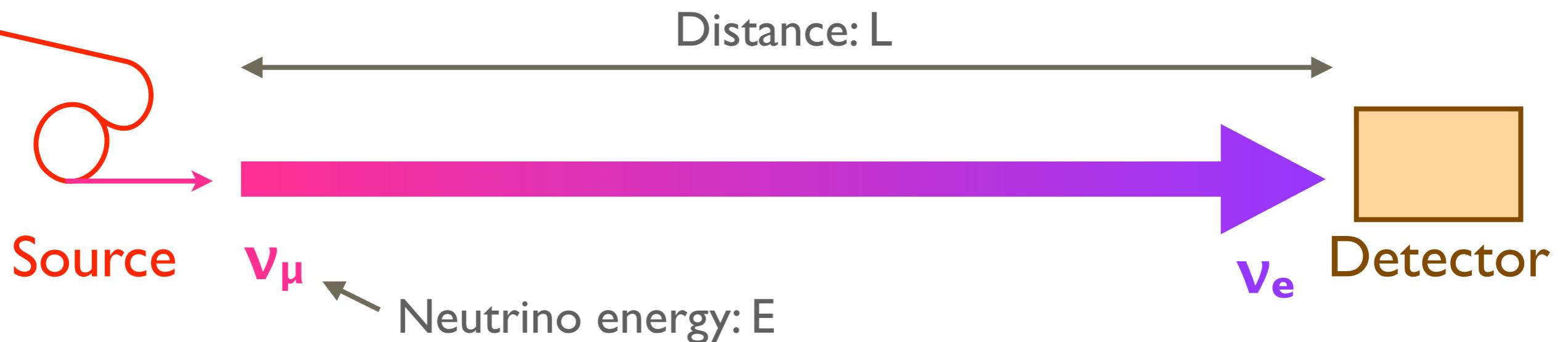
- Two sets of eigenstates: flavour (interaction) and mass (propagation)



- Related by **PMNS\*** matrix
  - 3 mixing angles  $\theta_{12}, \theta_{23}, \theta_{13}$
  - 1 CP-violating phase,  $\delta_{CP}$

- Allows for neutrino flavour change

\*Named after physicists Pontecorvo, Maki, Nakagawa, and Sakata

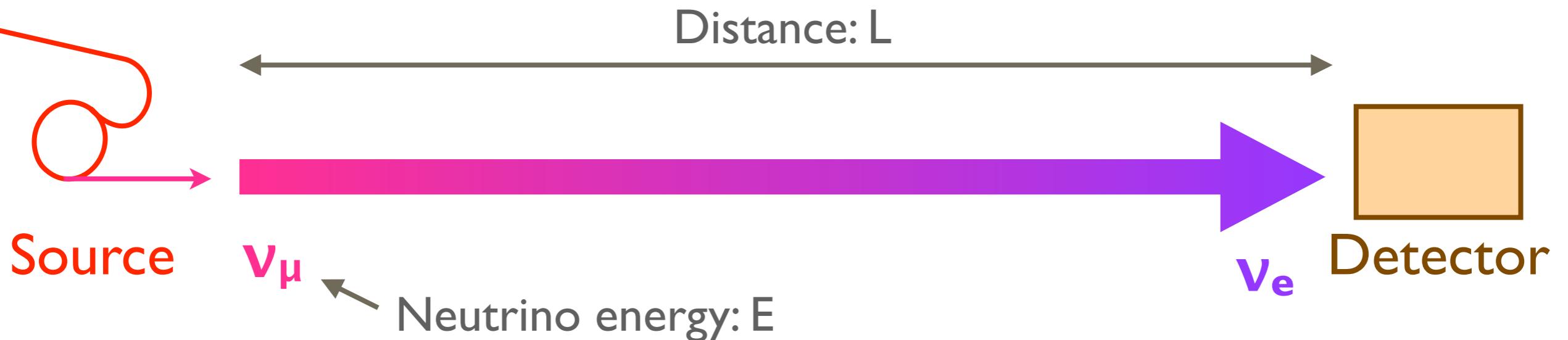


## Muon neutrino disappearance



## Electron neutrino appearance

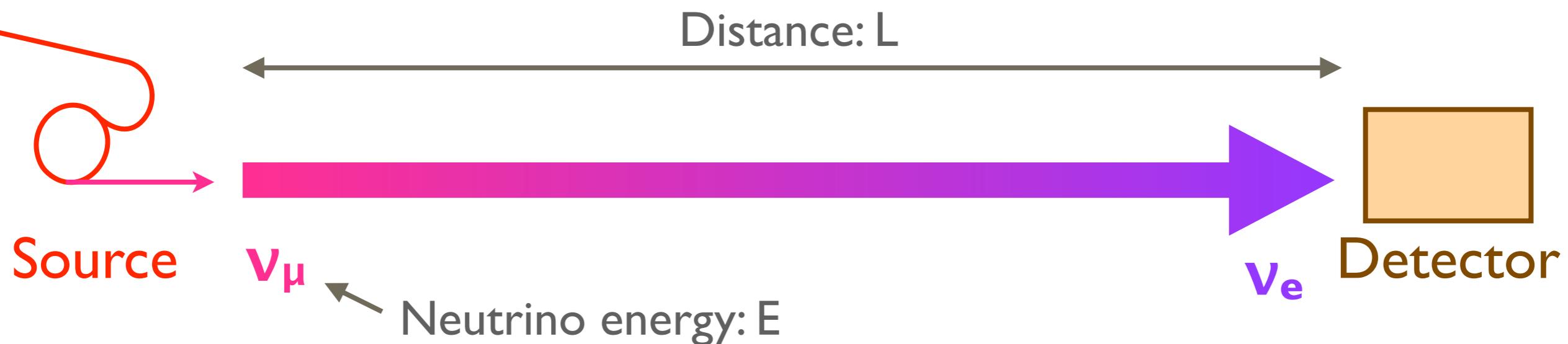




Probability to detect a neutrino of a given flavour **oscillates** as:

$$\sin^2 \left( \frac{\Delta m_{ij}^2 L}{4E} \right)$$

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$



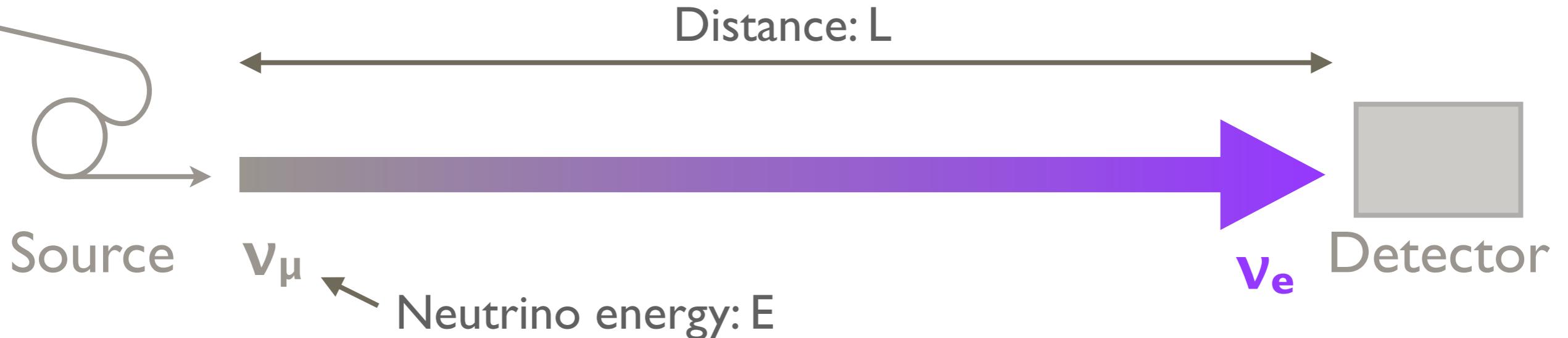
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## Reason number 1 why neutrinos are exciting:

- Neutrino oscillation
- Neutrinos have mass
- Physics beyond the Standard Model!



## Electron neutrino appearance

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \simeq \text{Something} + (\text{Something else} \times \sin \delta_{CP}) + (\text{CP-even, solar, matter effect terms})$$

$\bar{\nu}$   $\nu$   $\times$   $L$   $\sin \delta_{CP}$

## Reason number 2 why neutrinos are exciting:

CP violation in neutrinos might explain the matter-antimatter asymmetry we see in the universe

→ neutrinos could explain why we exist!

## Electron neutrino app

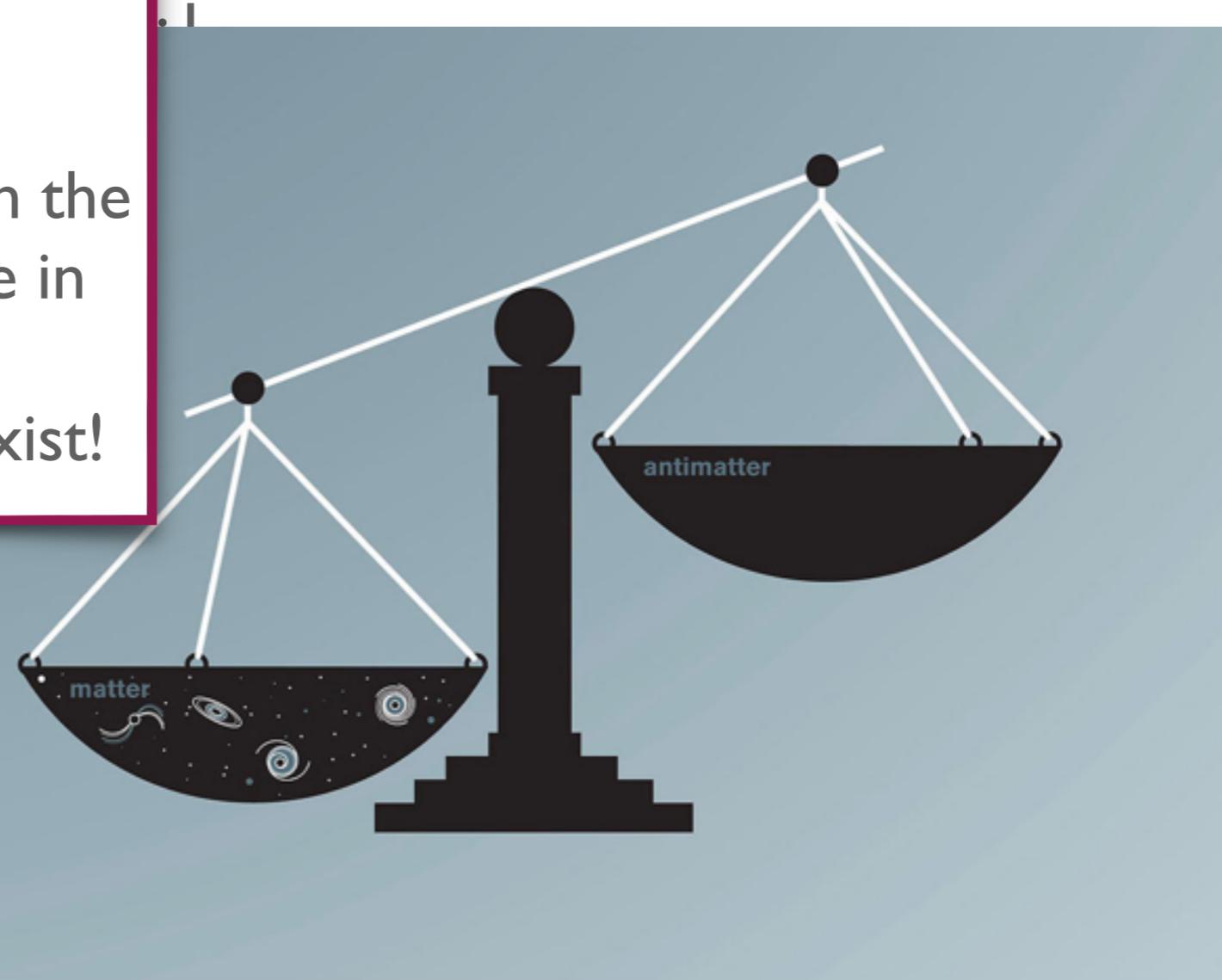
$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \simeq$$

$$\bar{\nu} \xrightarrow{(+)-} v$$

**Something  
else**

$$\times \sin \delta_{CP}$$

+ (CP-even, solar, matter effect terms)



# CURRENT MEASUREMENTS

**NOvA** and **T2K** experiments are measuring neutrino oscillations over long baselines

**Both have large uncertainties due to cross-section modeling**

Type of Uncertainty	$\nu_e/\bar{\nu}_e$ Candidate Relative Uncertainty (%)
Super-K Detector Model	1.5
Pion Final State Interaction and Rescattering Model	1.6
Neutrino Production and Interaction Model Constrained by ND280 Data	2.7
Electron Neutrino and Antineutrino Interaction Model	3.0
Nucleon Removal Energy in Interaction Model	3.7
Modeling of Neutral Current Interactions with Single $\gamma$ Production	1.5
Modeling of Other Neutral Current Interactions	0.2
Total Systematic Uncertainty	6.0

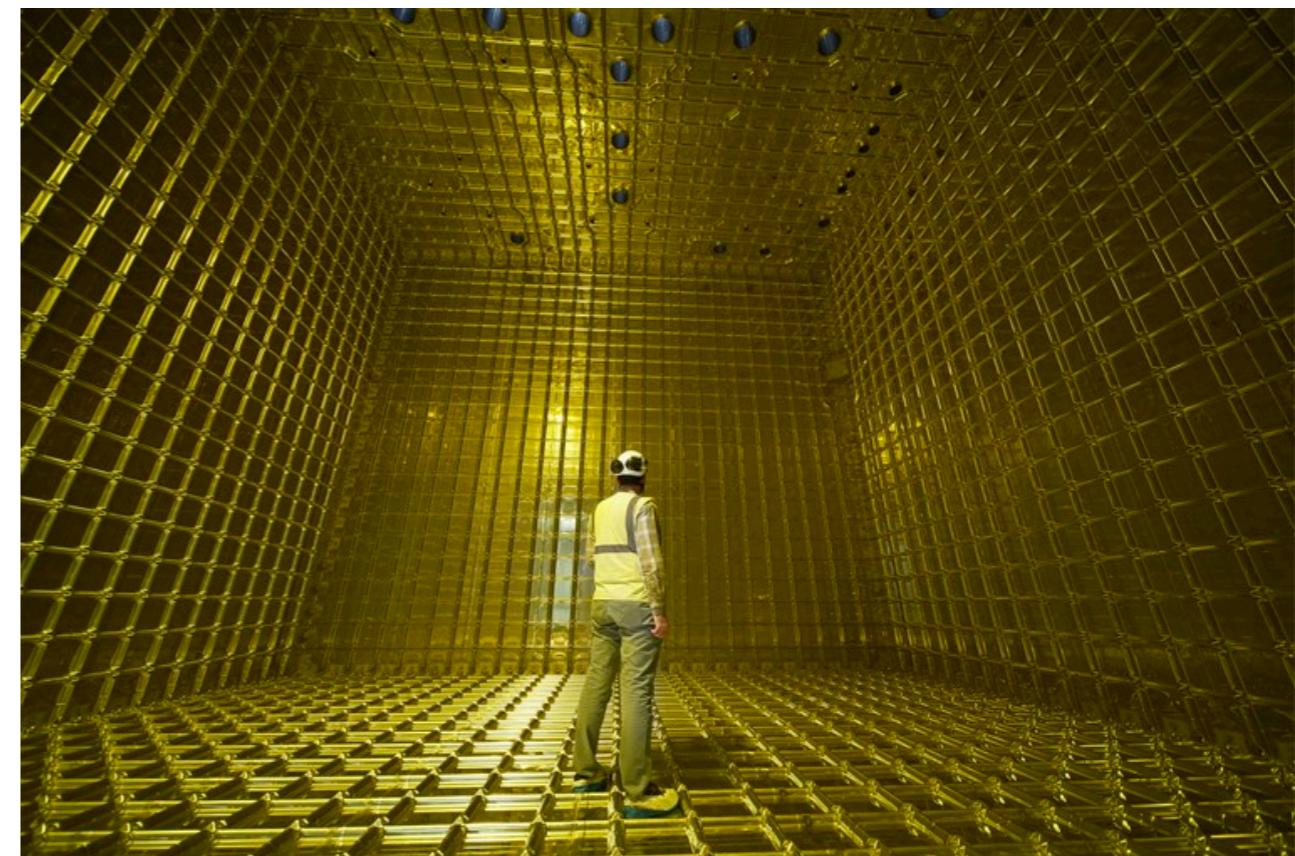
**T2K** Nature 580, 339-344 (2020)

# LOOKING FORWARD

Next-generation experiments **DUNE** and **Hyper-Kamiokande** will measure neutrino oscillations with unprecedented precision

## Deep Underground Neutrino Experiment

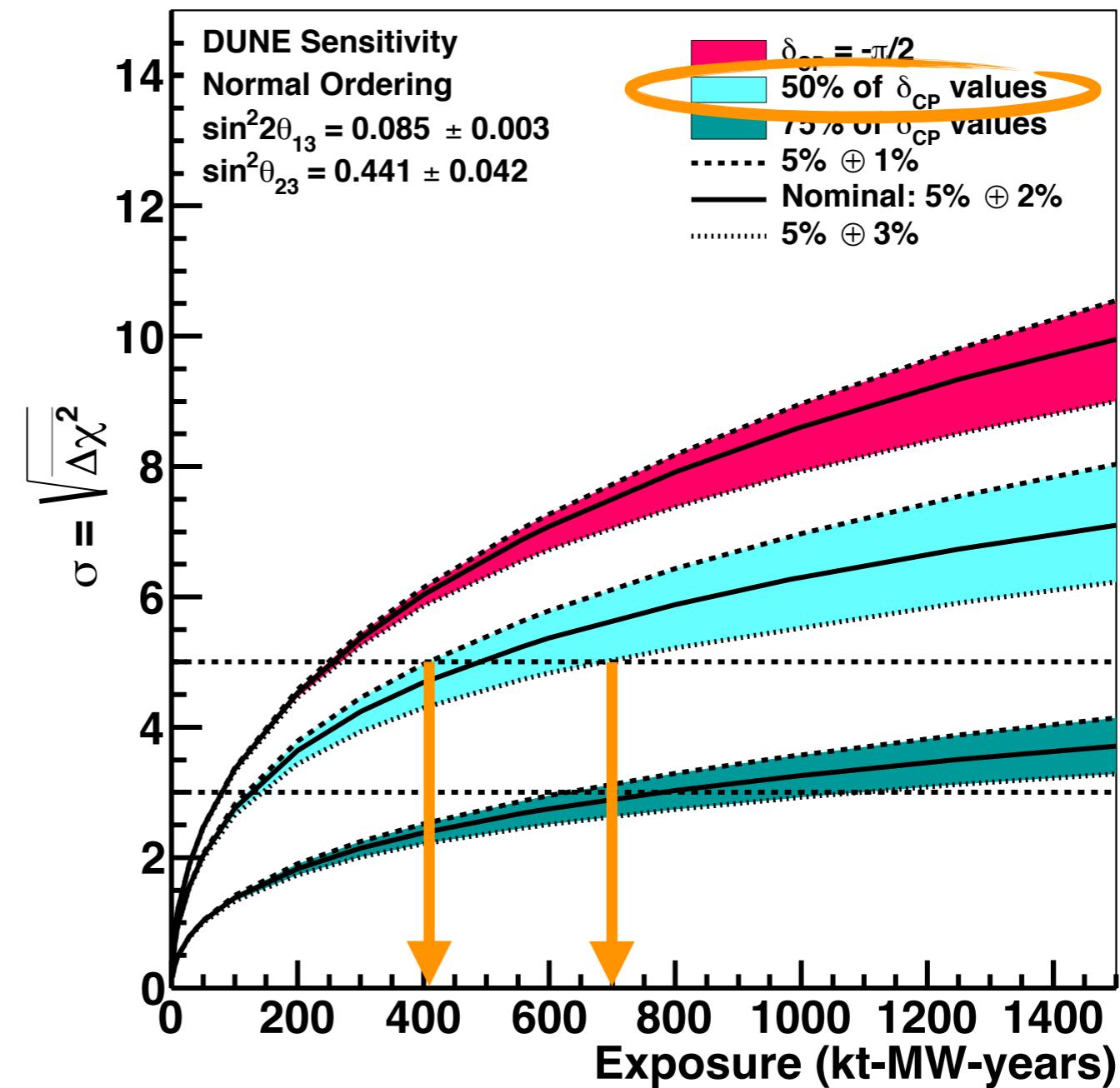
- 40kton **Liquid Argon**
- build on current and future US LAr program:
  - ArgoNeuT, LArIAT
  - **MicroBooNE**
  - ProtoDUNEs
  - SBND, ICARUS



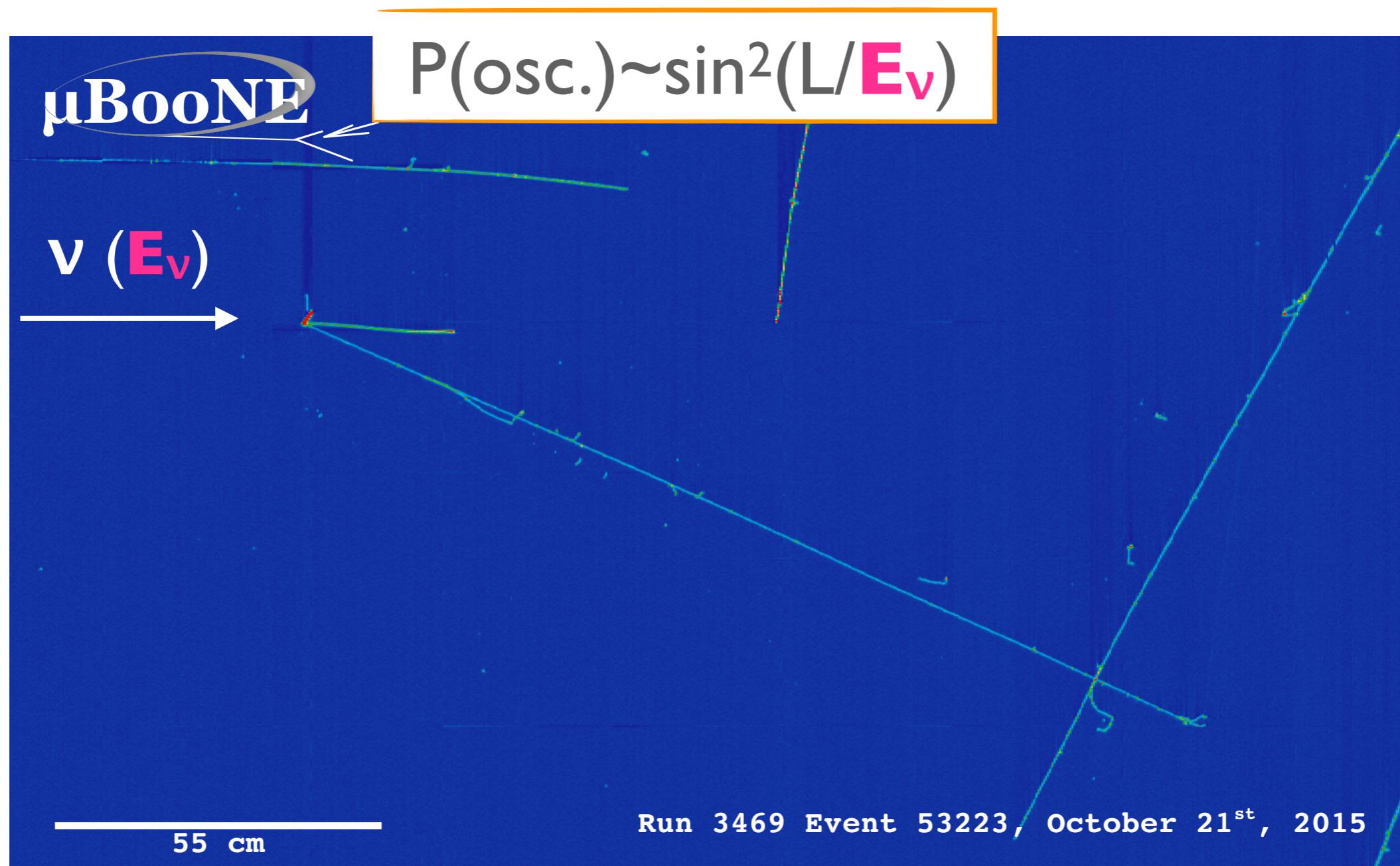
# LOOKING FORWARD

A **better understanding** of neutrino interactions is **vital for DUNE's success**

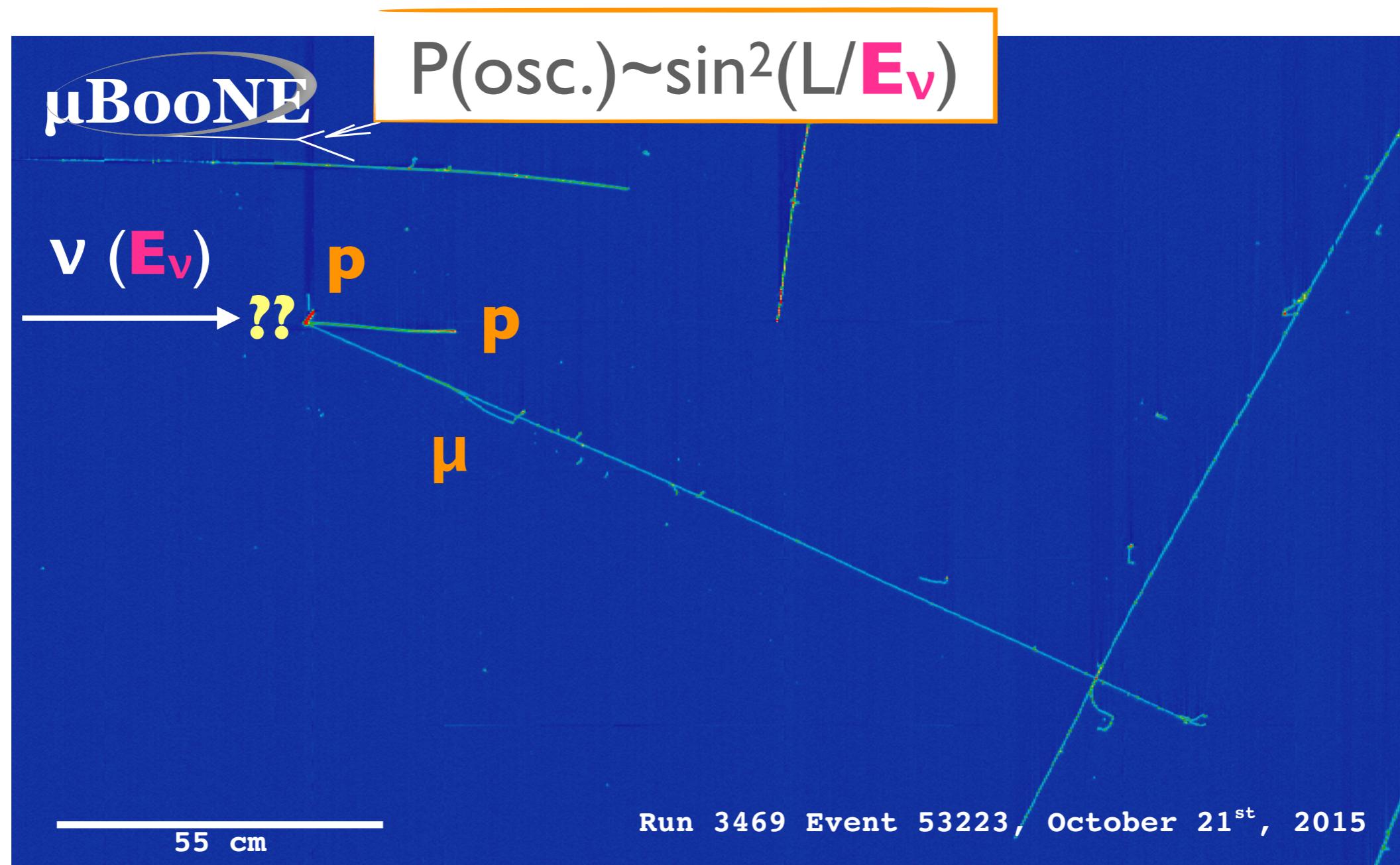
2% reduction in uncertainty  
 → 300 kt-MW-years  
 → ~6 years real time  
 → \$150m operations cost



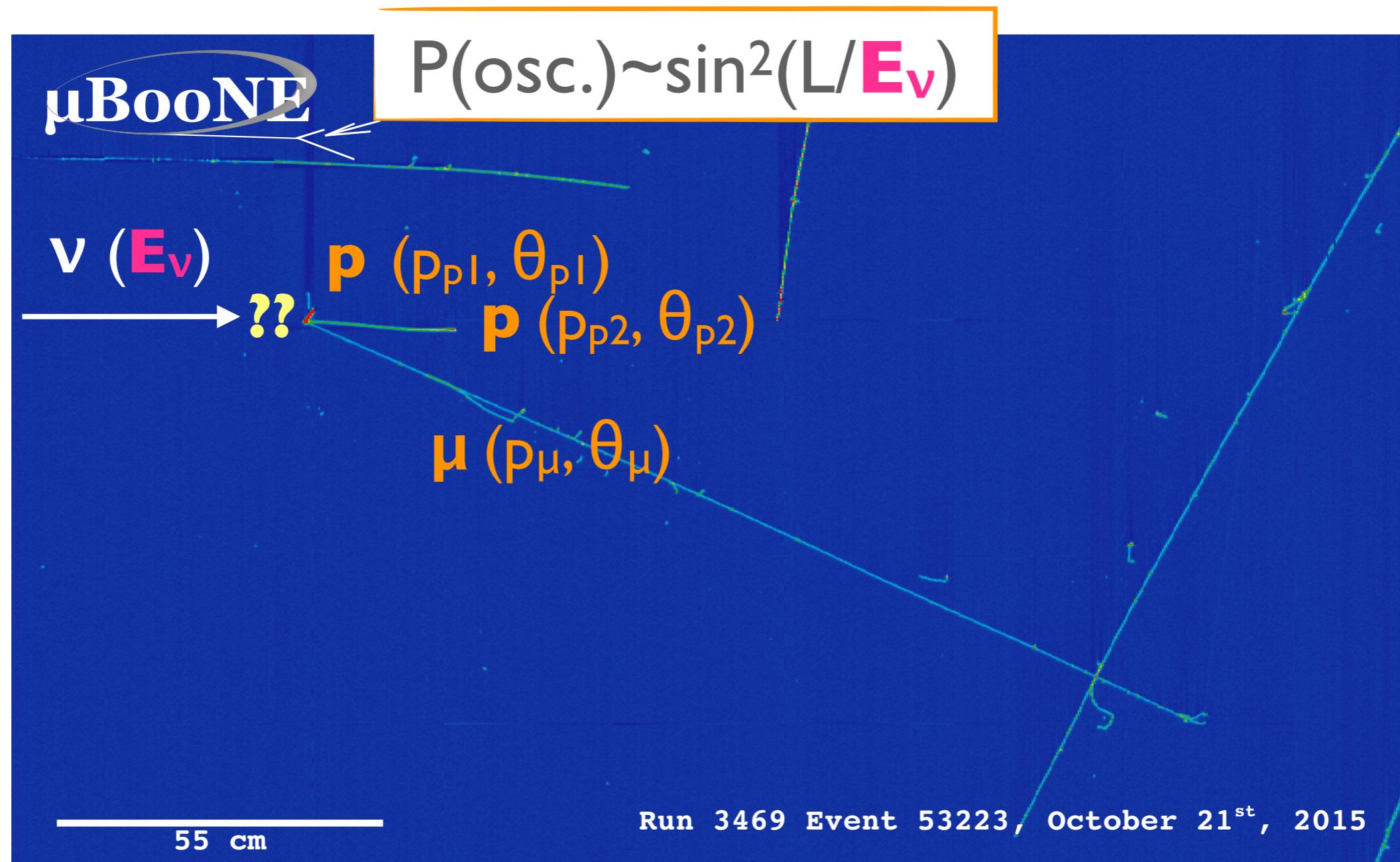
# NEUTRINO INTERACTIONS



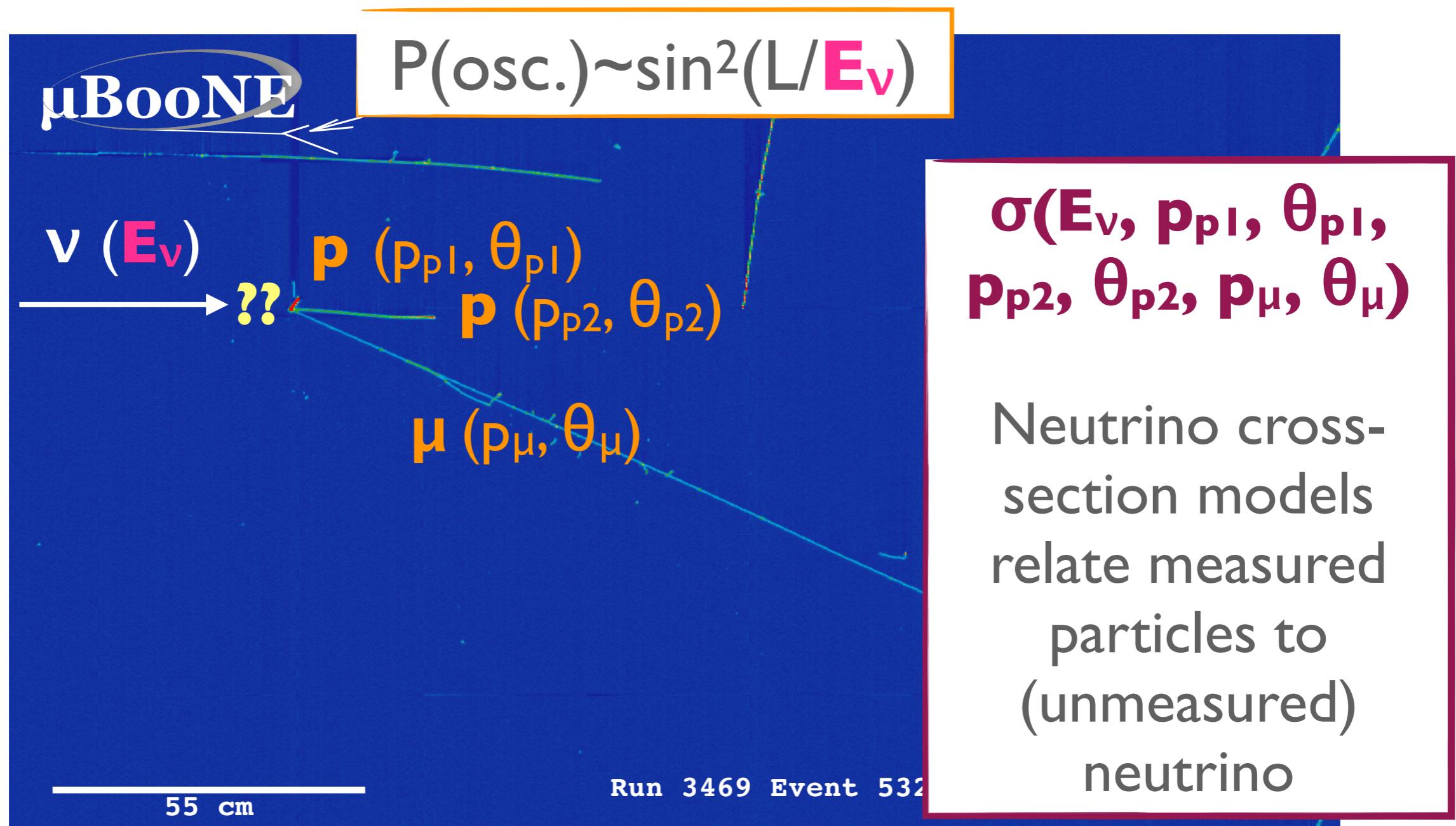
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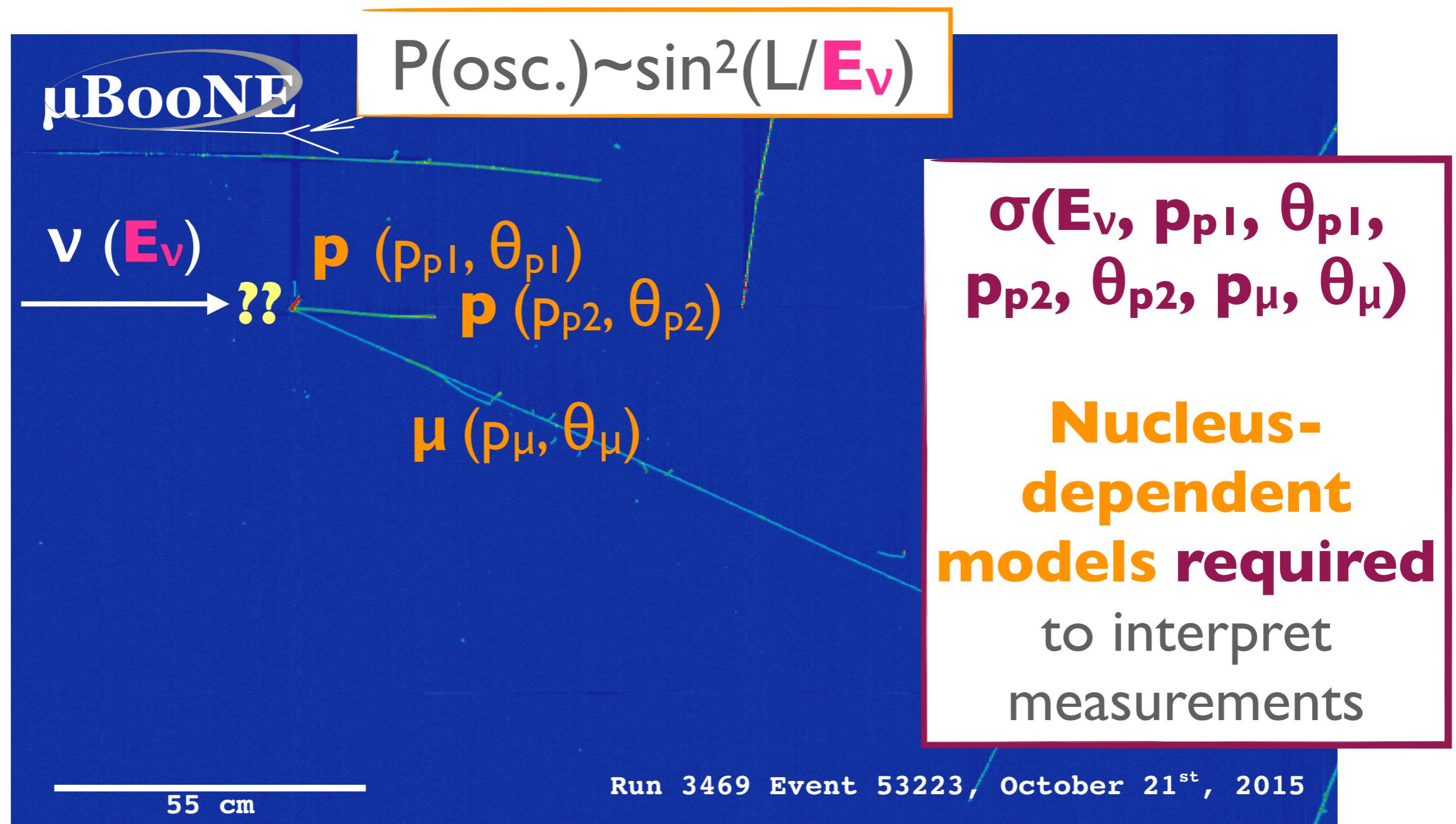
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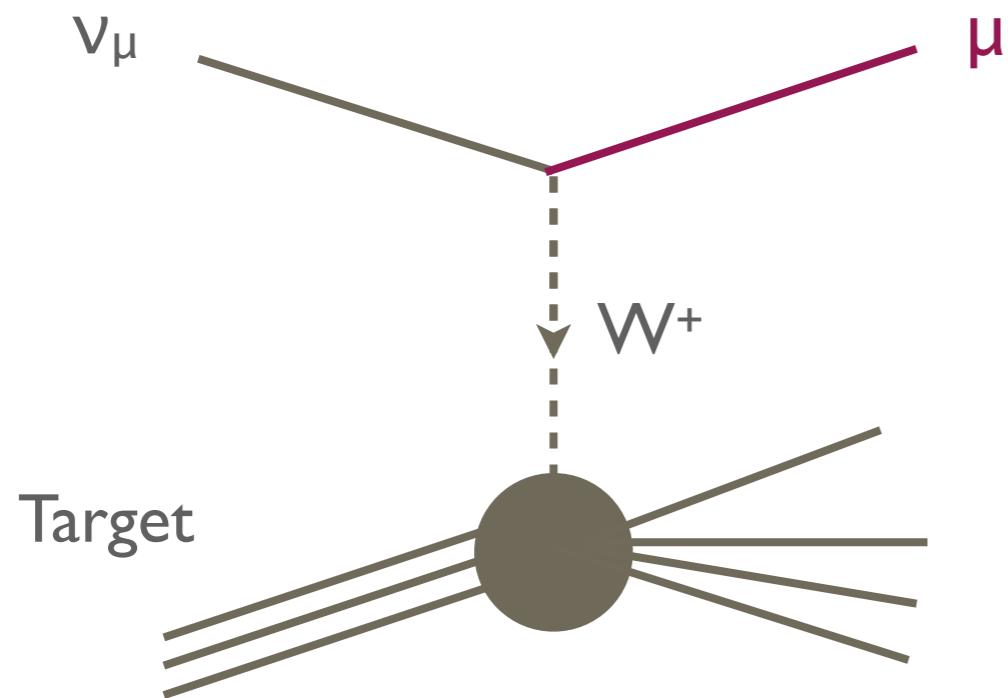
# NEUTRINO INTERACTIONS



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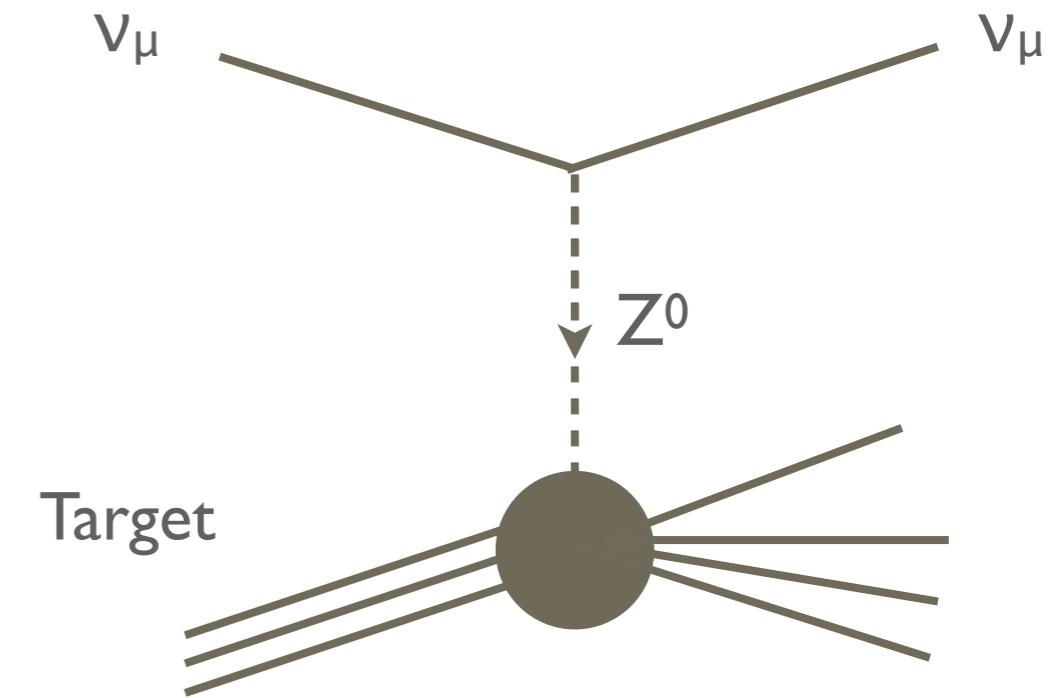


# NEUTRINO INTERACTIONS



## Charged-current

- Exchange of  $W$  boson
- Lepton produced with same flavour as  $\nu$



## Neutral-current

- Exchange of  $Z$  boson
- Independent of  $\nu$  flavour

# NEUTRINO INTERACTIONS

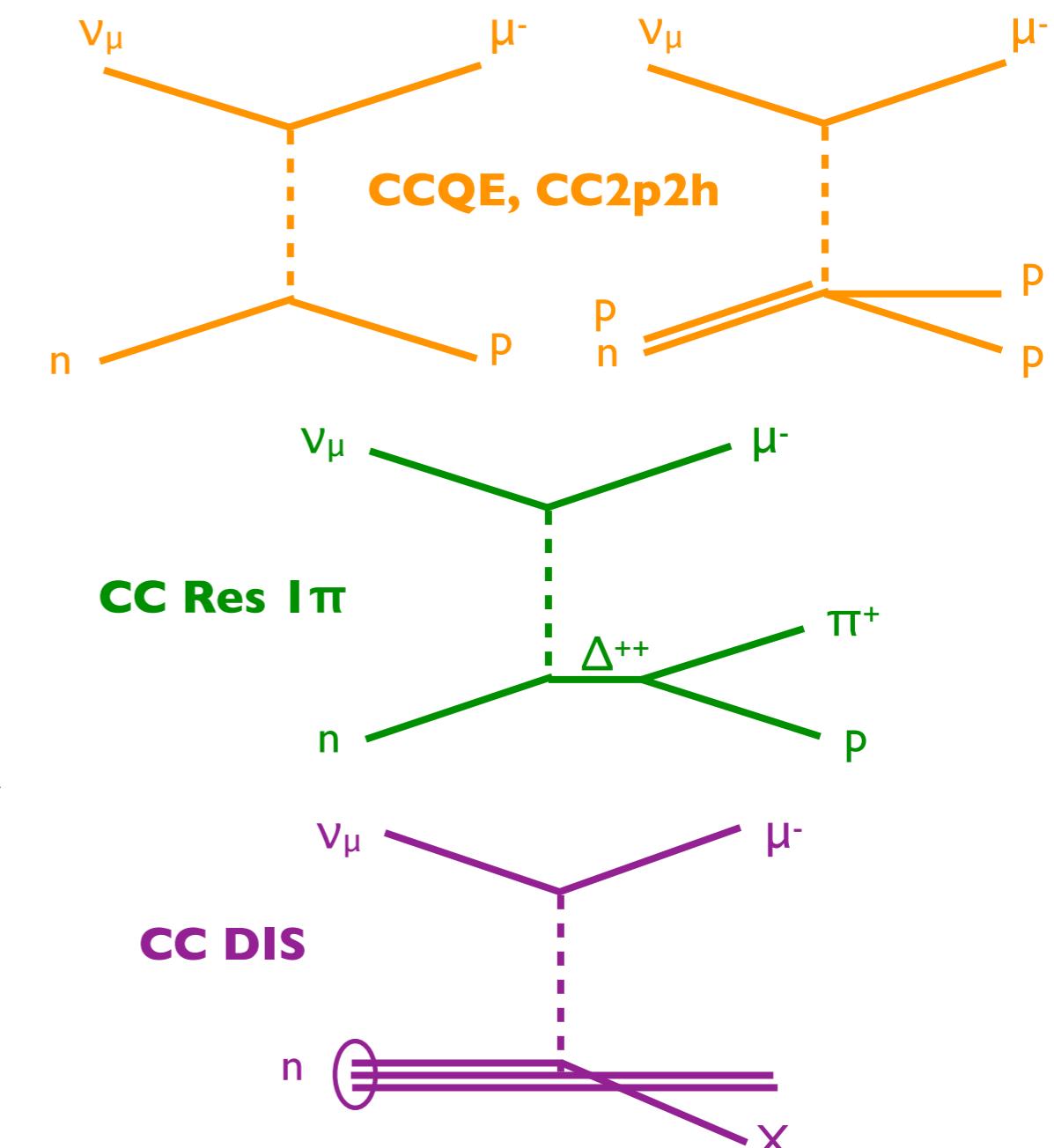
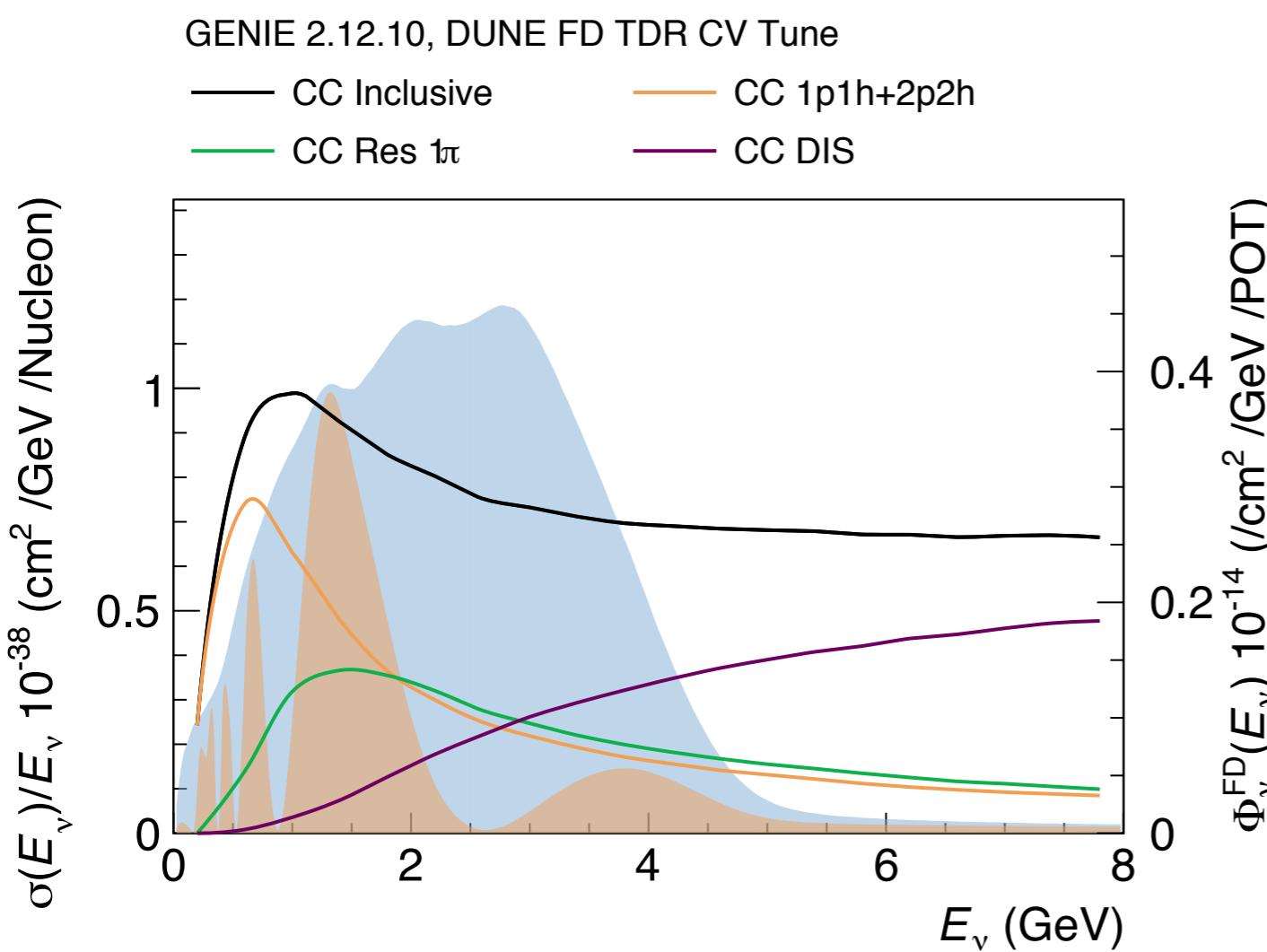
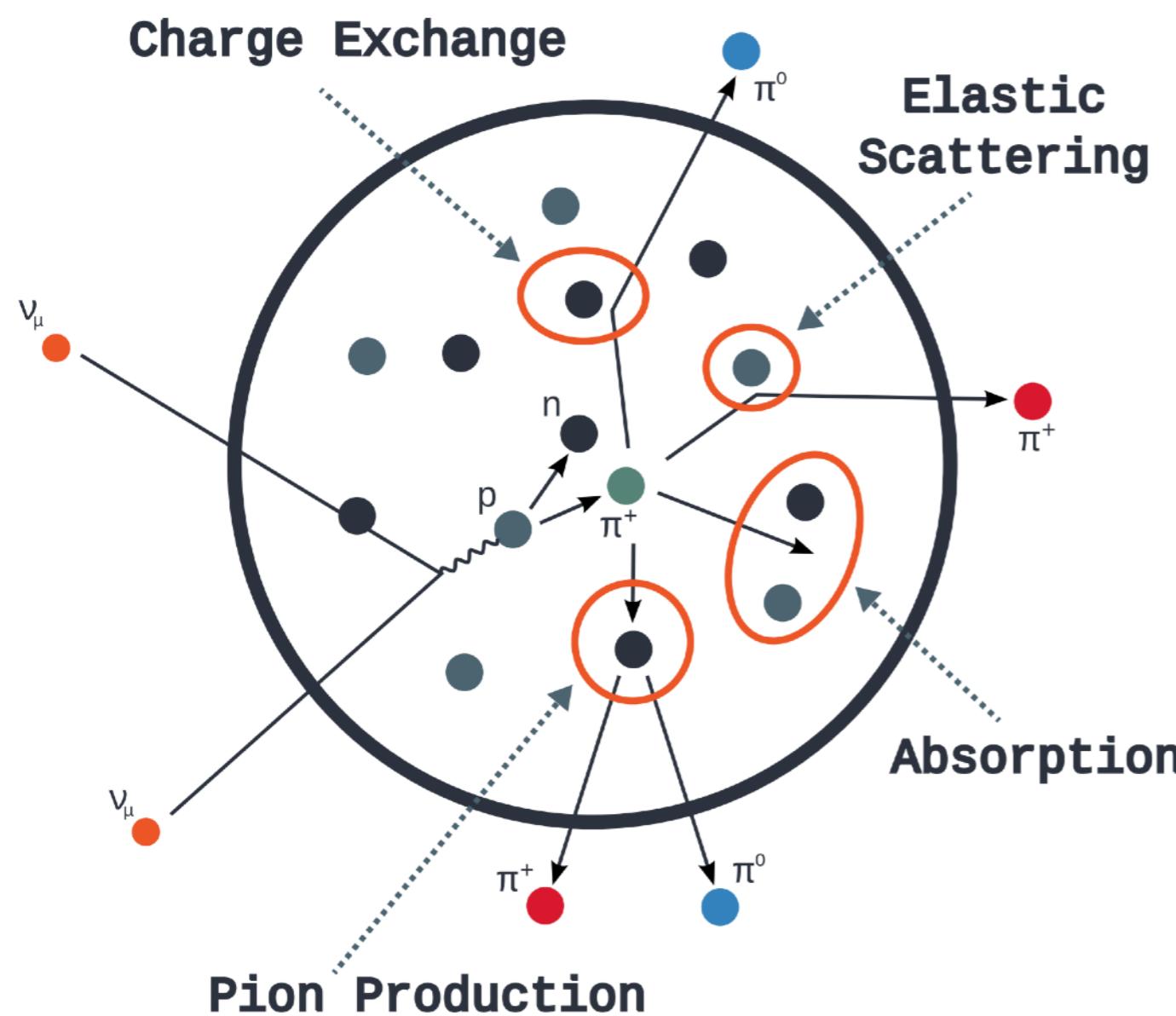


Figure: L. Pickering  
Oscillation parameters from NuFit 4.1

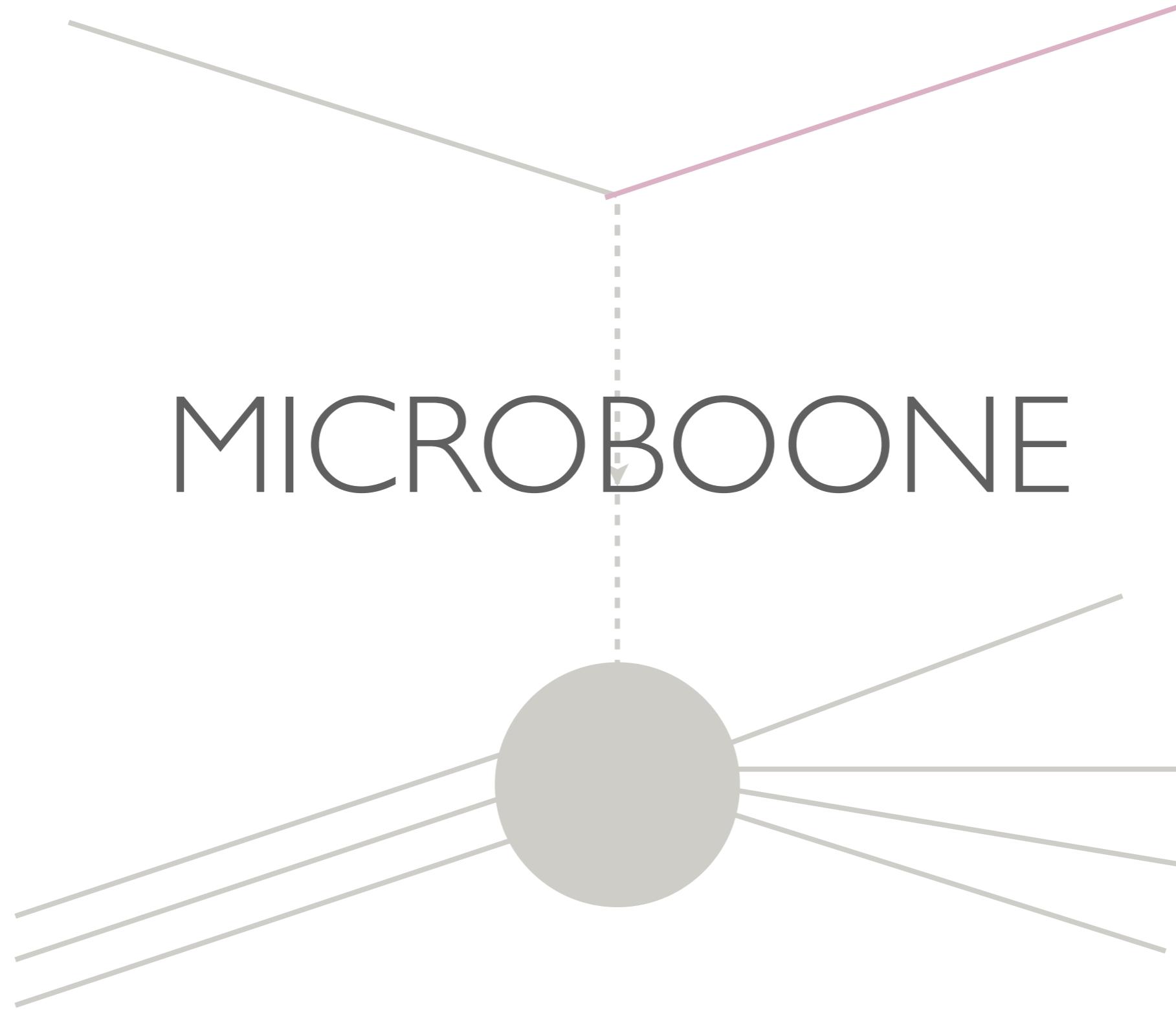
# NUCLEAR EFFECTS

Image: T. Golan



- Particles may re-interact as they exit nucleus  
→ particles can be lost  
→ particles you see may not be from neutrino interaction
- Neutrino isn't interacting with a nucleon at rest → initial state of interaction unknown
- Nuclear effects depend on nucleus — argon data is vital!

# MICROBOONE









**MicroBooNE** is a 170 ton Liquid Argon Time Projection Chamber (LArTPC)

Main physics goals:

- LArTPC detector R&D for future experiments
- Measure neutrino-Ar cross sections
- Search for new neutrino oscillations at short distances
- Exotic beam-dump/astrophysical neutrino signals



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# FERMILAB'S NEUTRINO BEAMS

## Booster $\nu$ beam

MicroBooNE, SBN program



## Booster

proton energy: 8 GeV

## NuMI $\nu$ beam

NOvA, MINERvA, MINOS+

# FERMILAB'S NEUTRINO BEAMS

## Booster $\nu$ beam

*MicroBooNE, SBN program*

 **MicroBooNE**

 **NuMI  $\nu$  beam**

*NOvA, MINERvA, MINOS+*

 **DUNE  $\nu$  beam**  
*(planned)*

## Booster

proton energy: 8 GeV

## Main Inje

proton energy: 1

## Booster Neutrino Beam (BNB): 463m

>99%  $\nu_\mu/\bar{\nu}_\mu$  at peak

$\langle E_\nu \rangle = 850$  MeV

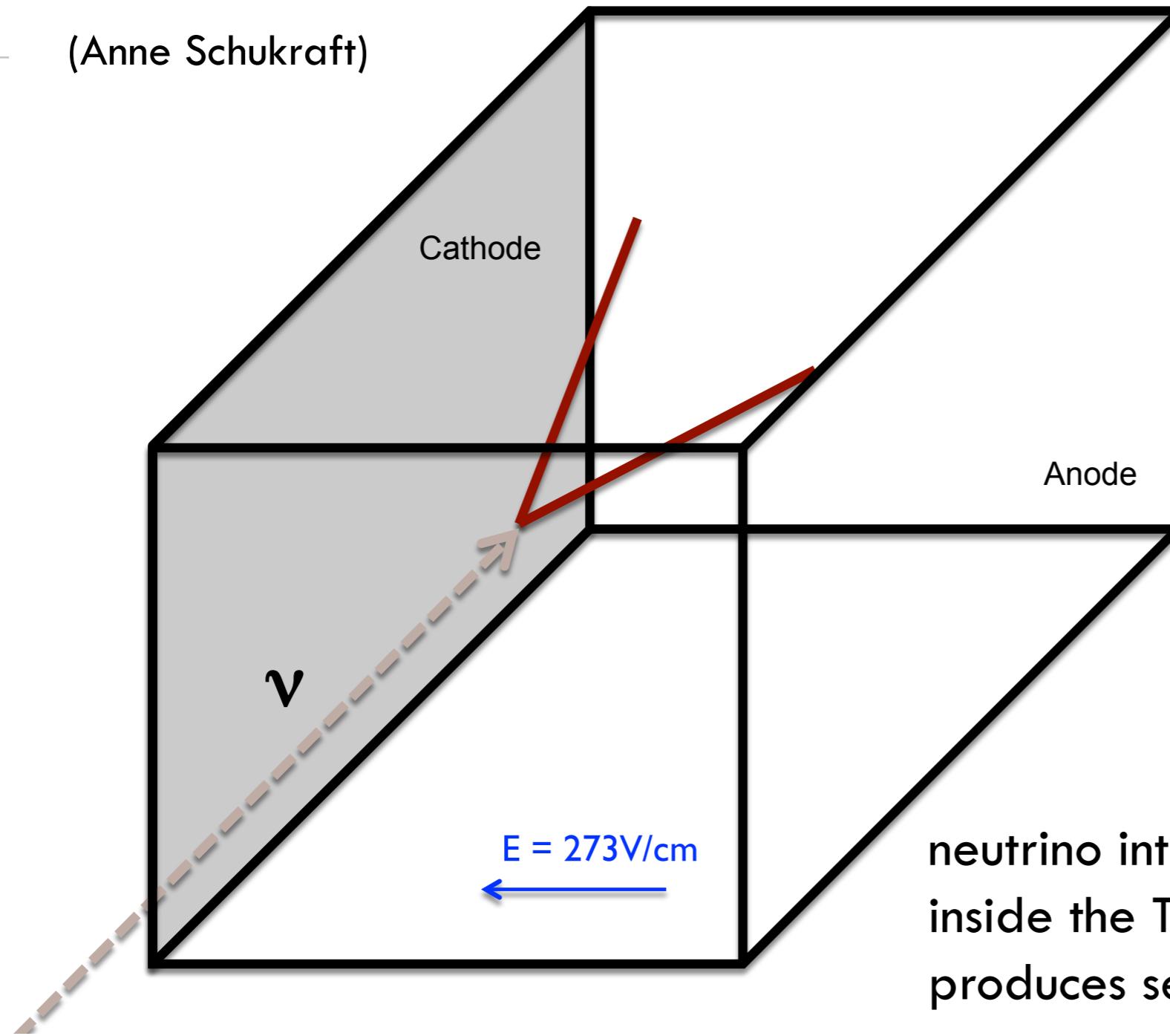
## NuMI Neutrino Beam (NuMI): ~680m

8° off axis → 5%  $\nu_e$

Image: G. Zeller

# LIQUID ARGON TPC

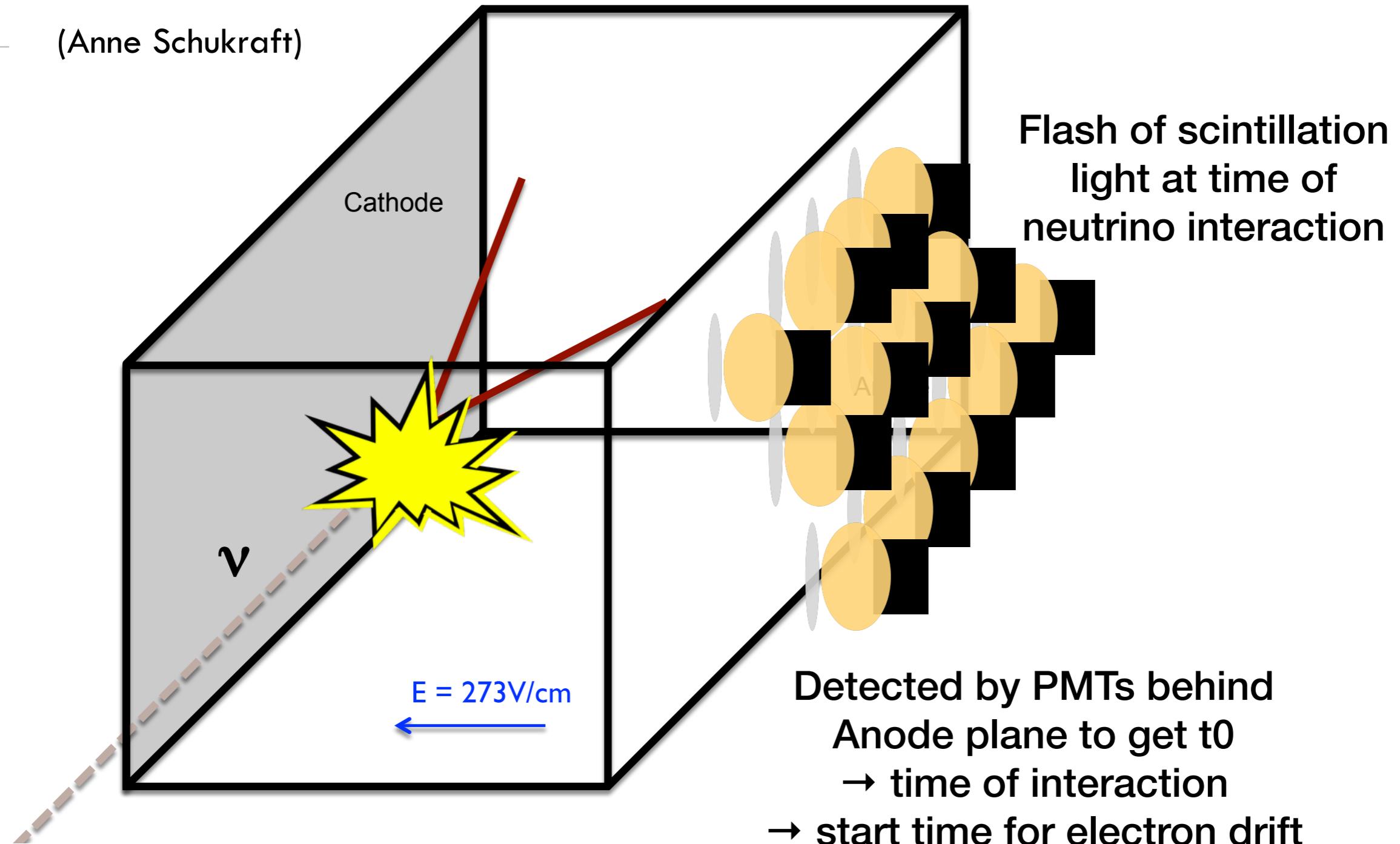
(Anne Schukraft)



neutrino interacts with the argon  
inside the TPC volume and  
produces secondary particles

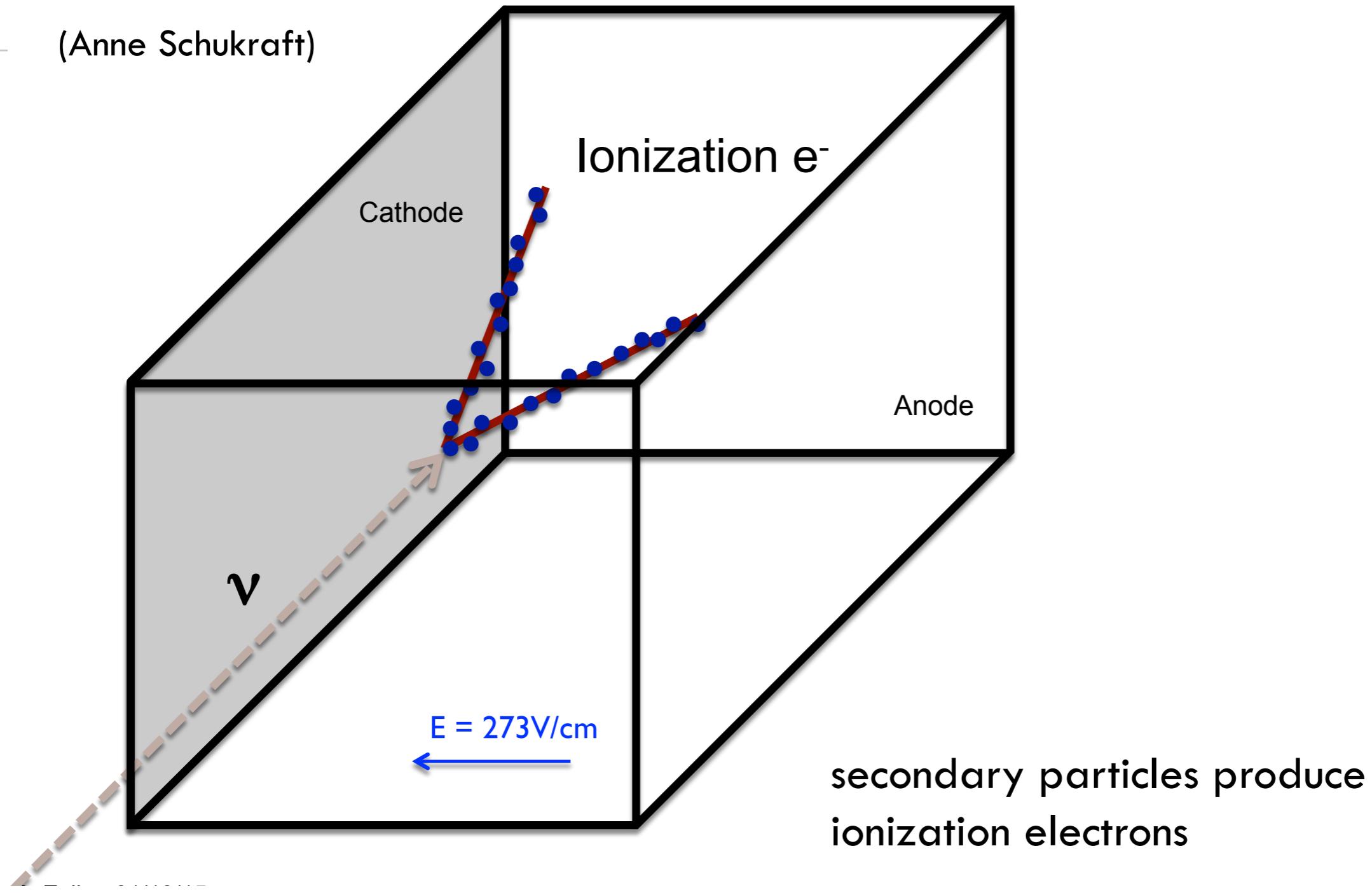
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(Anne Schukraft)



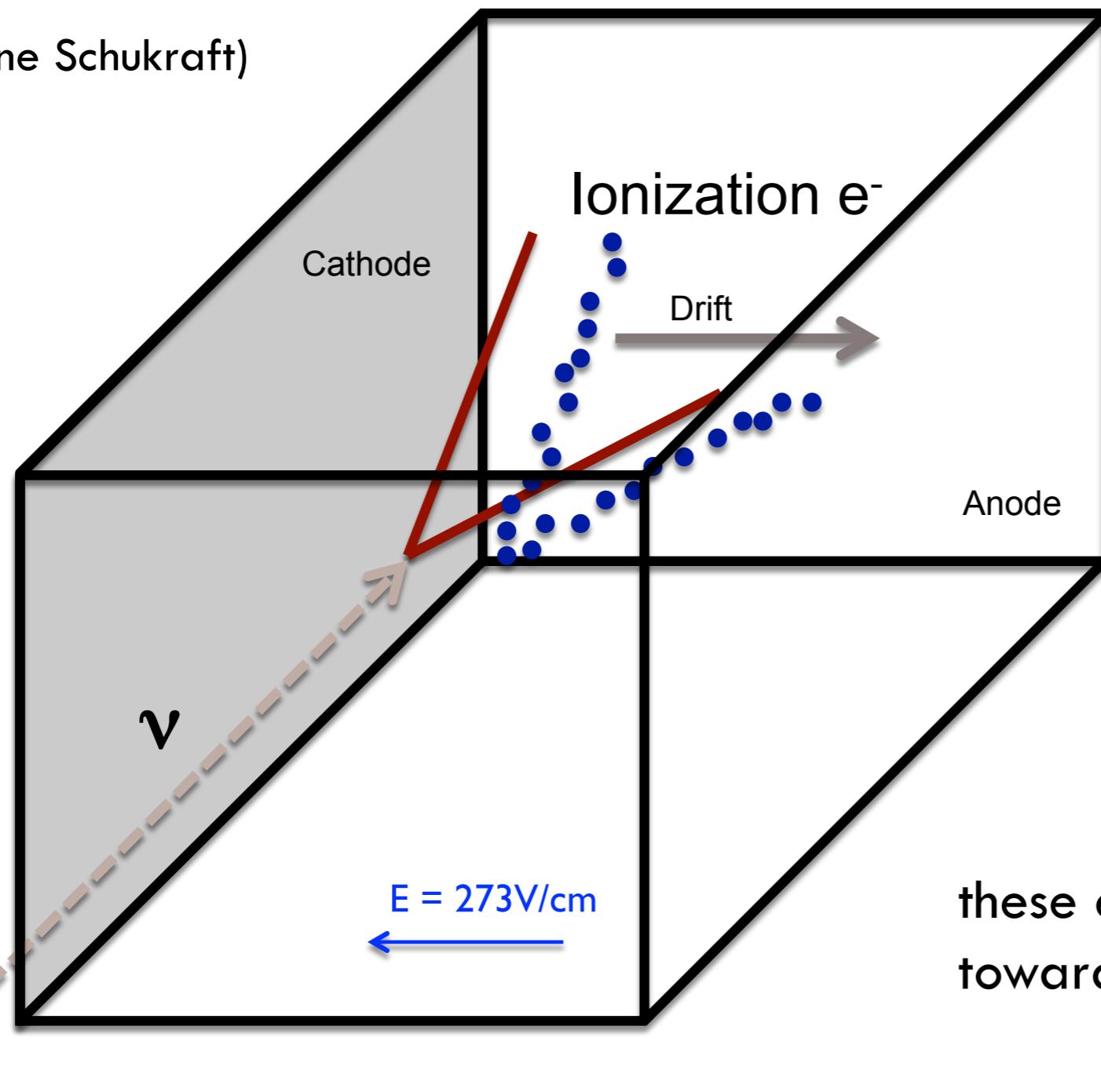
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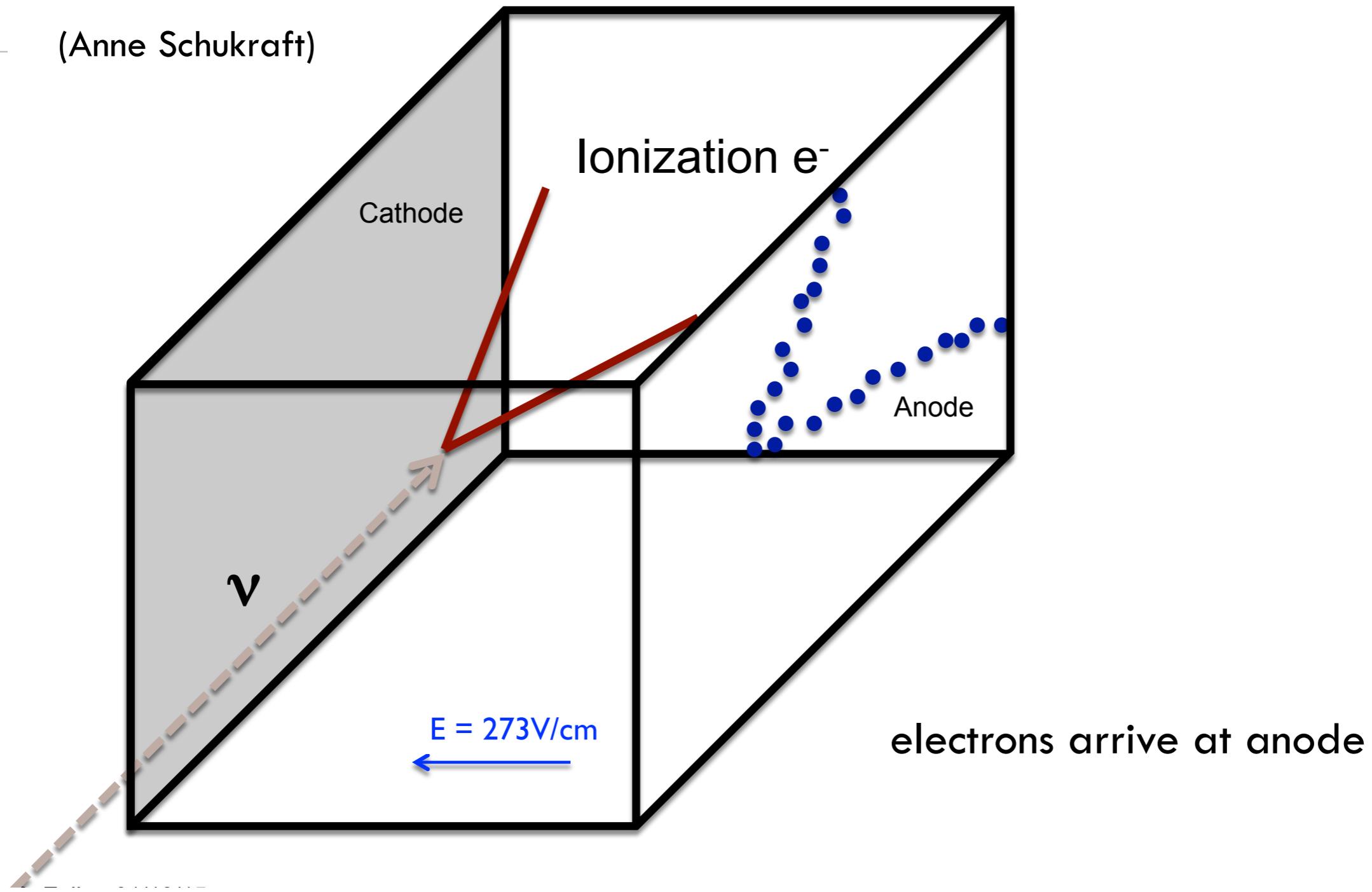
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(Anne Schukraft)



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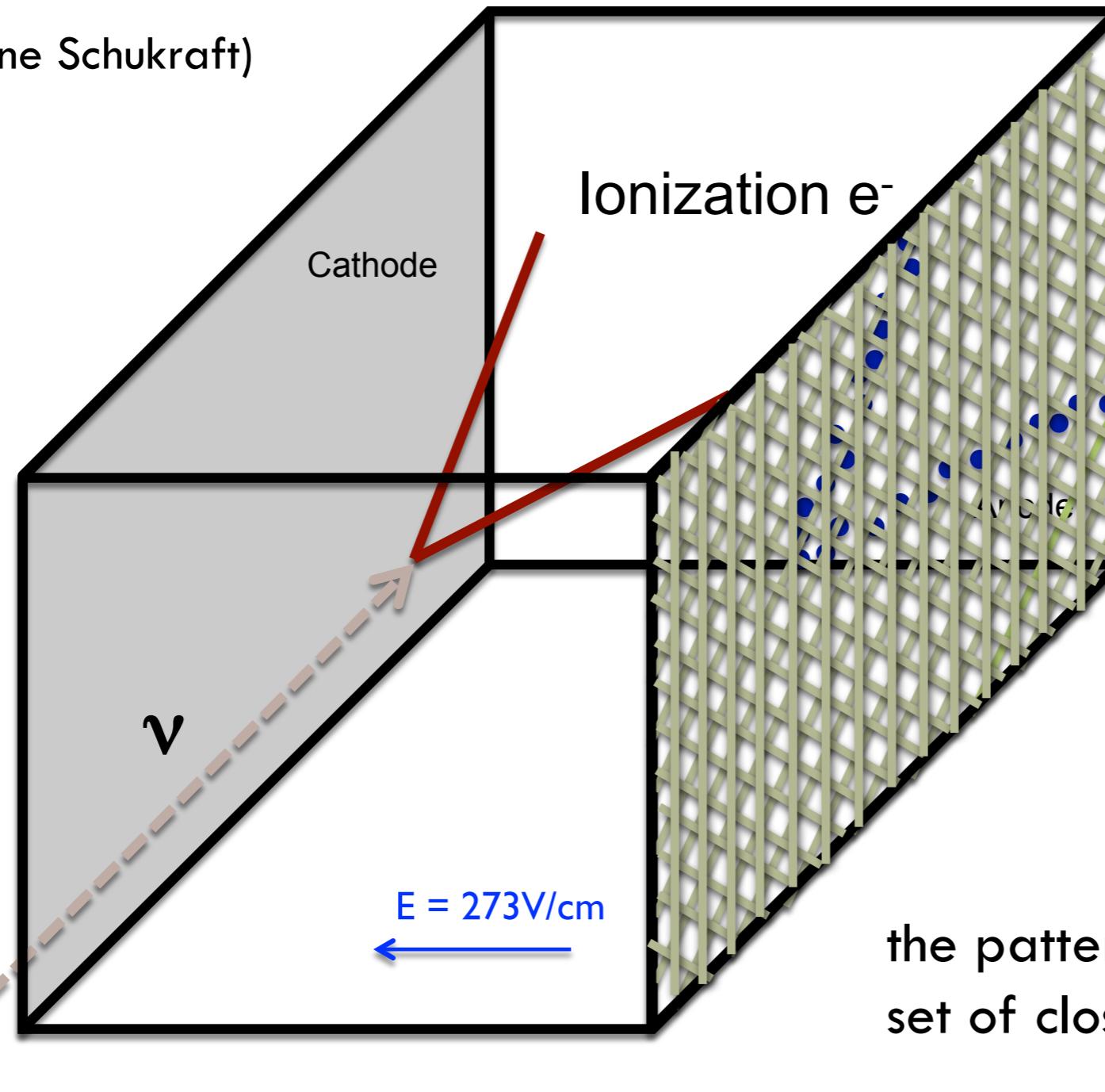
(Anne Schukraft)



# LIQUID ARGON TPC

(Anne Schukraft)

wire planes

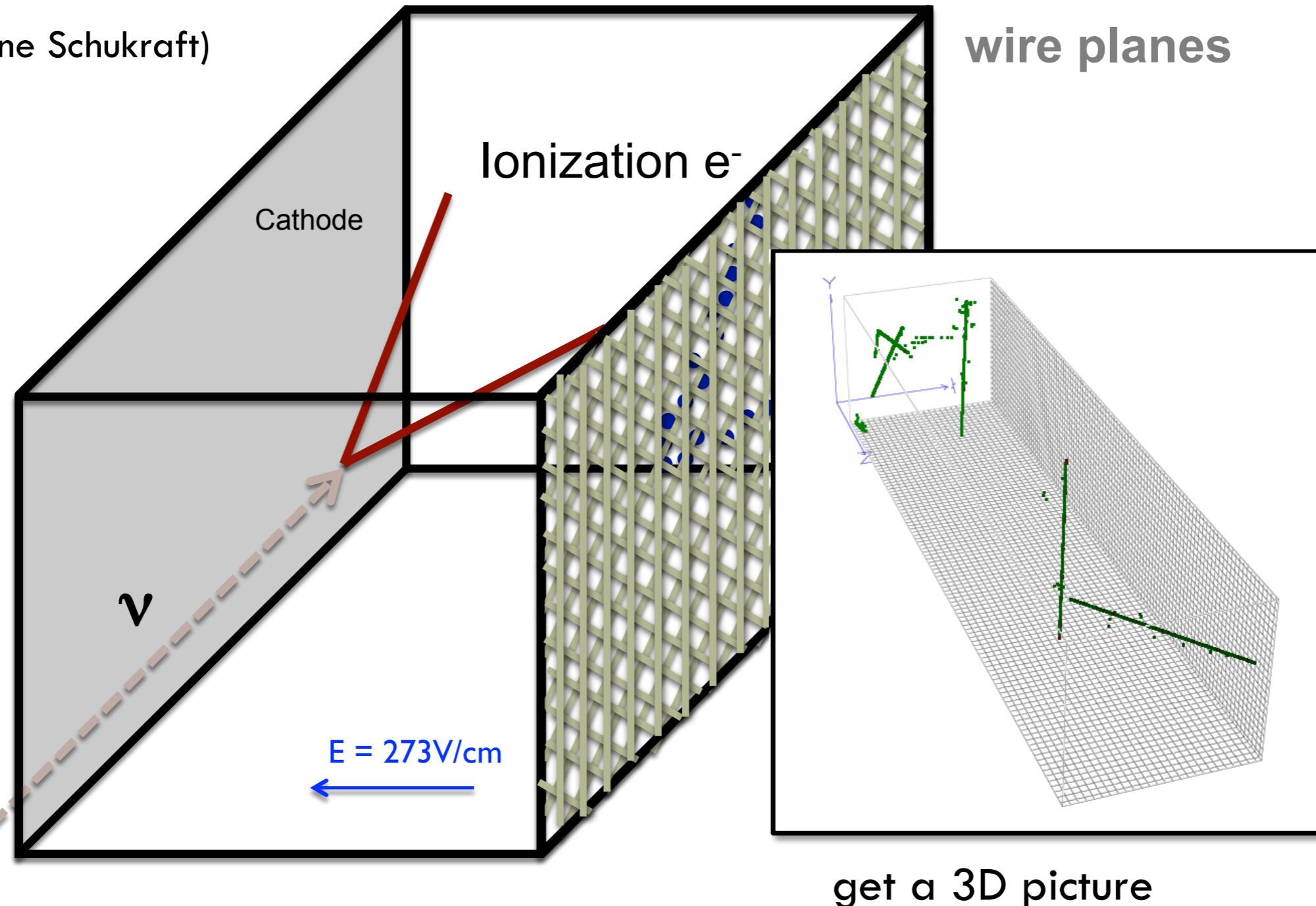


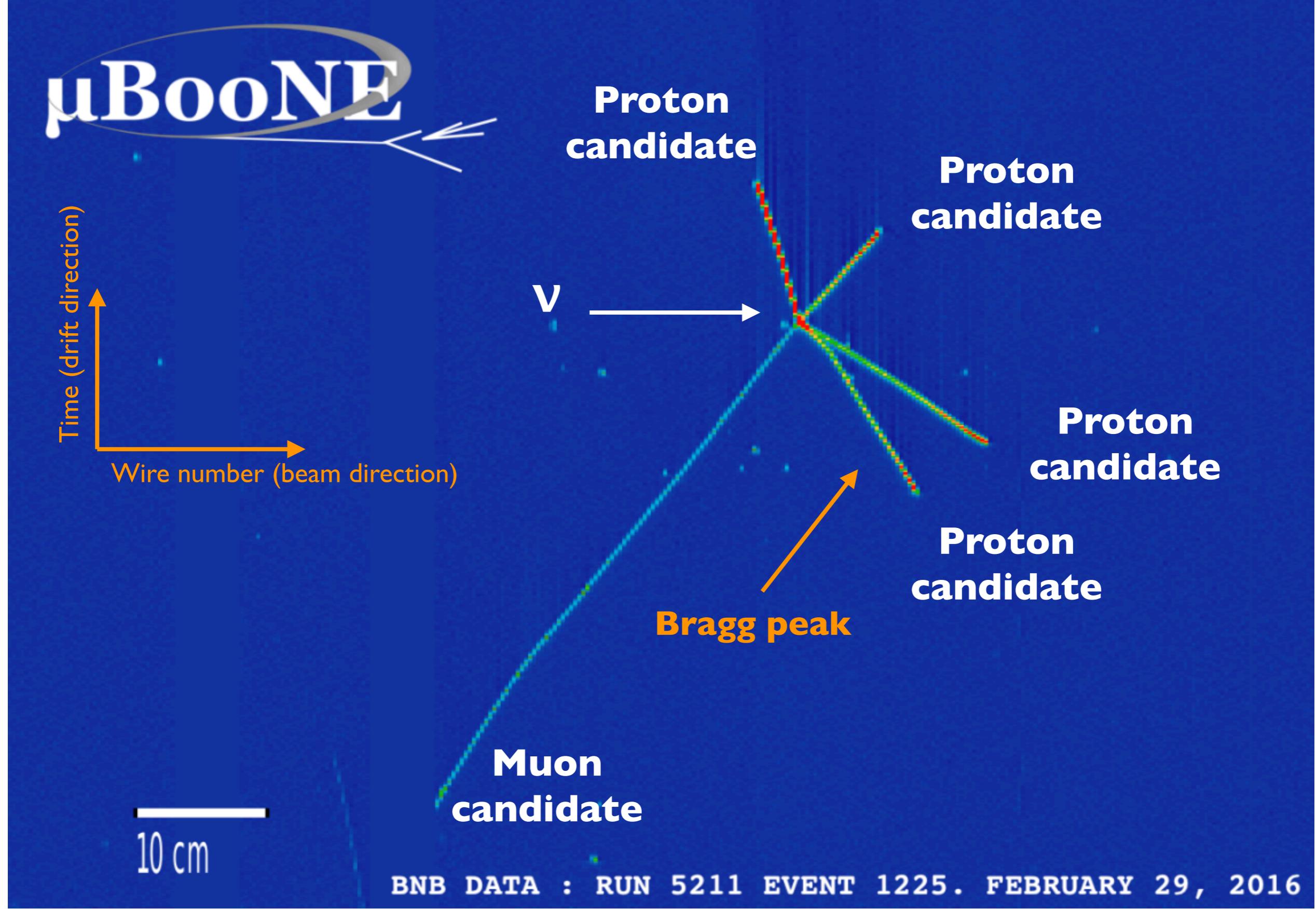
the pattern is recorded on a  
set of closely spaced wires

# LIQUID ARGON TPC

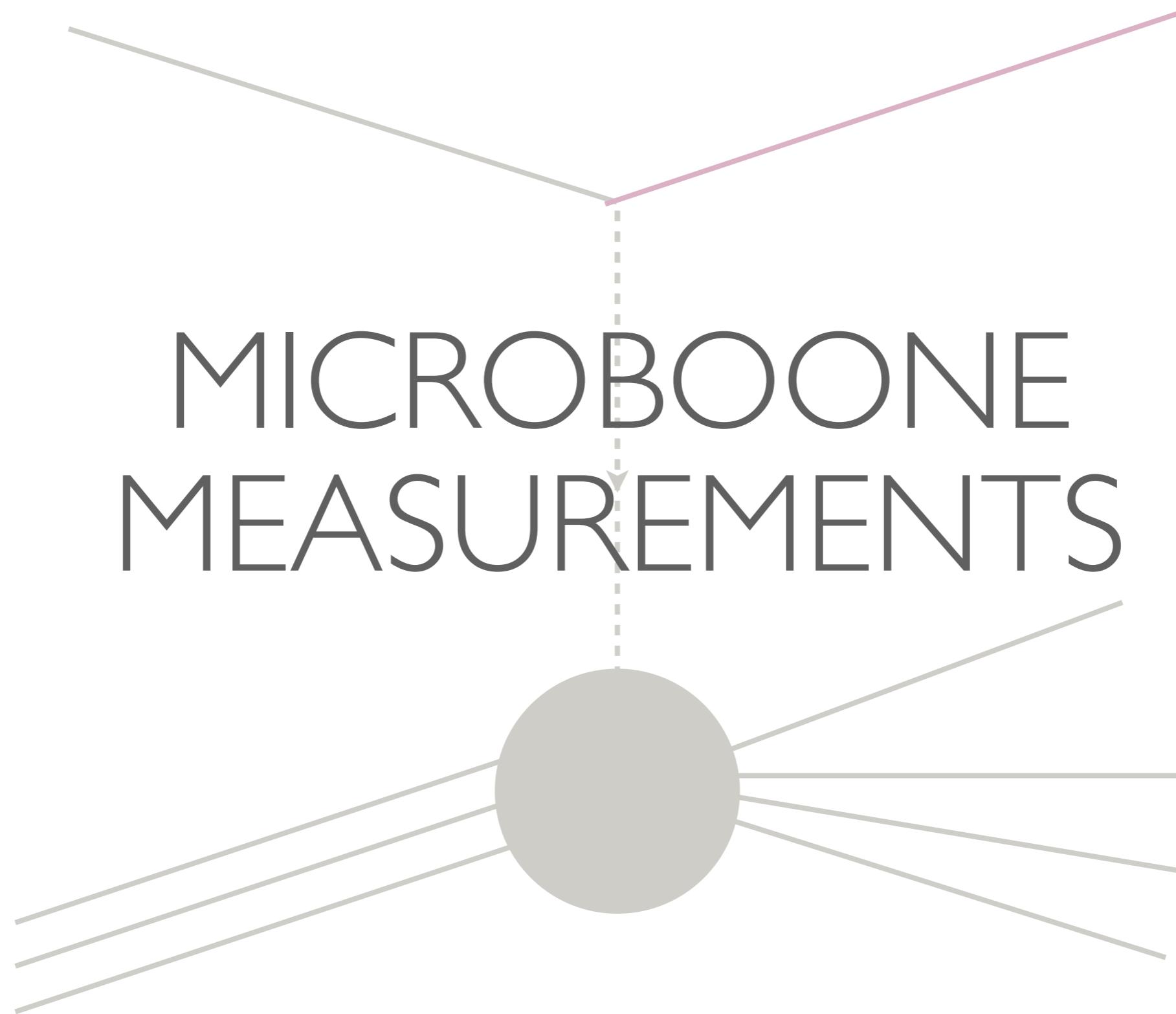
(Anne Schukraft)

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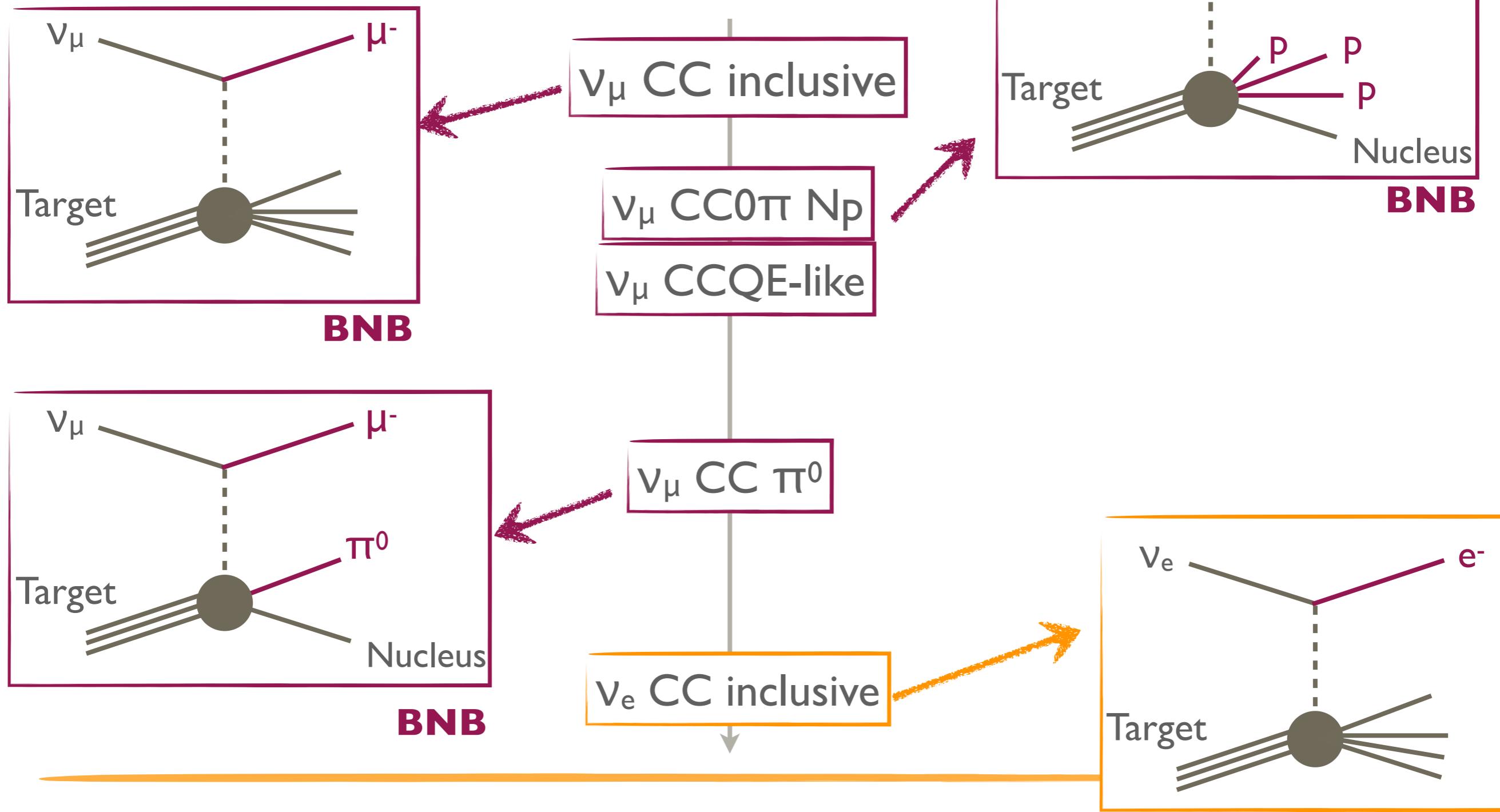




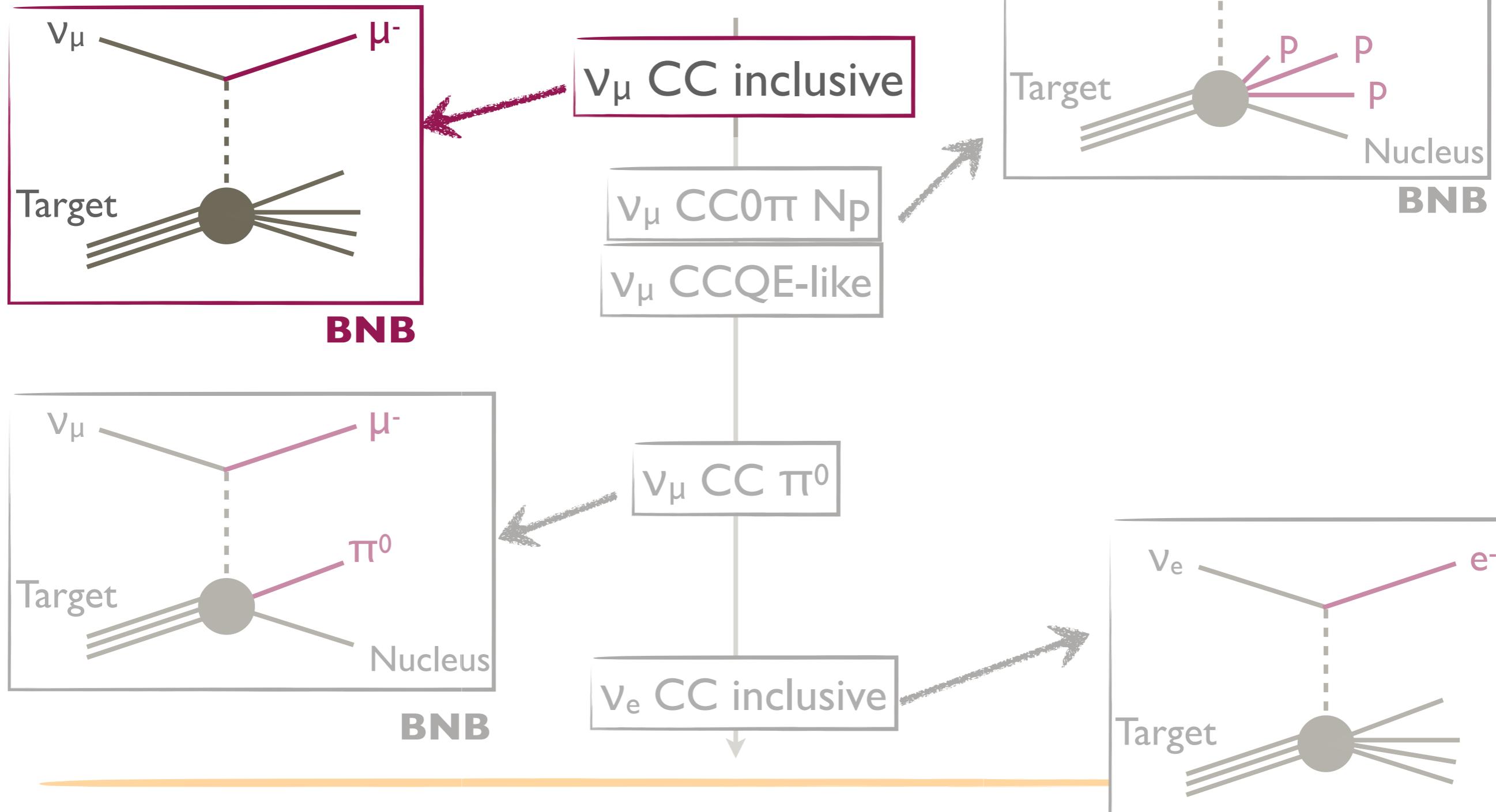
# MICROBOONE MEASUREMENTS



# TODAY'S CROSS SECTION MEASUREMENTS



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# CC-INCLUSIVE

Phys. Rev. Lett. 123, 131801 (2019)

- Measure cross-section for **charged-current (CC) inclusive** muon neutrino interactions

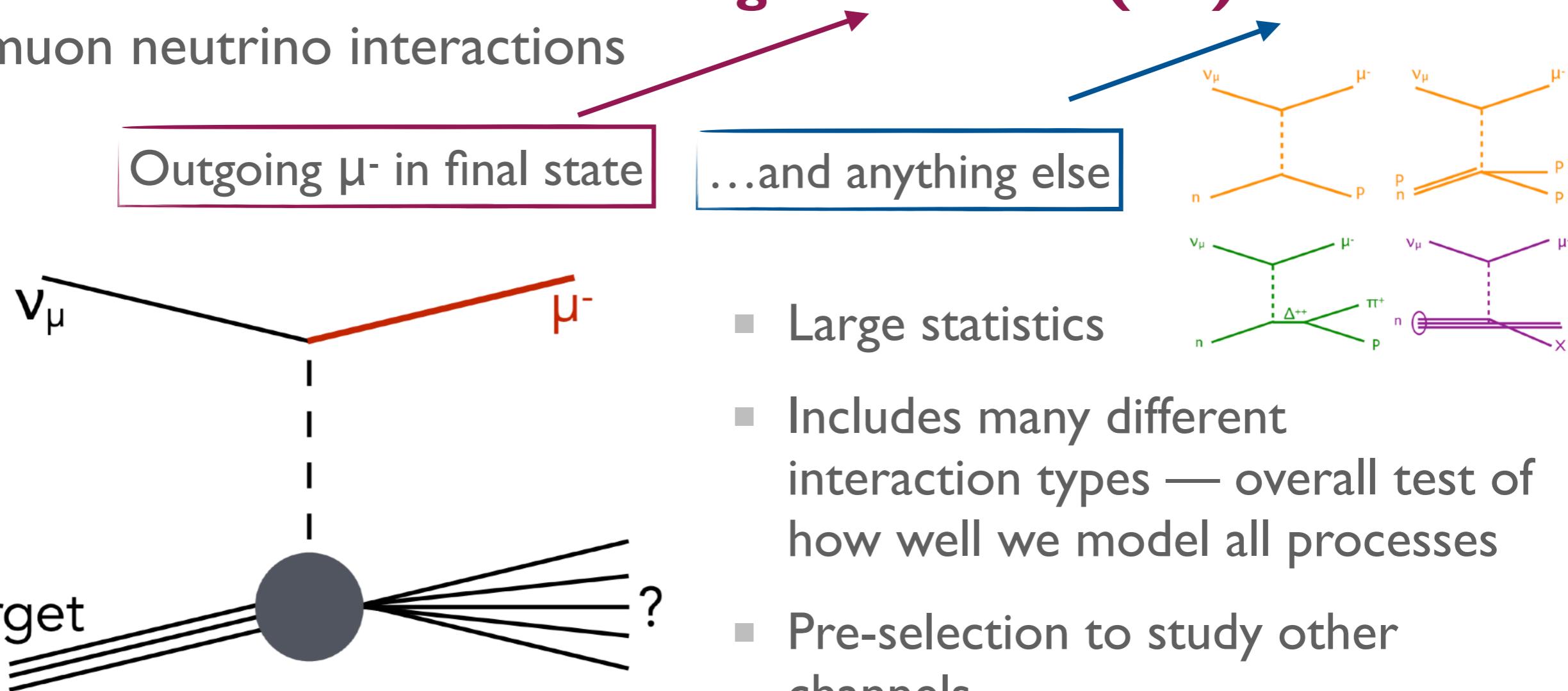


Figure from [M. del Tutto](#)

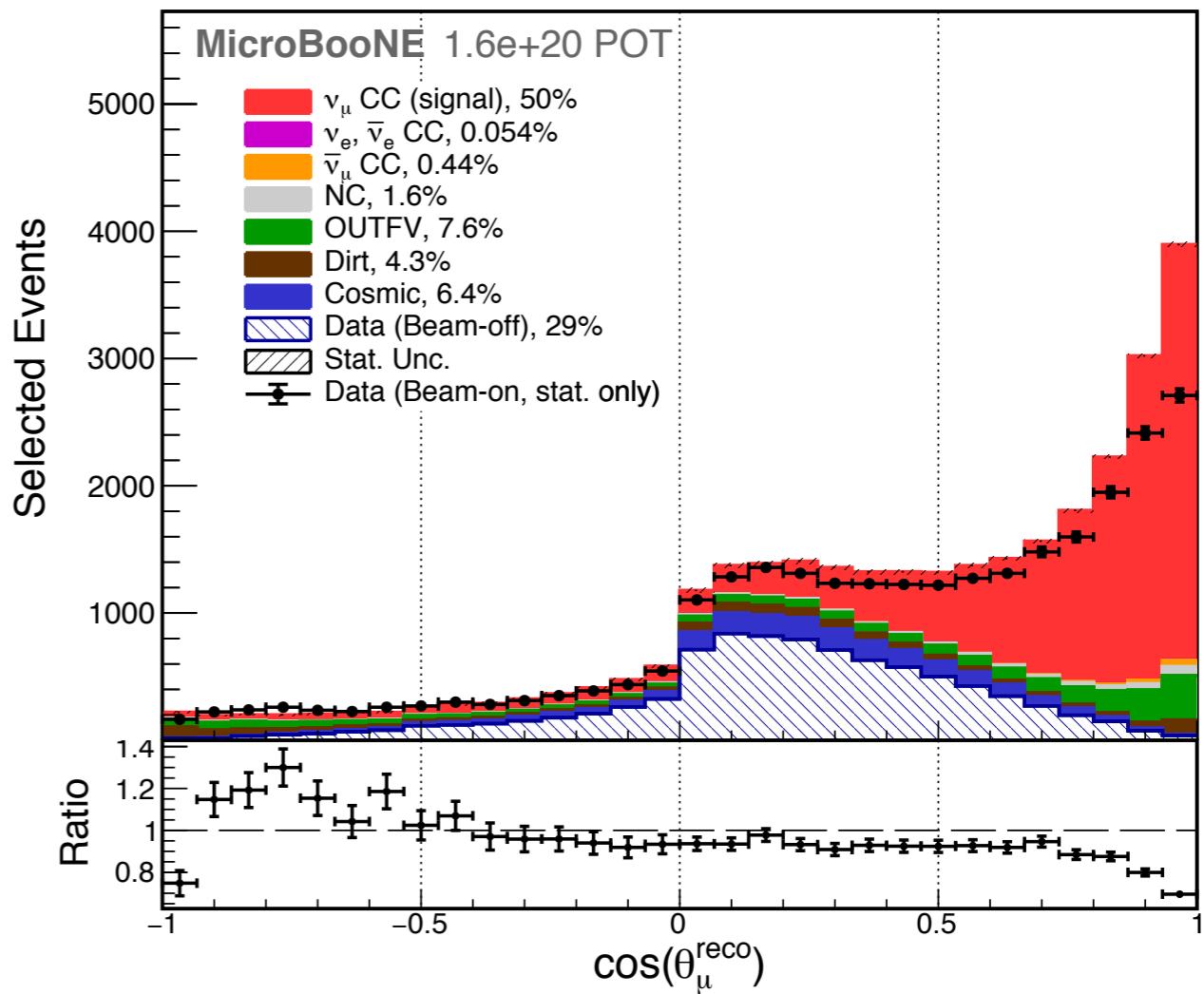
# CC INCLUSIVE CROSS SECTION MEASUREMENT

Phys. Rev. Lett. 123, 131801 (2019)

**Signal (CC-inclusive) events:**  
50.4%

Largest background: **cosmic rays** (29%) → directly measured with beam-off data

**First ever** double-differential cross section measurement on argon: compare to worldwide interaction generators



# CROSS SECTION

Phys. Rev. Lett. 123, 131801 (2019)

Double-differential cross section defined as:

$$\left\langle \frac{d^2\sigma}{dp_\mu^{\text{reco}} d\cos\theta_\mu^{\text{reco}}} \right\rangle_i = \frac{N_i - B_i}{\tilde{\epsilon}_i \cdot N_{\text{target}} \cdot \Phi_{\nu_\mu} \cdot (\Delta p_\mu \cdot \Delta \cos\theta_\mu)_i}$$

The diagram illustrates the components of the double-differential cross section formula. The formula is:

$$\left\langle \frac{d^2\sigma}{dp_\mu^{\text{reco}} d\cos\theta_\mu^{\text{reco}}} \right\rangle_i = \frac{N_i - B_i}{\tilde{\epsilon}_i \cdot N_{\text{target}} \cdot \Phi_{\nu_\mu} \cdot (\Delta p_\mu \cdot \Delta \cos\theta_\mu)_i}$$

The components are labeled as follows:

- No. selected events in bin i (data)
- No. background events in bin i (MC, cosmic data)
- Overall efficiency (selection and acceptance) in bin i (MC)
- No. target nucleons
- Muon neutrino flux (integrated over all energies)
- Width of bin i

# UNCERTAINTIES

Phys. Rev. Lett. 123, 131801 (2019)

Calculate uncertainty on **total** CC-inclusive cross section

Source of uncertainty	Relative uncertainty [%]
Beam flux	12.4
Cross section modeling	3.9
Detector response	16.2
Dirt background	10.9
Cosmic ray background	4.2
MC statistics	0.2
Stat	1.4
Total	23.8

Flux uncertainty ~12%

Inclusive measurement → not strongly dependent on cross section model

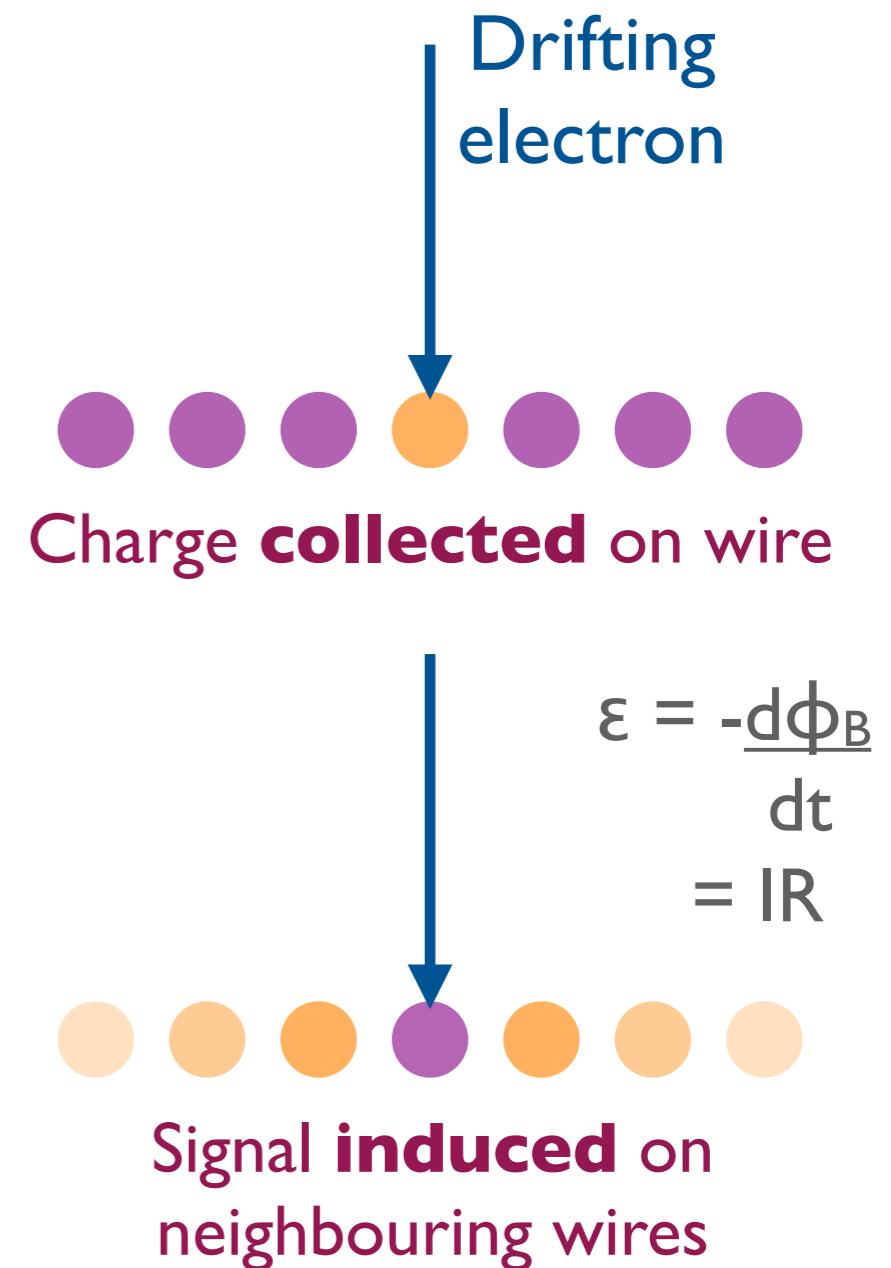
Total uncertainty dominated by detector response

# INDUCED CHARGE

JINST 13 P07006 (2018)

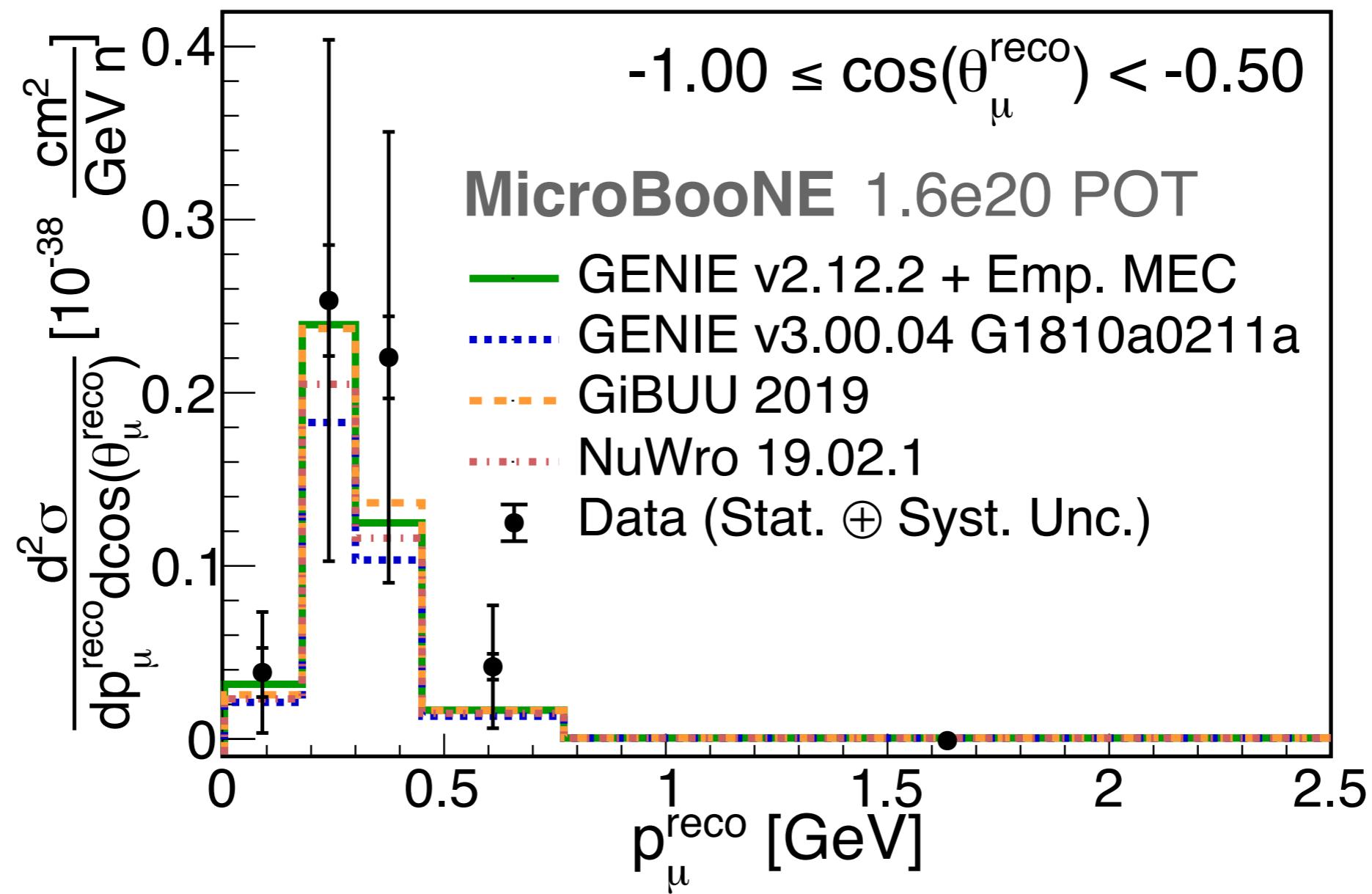
JINST 13 P07007 (2018)

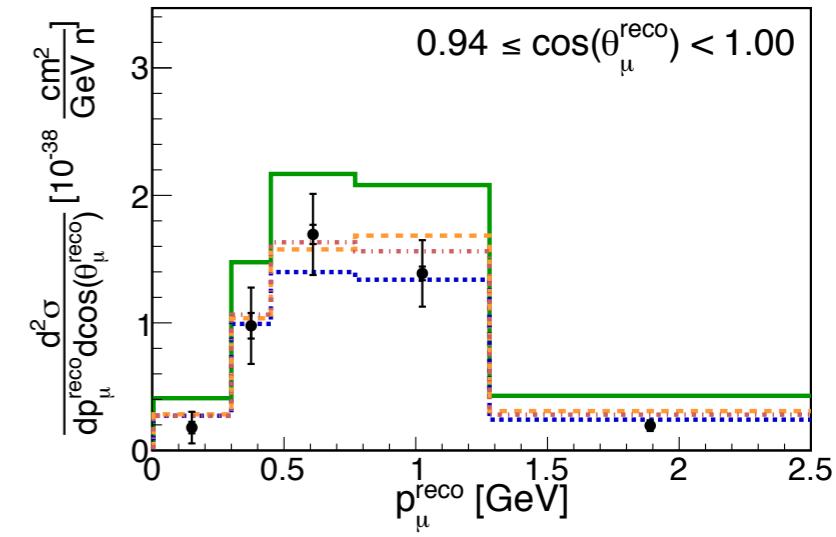
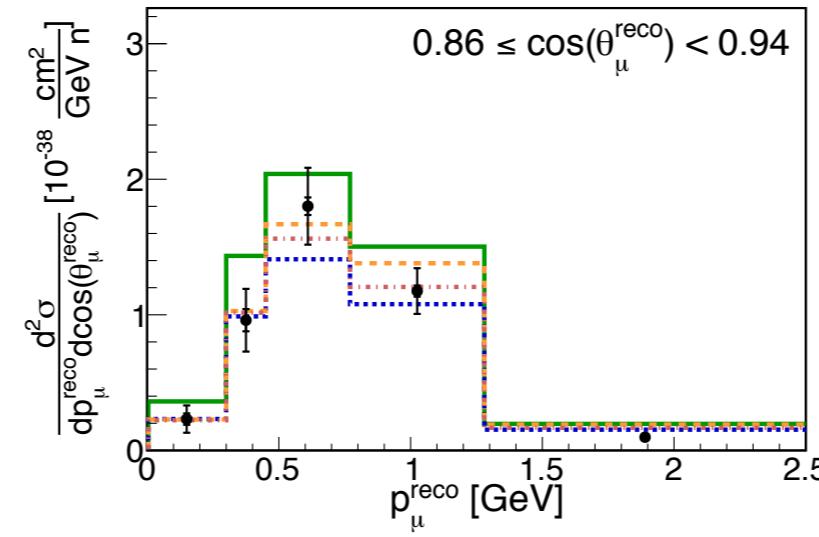
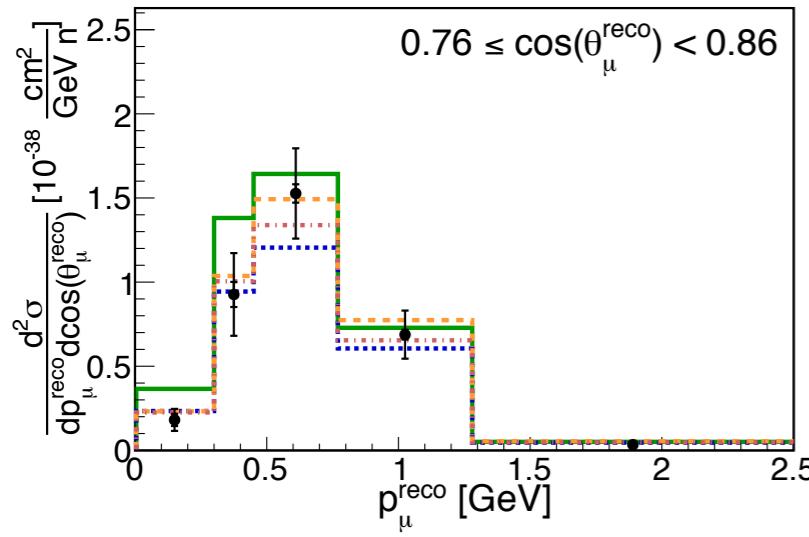
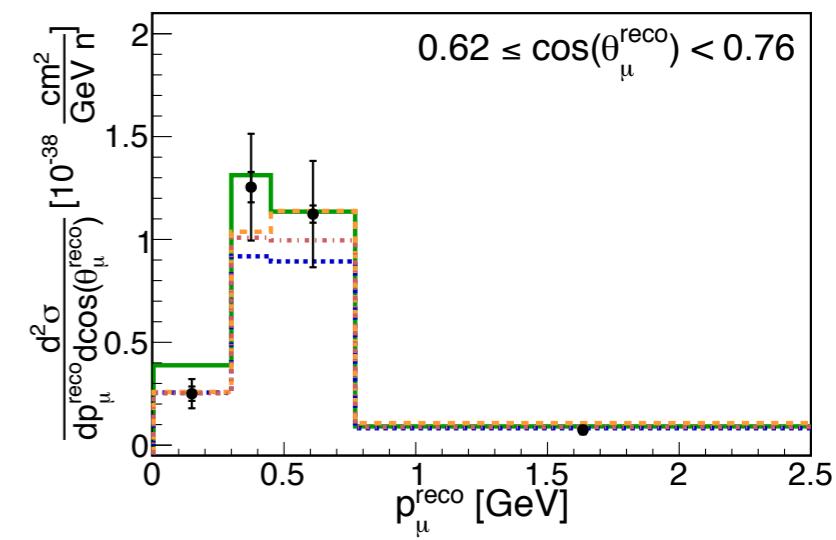
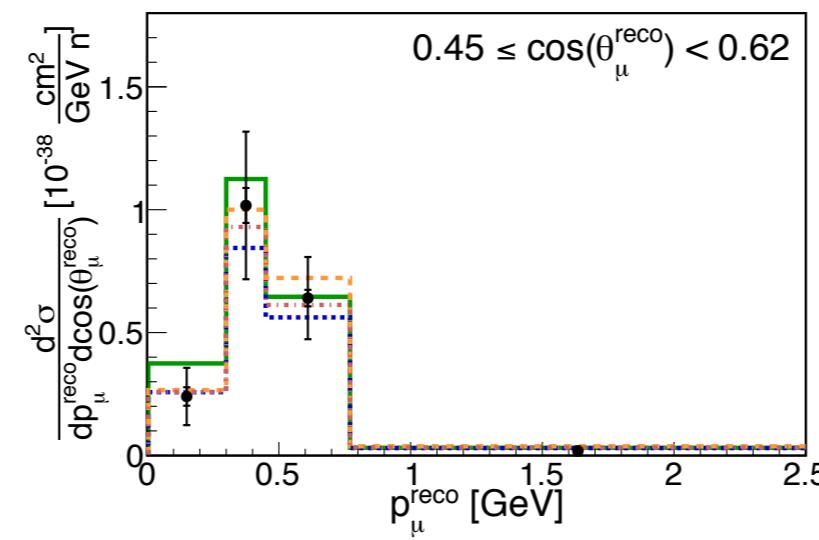
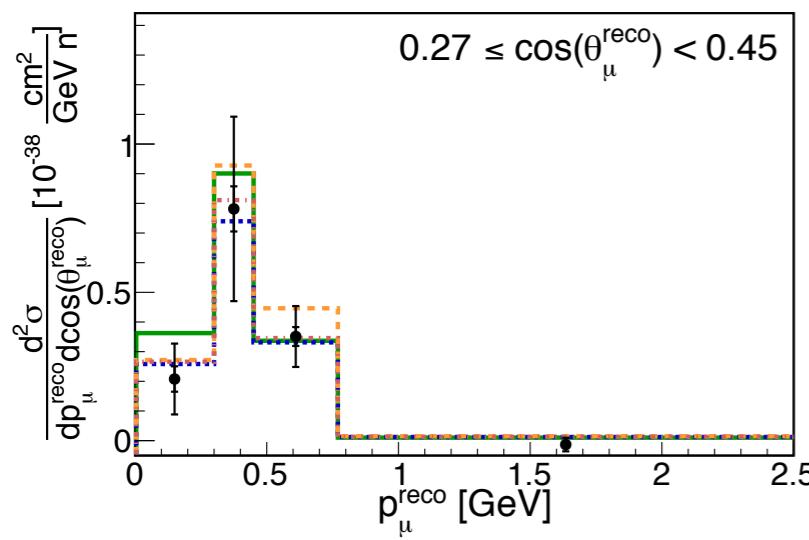
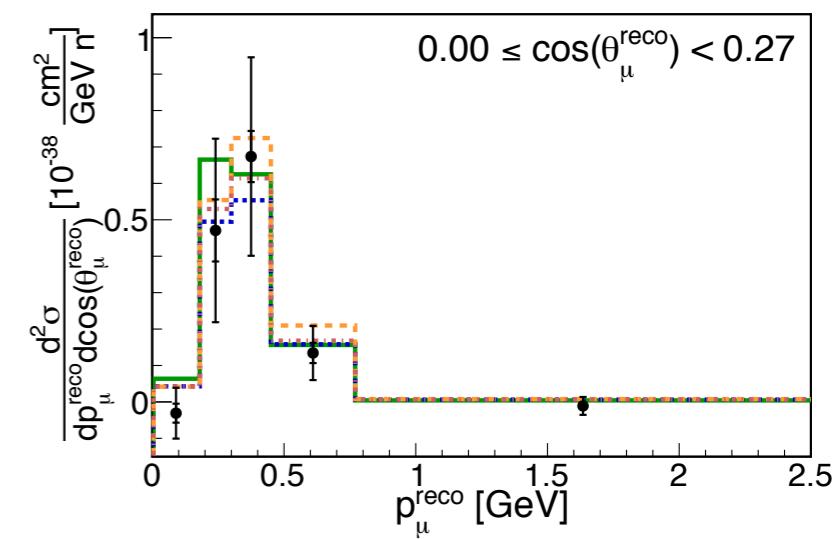
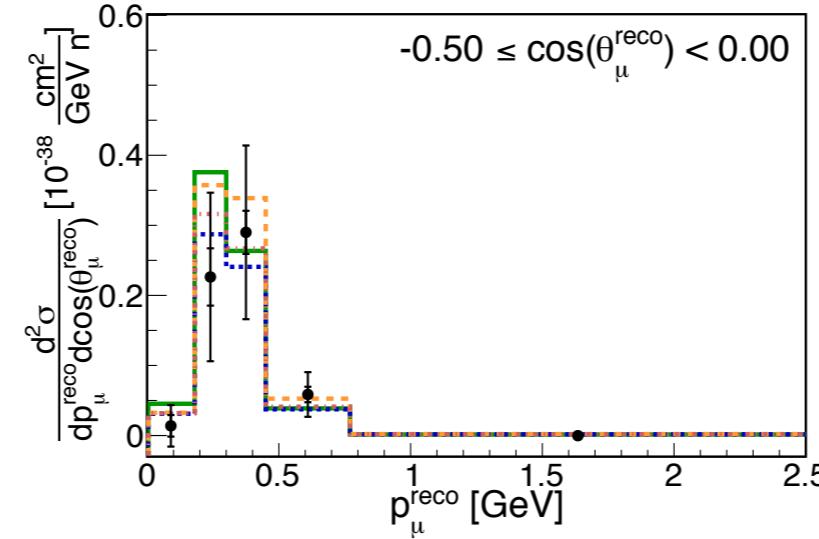
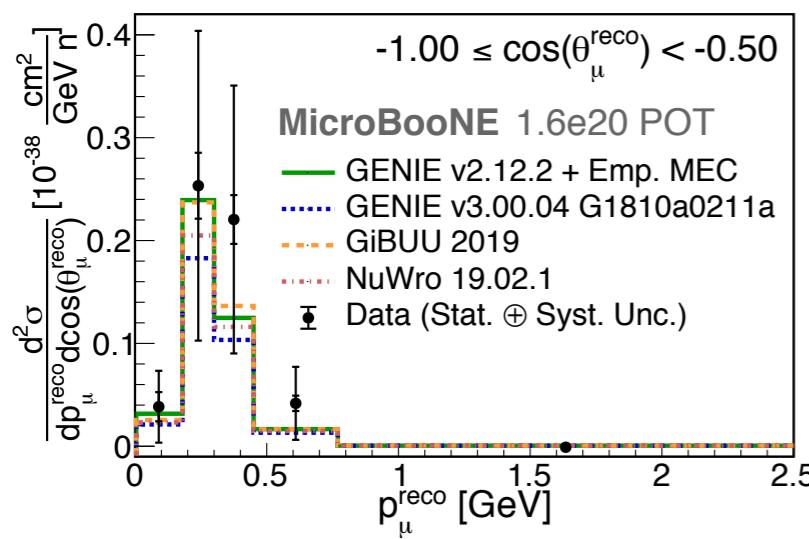
- Detector response → 16% uncertainty on total cross section
- Largest single contribution: 13% due to modeling of **induced charge**
  - MC assumes drift electrons cause signal on only one wire
  - Not true! Nearby wires see **charge by induction**
  - Because of this, **detector response depends on track angle**
  - Easy to fix — include in simulation and account for in reconstruction!

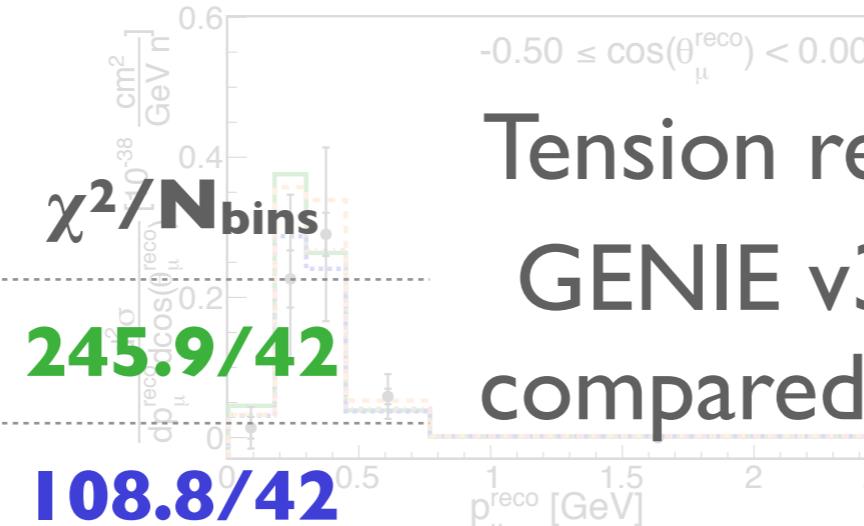
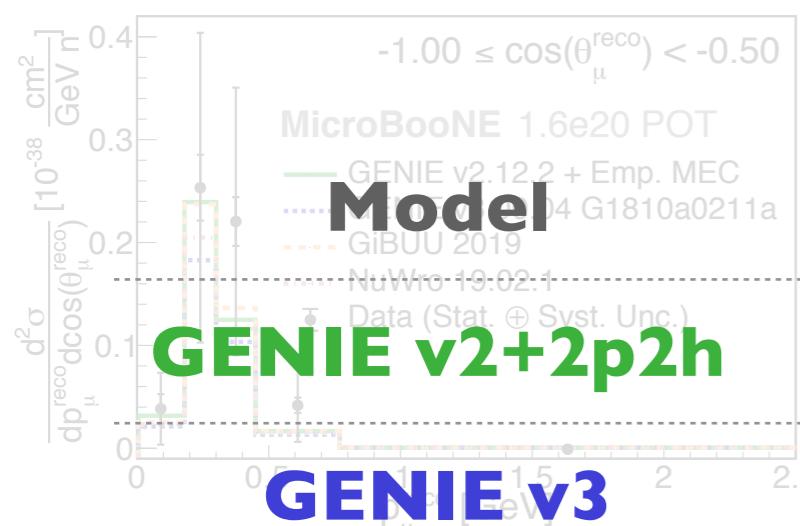


# RESULTS

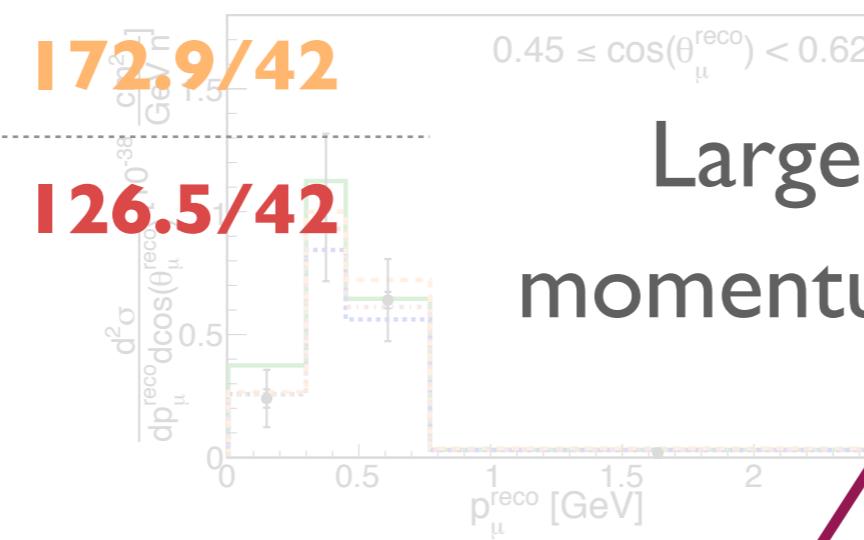
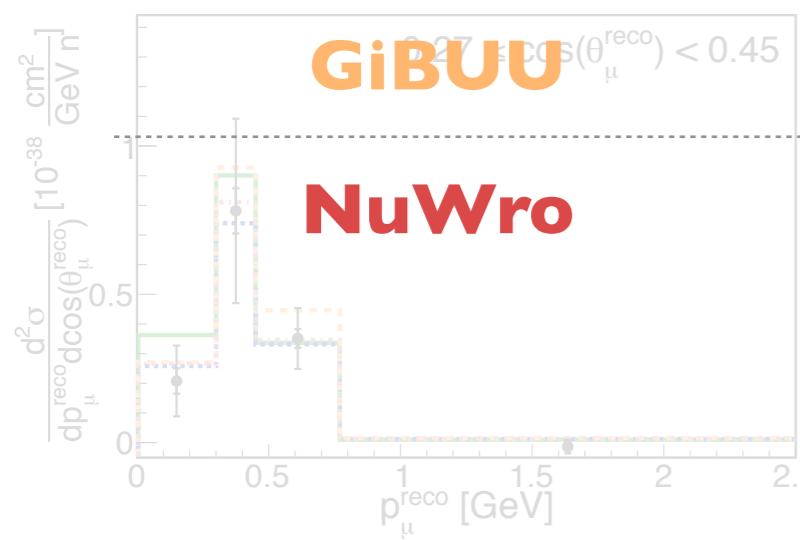
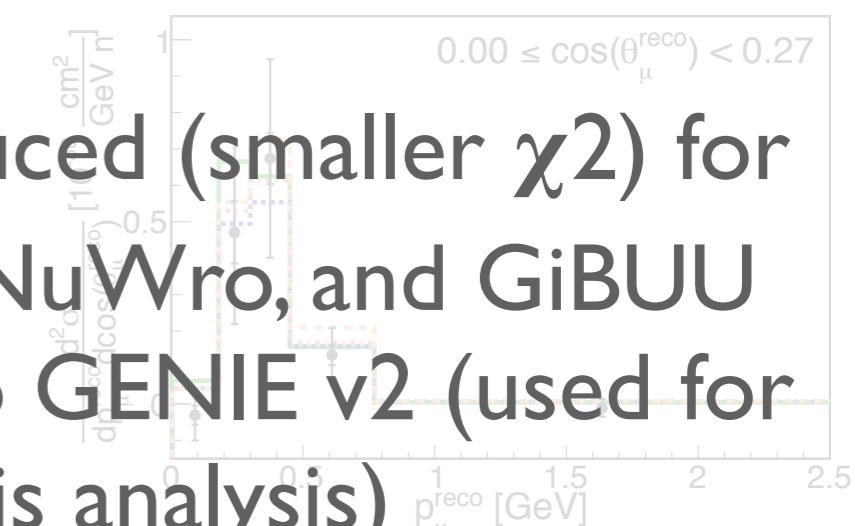
Phys. Rev. Lett. 123, 131801 (2019)



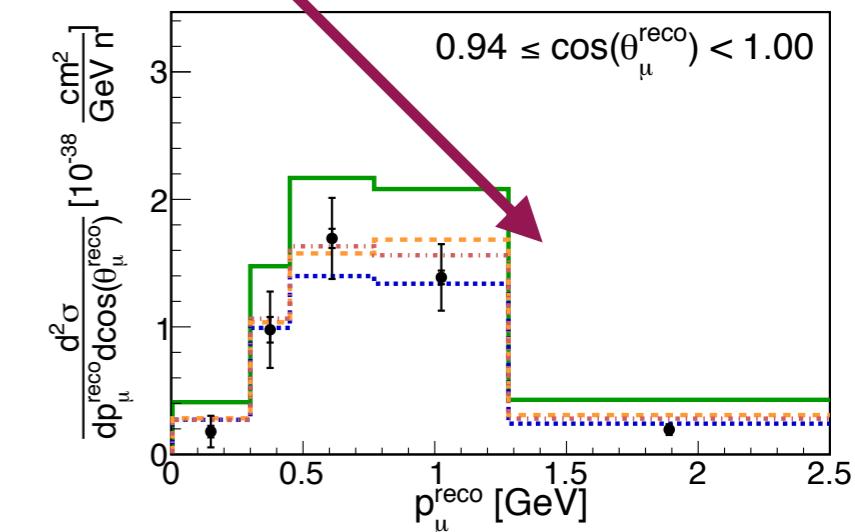
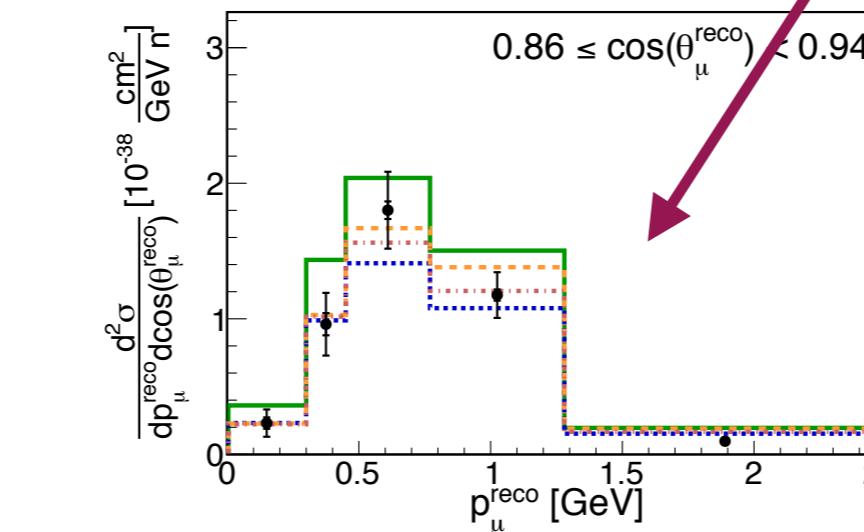
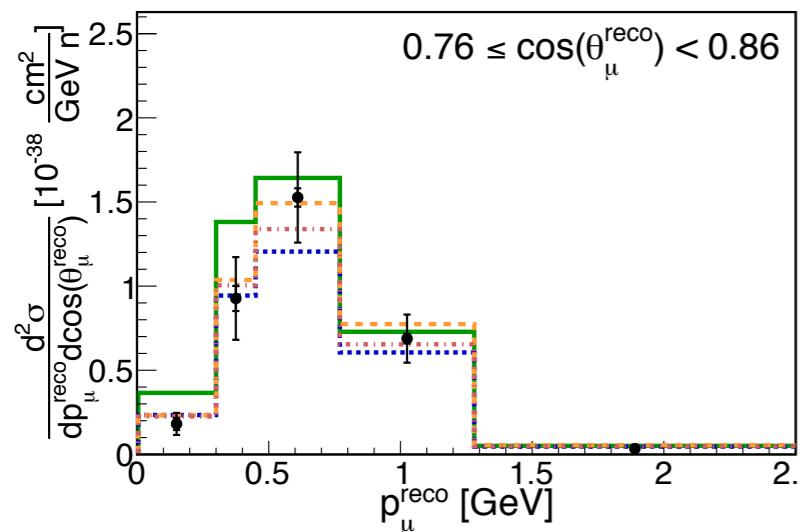
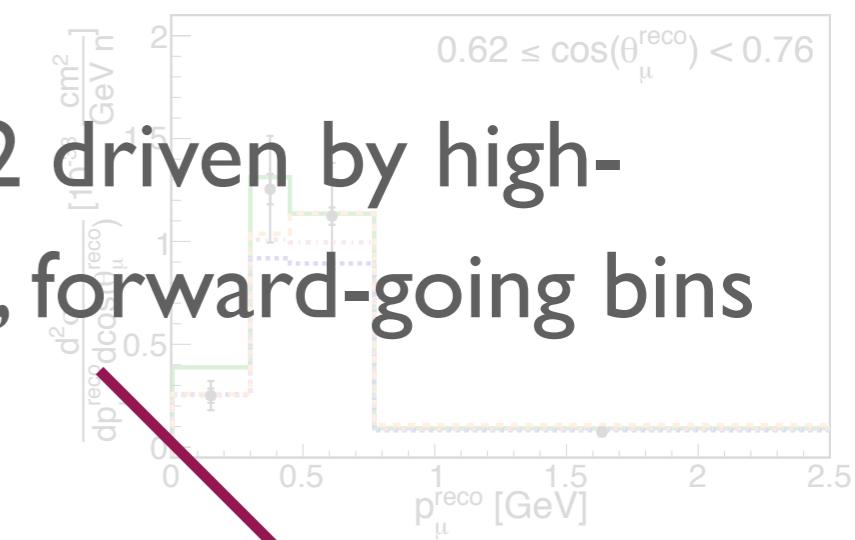




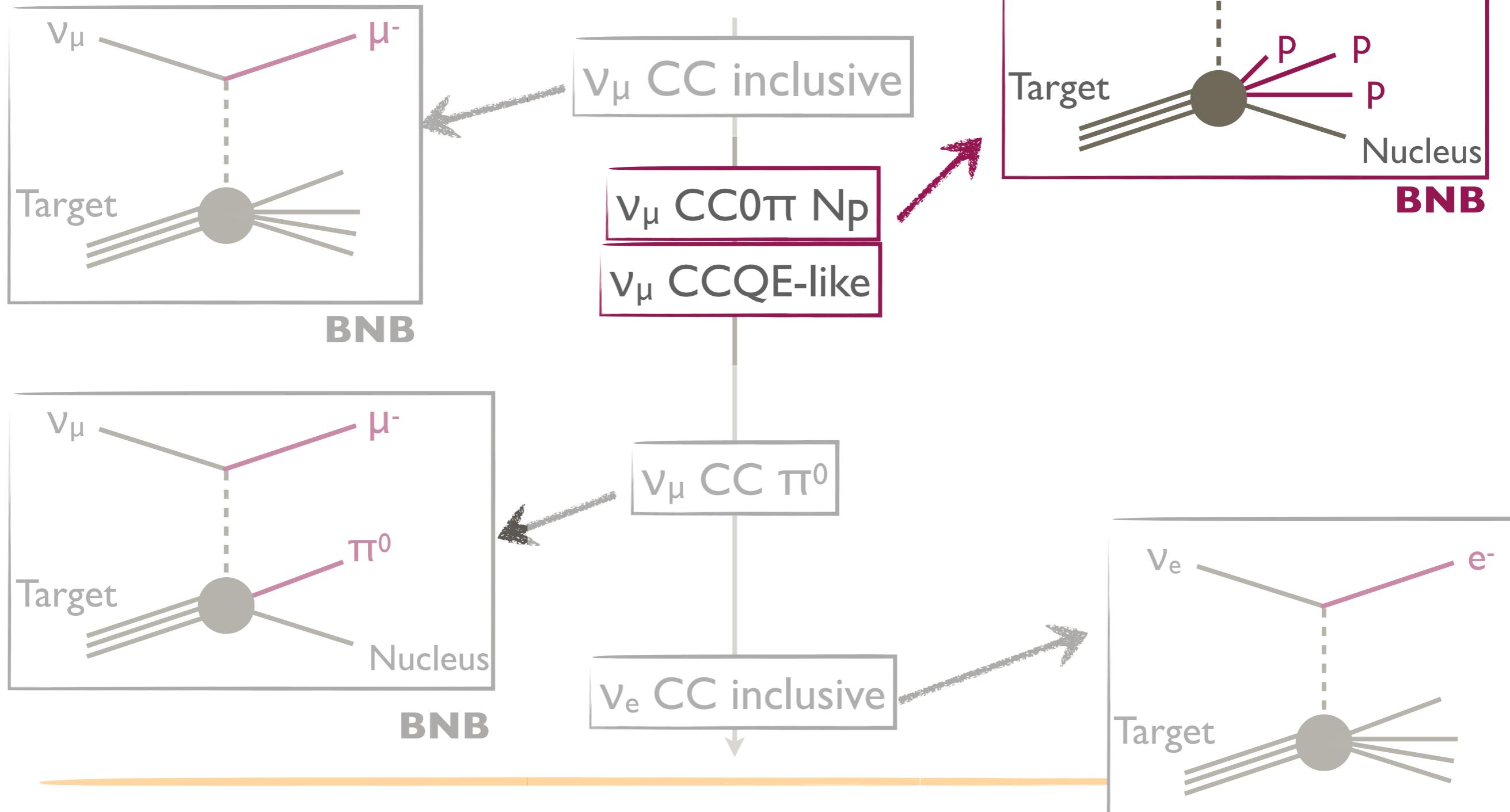
Tension reduced (smaller  $\chi^2$ ) for GENIE v3, NuWro, and GiBUU compared to GENIE v2 (used for this analysis)



Large  $\chi^2$  driven by high-momentum, forward-going bins



# TODAY'S CROSS SECTION MEASUREMENTS



# LArTPC STRENGTH: LOW PROTON THRESHOLDS

Phys. Rev. Lett. 125, 201803 (2020)  
Phys. Rev. D 102, 112013 (2020)  
JINST 15, P03022 (2020)

- Measuring proton kinematics gives us more information about the interaction
- **Low thresholds** → access to new information about nuclear effects, probe e.g. 2p2h scattering

- MicroBooNE: **300 MeV/c**
- ArgoNeuT: **200 MeV/c**

Phys. Rev. D 90, 012008 (2014)

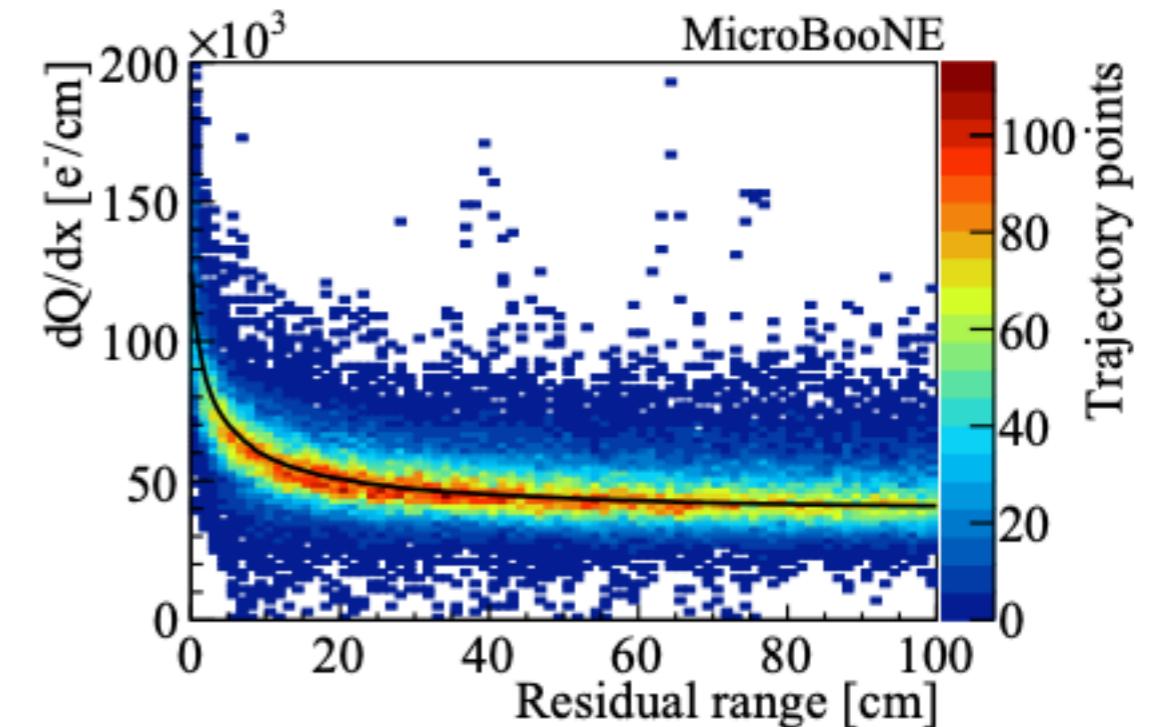
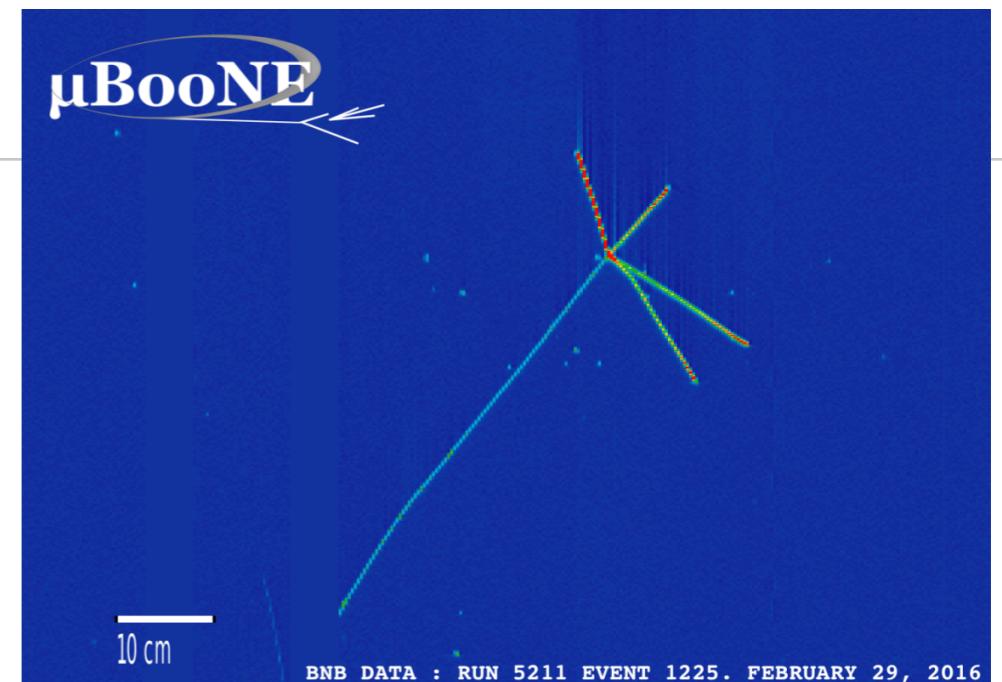
Phys. Rev. D 98, 032003 (2018)

T2K: 500 MeV/c

MINERvA: 450 MeV/c

Phys. Rev. D 99, 012004 (2019)

- Protons **identified by Bragg peak** in last 30 cm of track

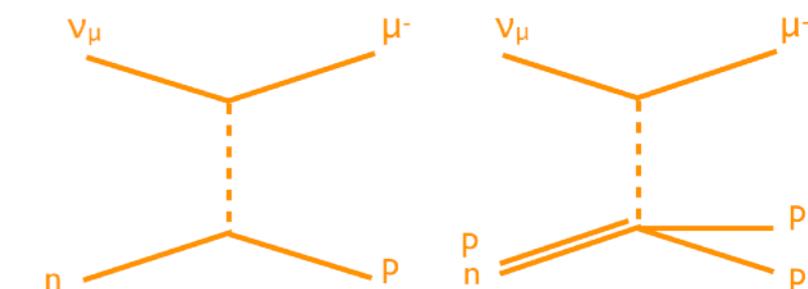


# MEASUREMENTS WITH PROTONS

$\nu_\mu$  CC interaction



No pions

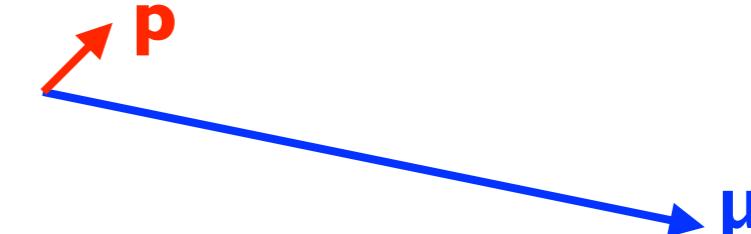
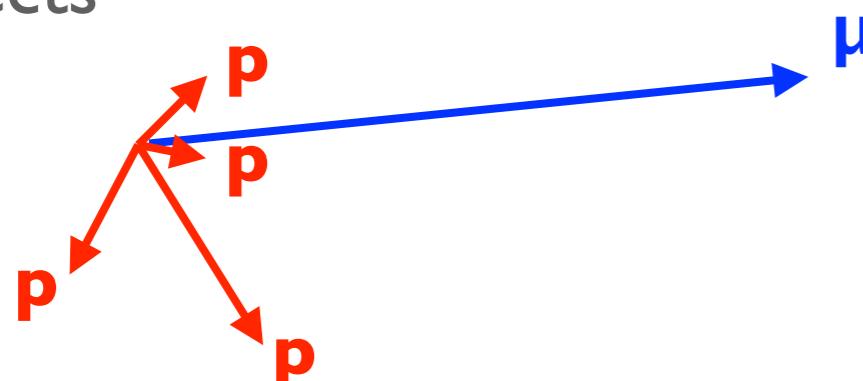


I or more proton candidates

- CCQE and CC2p2h (and other processes with FSI)
- general probe of nuclear effects

Exactly 1 proton candidate:  
CCQE-like kinematics

- try to isolate CCQE or CCQE-like interactions

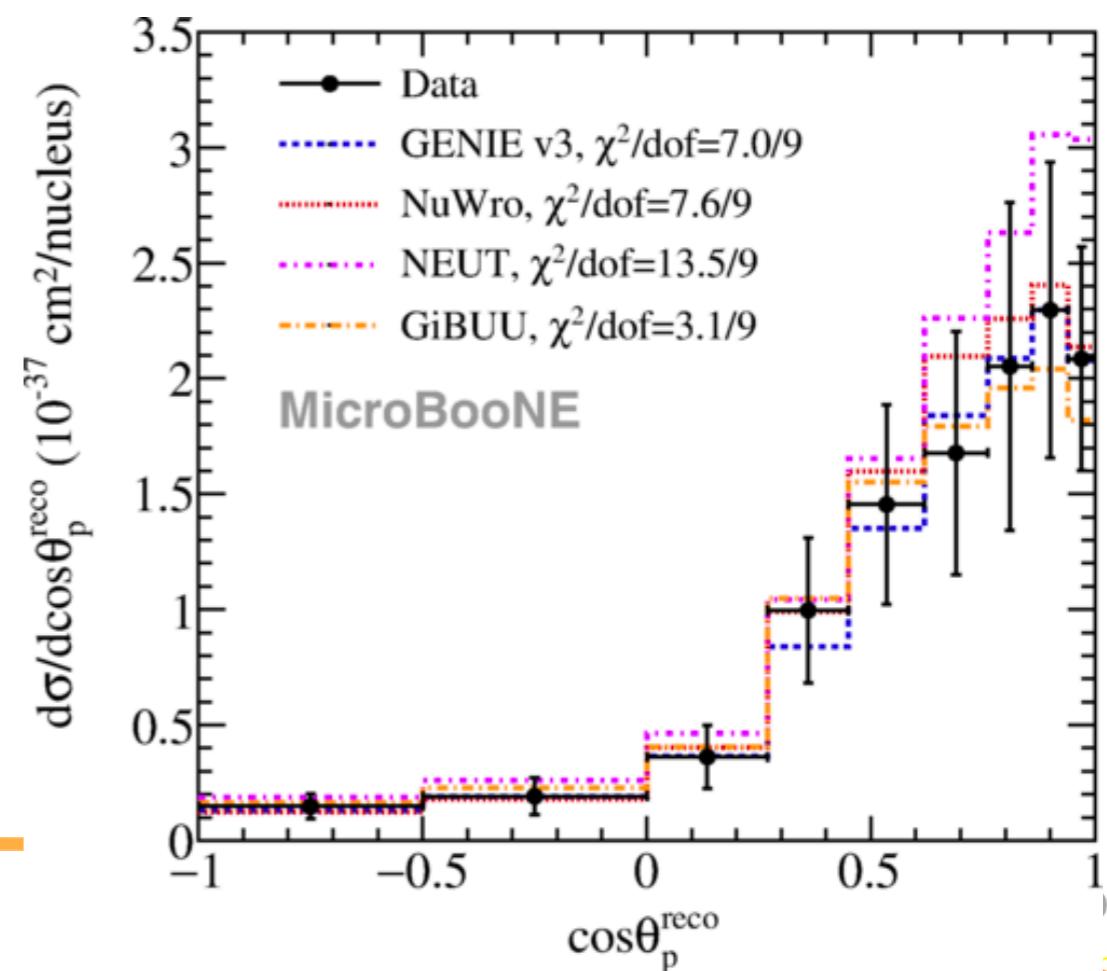
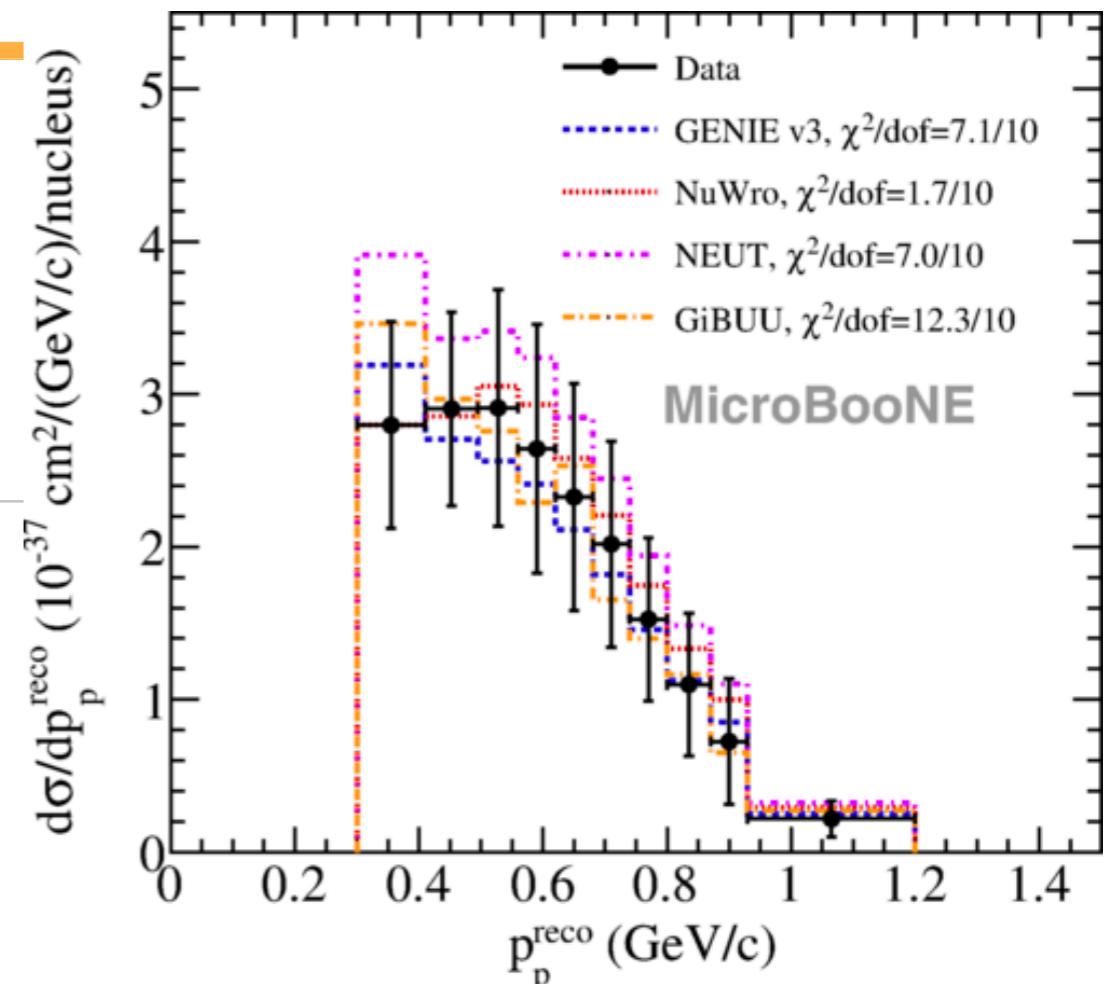
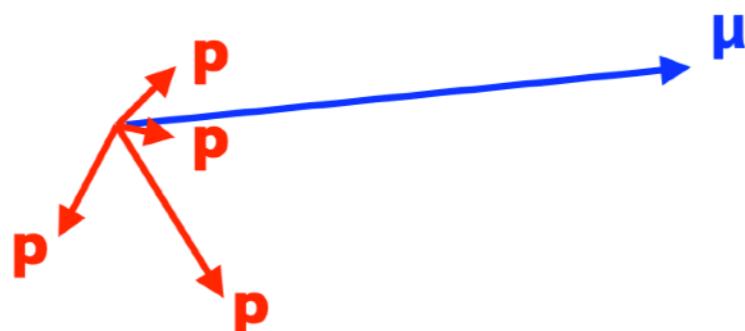


# CC0 $\pi$ Np ( $N \geq 1$ ) CROSS SECTION

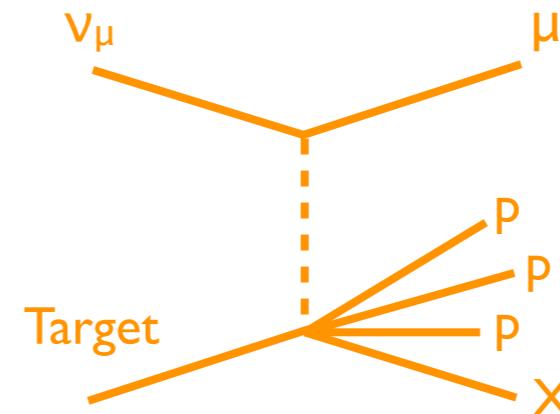
Phys. Rev. D 102, 112013 (2020)

Signal:

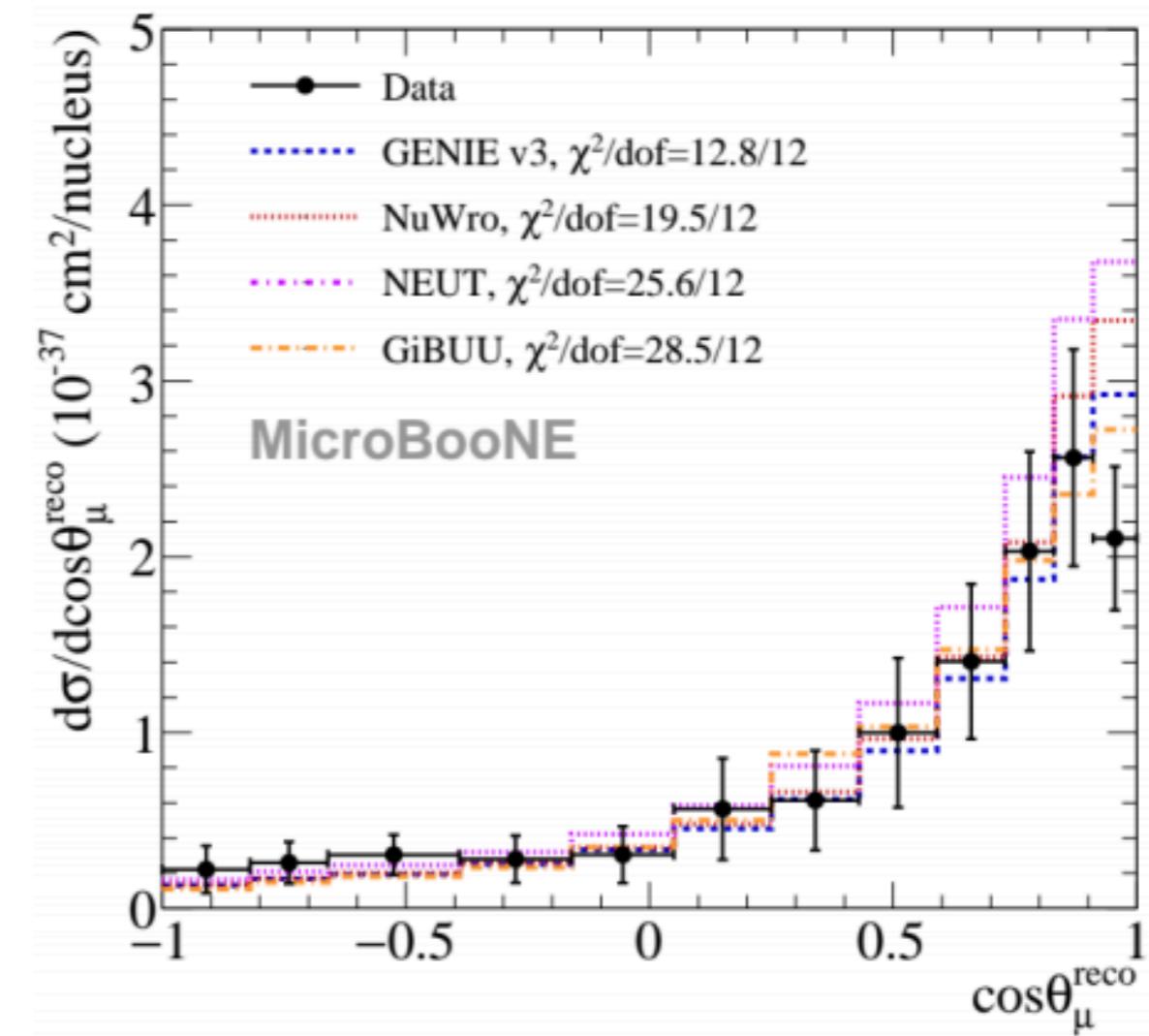
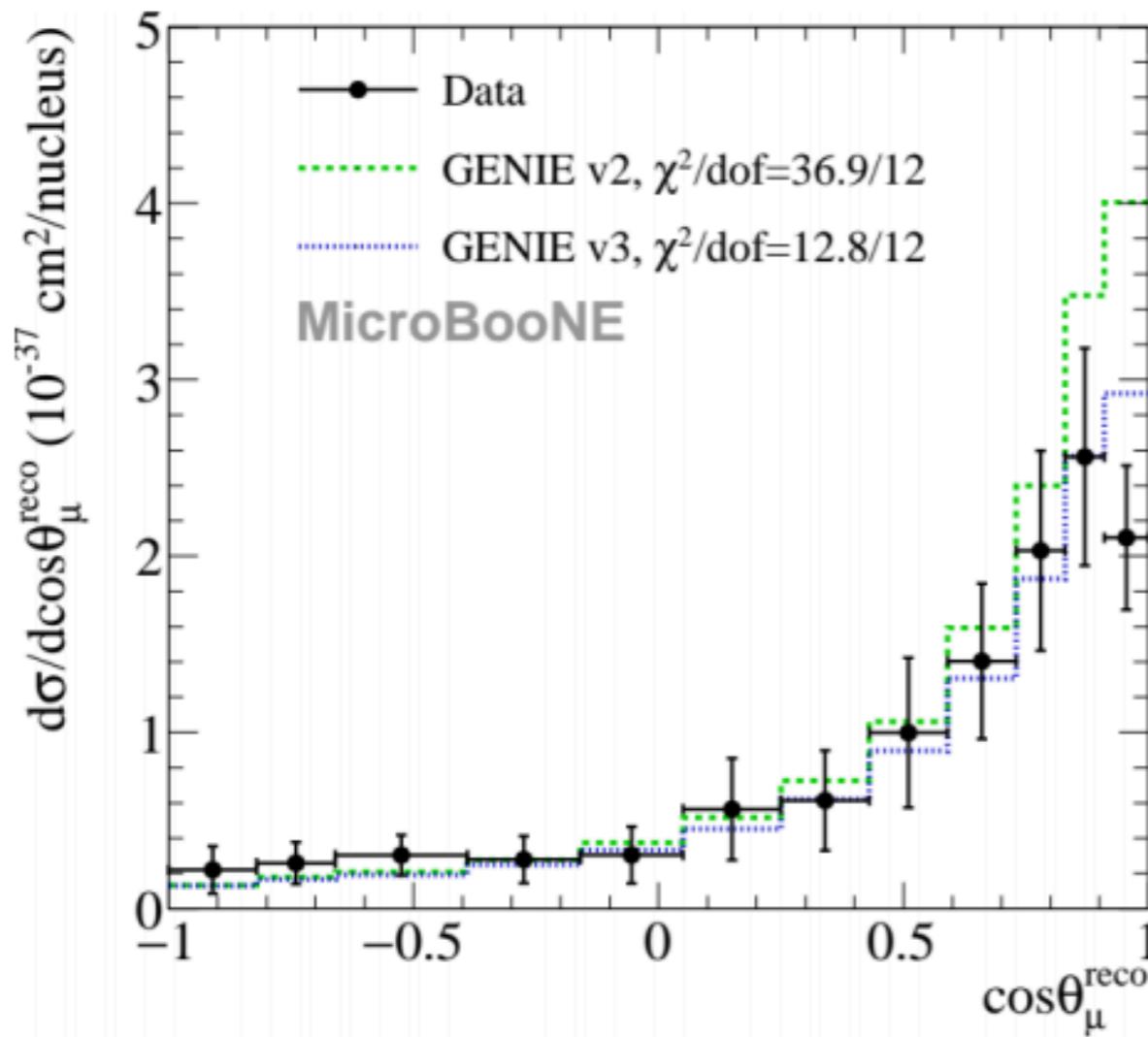
- 1 muon ( $p_\mu > 100$  MeV/c)
  - At least 1 proton  
( $300 < p_p < 1200$  MeV/c)
  - No pions
- **71% purity, 29% efficiency**
- Proton momentum and angle show **reasonable agreement** with generators
  - Lowest bin in proton momentum has not been seen before — **Low thresholds = new information** about proton kinematics



# CC0 $\pi$ Np ( $N \geq 1$ ) CROSS SECTION



Phys. Rev. D 102, 112013  
(2020) also includes  
measurement as a function  
of muon momentum,  
muon-proton opening angle



- Big over-prediction at forward-going angles
- Models with RPA do much better, but not quite enough

# CCQE-LIKE CROSS SECTION

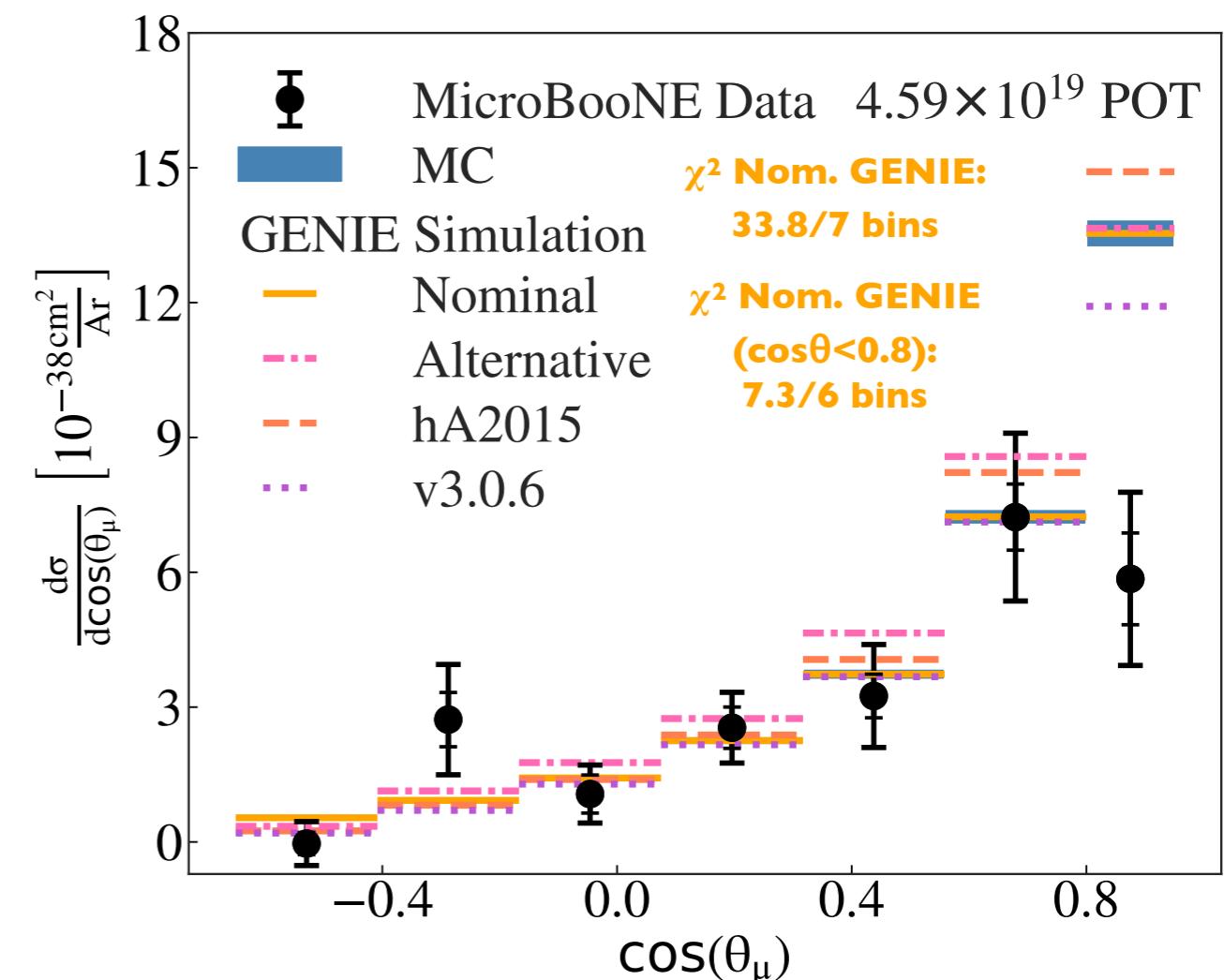
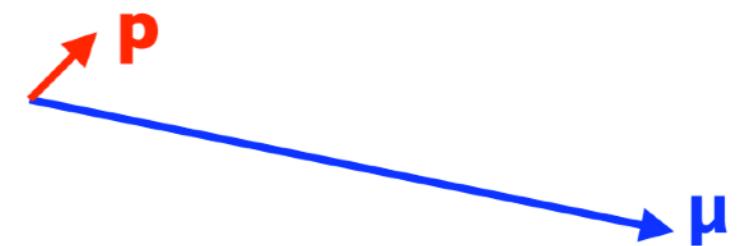
Eur. Phys. J. C 79 673 (2019) Phys. Rev. Lett. 125, 201803 (2020)

**First extraction of  $\nu_\mu$ - $^{40}\text{Ar}$  CCQE-like cross section using a surface LArTPC**



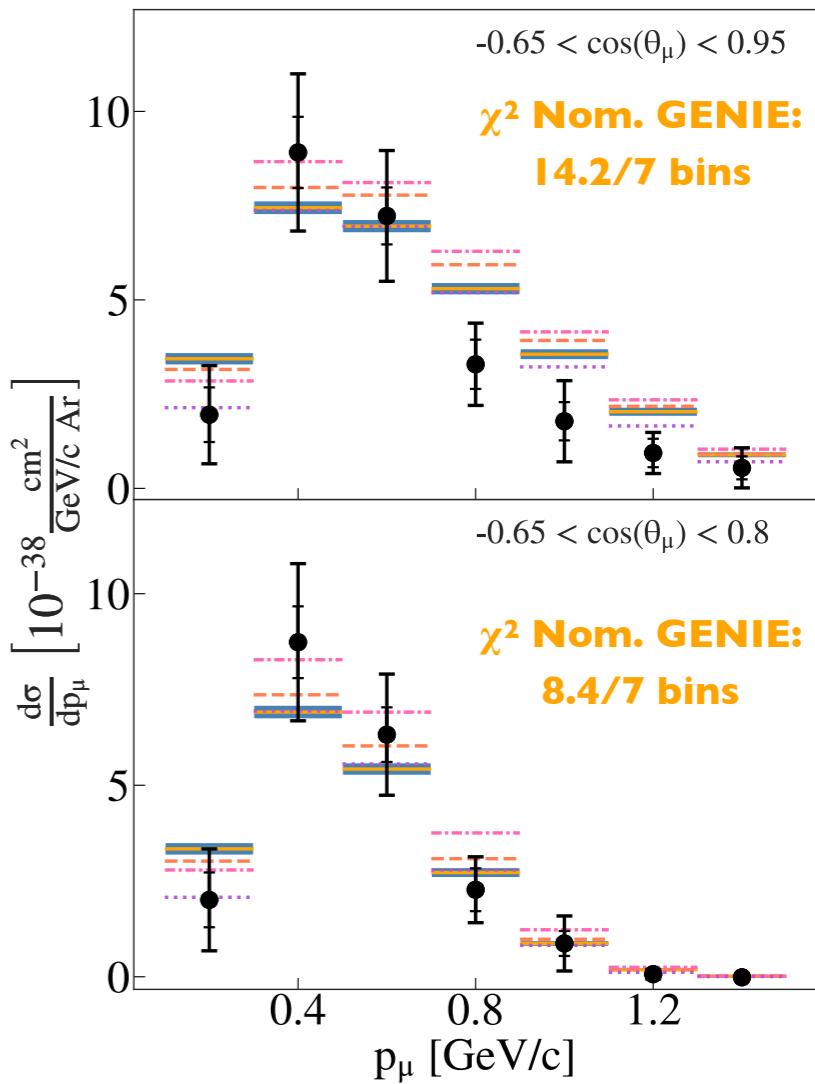
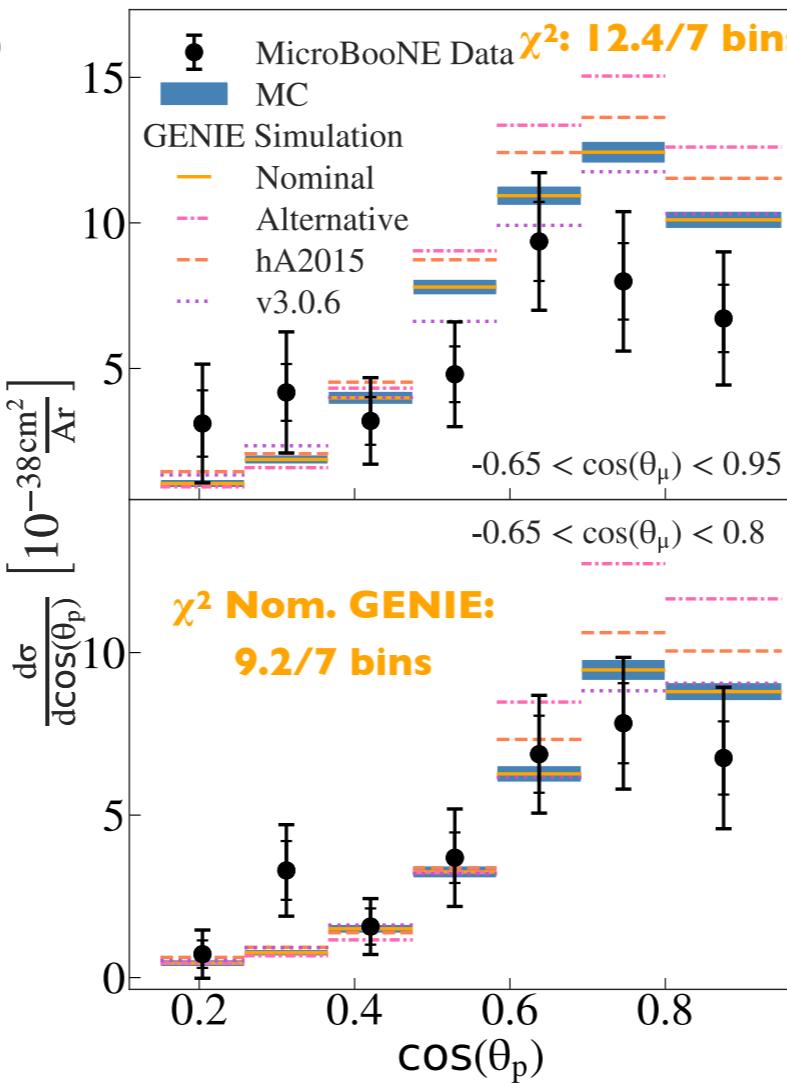
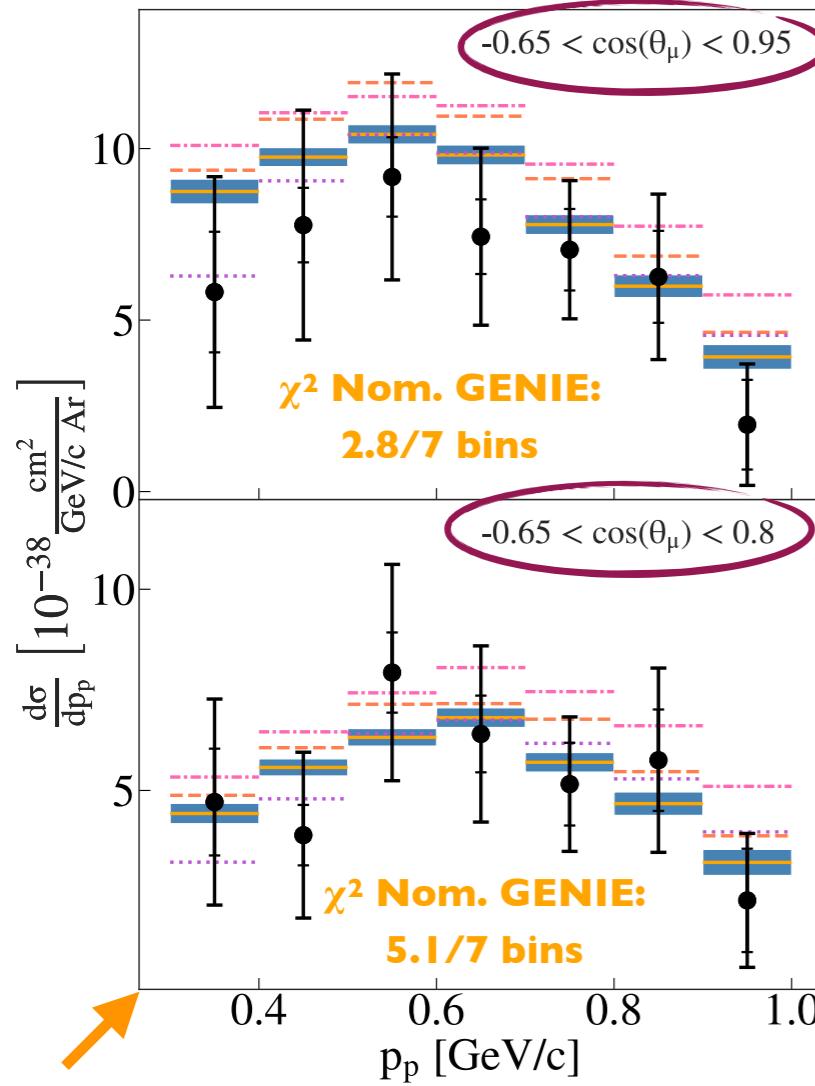
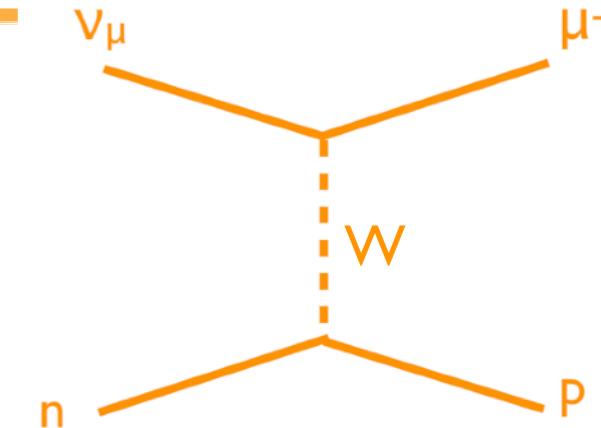
→ ~84% CC1p0π (~81% CCQE) purity,  
~20% efficiency

Good agreement with models, except at  
very **forward muon scattering angles**



# CCQE-LIKE CROSS SECTION

Eur. Phys. J. C 79 673 (2019) Phys. Rev. Lett. 125, 201803 (2020)



Proton momentum threshold 300 MeV/c

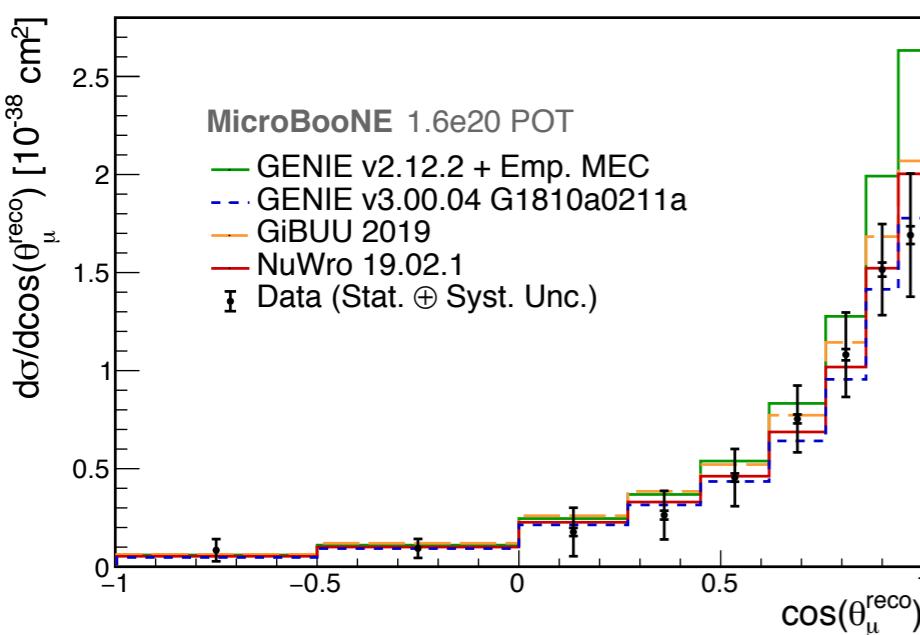
**Across all kinematic variables, agreement is improved if forward muon angles are excluded**

# THAT FORWARD-ANGLE BIN: A CONSISTENT STORY

Phys. Rev. Lett. 123, 131801 (2019) Phys. Rev. D 102, 112013 (2020) Eur. Phys. J. C 79 673 (2019) Phys. Rev. Lett. 125, 201803 (2020)

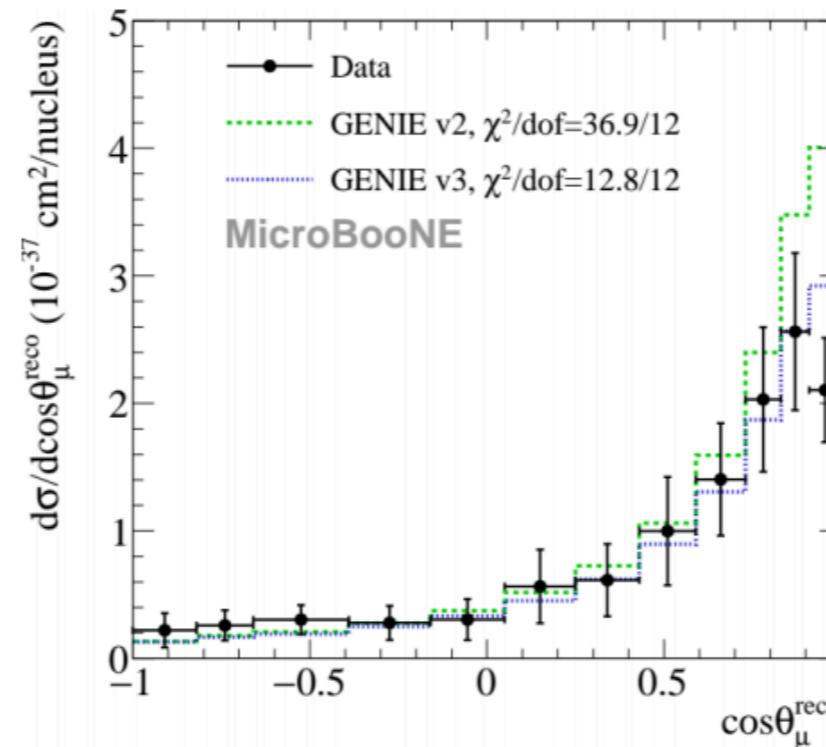
All three compare to the same GENIE models → cross-comparison

**GENIE v3**



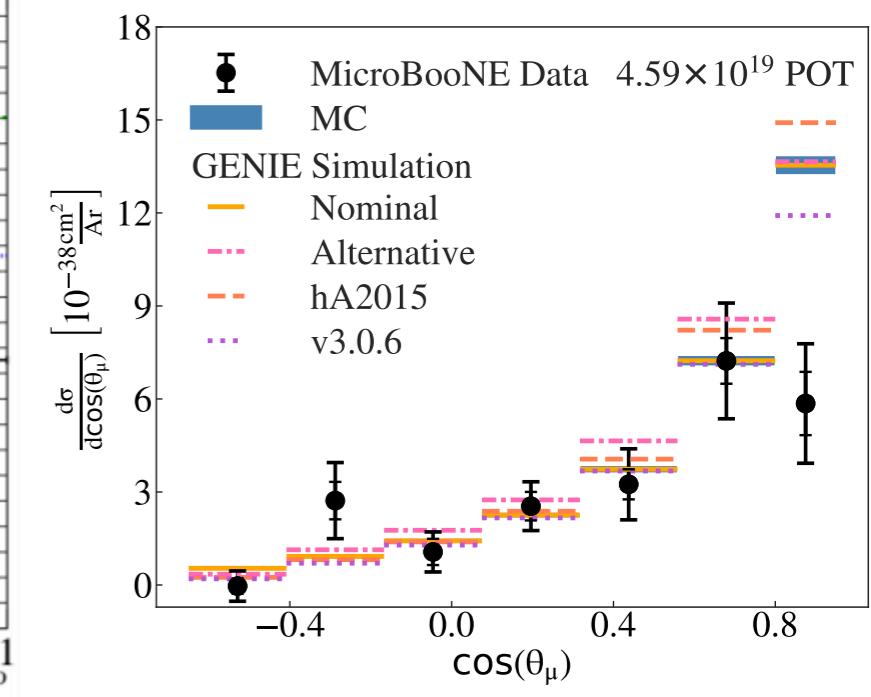
**CC Inclusive**  
Inclusive  
Some deficit

**GENIE v2 GENIE v3**



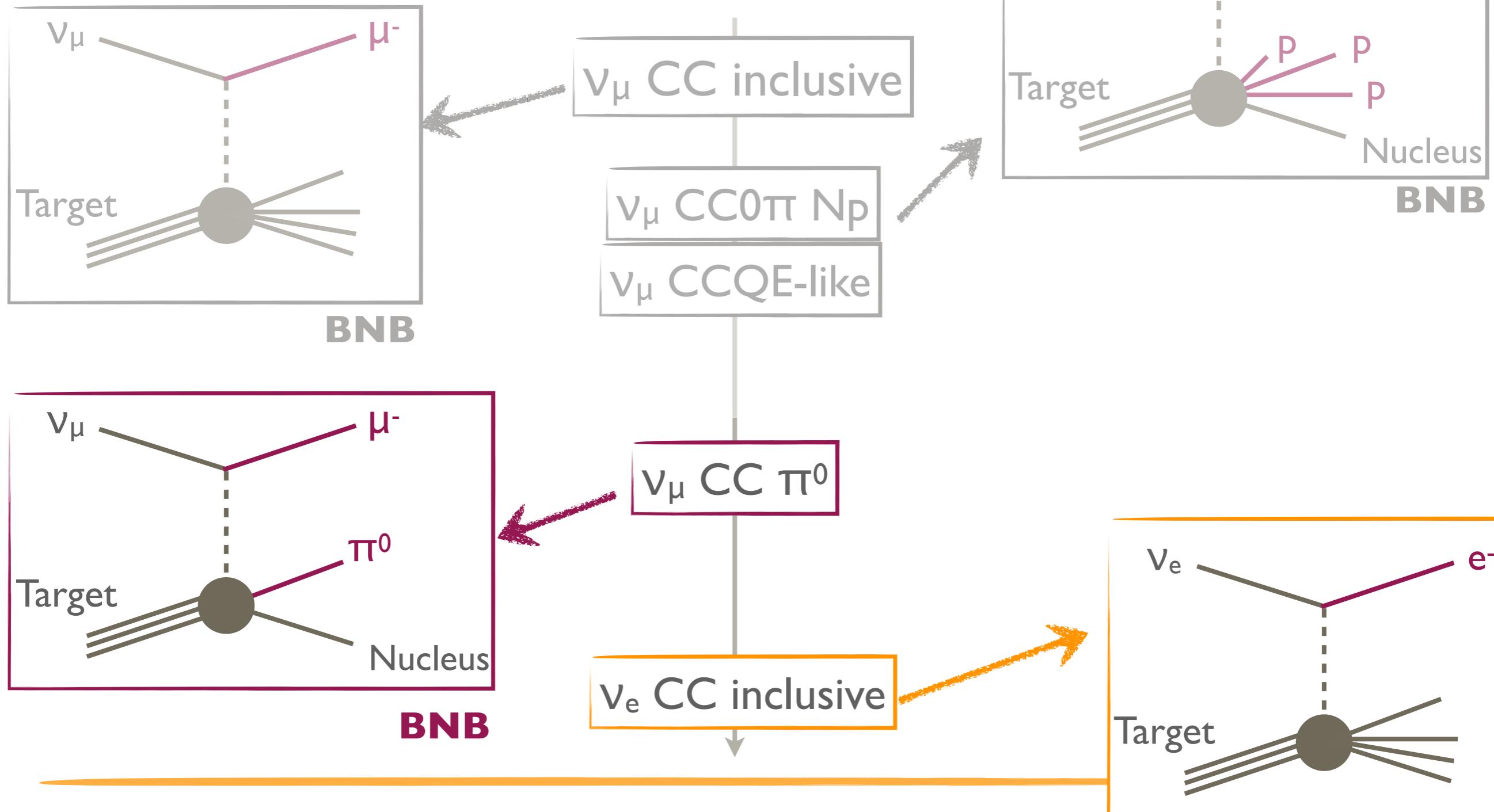
**CC0πNp**  
More exclusive  
Turnover in data

**GENIE v2 GENIE v3**



**CCQE-like**  
Even more exclusive  
Even more deficit

# TODAY'S CROSS SECTION MEASUREMENTS

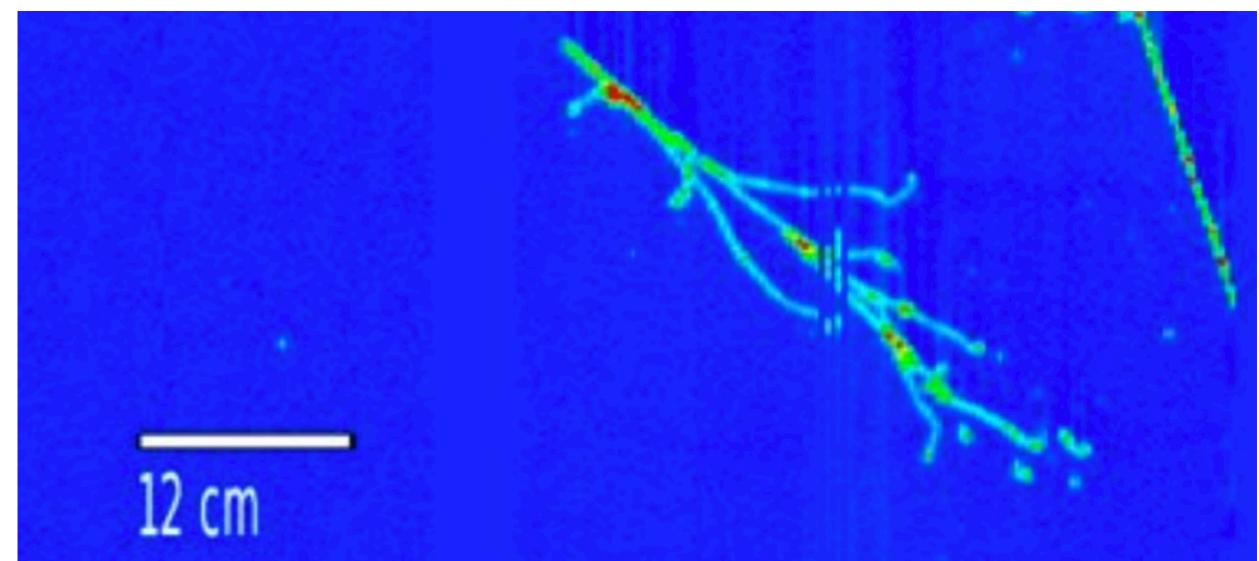


# LArTPC STRENGTH: ELECTRONS AND PHOTONS

Phys. Rev. D 99, 091102(R) (2019)

- **Electrons and photons produce showers in LArTPCs**

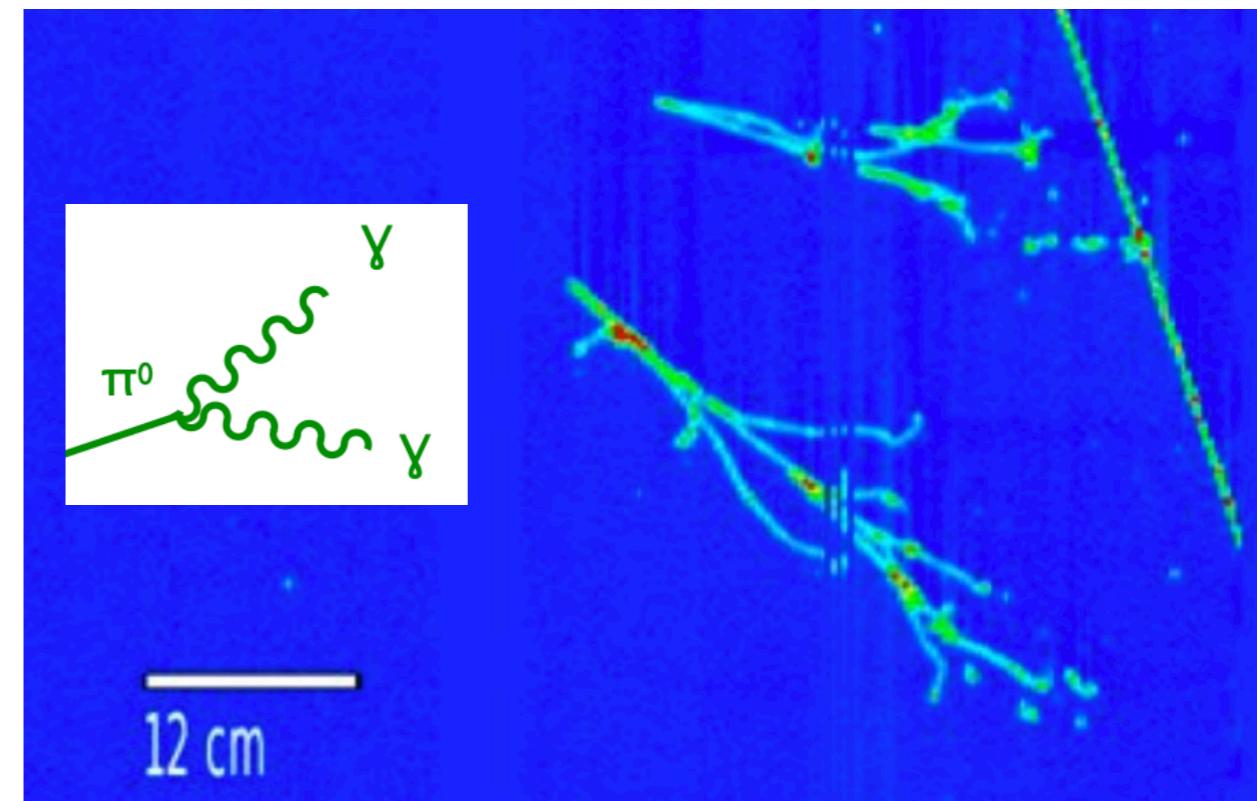
→ important to understand for  $\nu_e$  appearance searches in SBN and DUNE



# LArTPC STRENGTH: ELECTRONS AND PHOTONS

Phys. Rev. D 99, 091102(R) (2019)

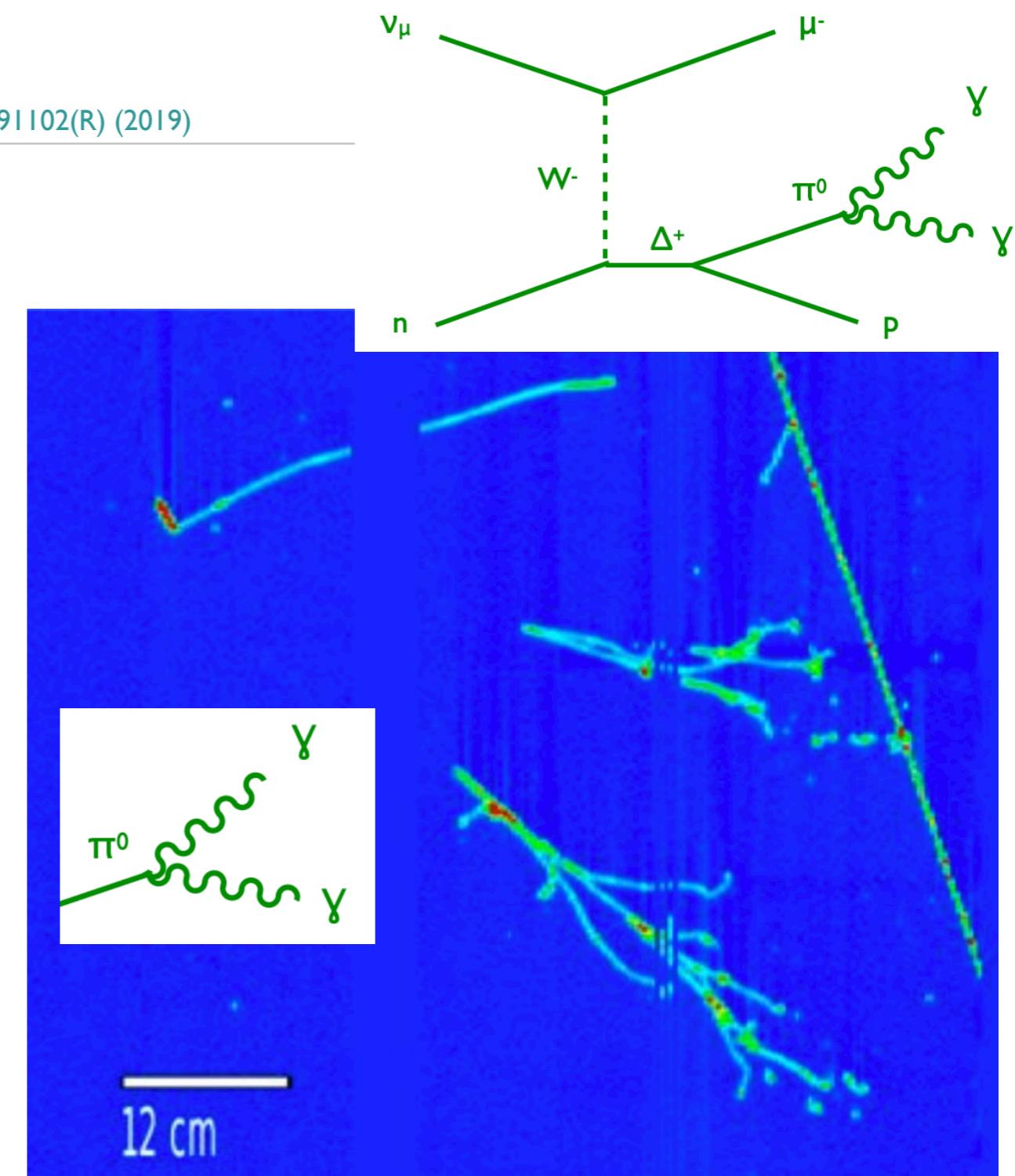
- **Electrons and photons produce showers in LArTPCs**  
→ important to understand for  $\nu_e$  appearance searches in SBN and DUNE
- $\pi^0$  interactions are a background (although often can be distinguished by energy deposition)



# LArTPC STRENGTH: ELECTRONS AND PHOTONS

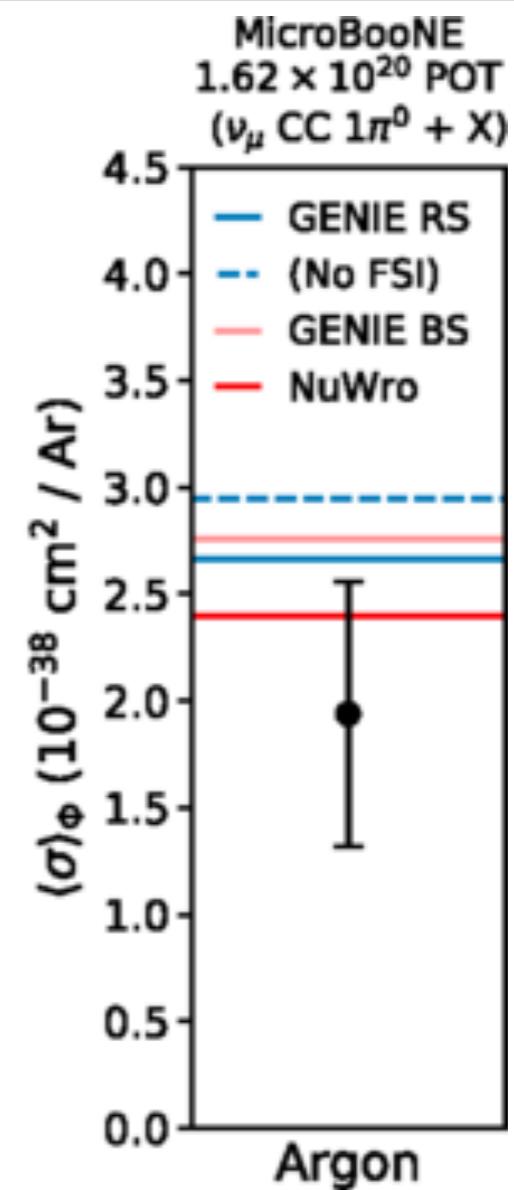
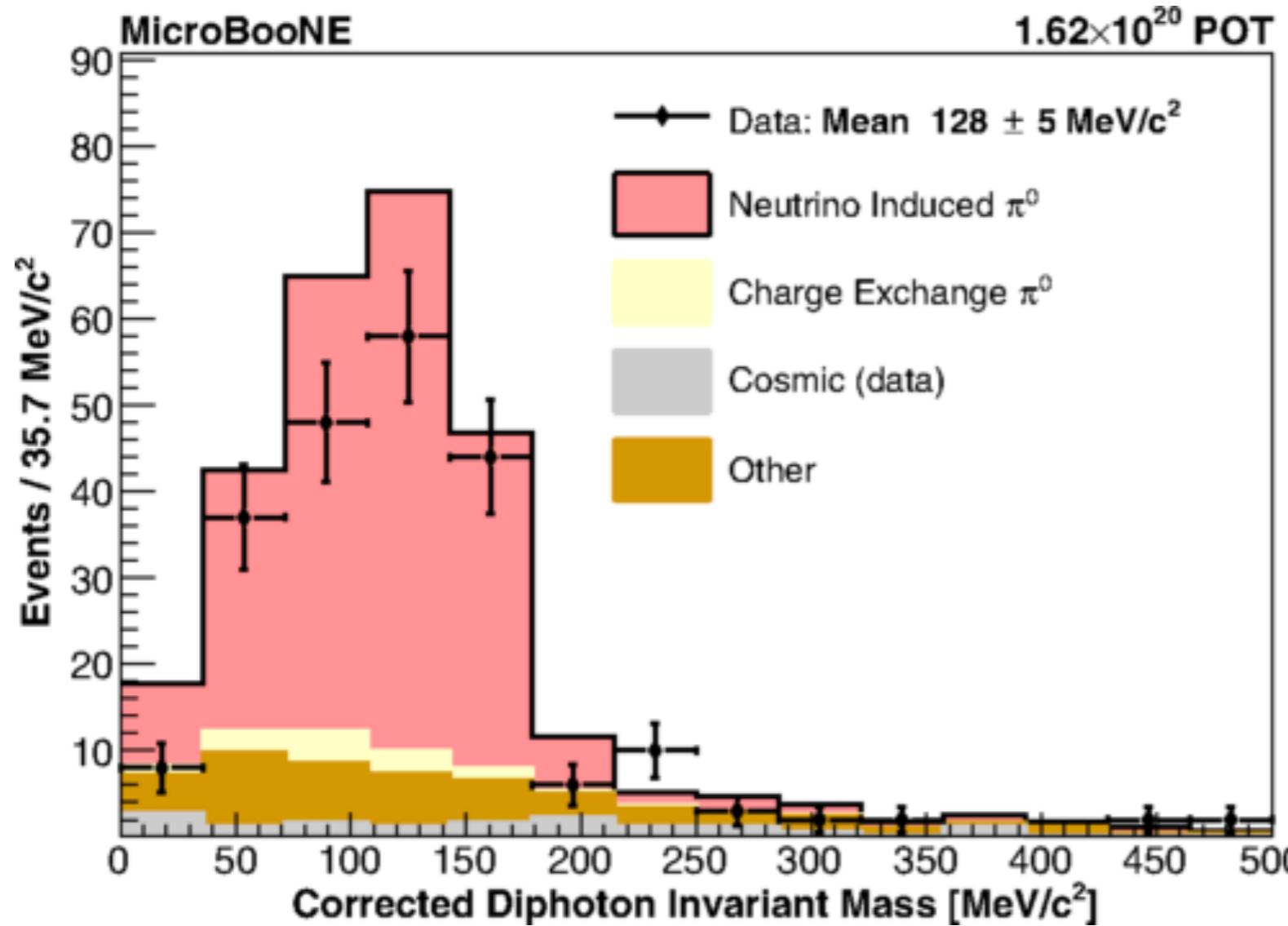
Phys. Rev. D 99, 091102(R) (2019)

- **Electrons and photons produce showers in LArTPCs**
  - important to understand for  $\nu_e$  appearance searches in SBN and DUNE
- $\pi^0$  interactions are a background (although often can be distinguished by energy deposition)
  - → **measurement of CC $\pi^0$  production cross section**
  - → **verify shower reconstruction** by reconstructing  $\pi^0$  mass peak



# $\text{CC}\pi^0$ CROSS SECTION MEASUREMENT

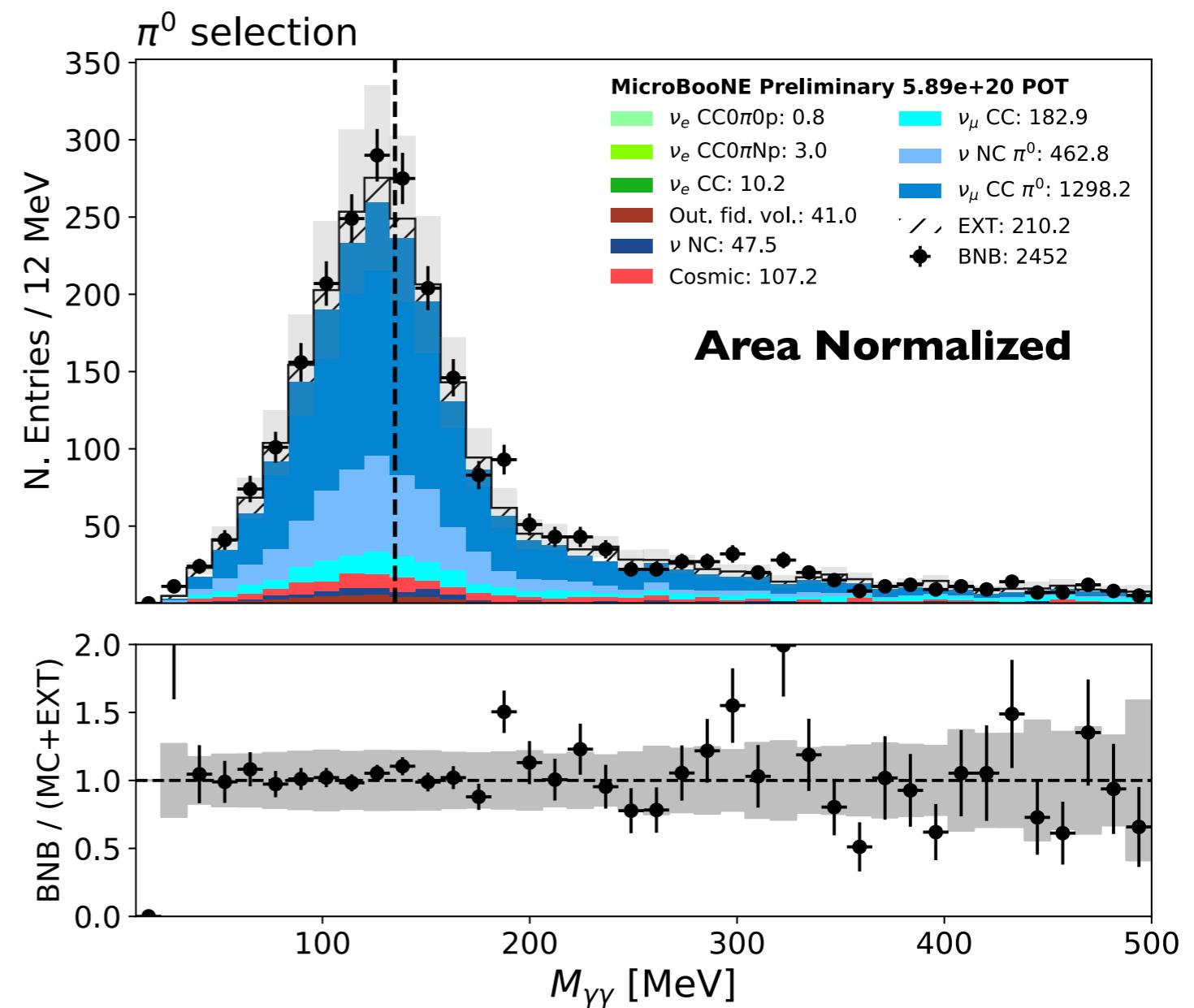
Phys. Rev. D 99, 091102(R) (2019)



# $\pi^0$ MEASUREMENTS: FUTURE PROSPECTS

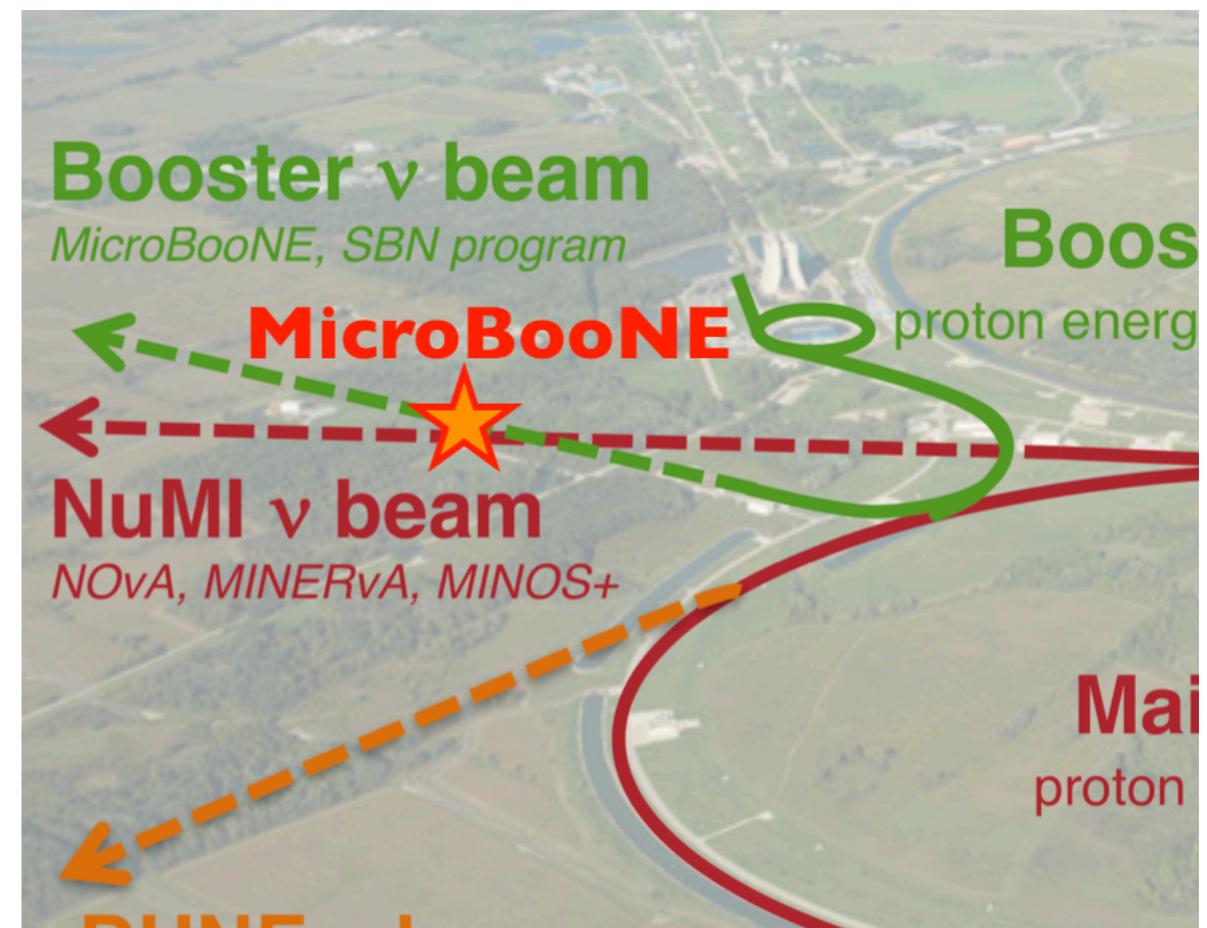
MICROBOONE-NOTE-1085-PUB

- With improvements in detector simulation and reconstruction, current selections have **higher statistics** and **improved resolution**
- Working towards:
  - **Differential CC $\pi^0$**  cross section measurement
  - **NC $\pi^0$**  cross section measurement
  - **CC/NC** comparisons



# CHANGING TRACK

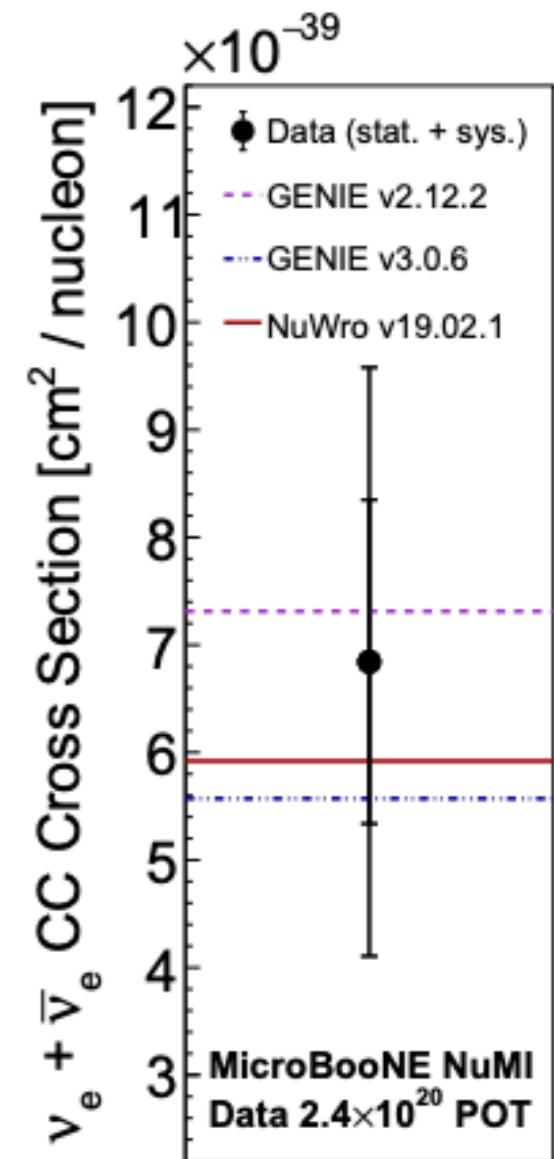
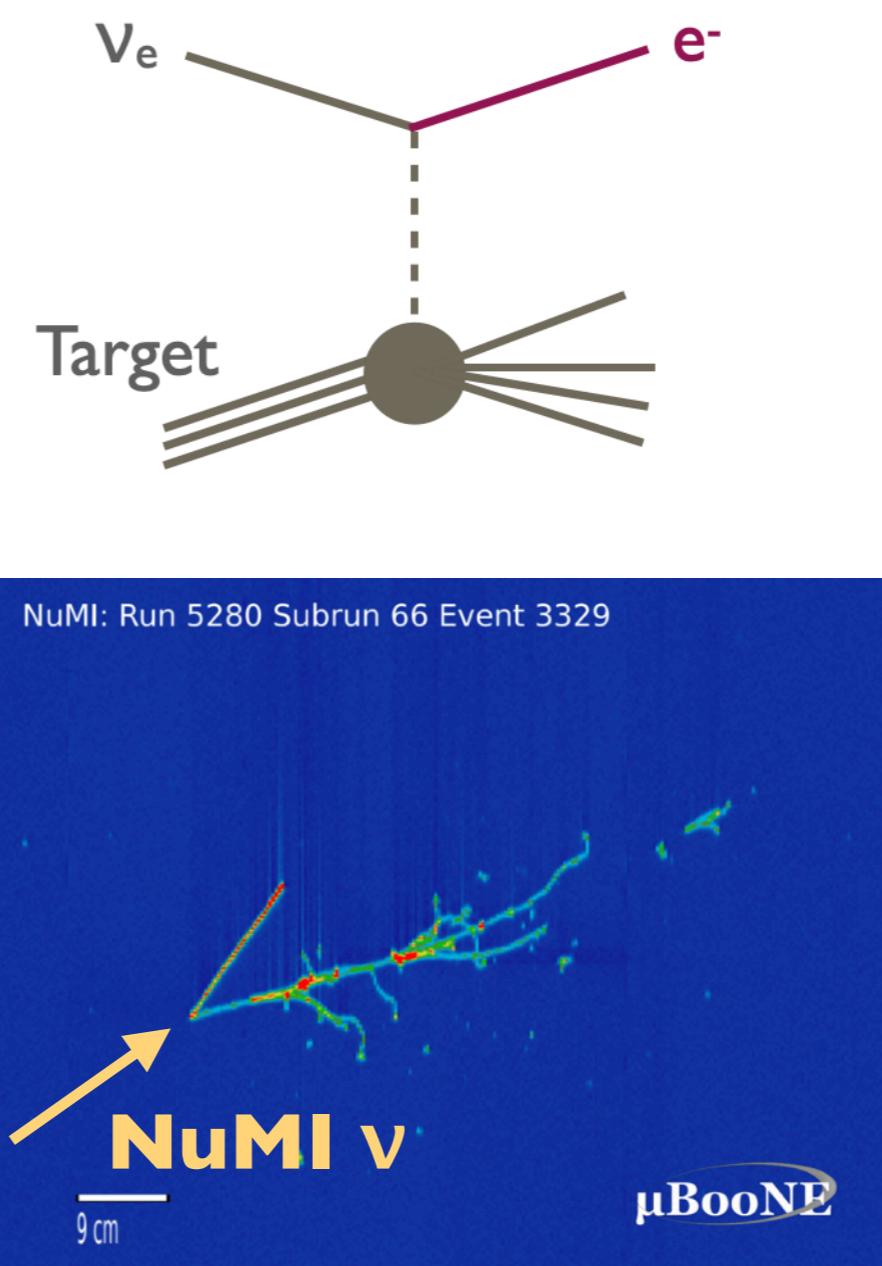
- Electron neutrinos are hard to study: oscillation experiments' neutrino beams are intentionally mostly  $\nu_\mu$
- But electron neutrinos are what we want to measure in an oscillation analysis!  
Important to study and understand
- Off-axis NuMI flux: ~5%  $\nu_e$



# ELECTRON NEUTRINO CROSS SECTION MEASUREMENT

arXiv:2101.04228[hep-ex]

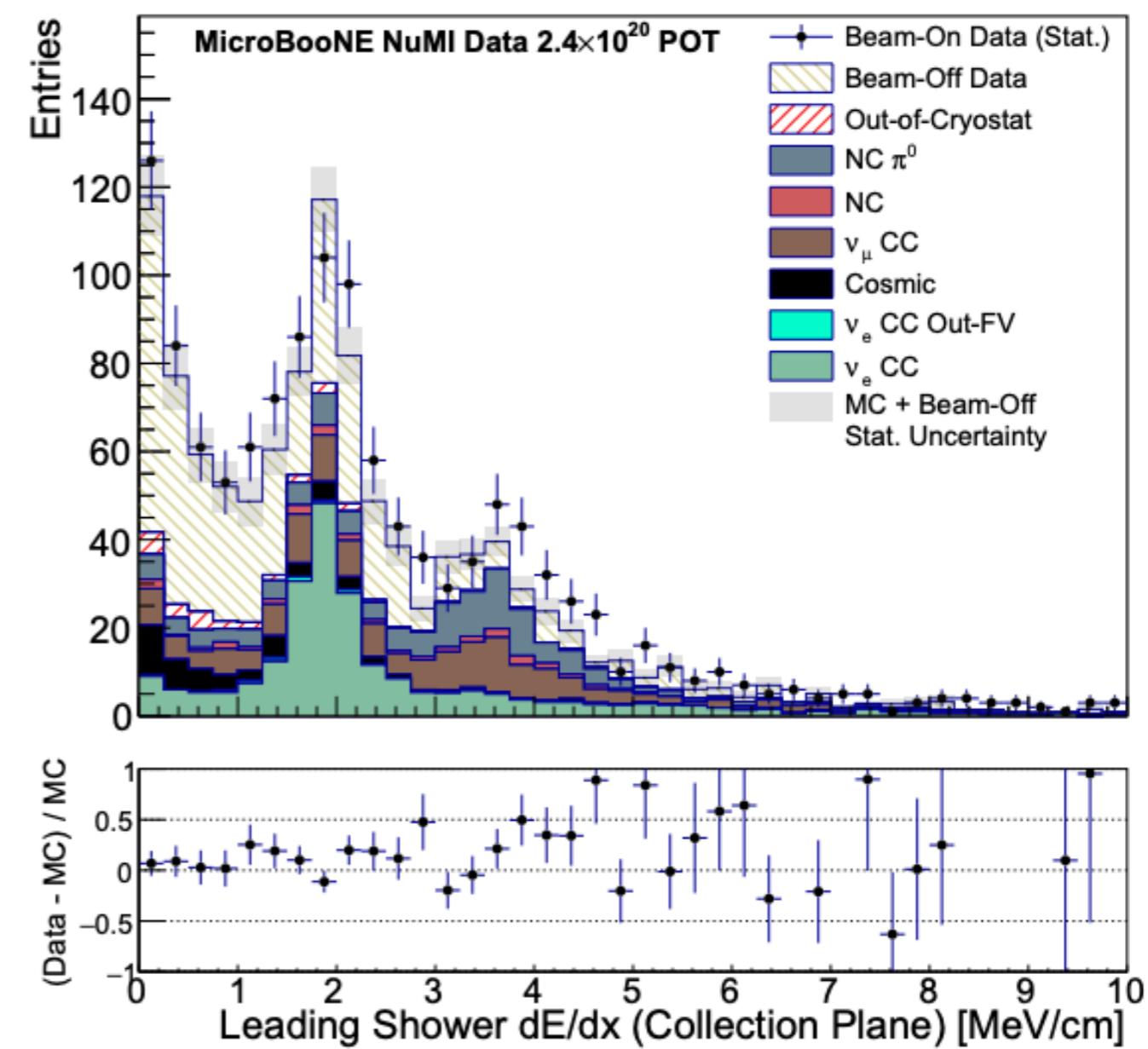
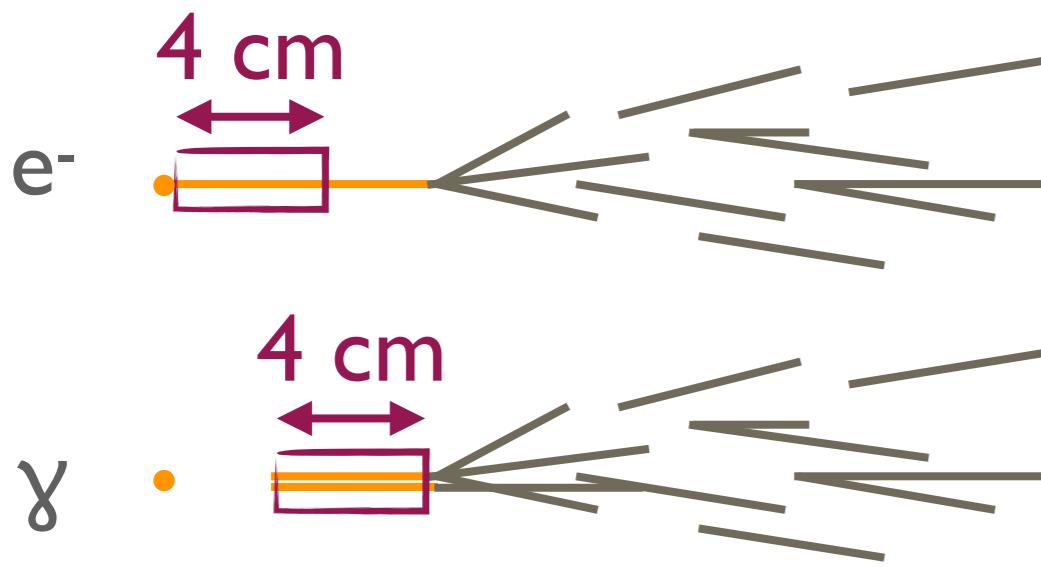
- **$\nu_e + \bar{\nu}_e$  inclusive cross-section** measurement with NuMI beam:
  - Purity 40%, efficiency 9%
  - Purity excluding cosmics > 65%
  - → **~214 events** in  $2.4 \times 10^{20}$  POT
  - First demonstration of selecting  $\nu_e$  in a surface LArTPC



# ELECTROMAGNETIC SHOWERS

arXiv:2101.04228[hep-ex]

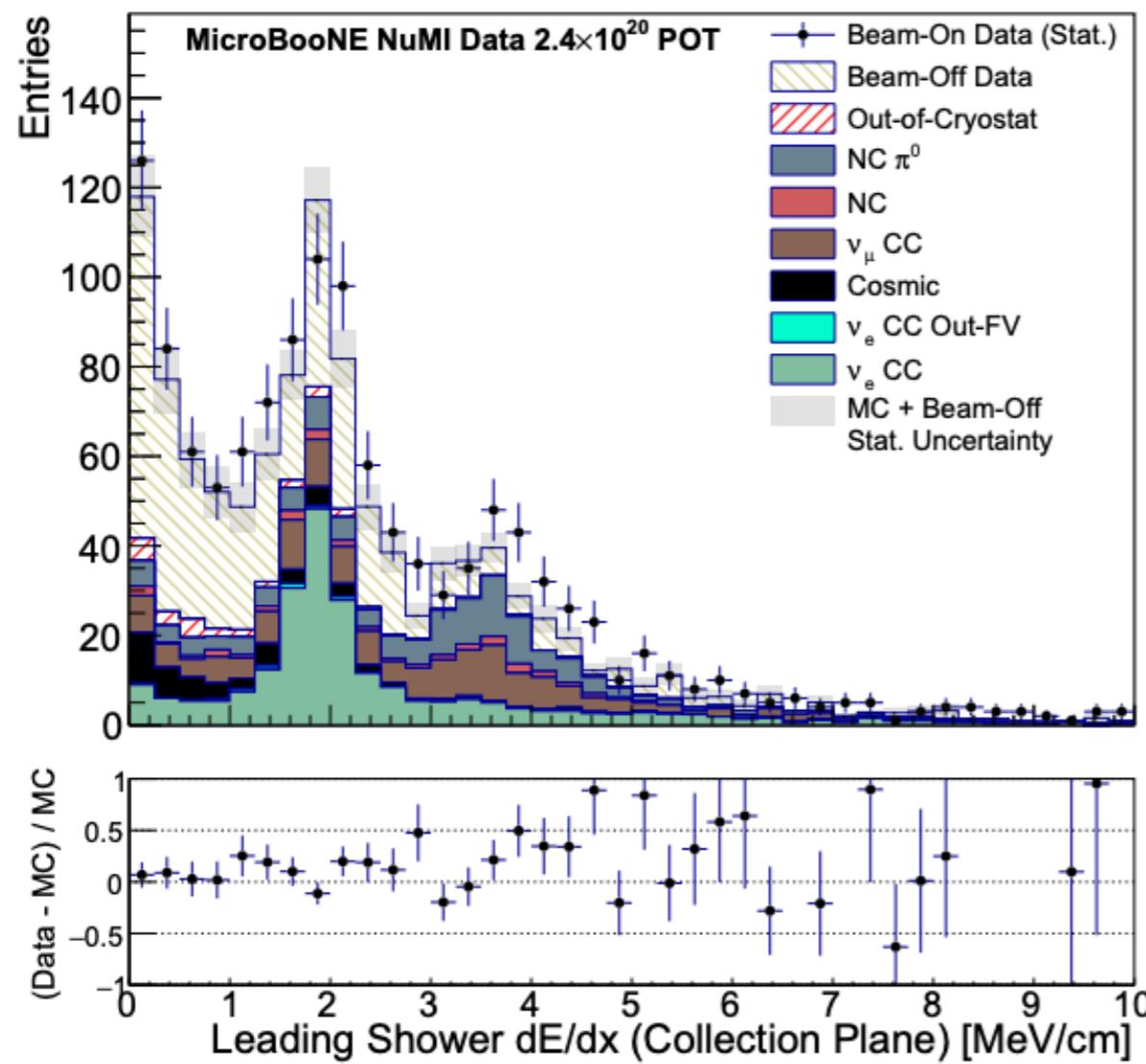
- Select  $\nu_e$  by looking for showers
- Distinguish from photons using  $dE/dx$  at start of shower
- One of the key capabilities of LAr detectors — in action!



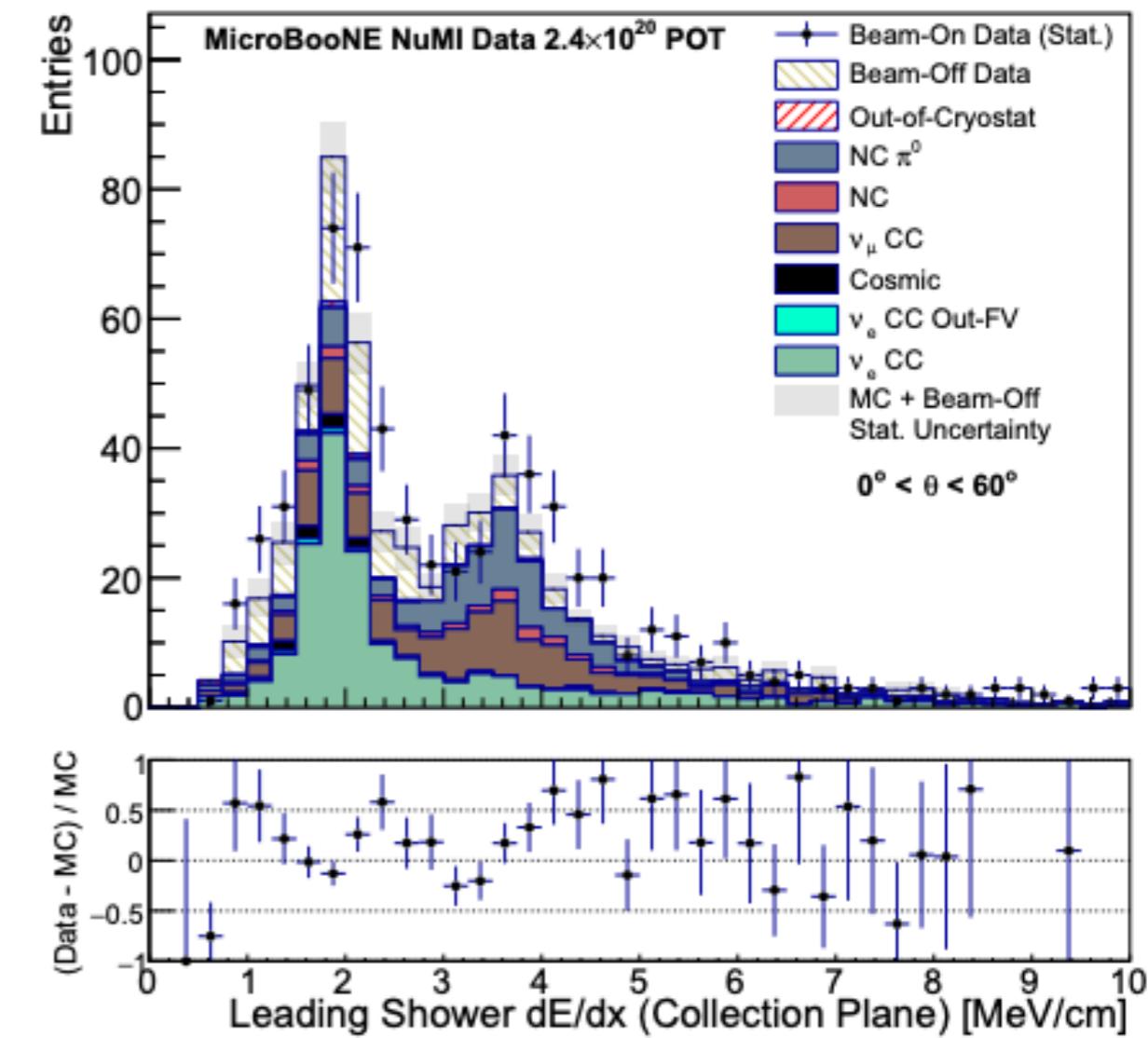
# ELECTROMAGNETIC SHOWERS

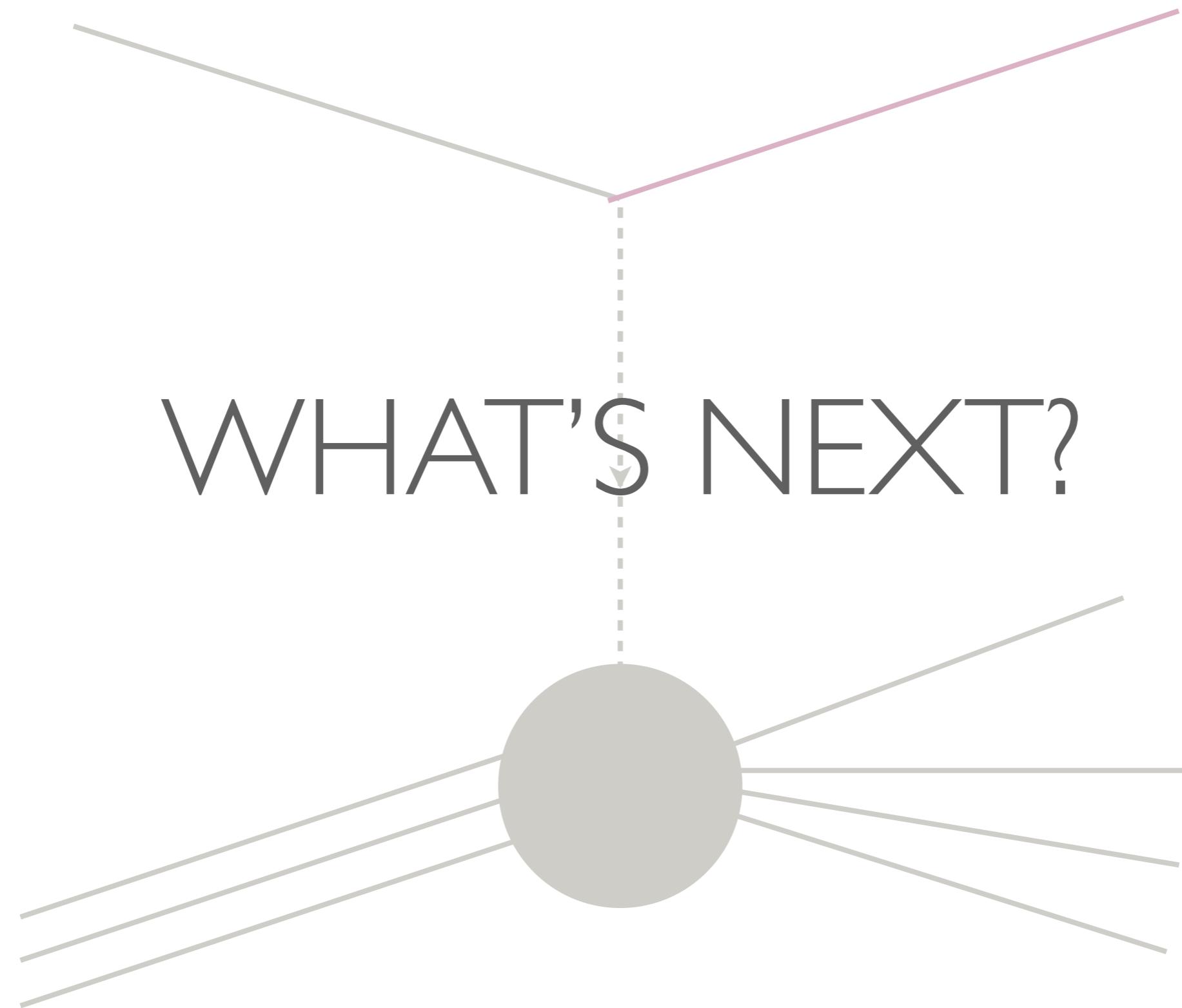
arXiv:2101.04228[hep-ex]

## All showers



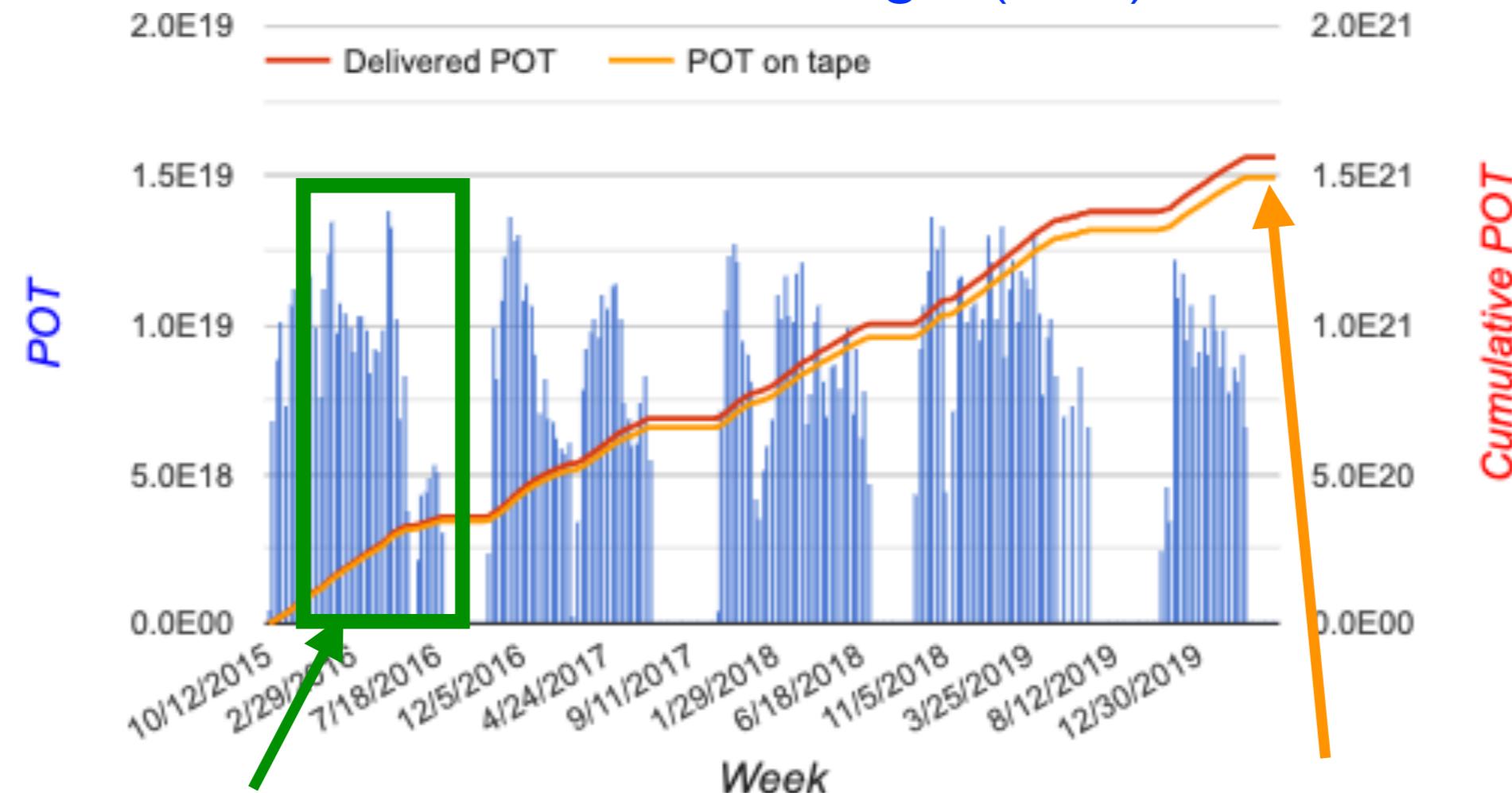
## $0 < \theta < 60^\circ$ : Perpendicular to wires





# (A LOT) MORE DATA

## BNB Data collection: Protons on Target (POT)



**These analyses:**  
 $\sim 1.6 \times 10^{20}$  POT  
(protons on target)

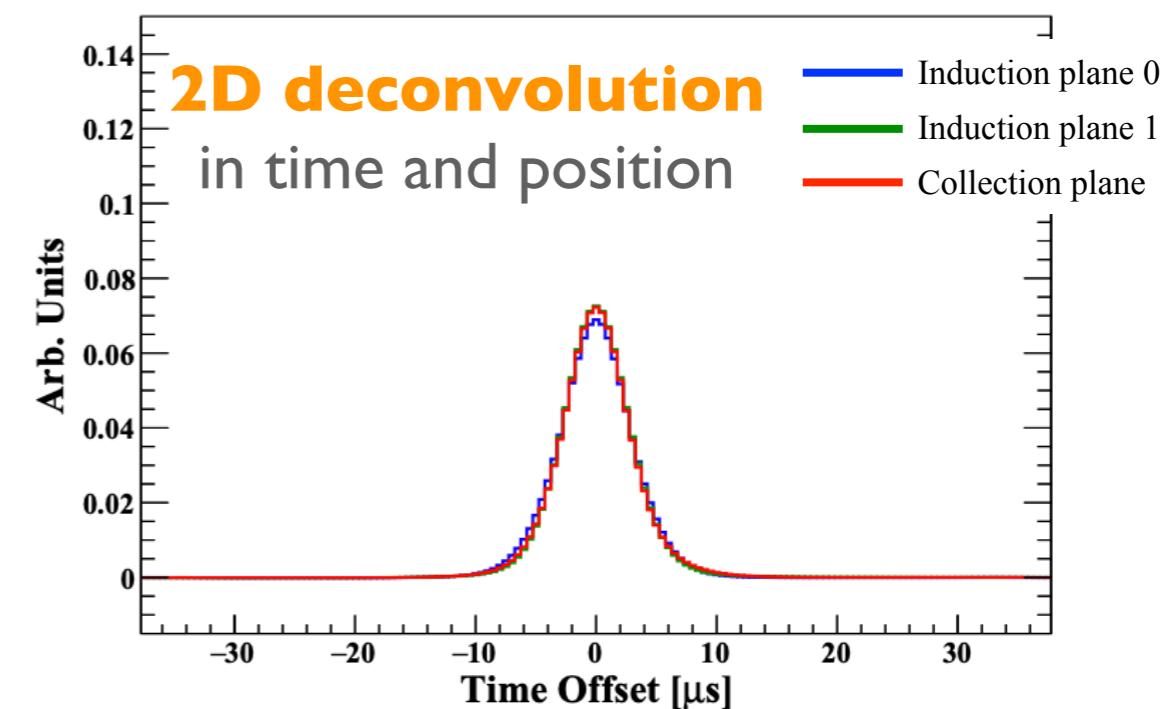
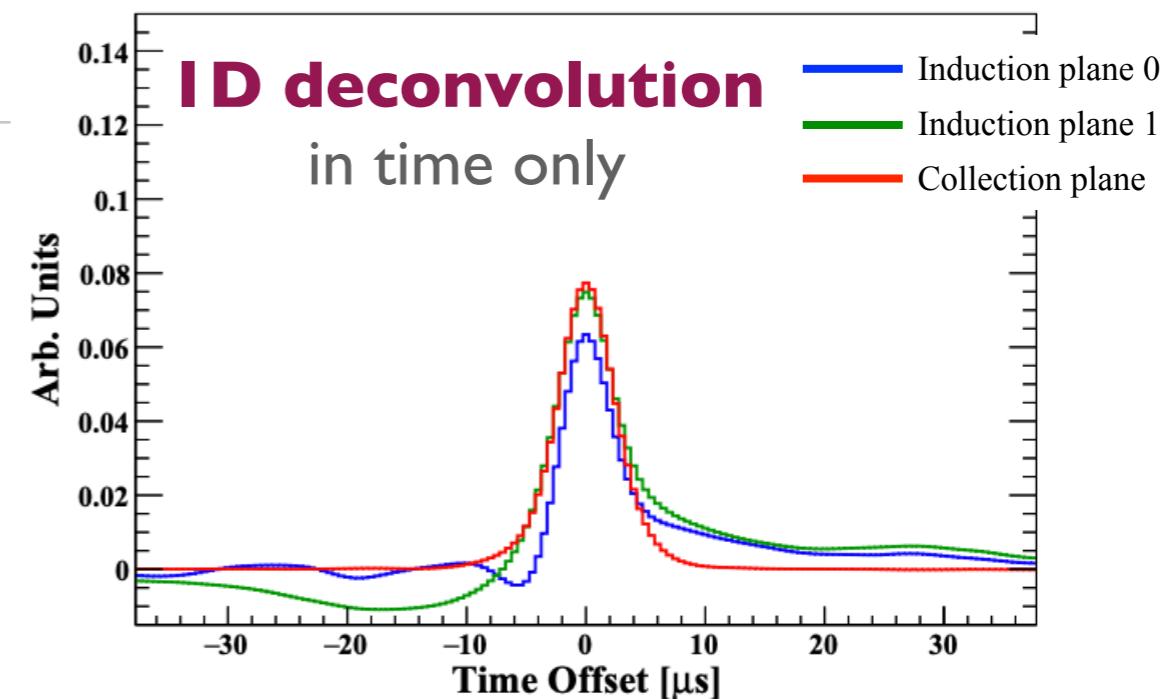
**Collected to date:**  
 $1.5 \times 10^{21}$  POT

# BETTER DATA AND MODELING

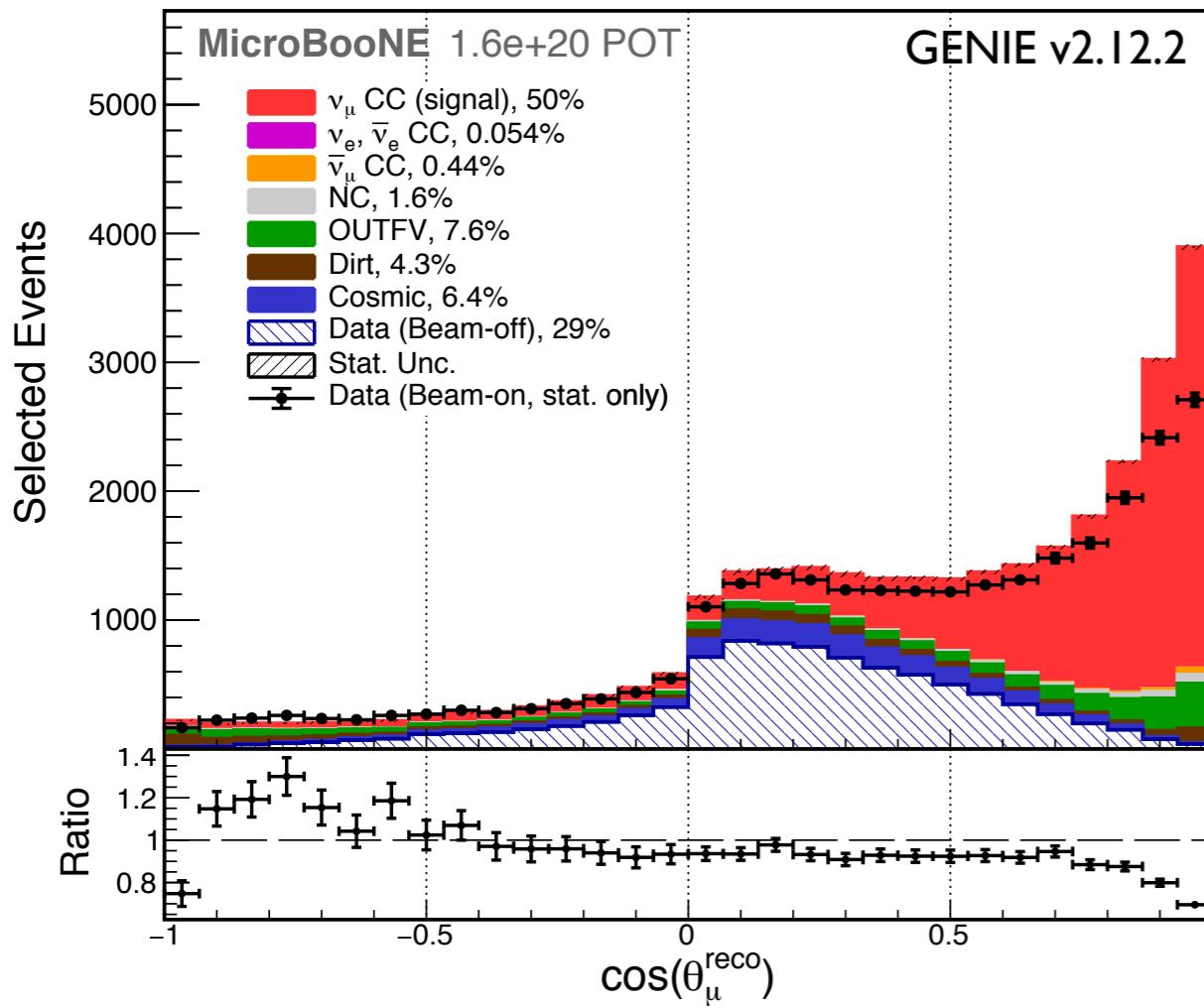
JINST 13 P07006 (2018)

JINST 13 P07007 (2018)

- MicroBooNE Collaboration has made **huge improvements** in our understanding of the detector in the past few years
- Detailed understanding of detector is **key to our R&D mission** for future LAr program
- **Improved signal processing** (2D deconvolution) accounts for interfering wire signals on all three planes
- Tracking is hard when particles go parallel to wires. Precise calorimetry on all planes → 3D tracking →  **$4\pi$  particle identification**



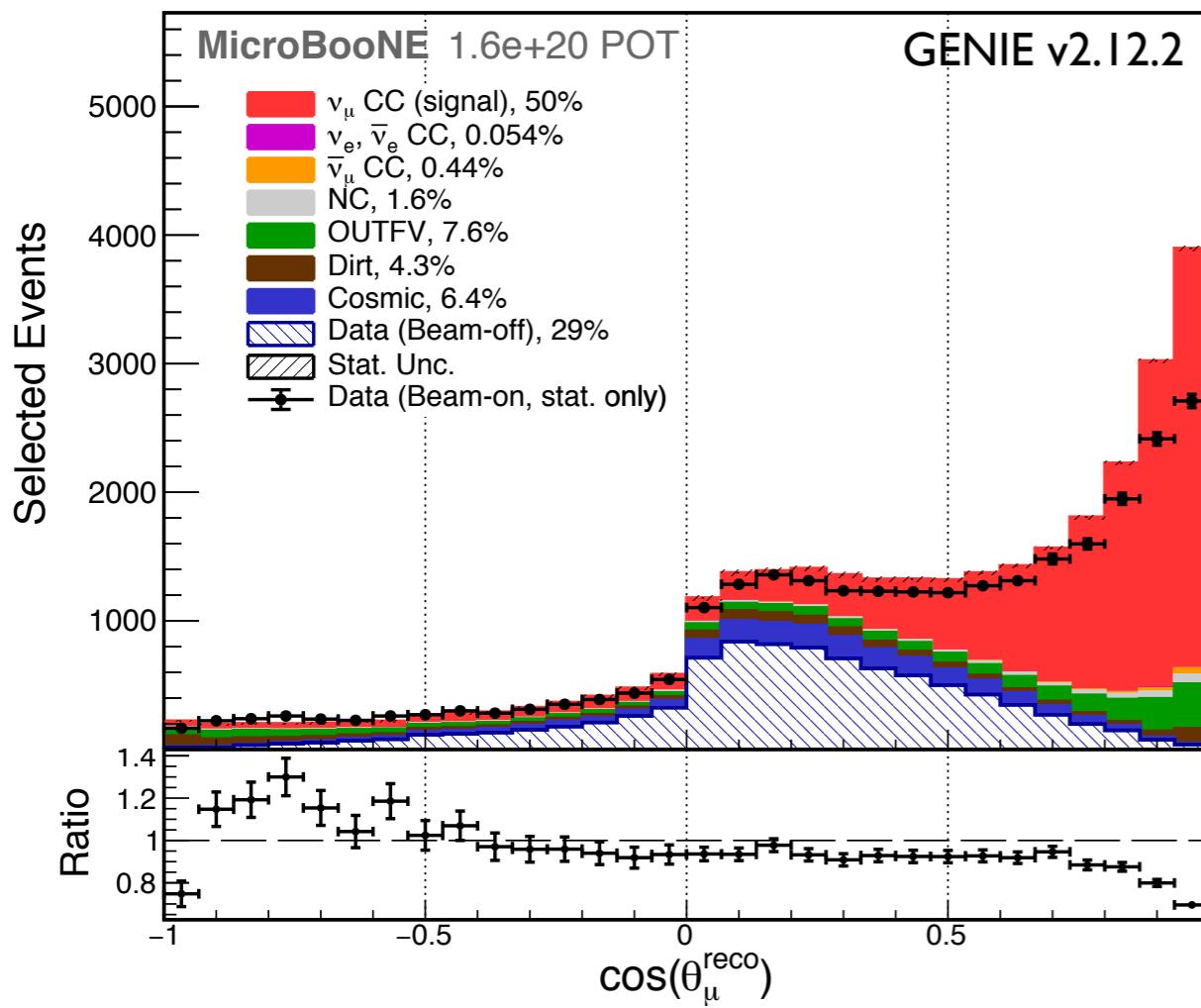
# IMPROVED DETECTOR UNDERSTANDING ENABLES BETTER MEASUREMENTS



## Previous Measurement

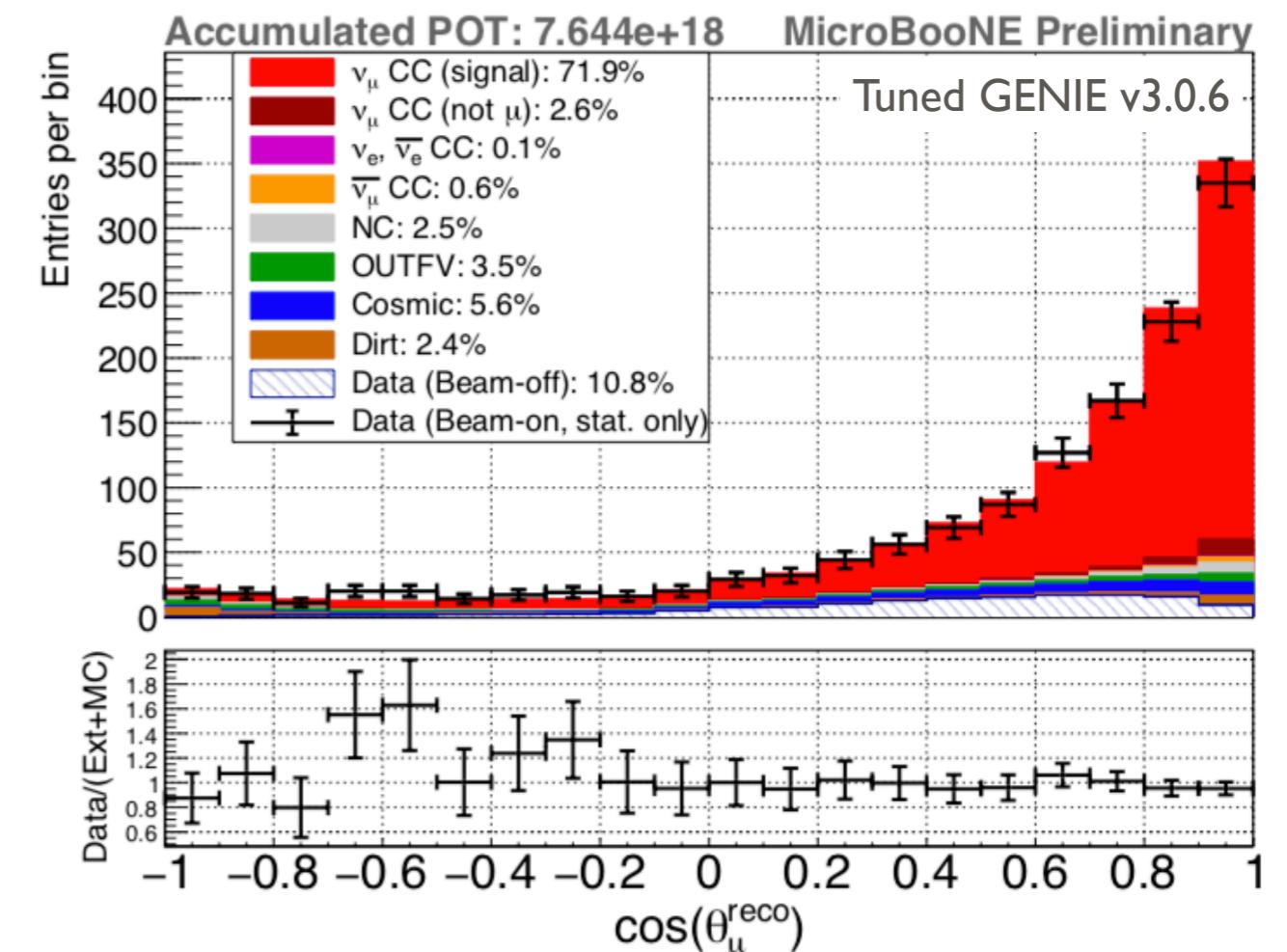
PRL 123, 131801 (2019)

# IMPROVED DETECTOR UNDERSTANDING ENABLES BETTER MEASUREMENTS



Previous Measurement

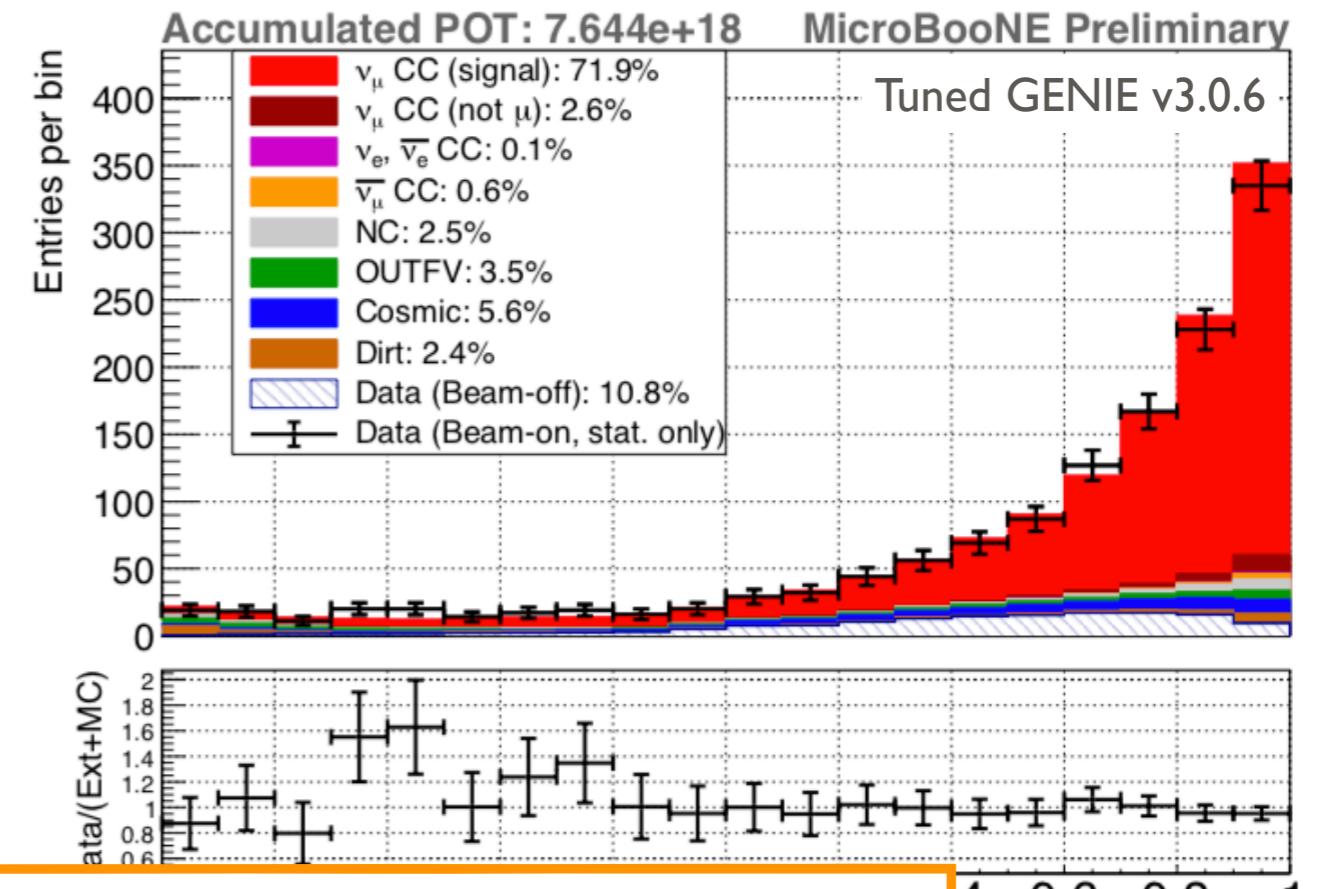
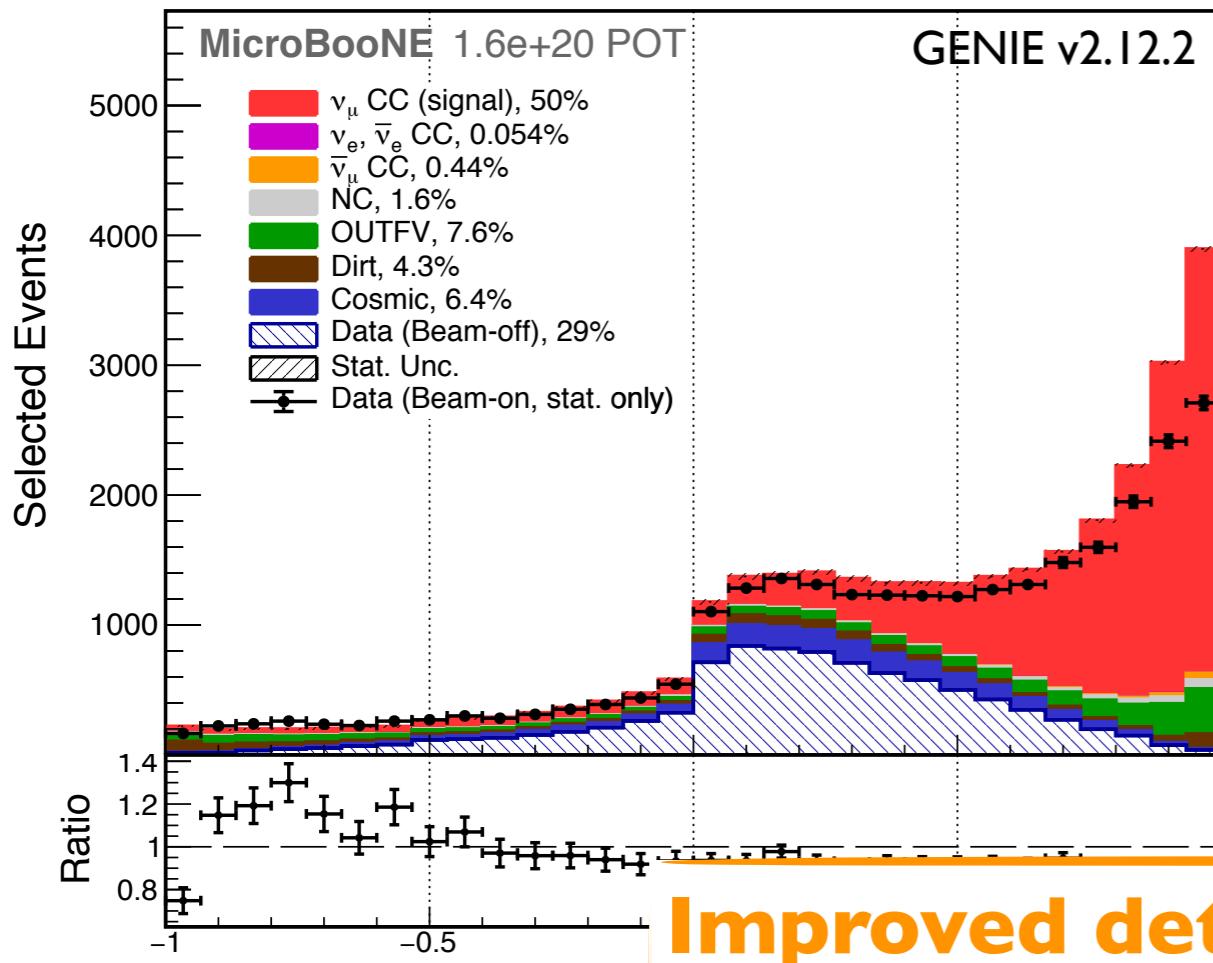
PRL 123, 131801 (2019)



Current Measurement

MICROBOONE-NOTE-1069-PUB

# IMPROVED DETECTOR UNDERSTANDING ENABLES BETTER MEASUREMENTS



Improved detector understanding,  
reconstruction → higher purity

Previous Meas.

PRL 123,

$\nu_\mu$  CC (signal) purity: 50% → 71.9%

Entering backgrounds: 33.3% → 13.2%

# IMPROVED INTERACTION MODELING ENABLES BETTER MEASUREMENTS

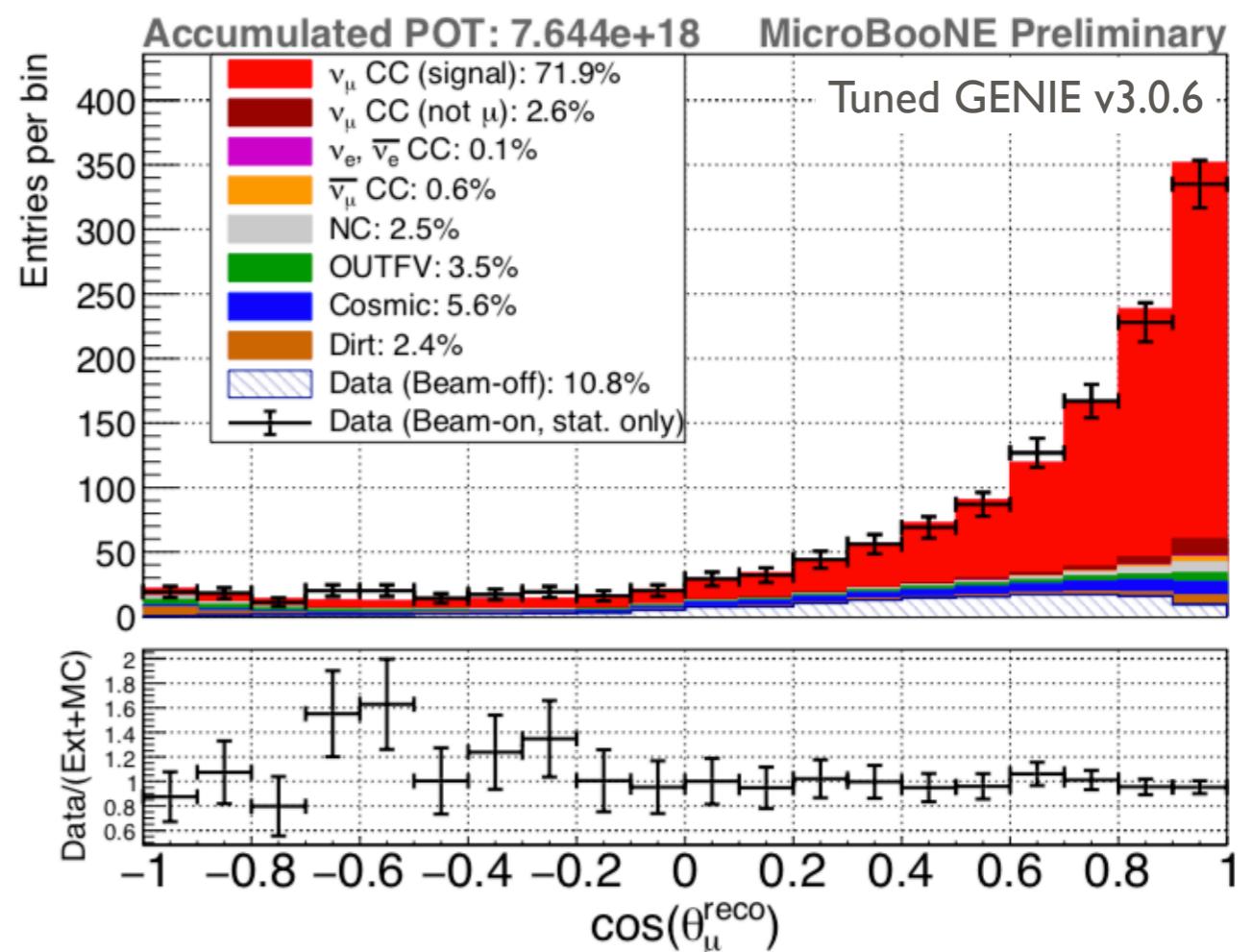
## Better data-simulation

**agreement** from improved neutrino interaction modeling

- GENIE v2.12.2 → GENIE v3.0.6
- **Tuned** CCQE and CC2p2h models to T2K  $\nu_\mu$  CC0 $\pi$  data
- T2K data is on a carbon target → tuning seems to give **good agreement with MicroBooNE's argon-target data**

PRL 123, 131801 (201)

MICROBOONE-NOTE-1074-PUB



GENIE v3.0.6 models used:

QE/2p2h → **J. Nieves, J.E. Amaro, M. Valverde** Phys. Rev. C 70, 055503 (2004) and

**R. Gran, J. Nieves, F. Sanchez, M. Vicente-Vacas** Phys. Rev. D 88, 113007 (2013)

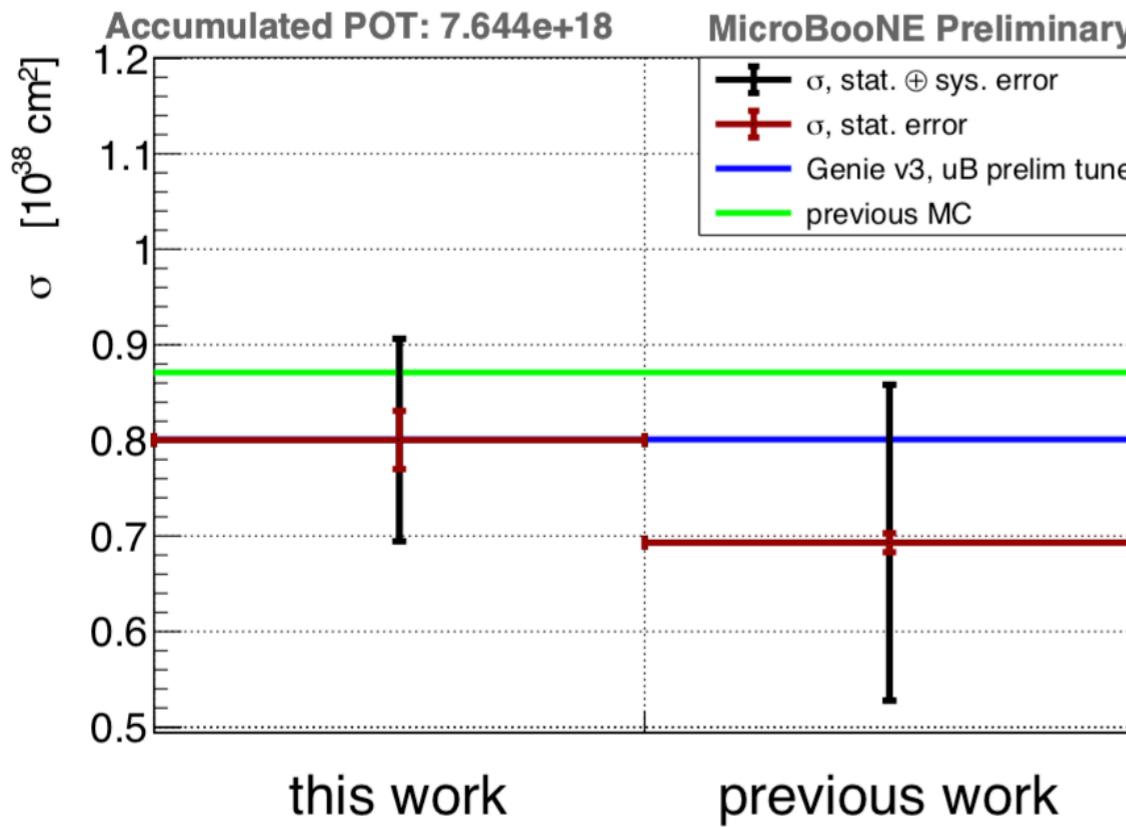
RES/COH → **C. Berger, L. Sehgal** Phys. Rev. D 76, 113004 (2007), Phys. Rev. D 79, 053003 (2009)

FSI → work by **L. Salcedo, E. Oset, M. Vicente-Vacas, C. Garcia-Recio**

Nucl. Phys. A 484, 557-592 (1988) and **V. Pandharipande, S.C. Pieper** Phys. Rev. C 45, 791-798 (1992)

# DRASTICALLY REDUCED SYSTEMATIC UNCERTAINTIES

MICROBOONE-NOTE-1075-PUB . MICROBOONE-NOTE-1069-PUB



Flux-integrated cross section  
**consistent with previous measurement**

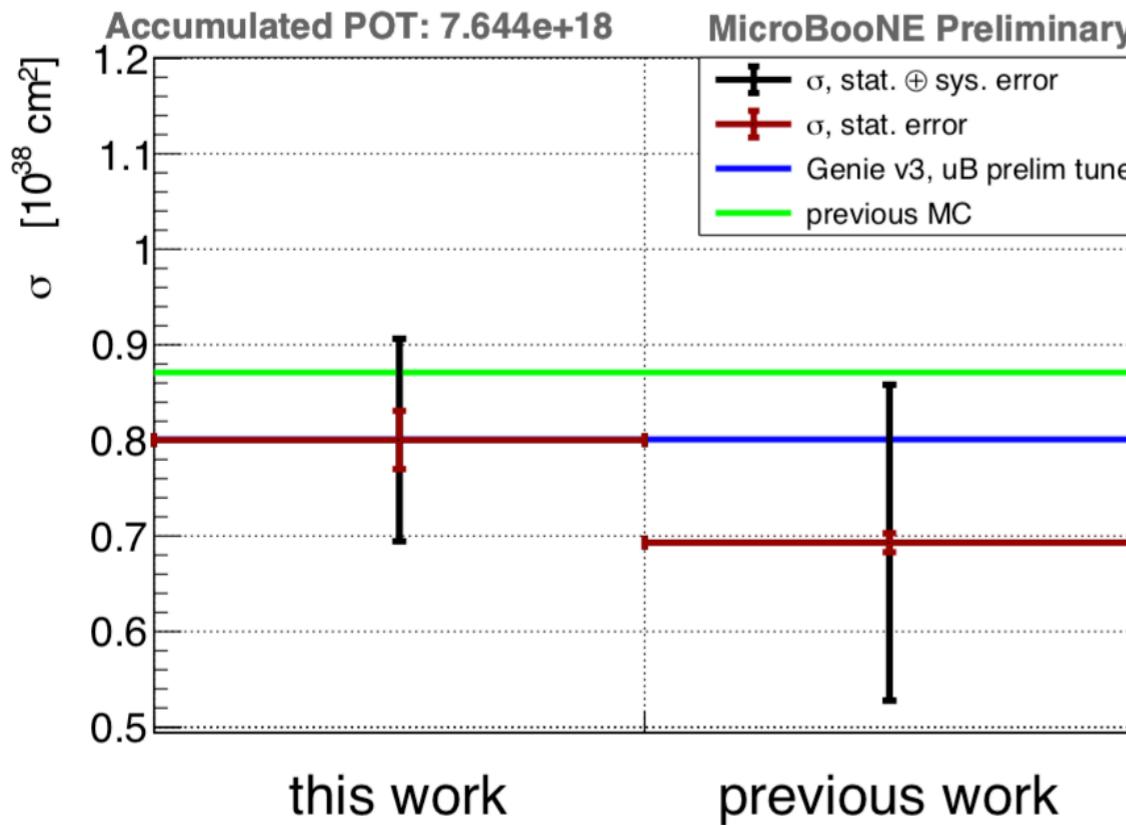
**Drastically reduced** systematic uncertainties

Source	Uncertainty	
	Previous Analysis	This Analysis
Detector response	16.2%	3.3%
Cross section	3.9%	2.7%
Flux	12.4%	10.5%
Dirt background	10.9%	3.3%
Cosmic ray background	4.2%	-
POT counting	2.0%	2.0%
CRT	N/A	1.7%
Total Sys. Error	23.8%	12.1%
Statistics	1.4%	3.8%
<b>Total (Quadratic Sum)</b>	<b>23.8%</b>	<b>12.7%</b>

PRL 123, 131801 (2019)

# DRASTICALLY REDUCED SYSTEMATIC UNCERTAINTIES

MICROBOONE-NOTE-1075-PUB . MICROBOONE-NOTE-1069-PUB



Flux-integrated cross section  
**consistent with previous measurement**

**Drastically reduced** systematic uncertainties

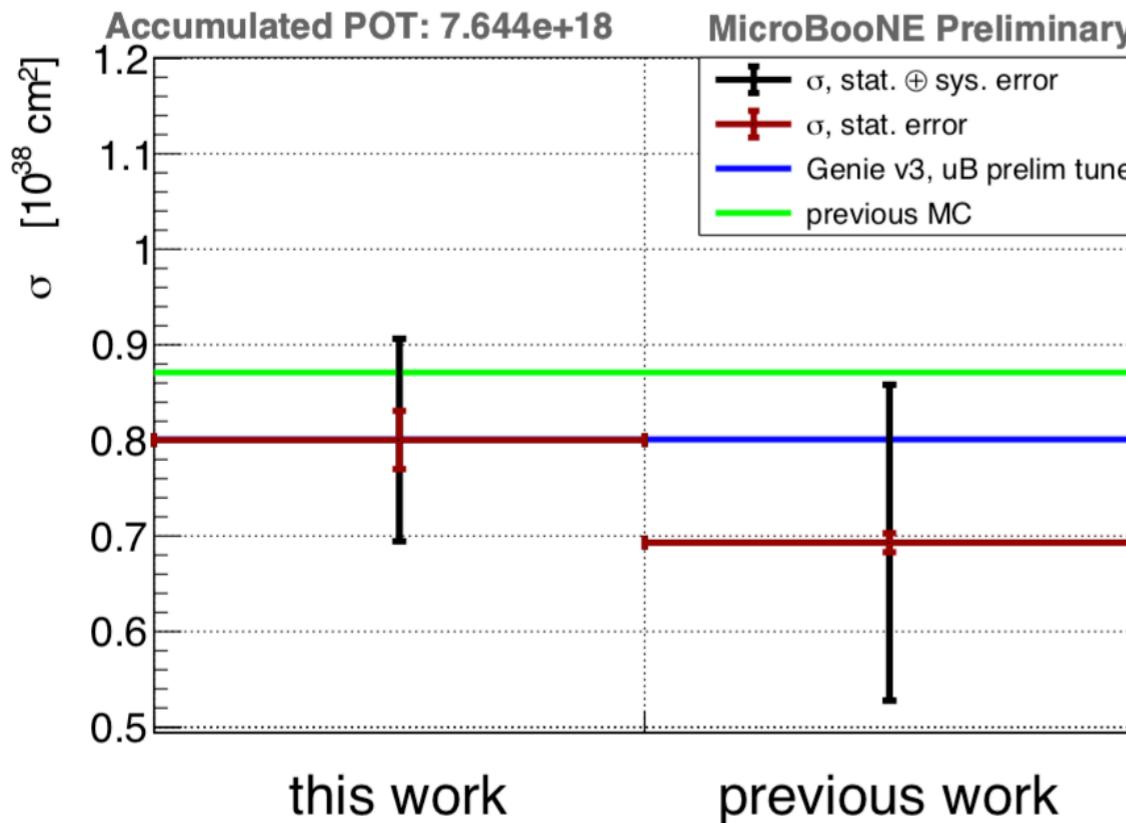
Largest reduction in uncertainties comes from improved detector understanding

Source	Uncertainty	
	Previous Analysis	This Analysis
Detector response	16.2%	3.3%
Cross section	3.9%	2.7%
Flux	12.4%	10.5%
Dirt background	10.9%	3.3%
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PRL 123, 131801 (2019)

# DRASTICALLY REDUCED SYSTEMATIC UNCERTAINTIES

MICROBOONE-NOTE-1075-PUB . MICROBOONE-NOTE-1069-PUB



Flux-integrated cross section  
**consistent with previous measurement**

**Drastically reduced** systematic uncertainties

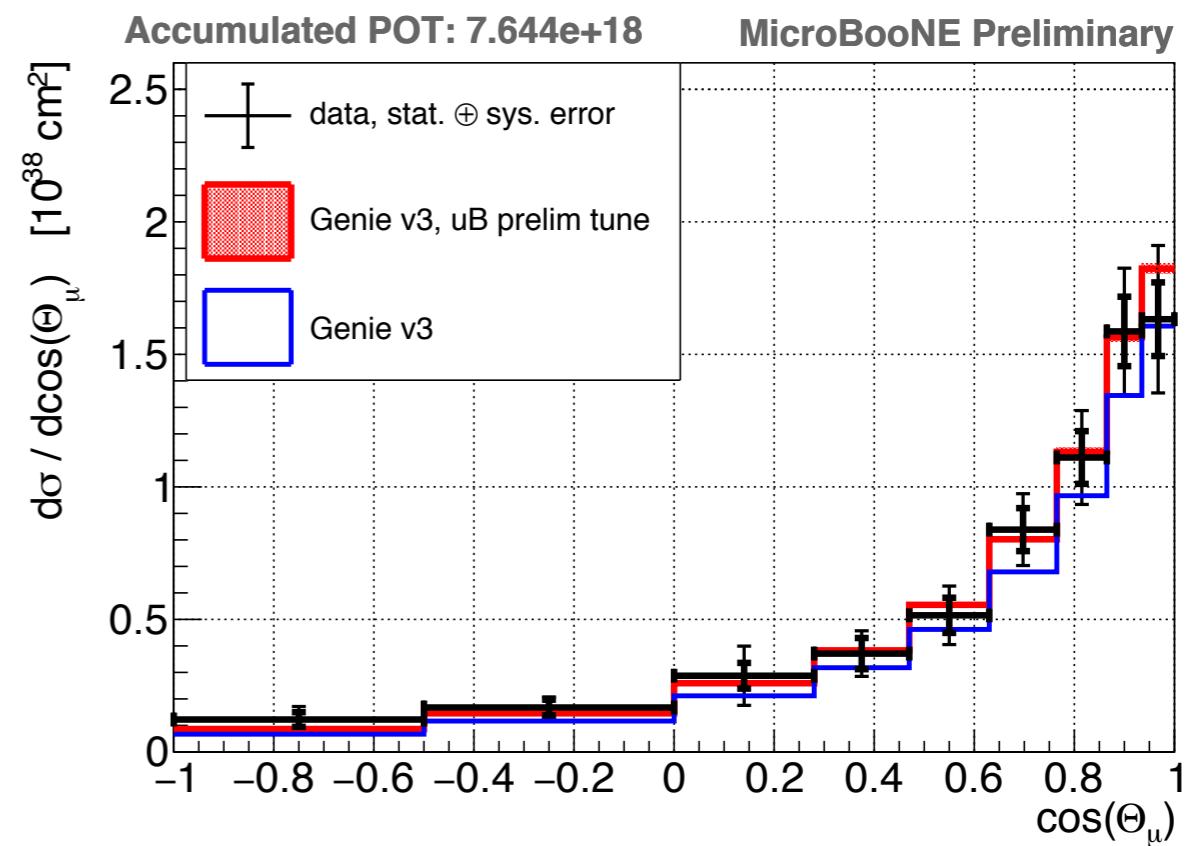
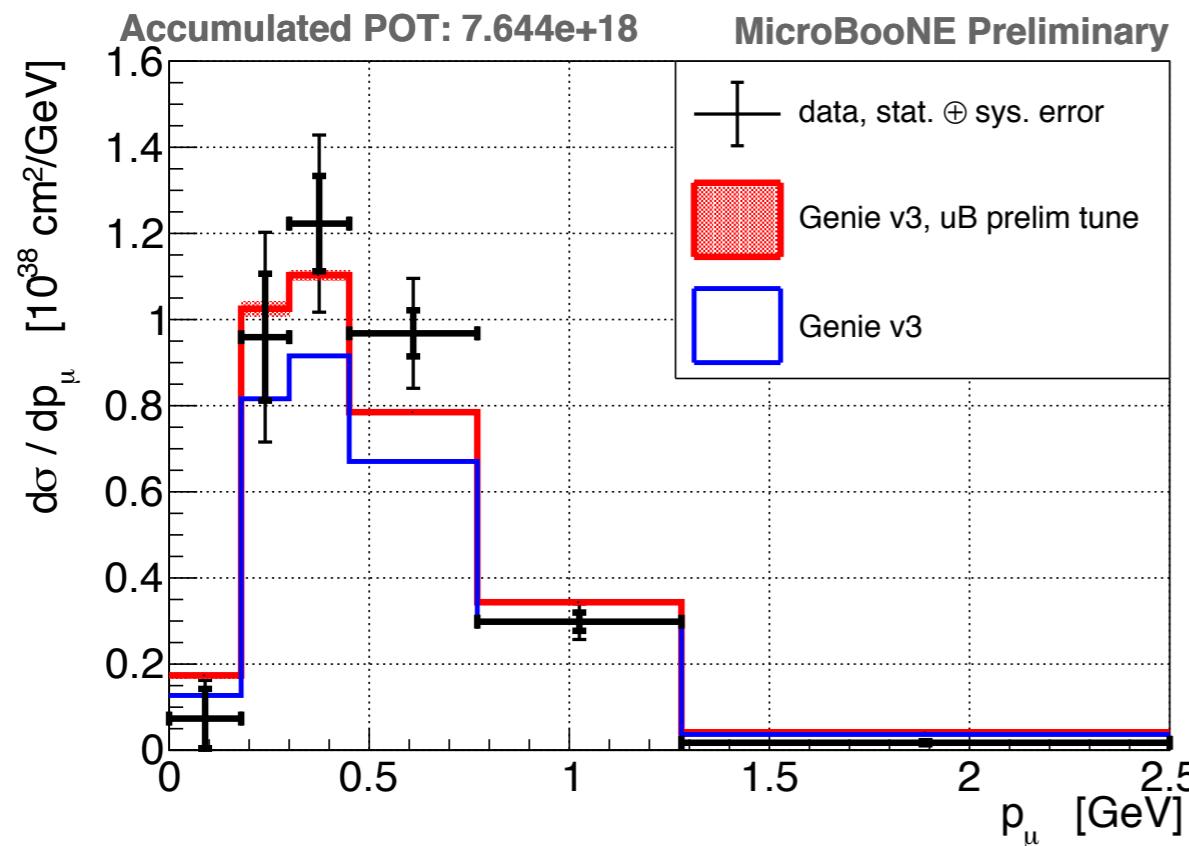
Instead of cosmic ray simulation, now use overlay: simulated neutrino interactions overlaid on real cosmic data → no uncertainty in cosmic ray model

Source	Uncertainty	
	Previous Analysis	This Analysis
Detector response	16.2%	3.3%
Cross section	3.9%	2.7%
Flux	12.4%	10.5%
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PRL 123, 131801 (2019)

# IMPROVED CROSS SECTION MEASUREMENT

MICROBOONE-NOTE-1074-PUB . MICROBOONE-NOTE-1075-PUB . MICROBOONE-NOTE-1069-PUB



**Single-differential cross section** as a function of reconstructed muon momentum and angle → **very good agreement with previous measurement**, but **reduced uncertainties**

Future development towards **double-differential** cross-section measurement

# Many measurements of $\nu$ -Ar scattering

## ■ $\nu_\mu$ CC inclusive cross section



Single-differential cross section

Phys. Rev. Lett. 108 161802 (2012)



Updated single-differential cross section

Phys. Rev. D 89, 112003 (2014)

## ■ $\nu_\mu$ exclusive channels



$\nu_\mu$  and  $\bar{\nu}_\mu$  NC $\pi^0$  production

Phys. Rev. D 96, 012006 (2017)



$\nu_\mu$  and  $\bar{\nu}_\mu$  CC $\pi^+$  production

Phys. Rev. D 98, 052002 (2018)



$\nu_\mu$  and  $\bar{\nu}_\mu$  Coherent CC $\pi^+$  production

Phys. Rev. Lett. 113, 261801 (2014)



$\nu_\mu$  and  $\bar{\nu}_\mu$  CC2p production

Phys. Rev. D 90, 012008 (2014)

## ■ Other measurements

# Many measurements of $\nu$ -Ar scattering

## $\nu_\mu$ CC inclusive cross section



Single-differential cross section

Phys. Rev. Lett. 108 161802 (2012)



Updated single-differential cross section

Phys. Rev. D 89, 112003 (2014)



Double-differential cross section

Phys. Rev. Lett. 123, 131801 (2019)



Single-differential cross section with updated detector and interaction models

MICROBOONE-NOTE-1069-PUB

## $\nu_\mu$ exclusive channels



Charged-particle multiplicity

Eur. Phys. J. C79, 248 (2019)



$\nu_\mu$  CCQE-like scattering

Eur. Phys. J. C 79 673 (2019)

Phys. Rev. Lett. 125, 201803 (2020)



$\nu_\mu$  CC $0\pi Np$  ( $N \geq 1$ ) scattering

Phys. Rev. D 102, 112013 (2020)



$\nu_\mu$  and  $\bar{\nu}_\mu$  CC2p production

Phys. Rev. D 90, 012008 (2014)



$\nu_\mu$  CC $\pi^0$  production

Phys. Rev. D99, 091102(R) (2019)



$\nu_\mu$  and  $\bar{\nu}_\mu$  NC $\pi^0$  production

Phys. Rev. D 96, 012006 (2017)



$\nu_\mu$  and  $\bar{\nu}_\mu$  CC $\pi^+$  production

Phys. Rev. D 98, 052002 (2018)



$\nu_\mu$  and  $\bar{\nu}_\mu$  Coherent CC $\pi^+$  production

Phys. Rev. Lett. 113, 261801 (2014)



$\nu_\mu$  CC kaon production

MICROBOONE-NOTE-1071-PUB



$\nu_\mu$  NC  $1p$  production

MICROBOONE-NOTE-1067-PUB



MeV-scale physics

MICROBOONE-NOTE-1076-PUB



Limits on millicharged particles

Phys. Rev. Lett. 124, 131801 (2020)



$\nu_e$  and  $\bar{\nu}_e$  scattering (inclusive)

Phys. Rev. D 102, 011101(R) (2020)



$\nu_e$  and  $\bar{\nu}_e$  total cross section (inclusive)

arXiv:2101.04228[hep-ex]



MeV-scale physics

Phys. Rev. D 99, 012002 (2019)



# MORE MEASUREMENTS!

A **large number** of **new measurements** are in progress, including:

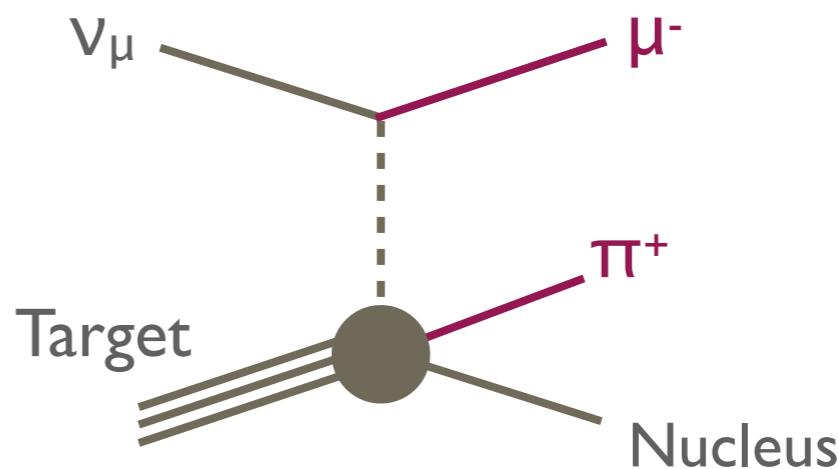
- $\nu_\mu$  CC  $0\pi 2p$
- $\nu_\mu$  CC  $0\pi Np$  transverse variables
- $\nu_\mu$  CC  $0\pi 1p$  transverse variables
- $\nu_\mu$  CC  $\pi^0$  (update)
- $\nu_\mu$  CC  $1\pi^+$
- $\nu_\mu$  CC-Coherent  $1\pi^+$
- $\nu_\mu$  NC  $\pi^0$
- $\nu_\mu$  NC  $1\gamma$
- $\nu_\mu$  CC  $0\pi 0p$
- $\nu_\mu$  CC inclusive (update)
- $\nu_\mu$  CC inclusive hadronic energy
- $\nu_\mu$  NC Elastic p
- $\nu_\mu$  CC  $1K^\pm$
- $\nu_\mu$  hyperon production
- $\nu_\mu$  CC eta production
- KDAR  $\nu_\mu$
- $\nu_e$  CC inclusive (update)
- $\nu_e$  CC  $0\pi Np$

# MORE MEASUREMENTS!

A **large number** of **new measurements** are in progress, including:

- $\nu_\mu$  CC 0 $\pi$ 2p
- $\nu_\mu$  CC 0 $\pi$ Np transverse variables
- $\nu_\mu$  CC 0 $\pi$ 1p transverse variables
- $\nu_\mu$  CC  $\pi^0$  (update)
- $\nu_\mu$  CC 1 $\pi^+$
- $\nu_\mu$  CC-Coherent 1 $\pi^+$
- $\nu_\mu$  NC  $\pi^0$
- $\nu_\mu$  NC 1 $\gamma$
- $\nu_\mu$  CC 0 $\pi$ 0p

- Charged pions  $\rightarrow$  dominant interaction modes at DUNE



- Development work focused on distinguishing charged pions from muons

- $\nu_e$  CC 0 $\pi$ Np

# MORE MEASUREMENTS!

- Rare interactions: charged kaon production
- Very little world data
- Background to proton decay search at DUNE:  $p \rightarrow K^+ \nu$

**Measurements** are in progress, including:

- $\nu_\mu$  CC inclusive (update)
- $\nu_\mu$  CC inclusive hadronic energy
- $\nu_\mu$  NC Elastic p
- $\nu_\mu$  CC  $IK^\pm$

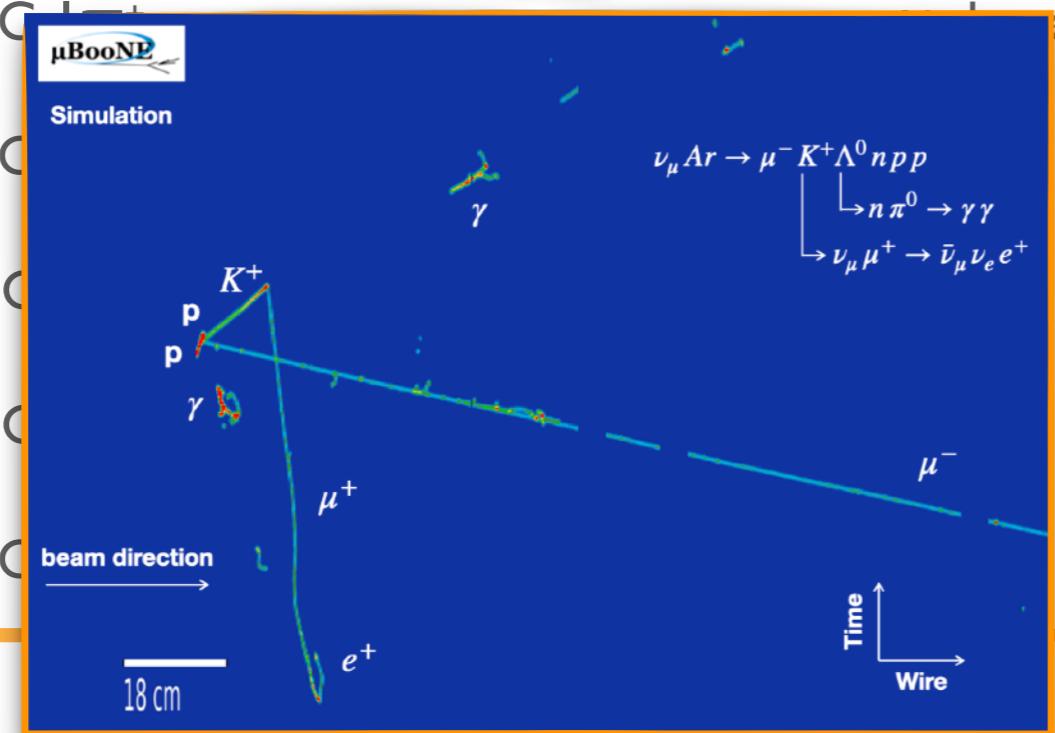
- $\nu_\mu$  CC  $I\bar{K}^\pm$

- $\nu_\mu$  CC  $I\eta$

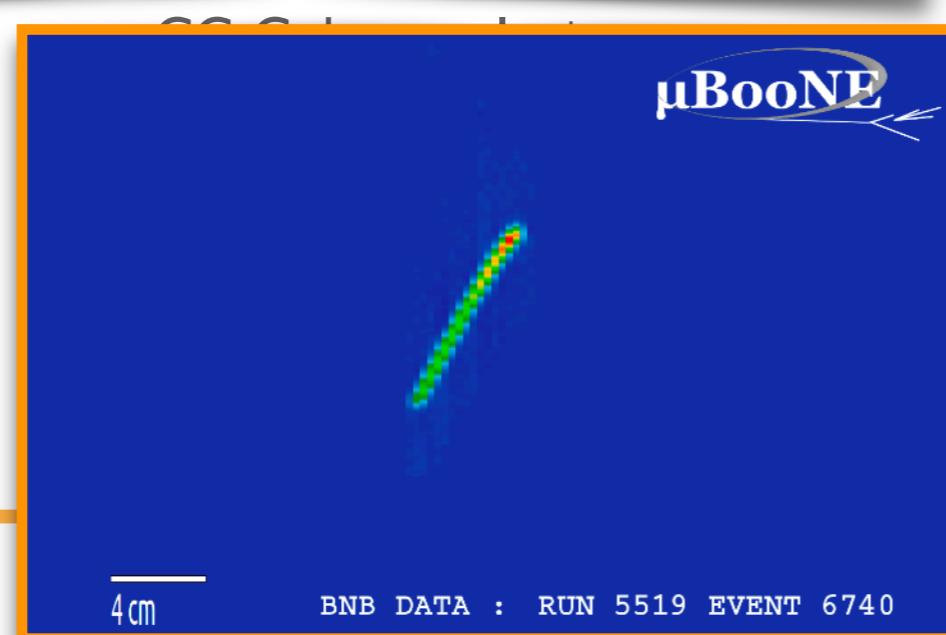
- $\nu_\mu$  NC  $I\eta$

- $\nu_\mu$  NC  $I\pi Np$

- $\nu_\mu$  CC  $I\pi Np$



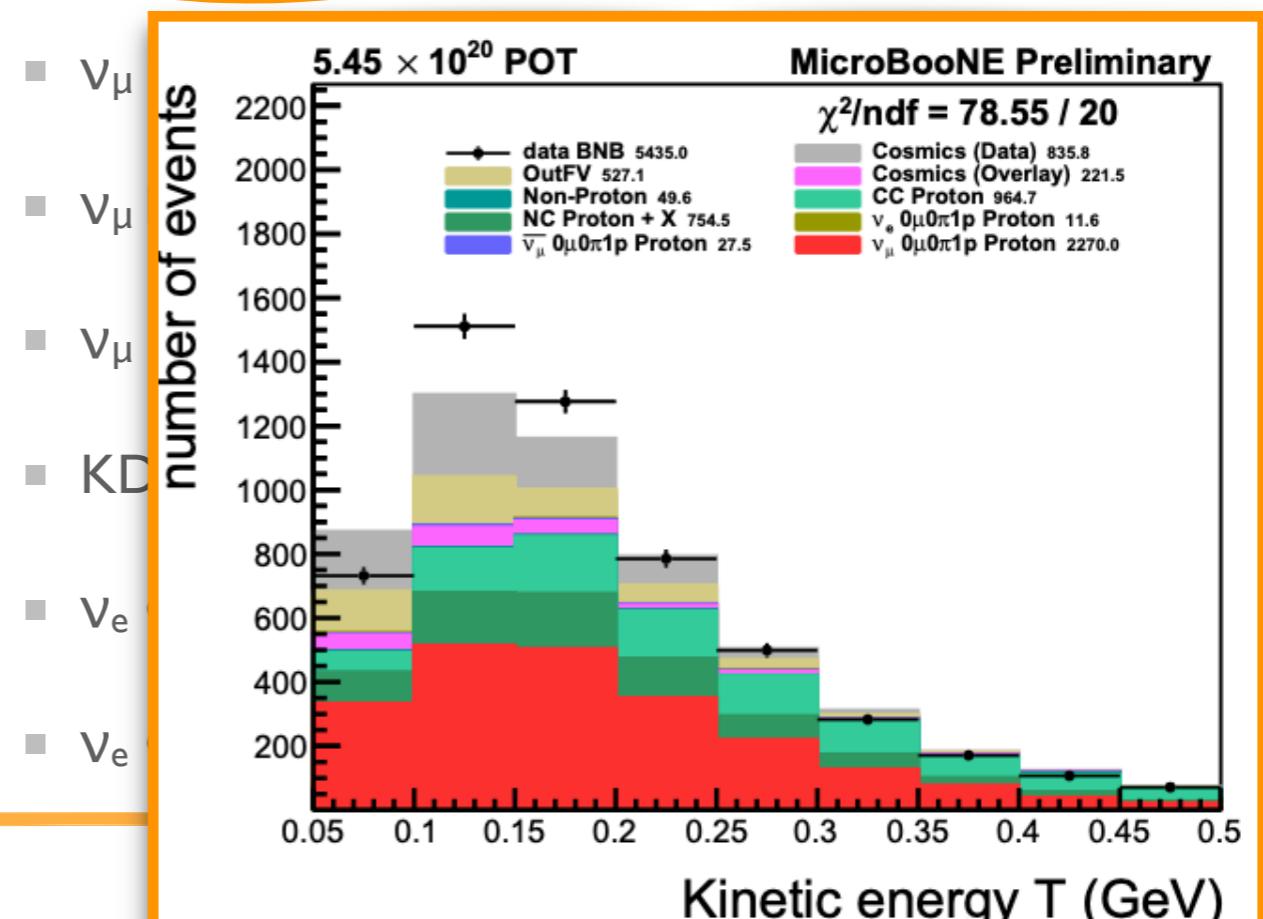
- Signal: single, isolated proton
- Already have an NC1p selection and preliminary cross section extraction
- Select events with  $Q^2 \sim 2m_p T_p = 0.1 \text{ GeV}^2$ , **significantly lower** than previous measurements
- Future development towards **NC elastic scattering** cross section  
→ measure strange component of neutral-current axial form factor



# RESULTS!

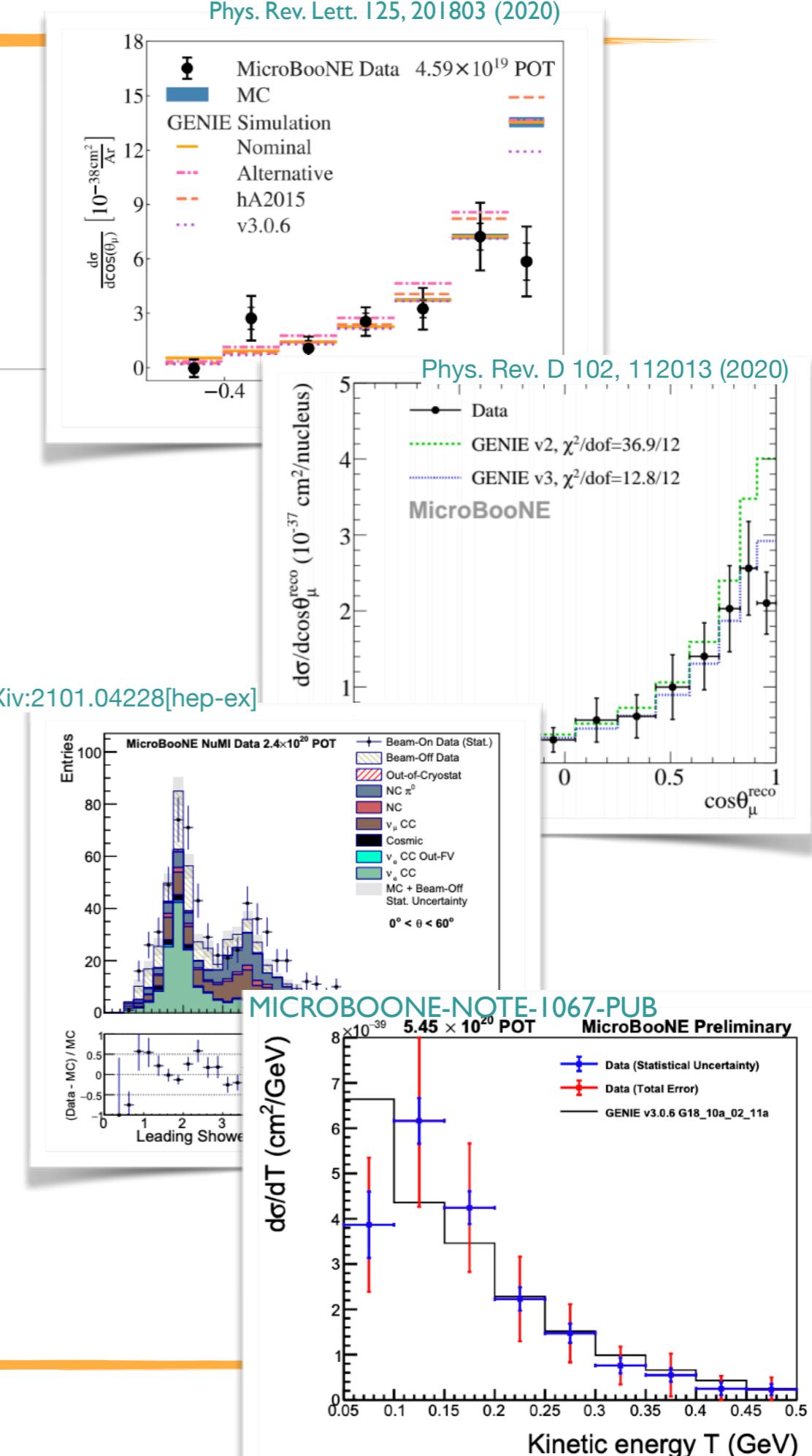
**Measurements** are in progress, including:

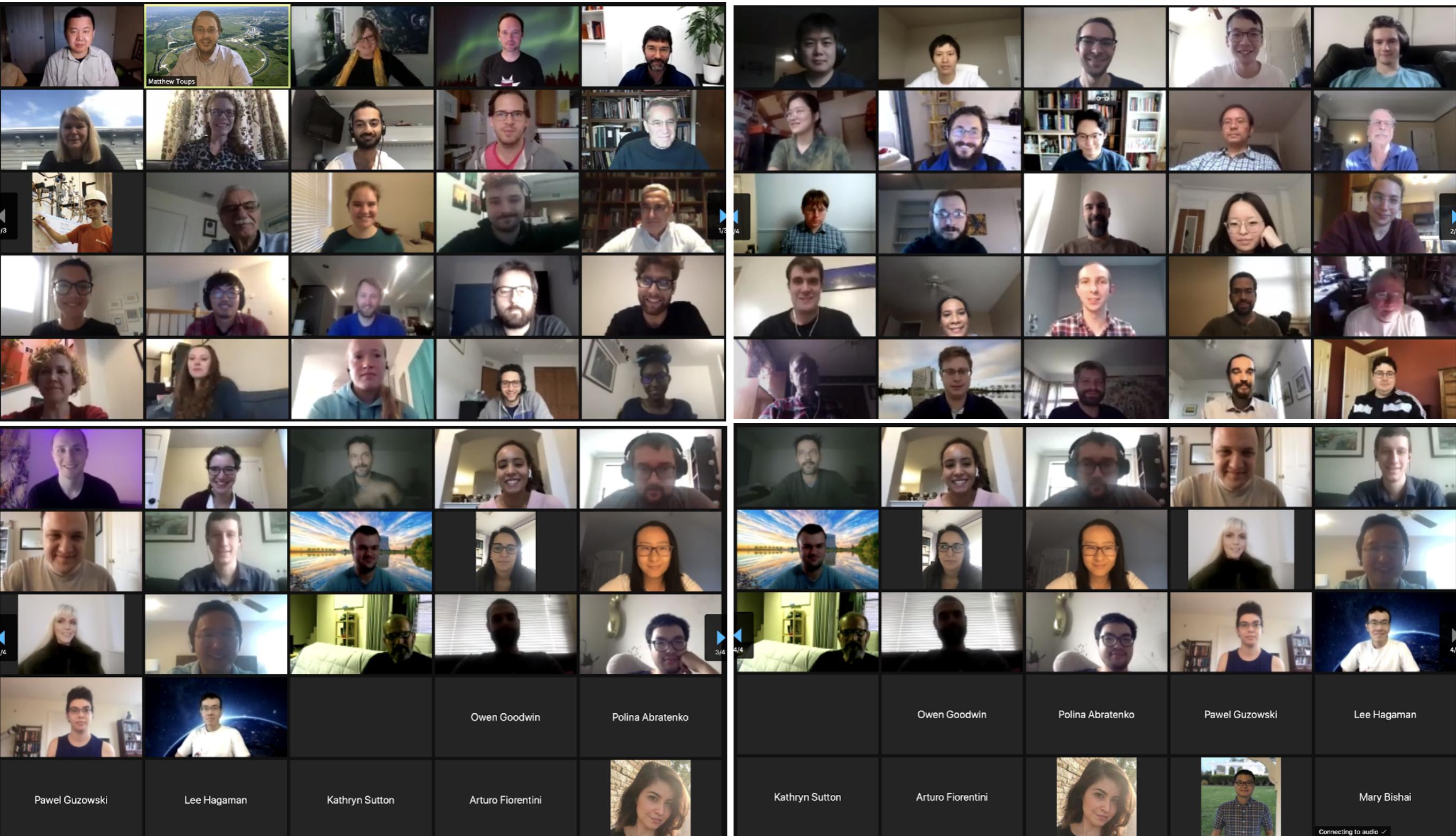
- $\nu_\mu$  CC inclusive (update)
- $\nu_\mu$  CC inclusive hadronic energy
- $\nu_\mu$  NC Elastic p



# SUMMARY

- Cross-section measurements on argon are **vital** for the success of the SBN program and eventually DUNE
- MicroBooNE has demonstrated  **$4\pi$  acceptance** and ability to measure particles with **very low thresholds**
- We are already able to make **precise, accurate measurements of exclusive final states**
- **Huge progress** over the past few years: first time we can confront **models tuned to carbon** with high-statistics argon data
- More (and more precise) measurements expected in the future → **stronger tests** of our models



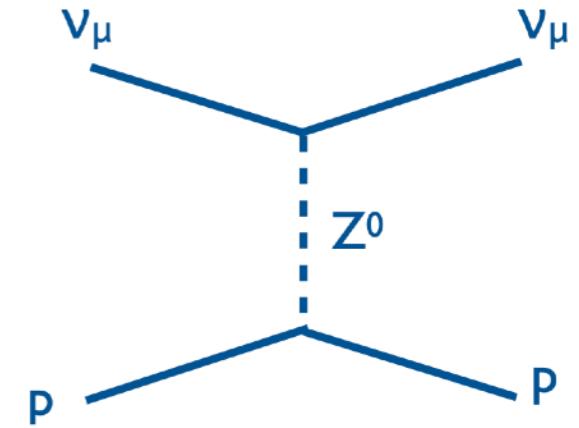


# THANK YOU

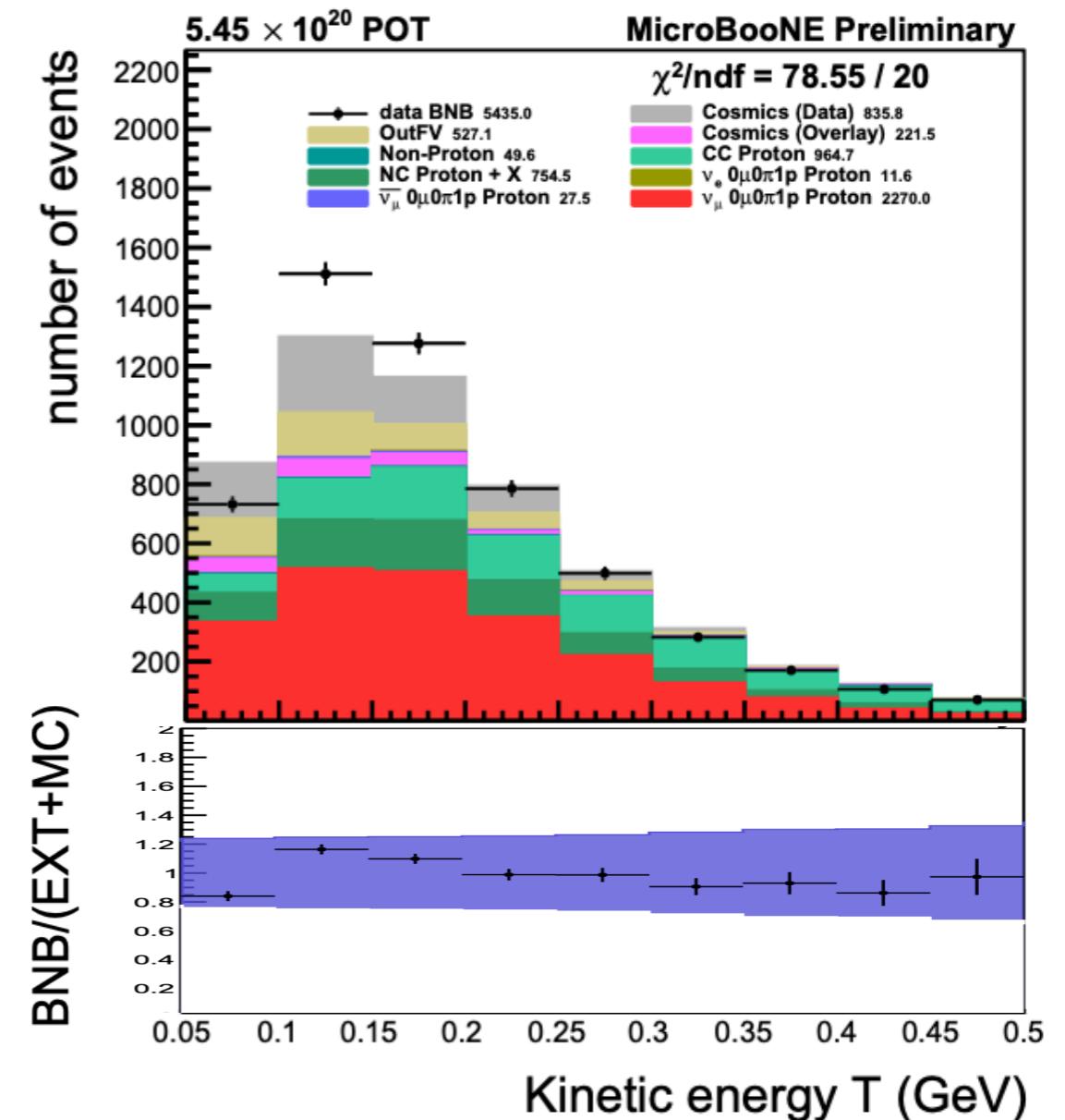
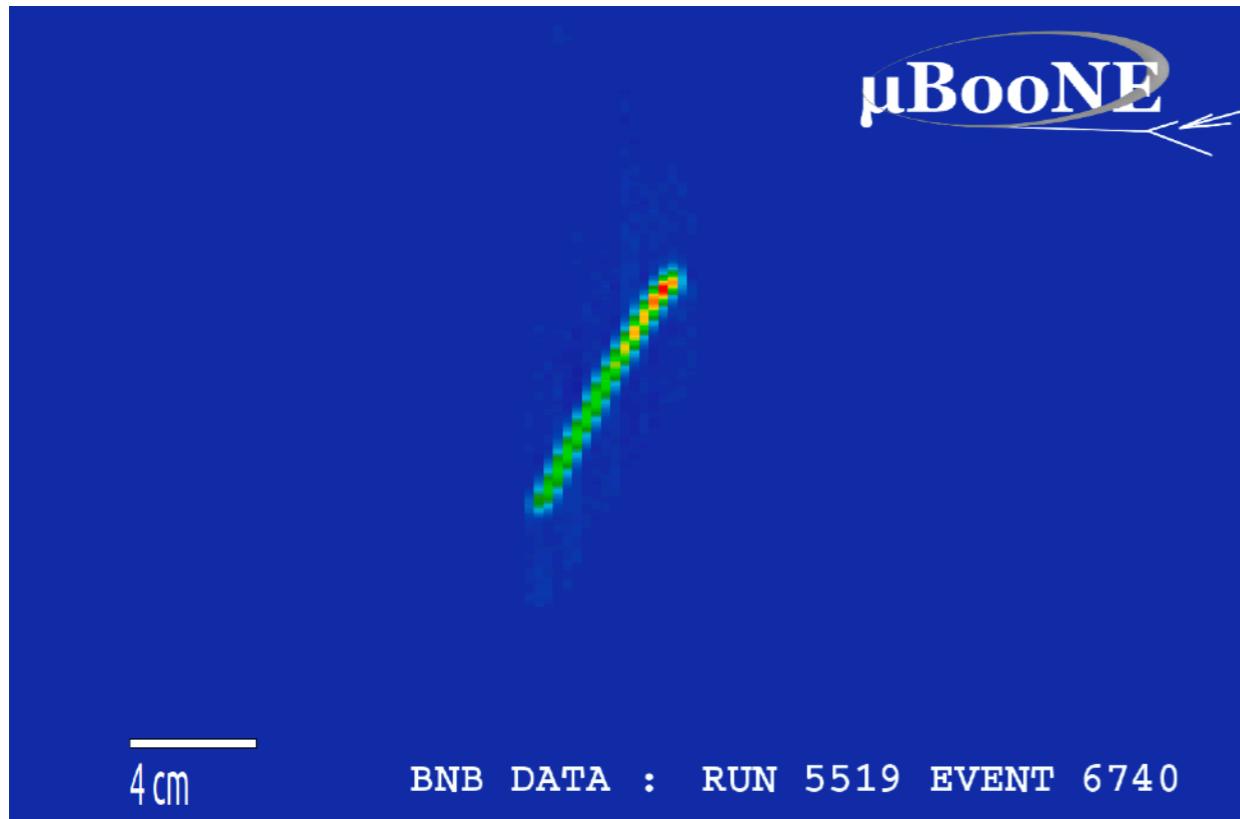


# NCIP CROSS SECTION MEASUREMENT

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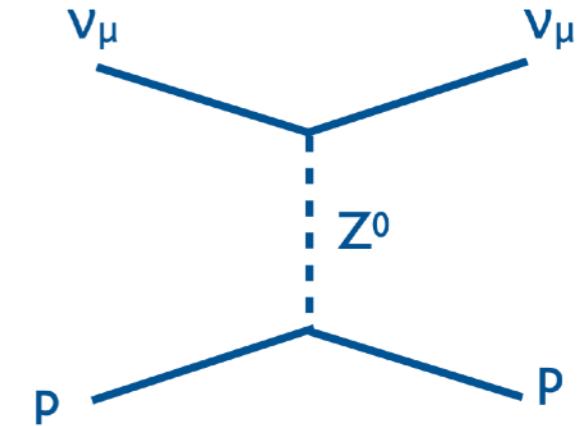


- Measure cross section for neutral-current single proton production
- Signal: 1 isolated proton

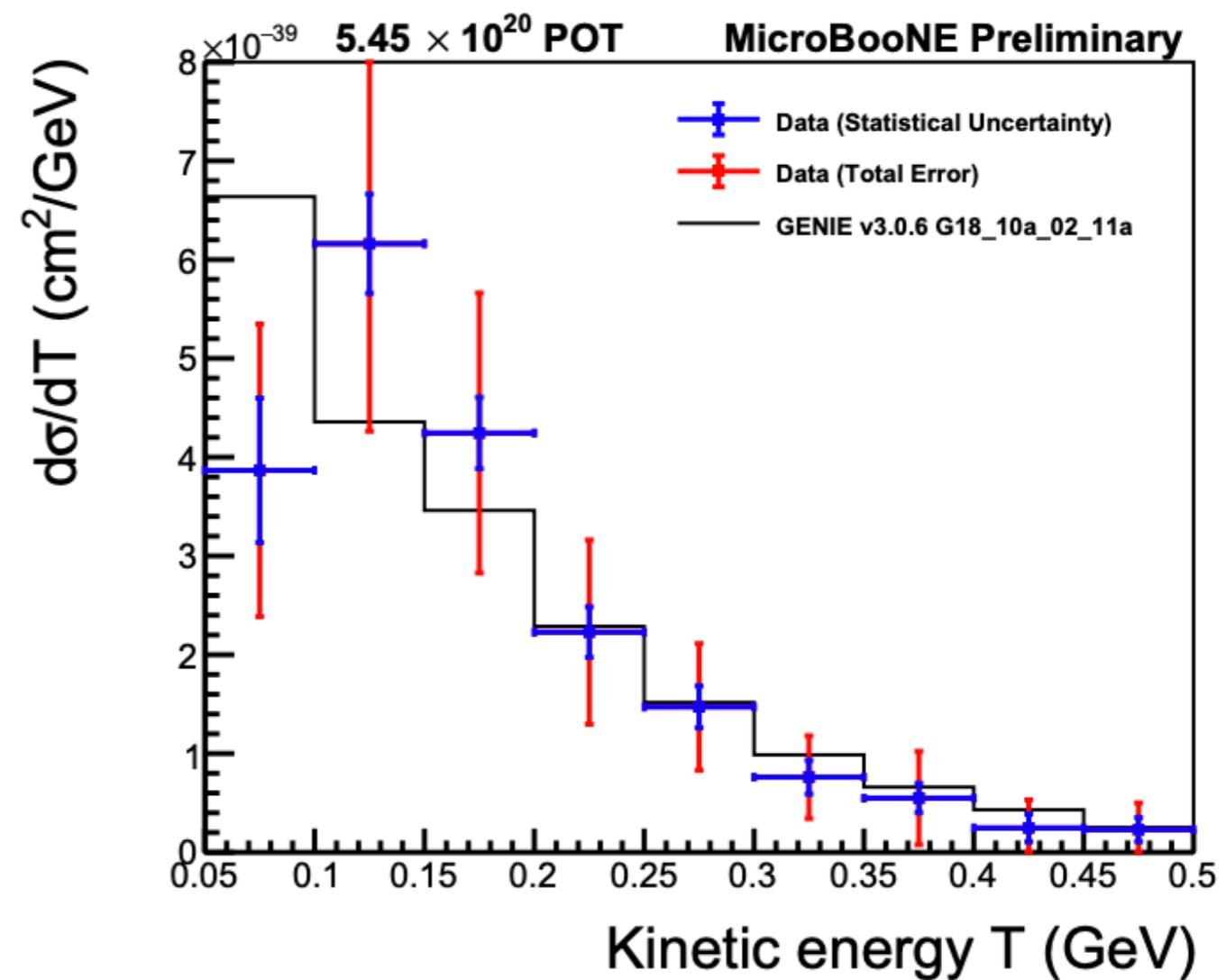


# NC IP CROSS SECTION MEASUREMENT

MICROBOONE-NOTE-1067-PUB



- Measurement includes events with  $Q^2 \sim 2m_p T_p = 0.1 \text{ GeV}^2$ , **significantly lower** than previous measurements
- Future development towards a measurement of **NC elastic scattering** cross section → measure strange component of neutral-current axial form factor

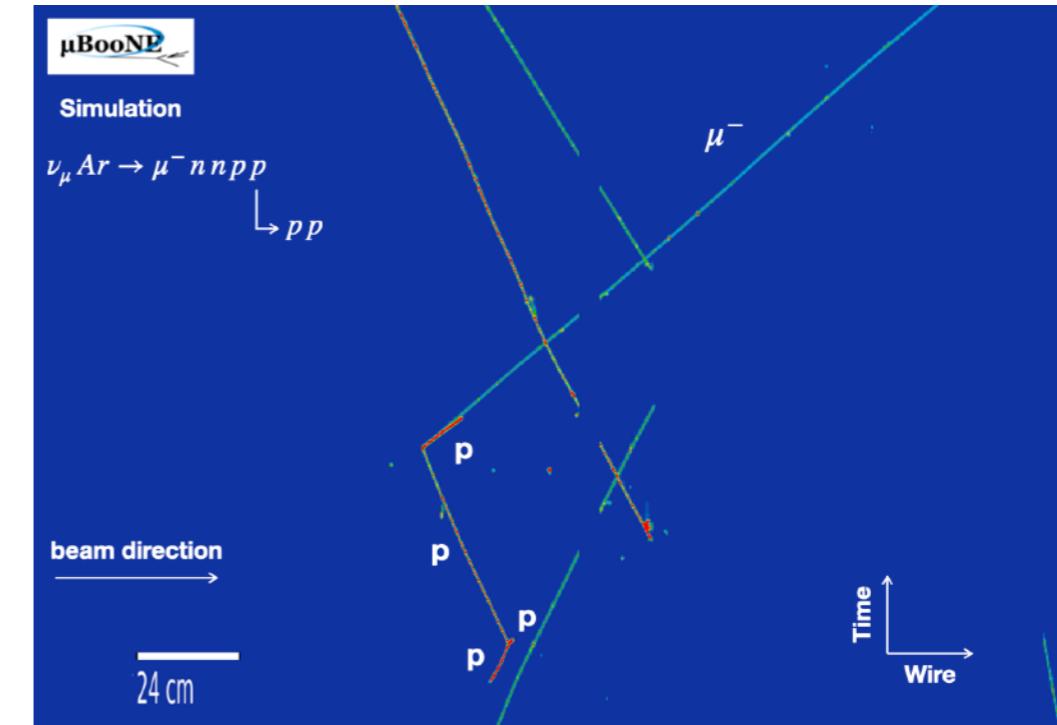
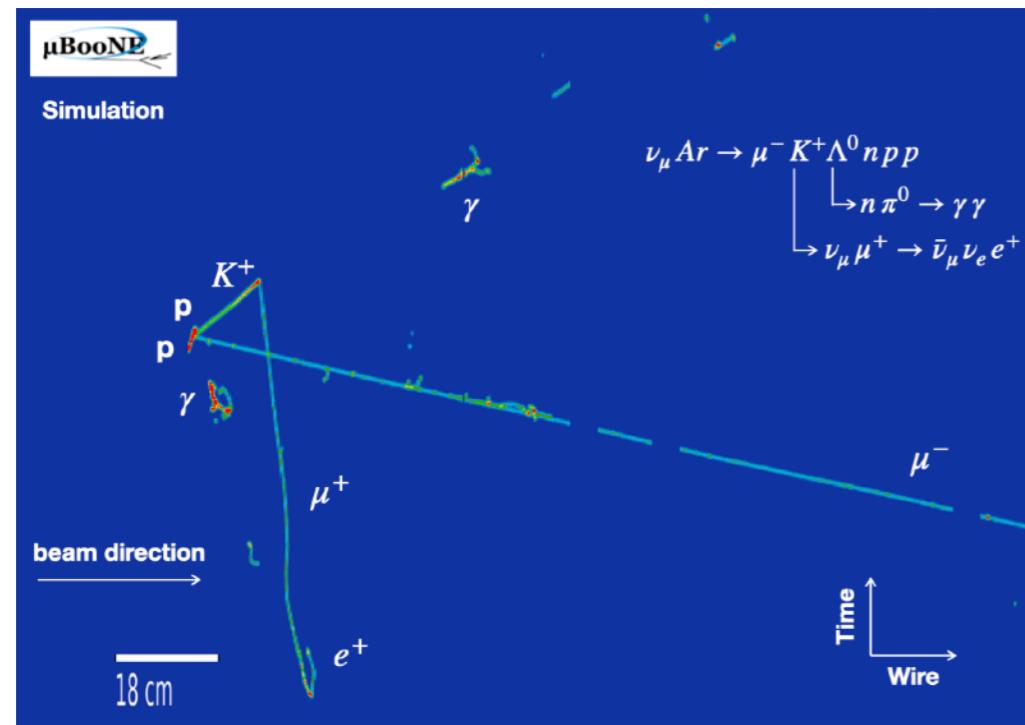


# CC KAON PRODUCTION SELECTION

A. Fiorentini, poster  
369, poster session 3

MICROBOONE-NOTE-1071-PUB

- **CC kaon production:** rare process, few existing measurements, background for **proton decay**  $p \rightarrow K^+ \nu$  searches in DUNE
- Selection developed on simulation: look for  $K^+$  track from neutrino interaction and  $\mu^+$  from  $K^+$  decay
- 67.7% purity and 7% efficiency → expect to select 12 candidate interactions in  $1.3 \times 10^{21}$  POT MicroBooNE data set
- Aim: cross section measurement and study of  $K^+$  in LArTPC



# PUSHING THE LIMITS

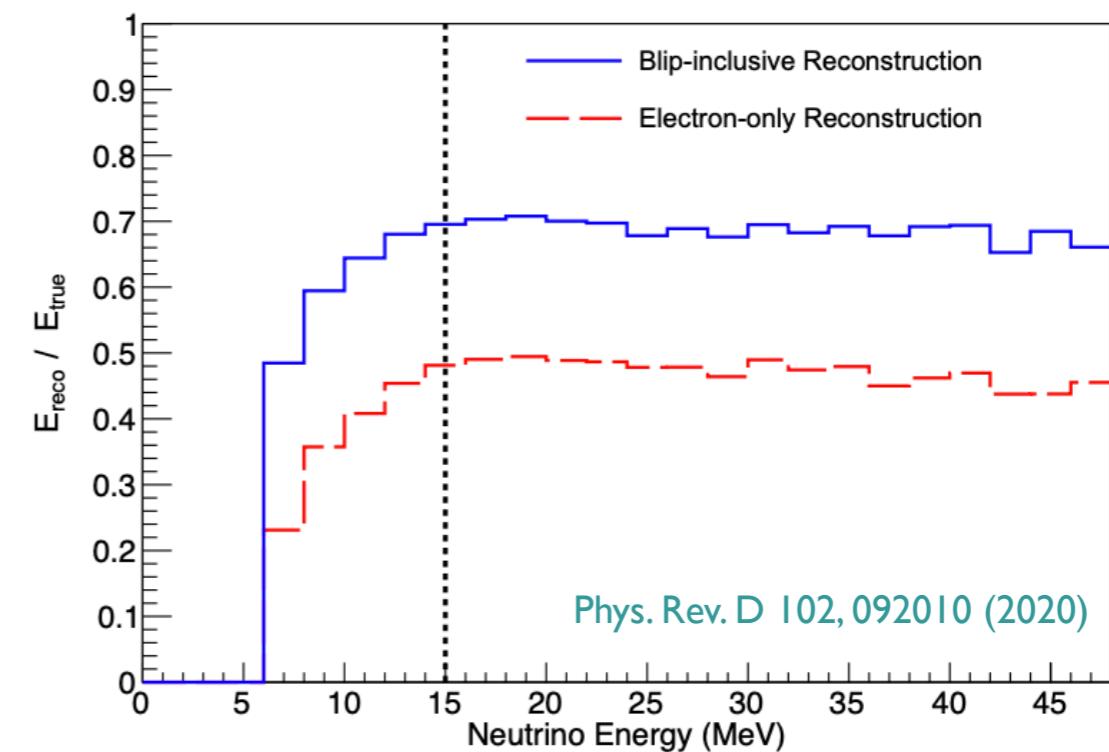
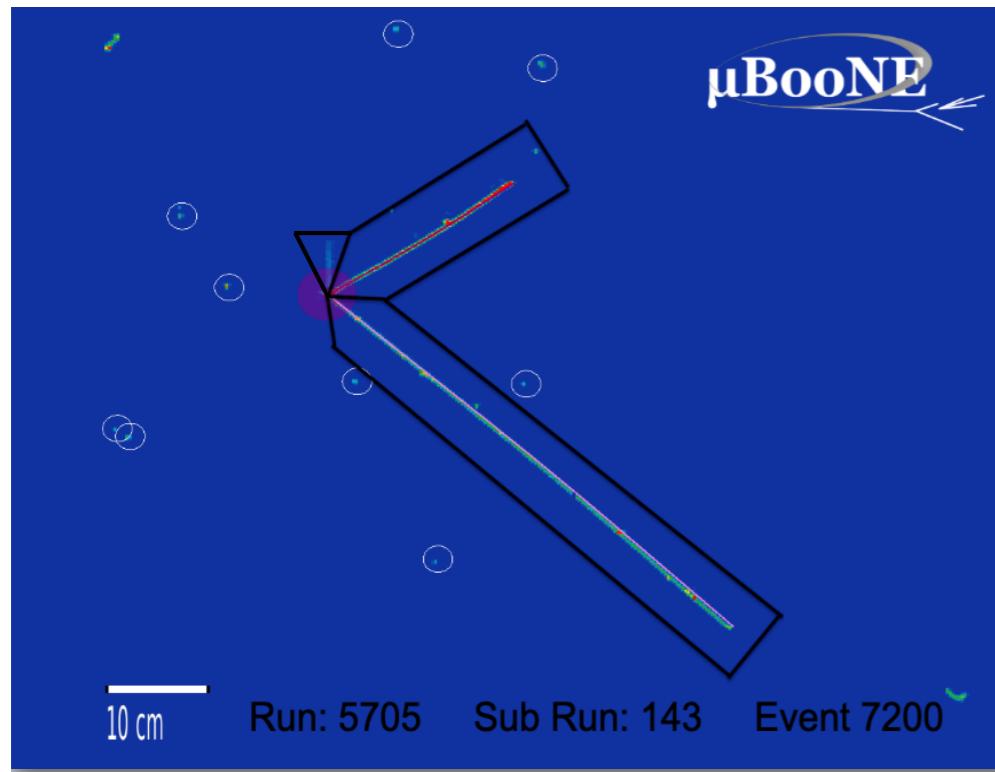
Phys. Rev. D 102, 092010 (2020)

Phys. Rev. Lett. 124, 131801 (2020)

Phys. Rev. D 99, 012002 (2019)

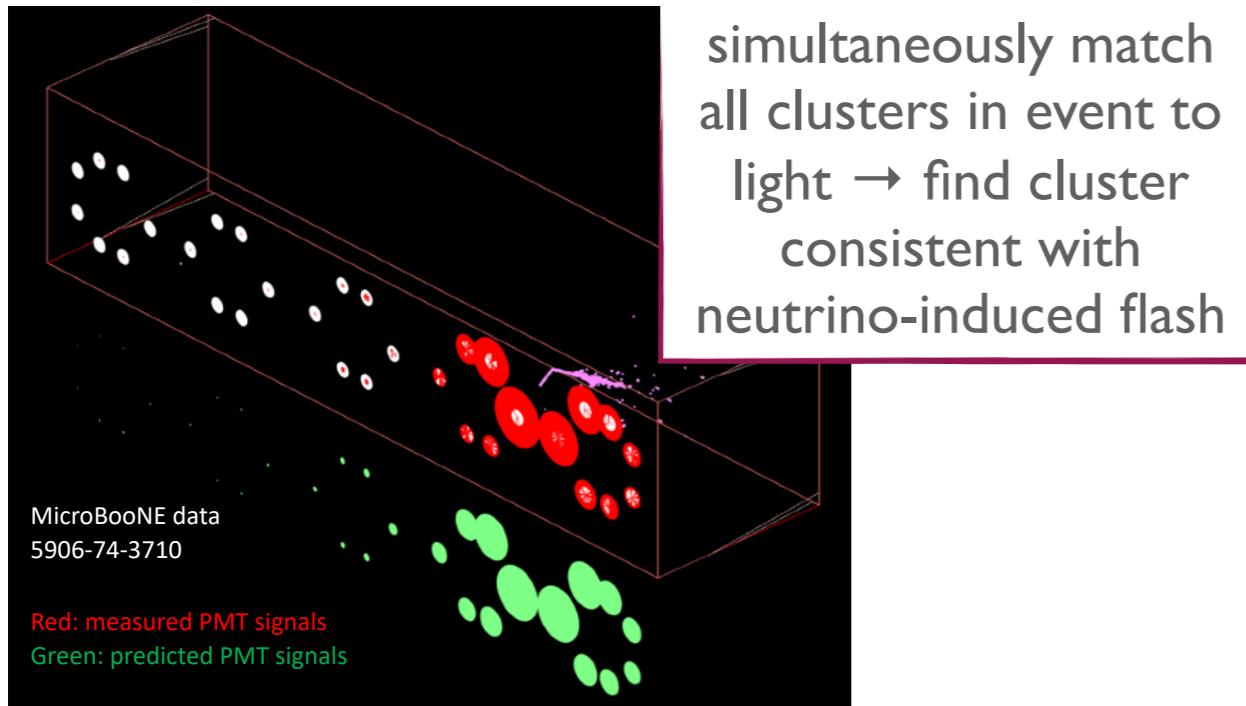
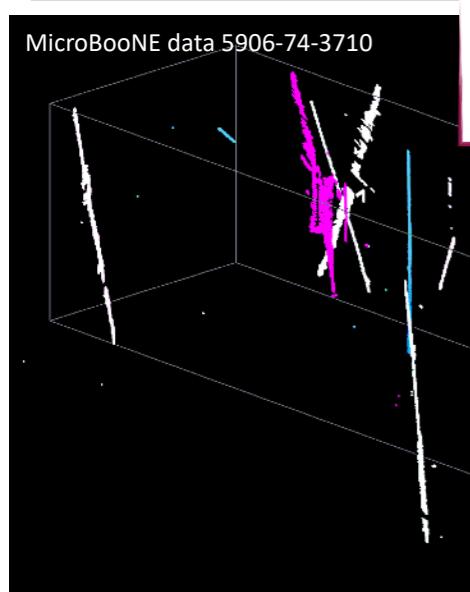
MICROBOONE-NOTE-1076-PUB

- Both ArgoNeuT and MicroBooNE have demonstrated ability to reconstruct **energy depositions from sub-MeV particles** (ArgoNeuT: 300 keV, MicroBooNE: 100 keV)
- Generally **photons** from nucleus de-excitation or **neutron** re-interactions → can give **substantial improvements** in calorimetry and energy reconstruction
- Used in ArgoNeuT to place constraints on BSM physics search for millicharged particles

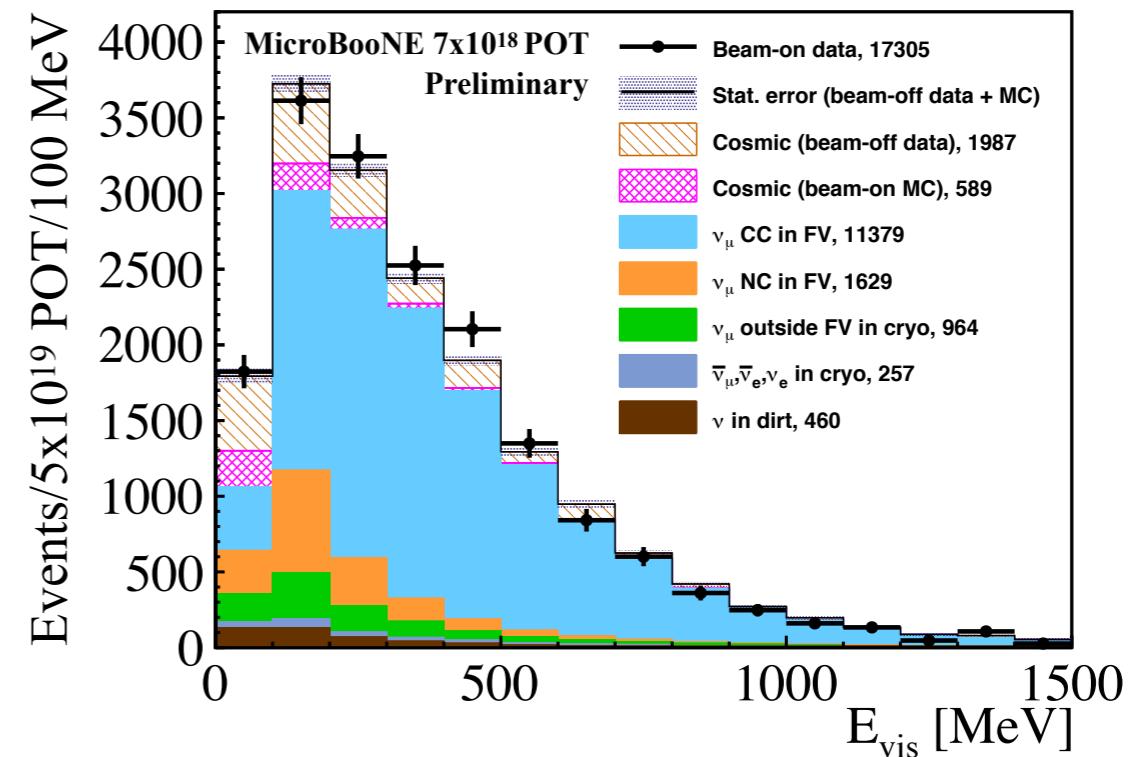


# IMPROVED DETECTOR UNDERSTANDING ENABLES BETTER MEASUREMENTS

arXiv:2011.01375[physics.ins-det] arXiv:2012.07928[hep-ex]



- Cosmic rejection power (without kinematic requirements) **increased by factor of 8** compared to previous publications
- **High efficiency:** 80.4% for  $\nu_\mu$ CC (87.6% for  $\nu_e$ CC)
- **Increased statistics:** 11.3k events, compared to 4.3k events in same data set for 2019 CC inclusive measurement



# COSMIC REJECTION

Phys. Rev. Lett.  
123, 131801  
(2019)

- **Check if in time with beam flash**
- **Check if compatible with beam flash in terms of position and light intensity**
- **Check if track goes all the way through the detector**
- **Check if track is a cosmic crossing anode/cathode (known  $t_0$ )**
- **Check for Bragg peak/ decay electron indicating muon enters and stops**

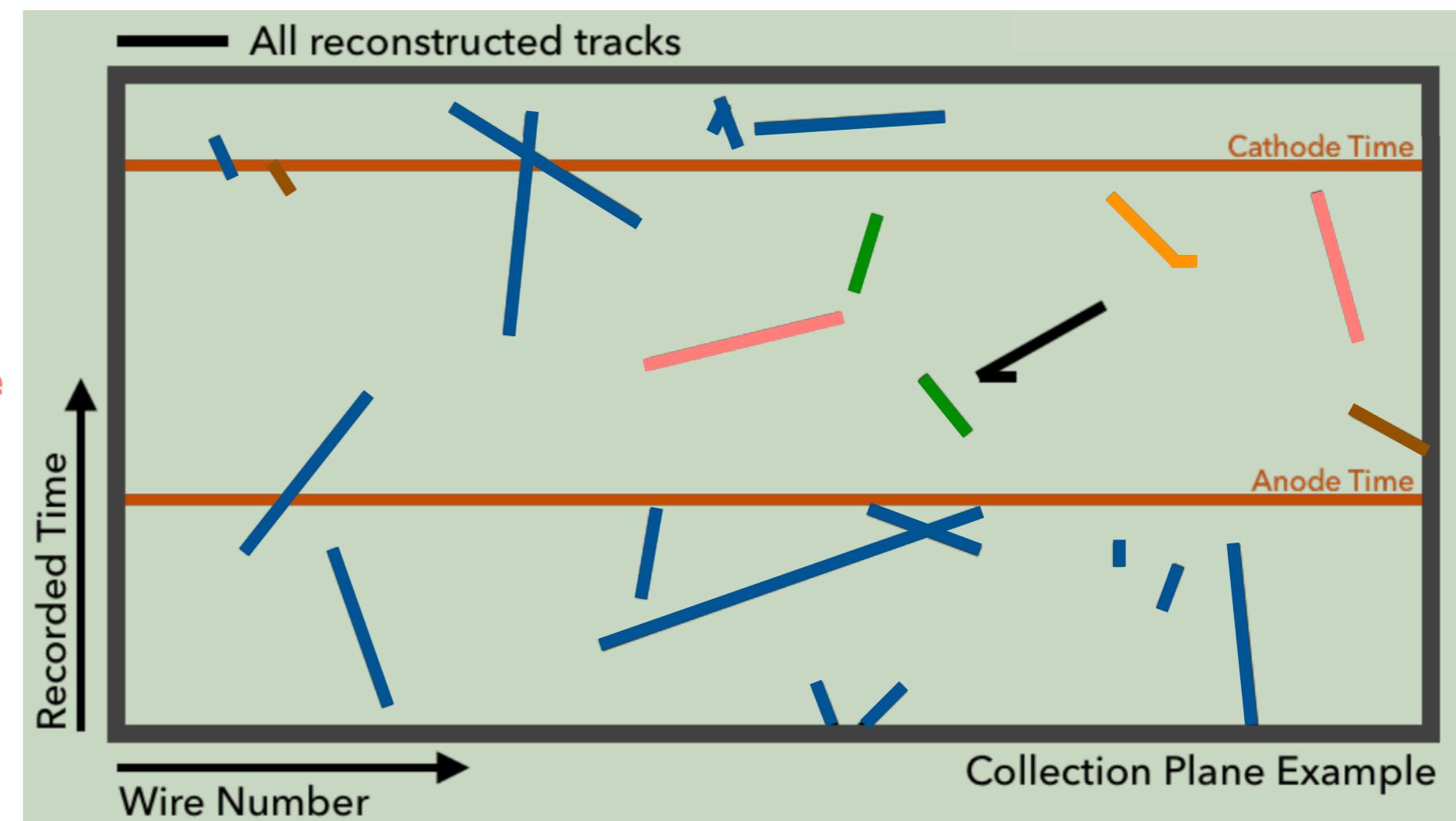
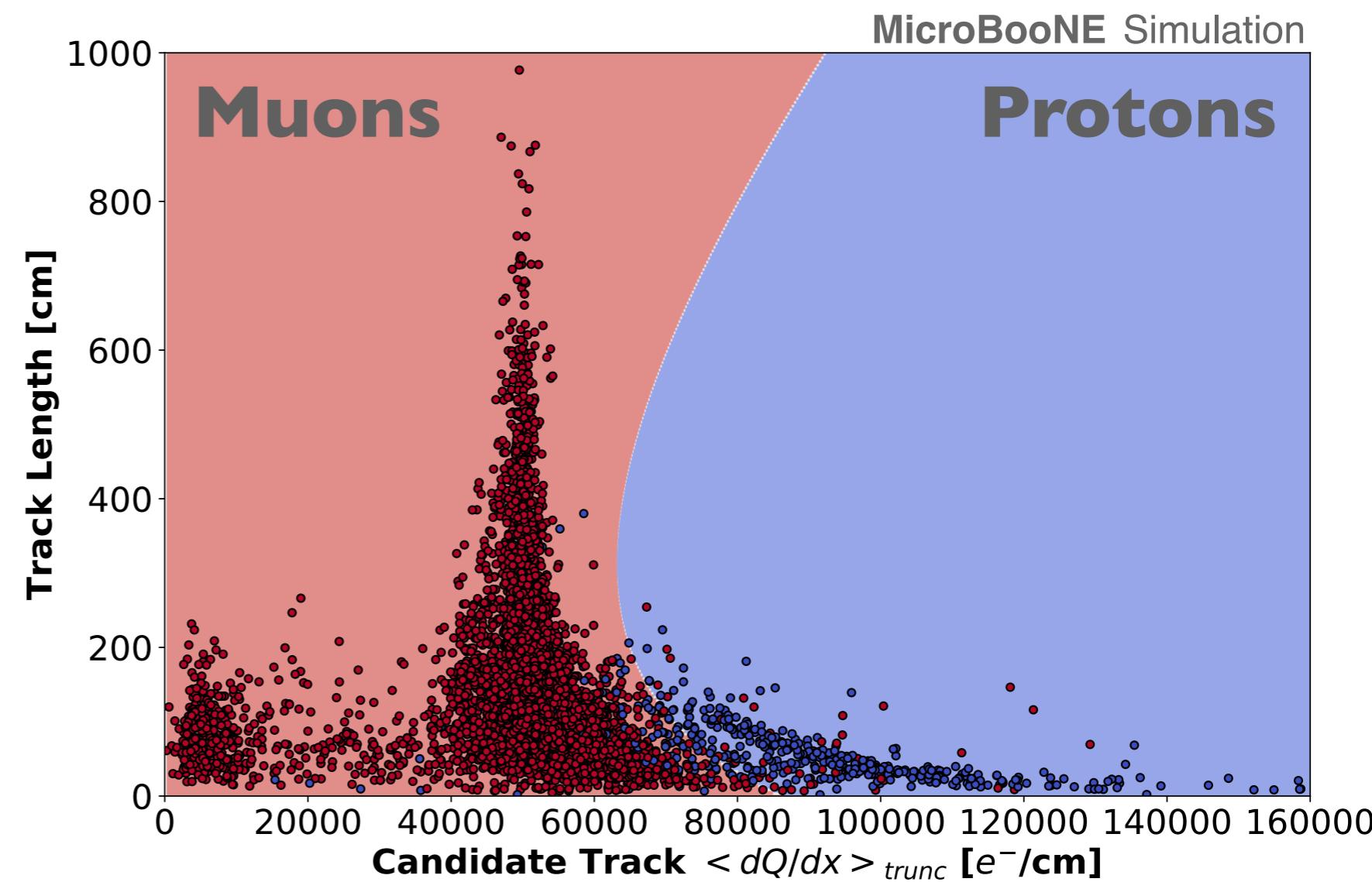


Figure from M. del Tutto

# SELECTING CC EVENTS

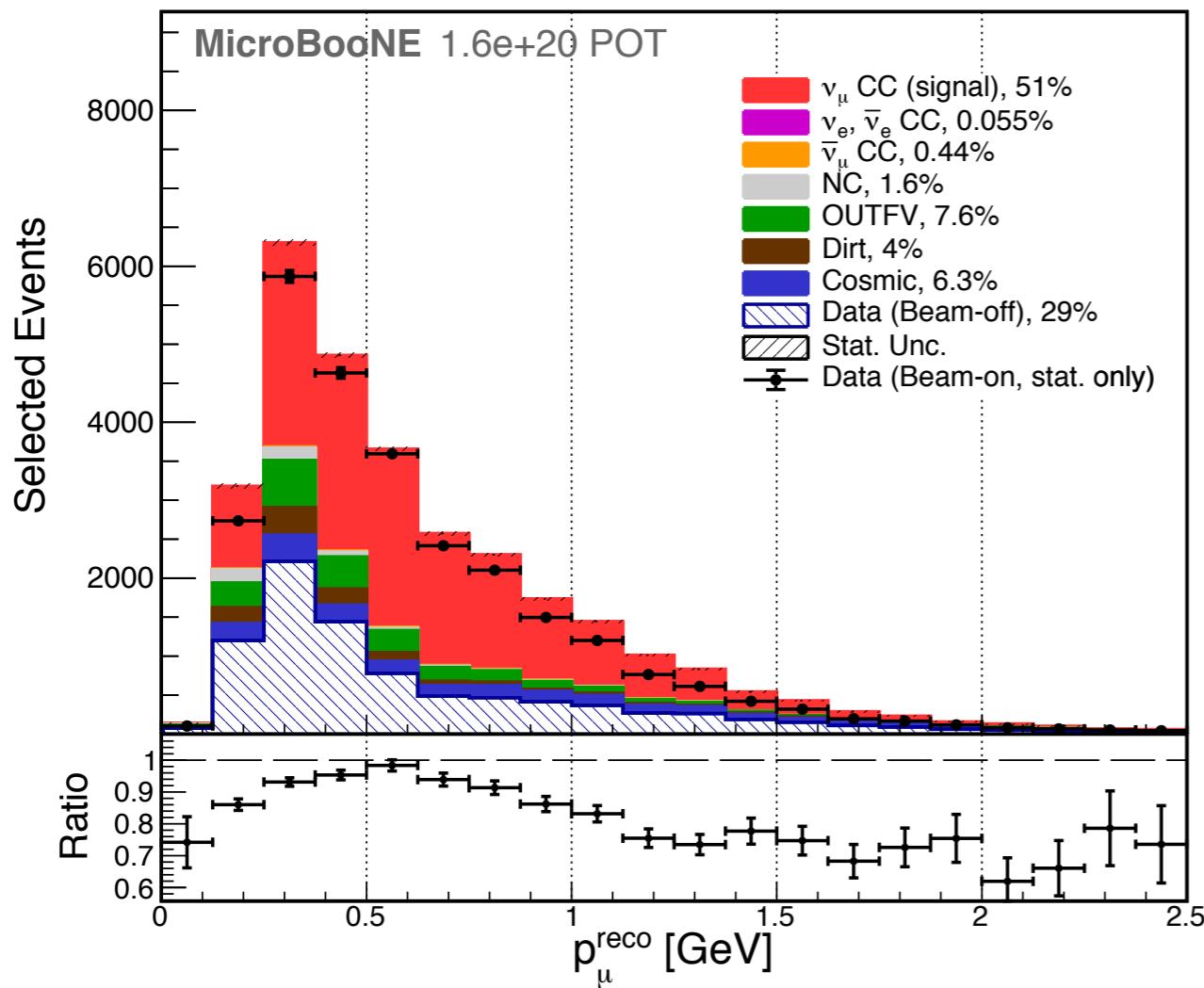
Phys. Rev. Lett.  
123, 131801  
(2019)

- Charged-current events selected by presence of muon
- Longest track in event is muon candidate → use  $dQ/dx$  and track length to reject protons



# SELECTED EVENTS

Phys. Rev. Lett.  
123, 131801  
(2019)



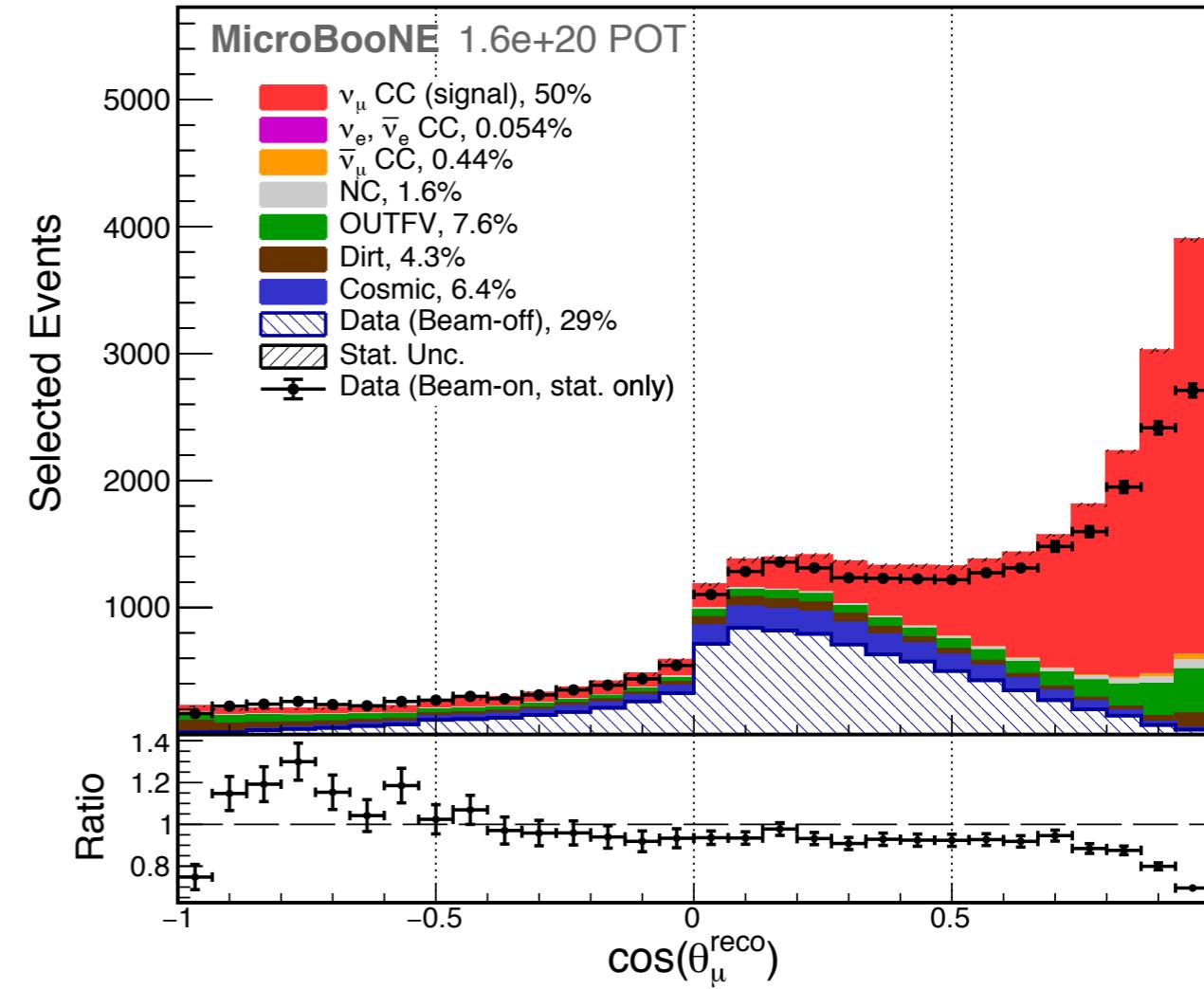
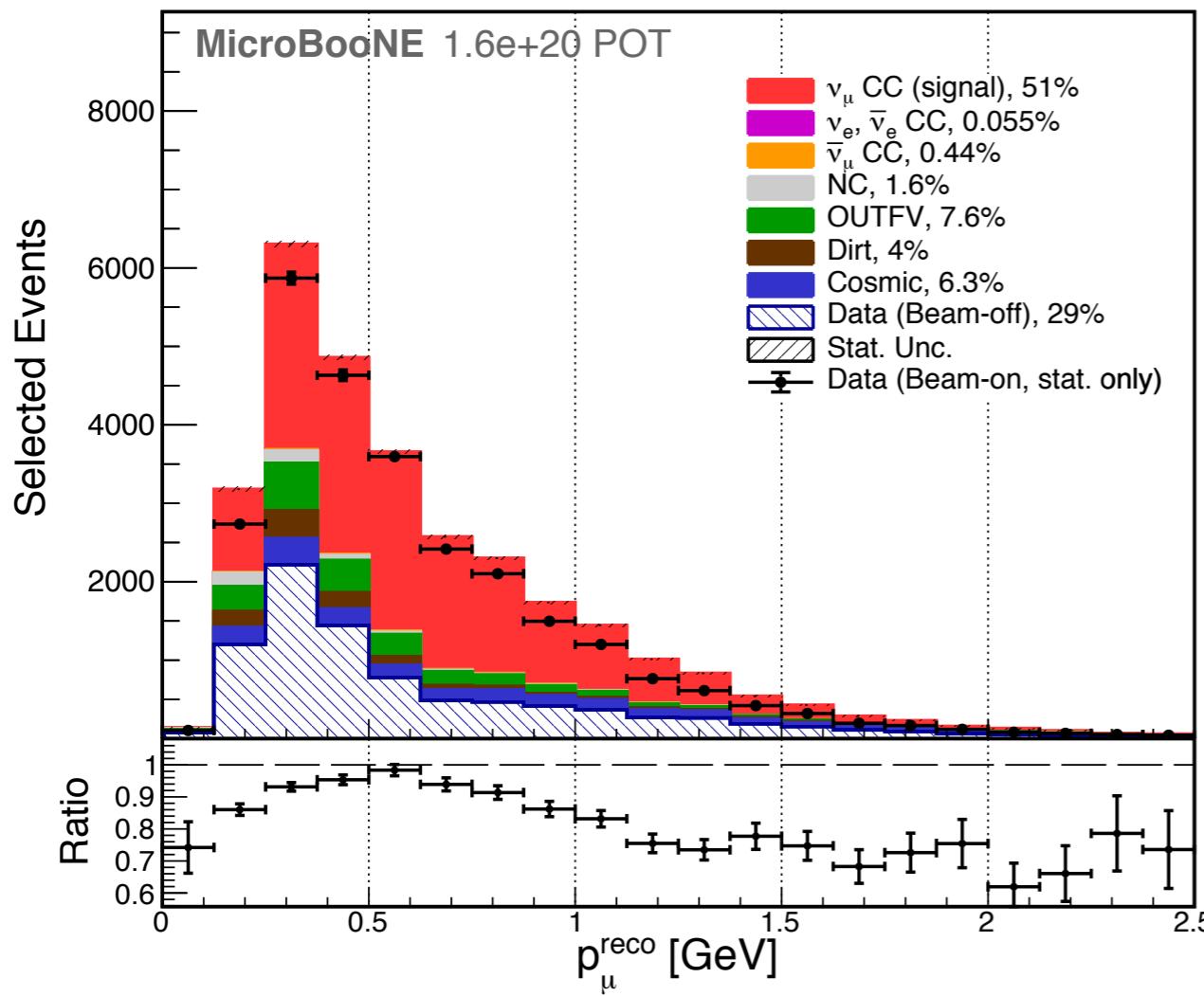
**Signal (CC-inclusive) events:**  
50.4%

Largest background: **cosmics** (29%)  
→ directly measured with beam-off data

Other backgrounds from **neutrino interactions outside fiducial volume** and **cosmic interactions in time with neutrino beam**

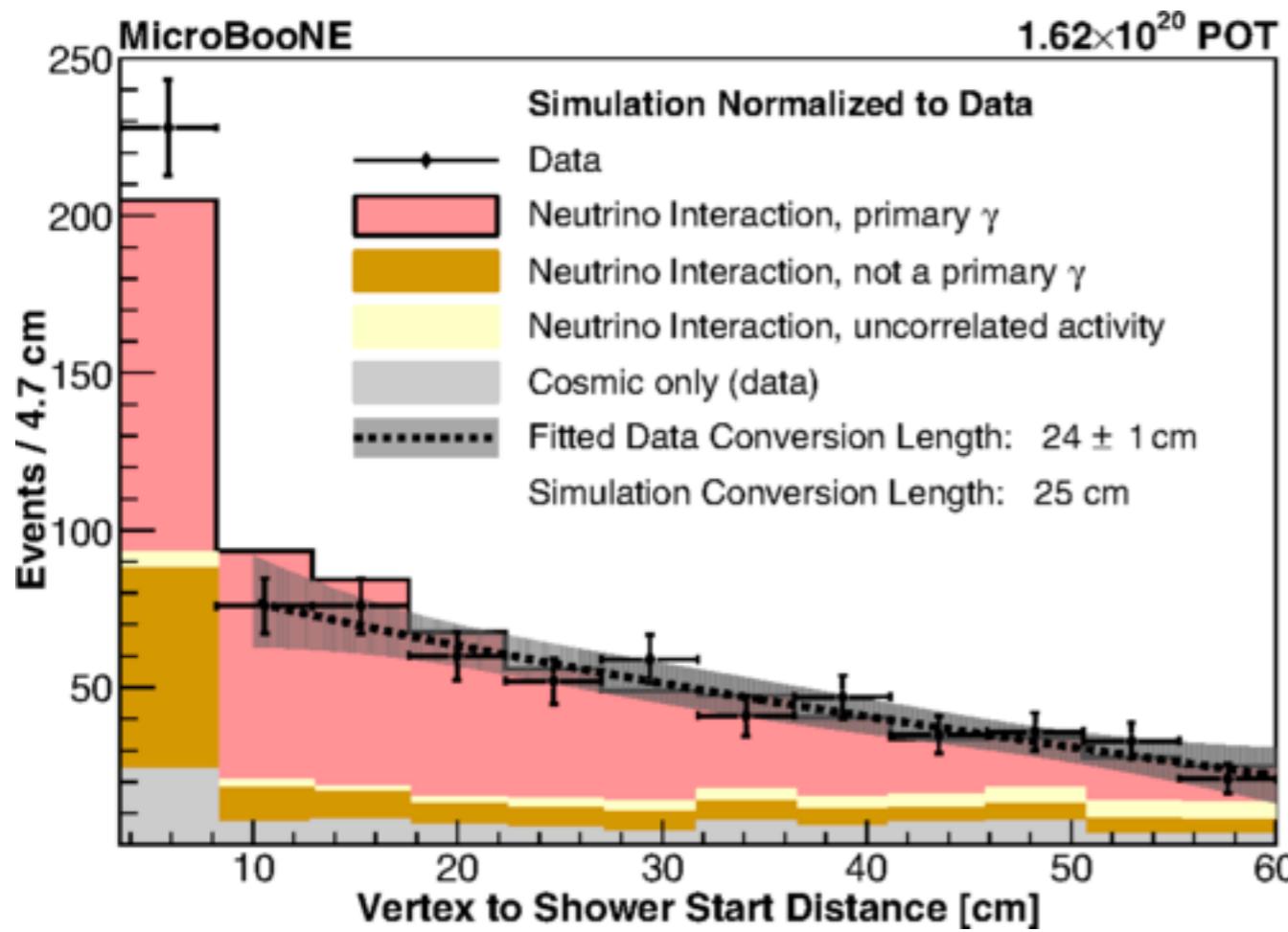
# SELECTED EVENTS

Phys. Rev. Lett.  
123, 131801  
(2019)



**Largest ever sample of neutrino interactions on argon**

# CCPI0: PHOTON CONVERSION



- Look at single photons passing CC $\pi^0$  selection
- Measure mean free path to validate:
  - 1) that we have selected photons
  - 2) that we have the interaction vertex right
- ✓ Data agrees with simulation within uncertainties

Phys. Rev. D 99, 091102(R) (2019)

# CCPIO SELECTIONS

Phys. Rev. D 99, 091102(R) (2019)

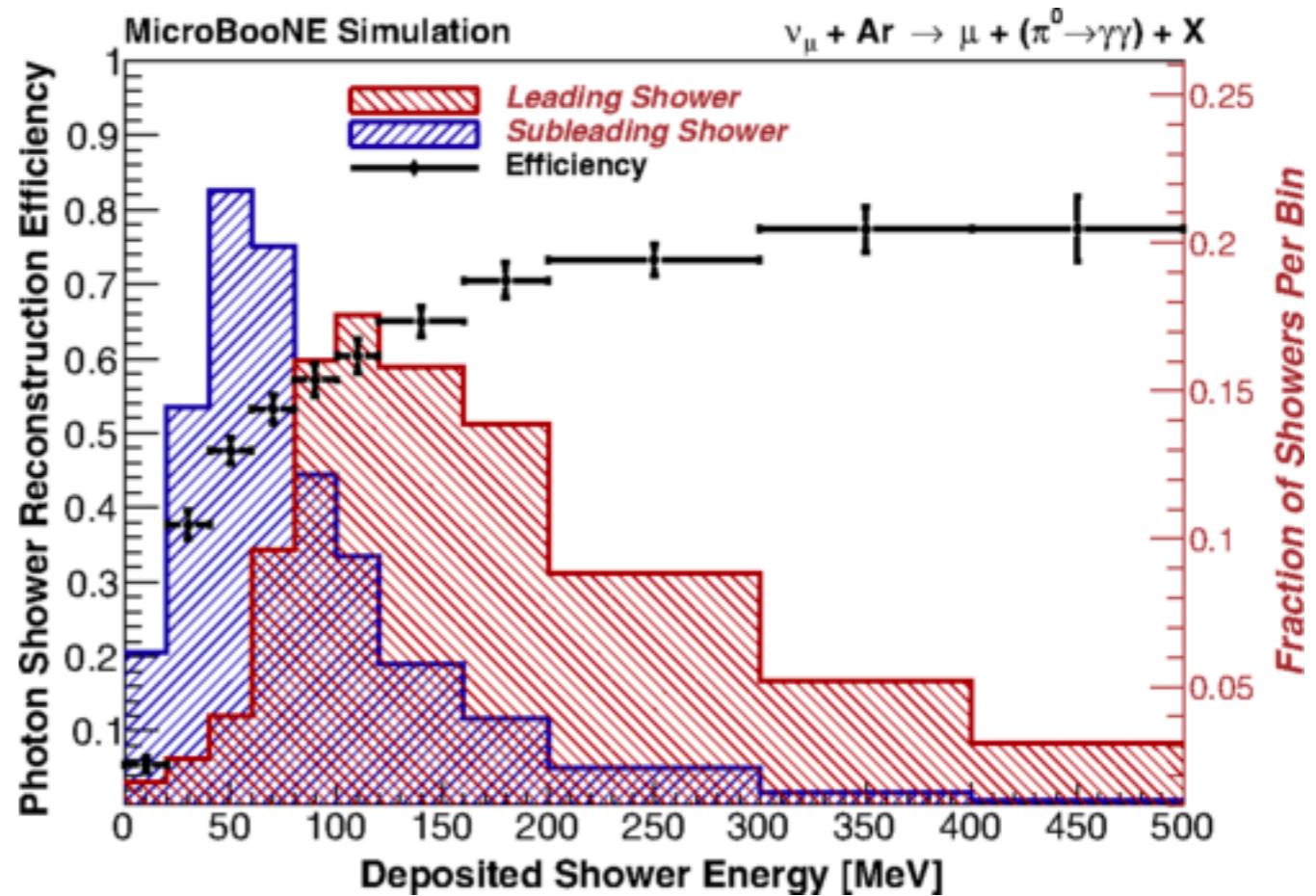
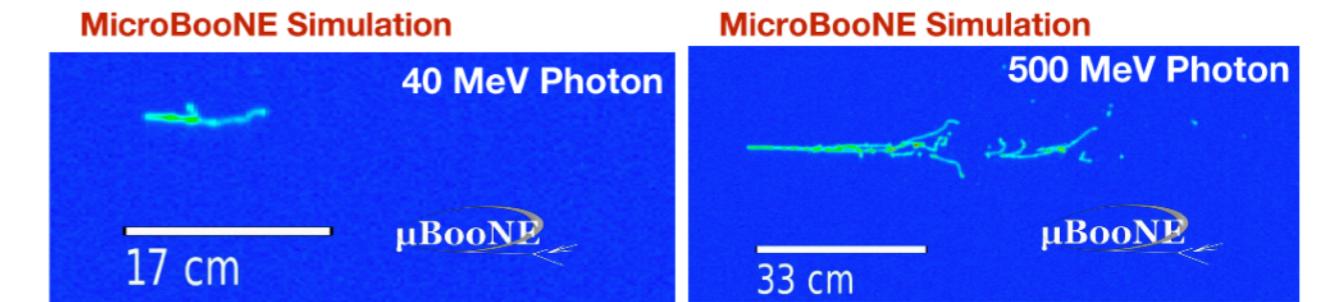
- Low-energy photons appear more track like  
→ low reconstruction efficiency  
→ requiring that we reconstruct both  $\pi^0$  photons limits statistics

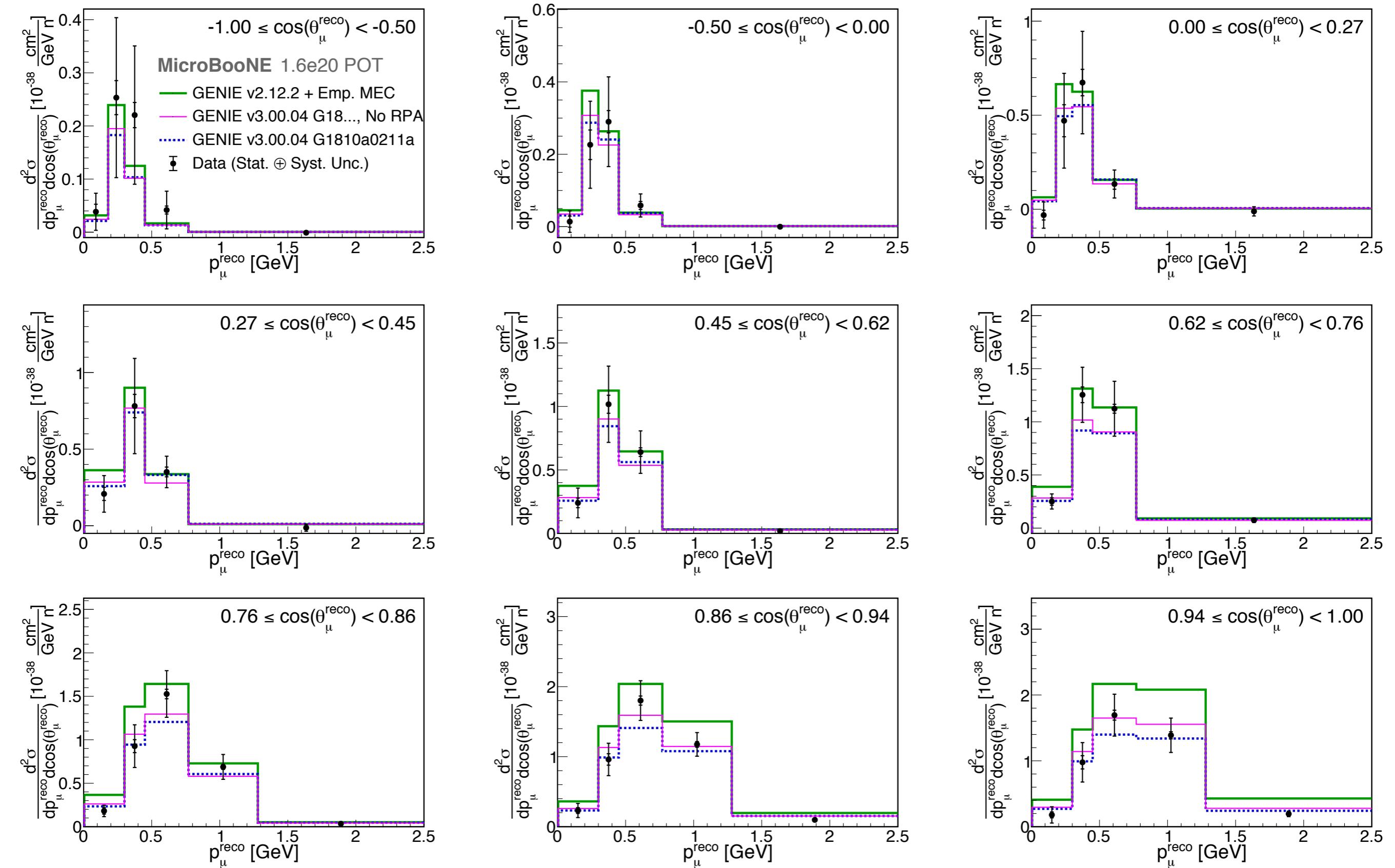
## Two-shower selection

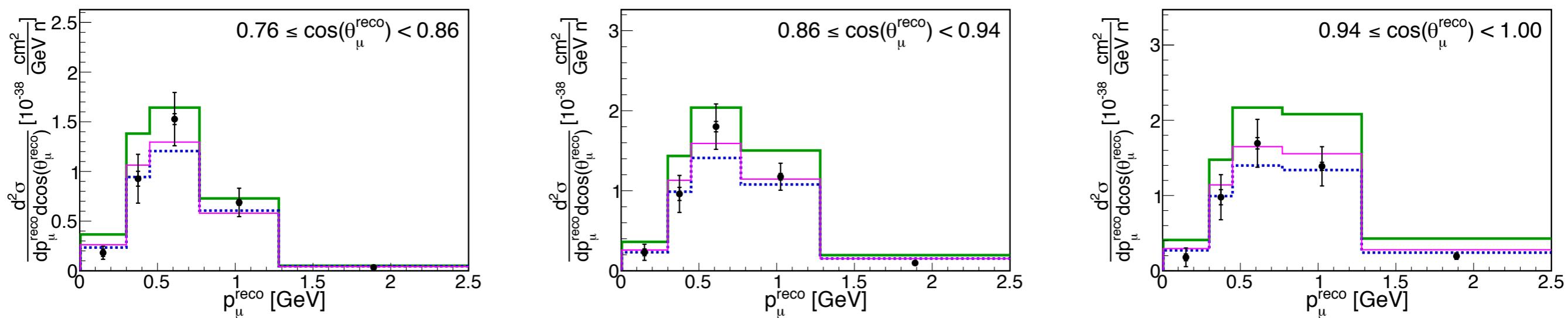
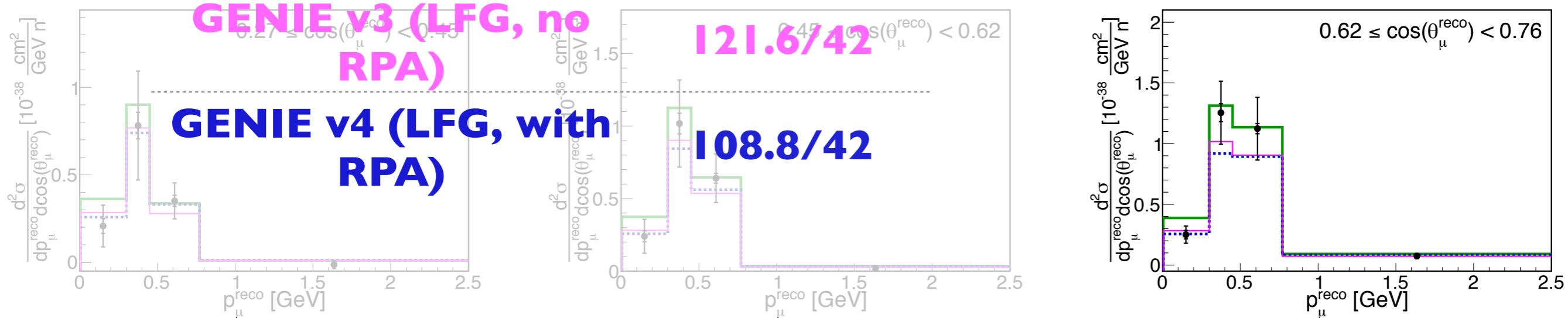
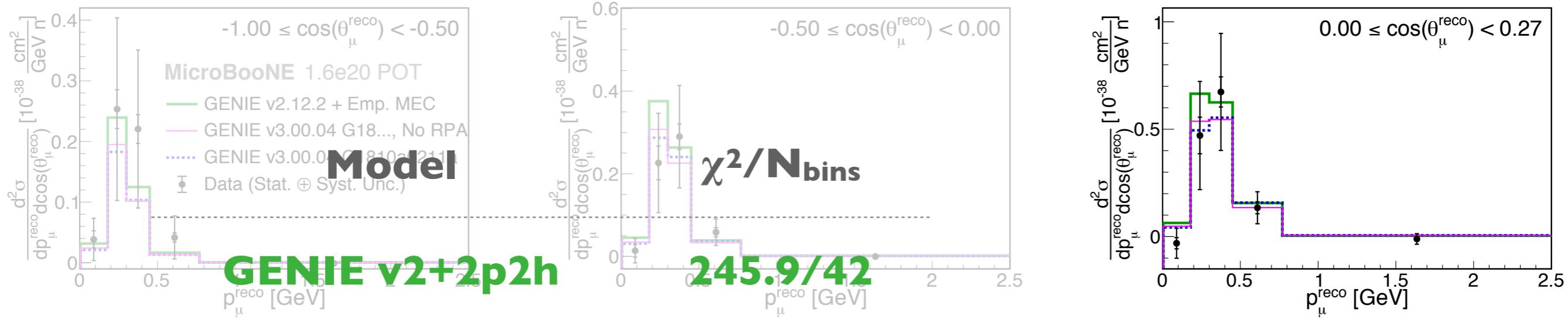
- validate  $\pi^0$  hypothesis by invariant diphoton mass

## Single shower selection

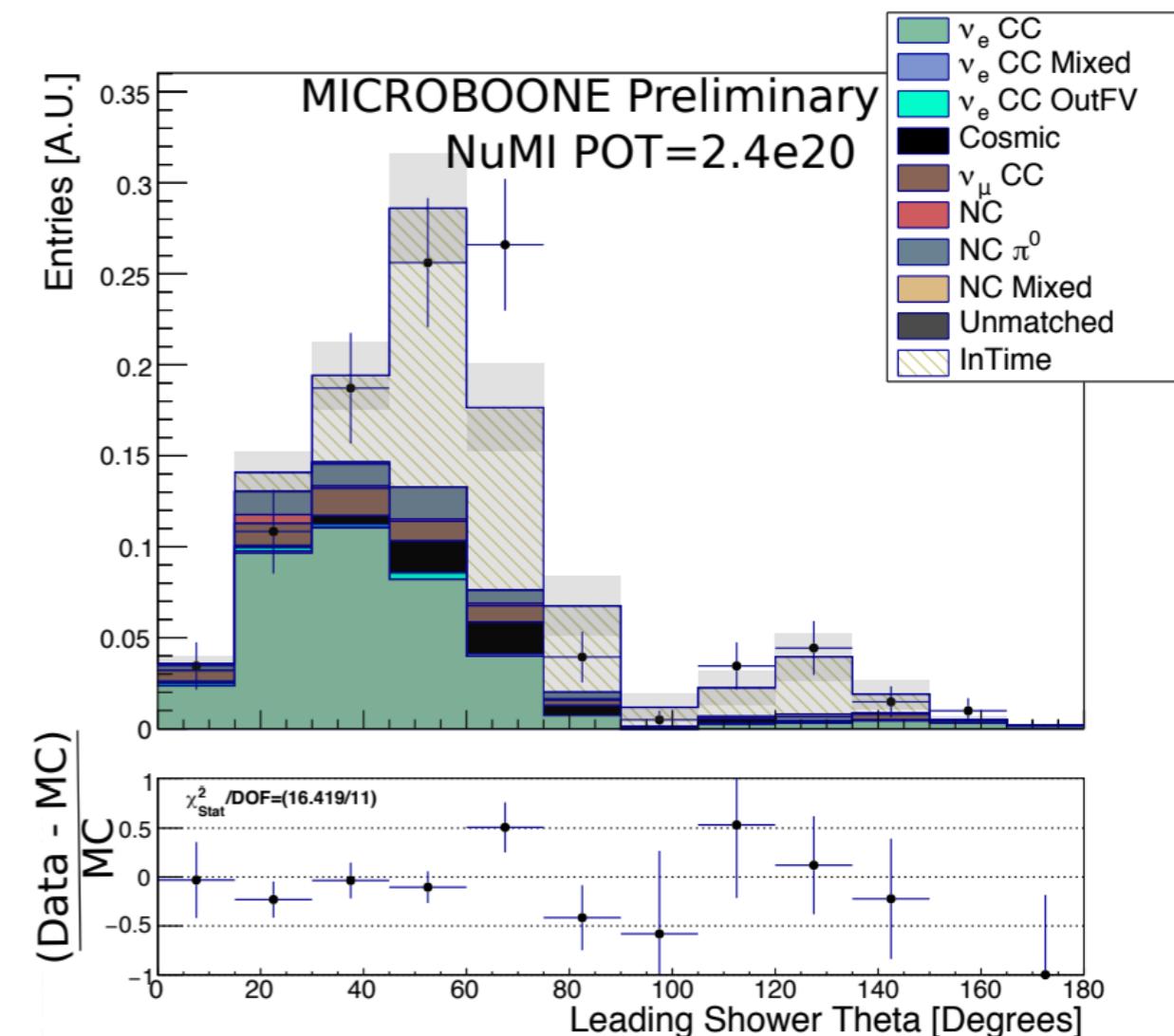
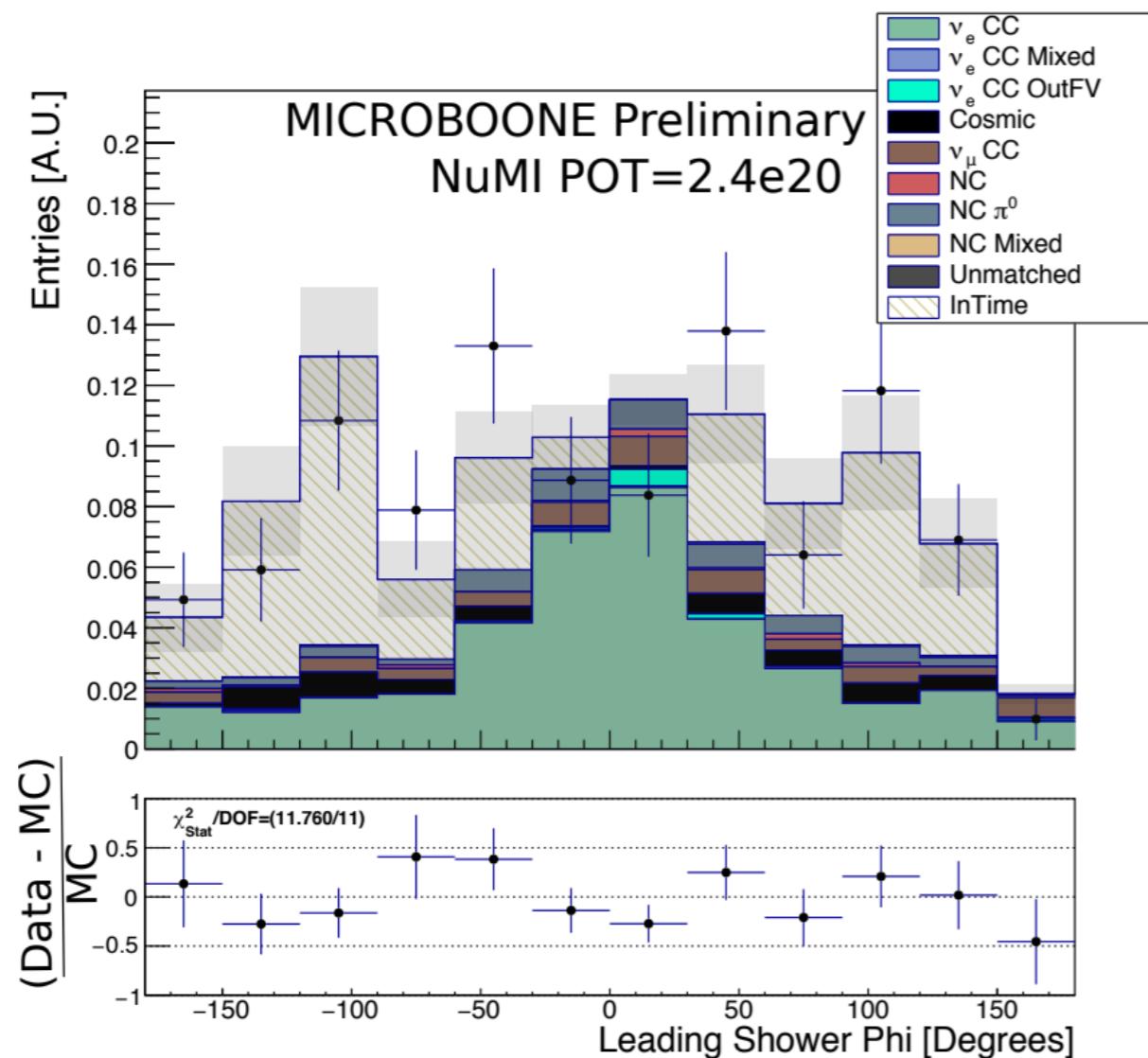
- validate photon hypothesis  
→ maximize statistics for cross section measurement







# NUE SELECTED EVENTS



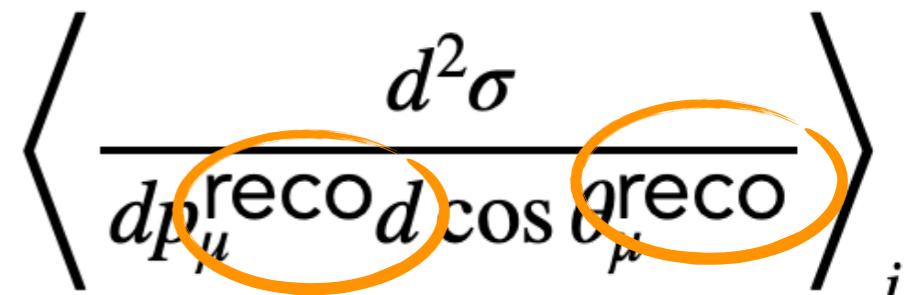
**$\nu_e$  selection efficiency: 9%, purity: 40%**

# CROSS SECTION

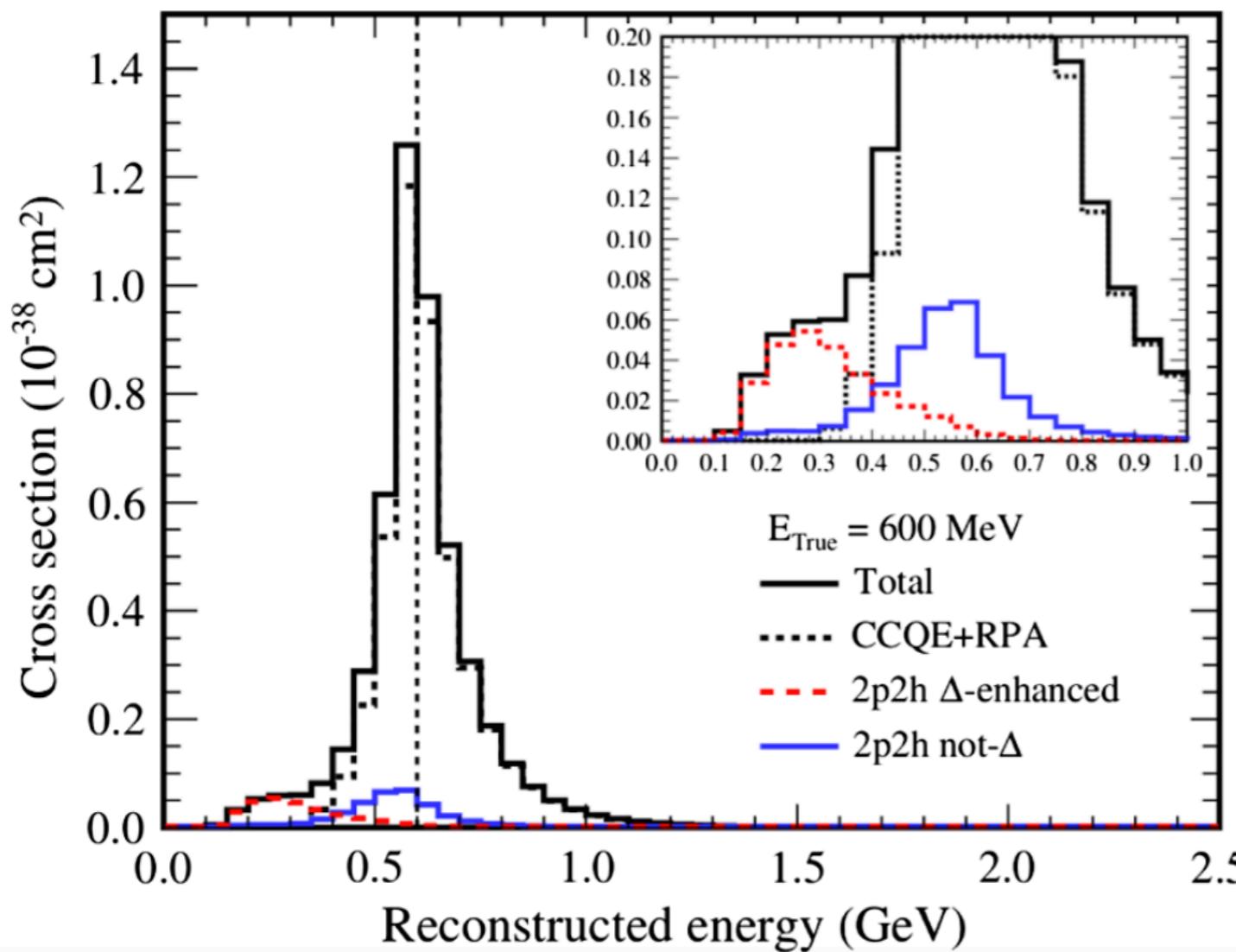
arXiv: 1905.09694  
 [hep-ex]  
 Accepted to PRL

Note: cross section presented in **reconstructed**  $p_\mu - \cos\theta_\mu$  space

- Unlike other cross section measurements we do not unfold to true muon momentum and angle
  - → unfolding introduces bias, inflates uncertainties
- Instead, present result in **reconstructed** muon momentum and angle
- Publish detector smearing and efficiency
  - → theoretical predictions can be **forward folded** (i.e. smeared by our known detector effects)
  - → produce a realistic prediction of what we expect to see in our detector
  - → directly compare to data

$$\left\langle \frac{d^2\sigma}{dp_\mu^{\text{reco}} d\cos\theta_\mu^{\text{reco}}} \right\rangle_i$$

 A mathematical expression for the differential cross section. It consists of a left-angle bracket, a fraction, and a right-angle bracket. The numerator is  $d^2\sigma$ . The denominator has two terms:  $dp_\mu^{\text{reco}}$  and  $d\cos\theta_\mu^{\text{reco}}$ . Both terms are enclosed in orange circles.

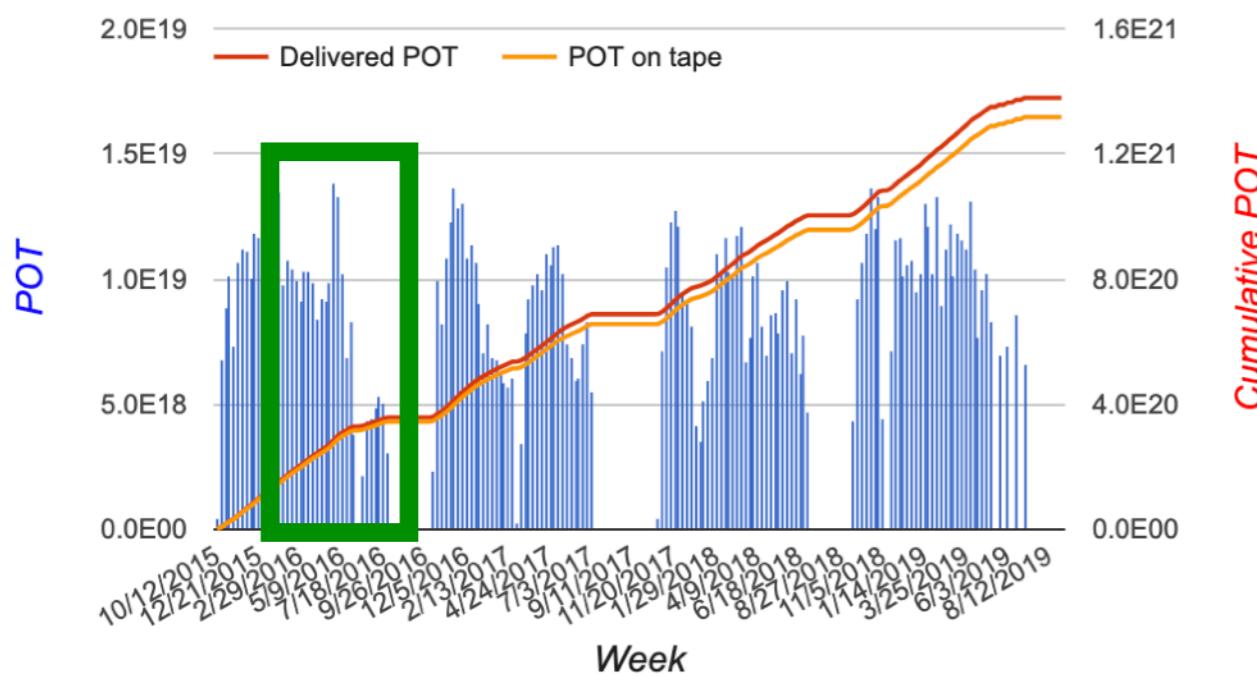
# NEUTRINO INTERACTIONS AND RECONSTRUCTING NEUTRINO ENERGY



Phys. Rev. D 96, 092006 (2017)

- Reconstructing neutrino energy requires assumptions
  - **CCQE formula:** assume all events are CCQE, reconstruct energy from lepton — what if it's not CCQE?
  - **Calorimetric:** measure deposited energy of all particles — what about neutrals/particles you miss?
- Assuming the wrong interaction model:
  - → **bias** in reconstructing neutrino energy
  - → **bias** in oscillation parameters

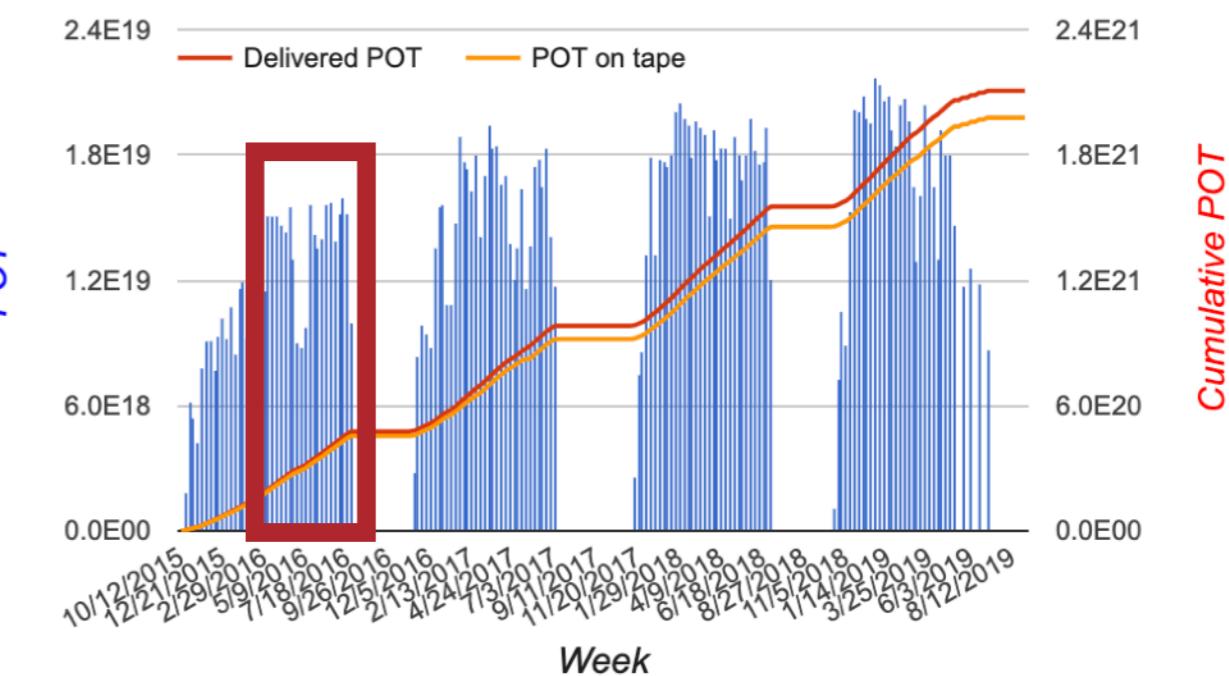
# MICROBOOONE'S DATA



## BNB

$1.38 \times 10^{21}$  POT delivered  
( $1.32 \times 10^{21}$  POT on tape)

This talk: Run I =  $1.6 \times 10^{20}$  POT

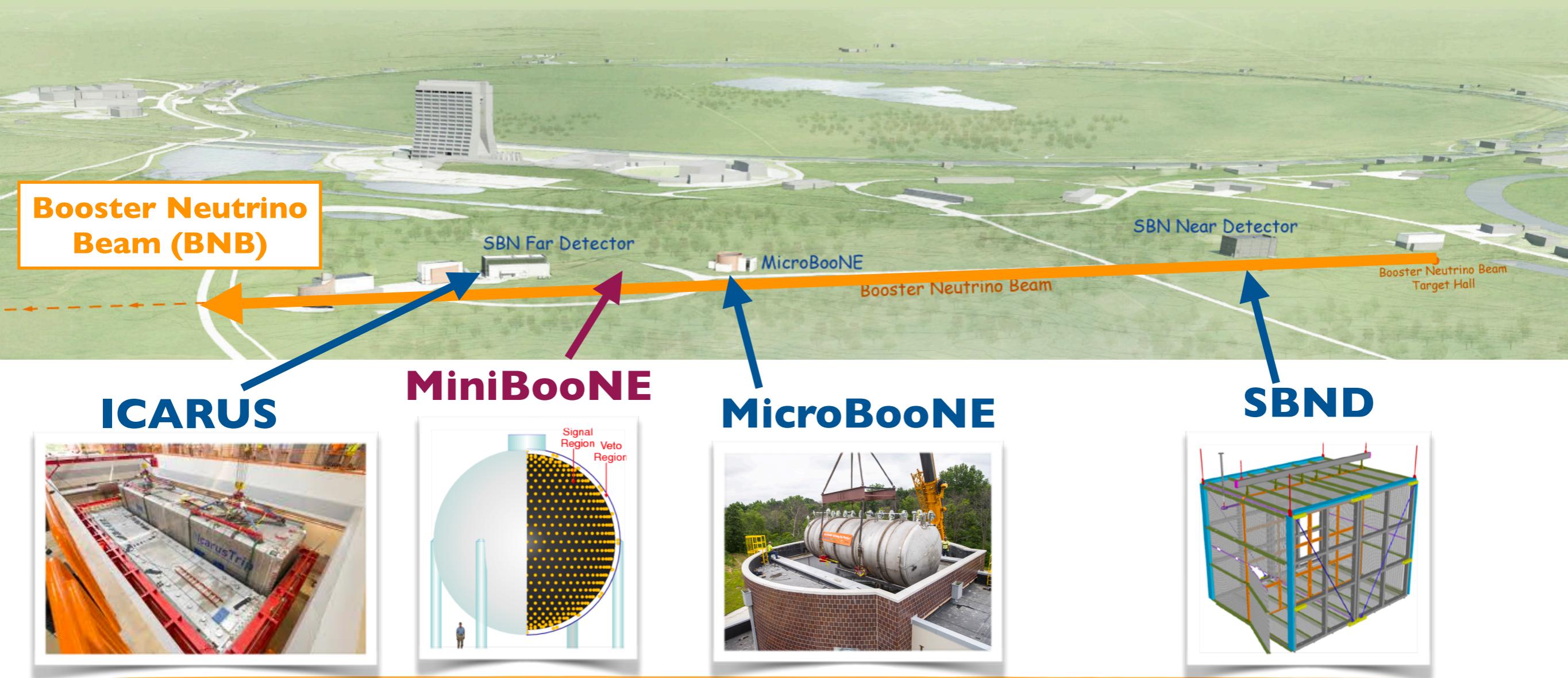


## NuMI

$2.1 \times 10^{21}$  POT delivered  
( $1.98 \times 10^{21}$  POT on tape)

This talk:  $2.4 \times 10^{20}$  POT

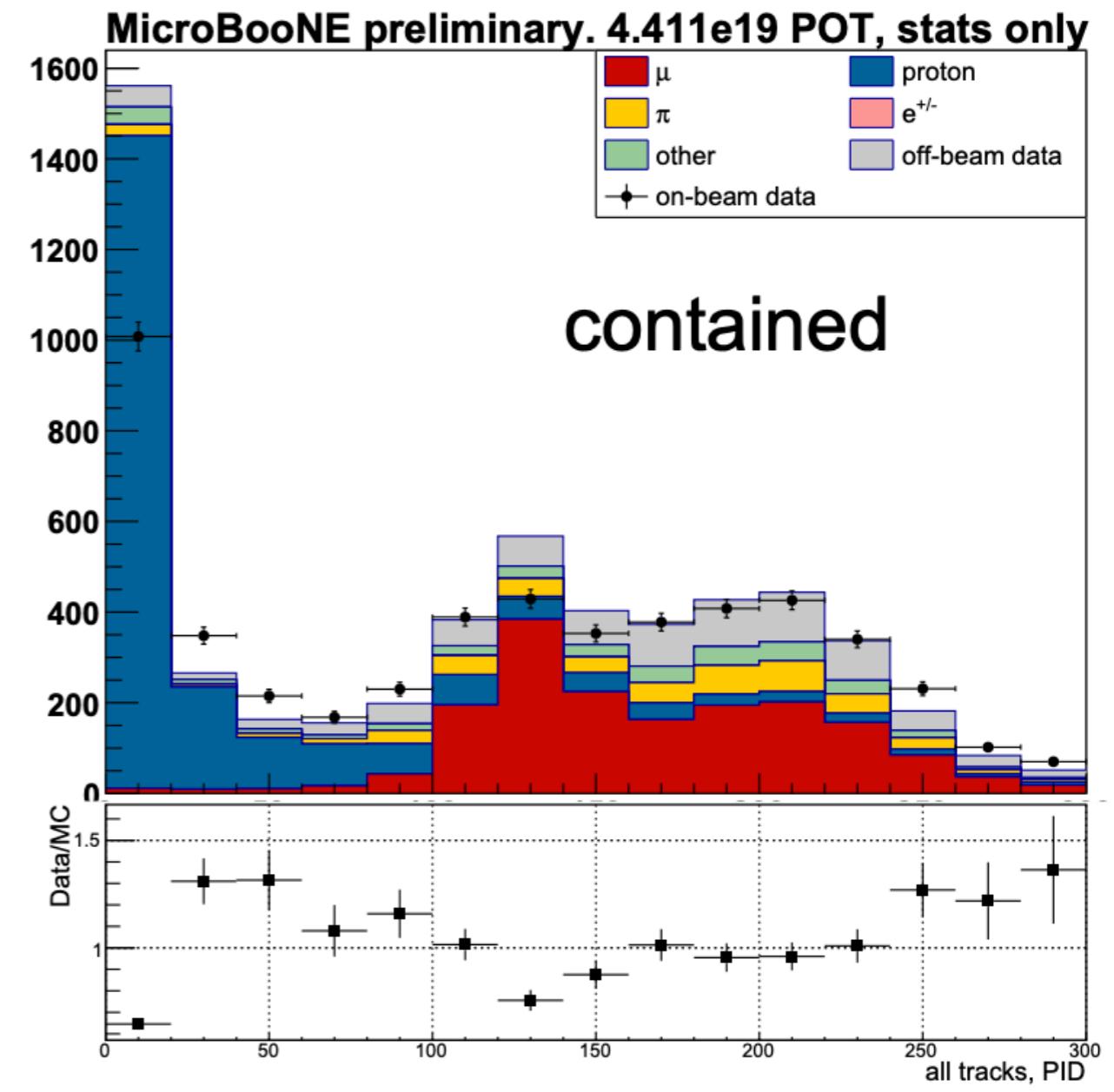
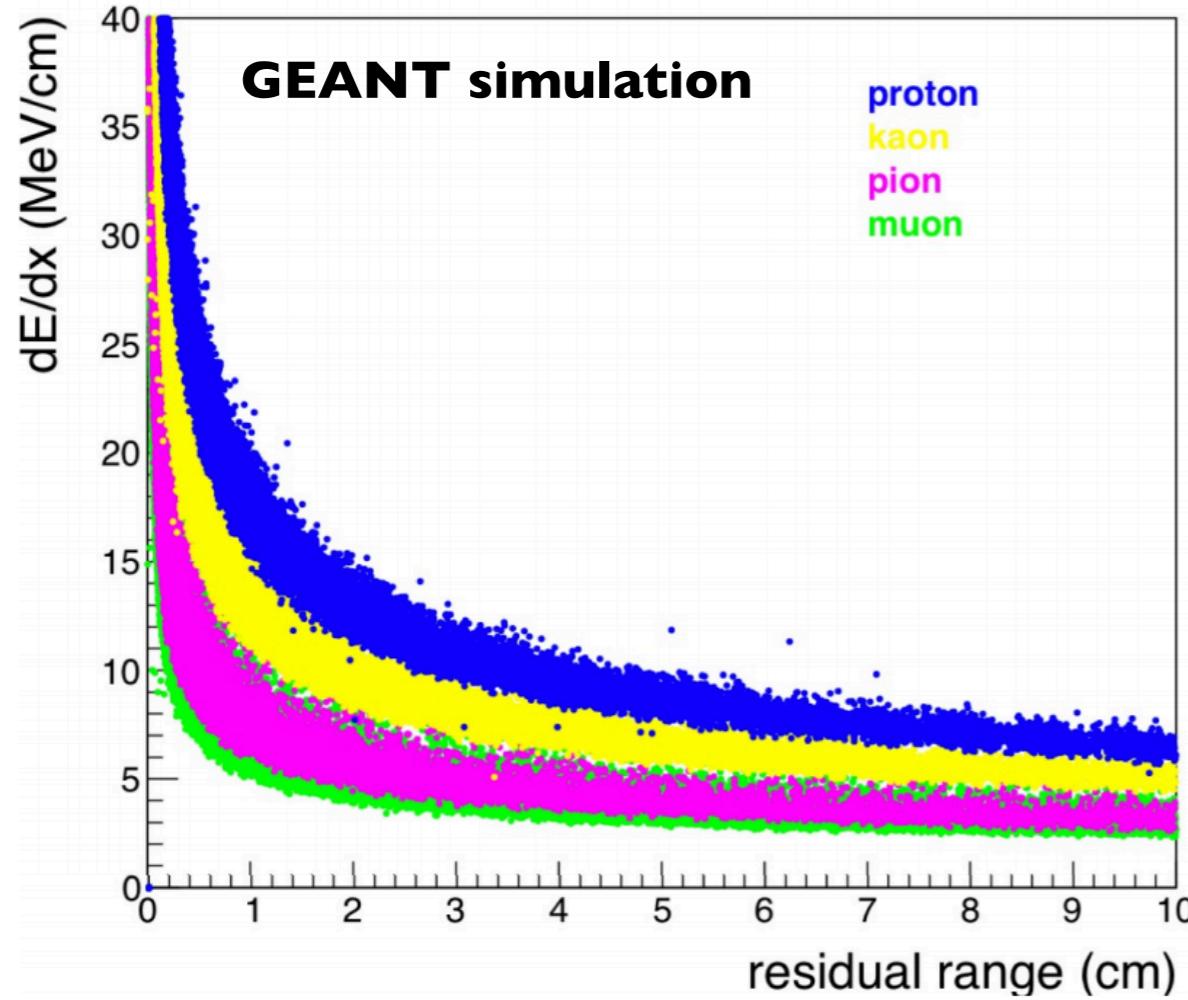
# SHORT-BASELINE NEUTRINOS AT FERMILAB



# IDENTIFYING $\pi^+$

MICROBOONE-  
NOTE-1056-PUB

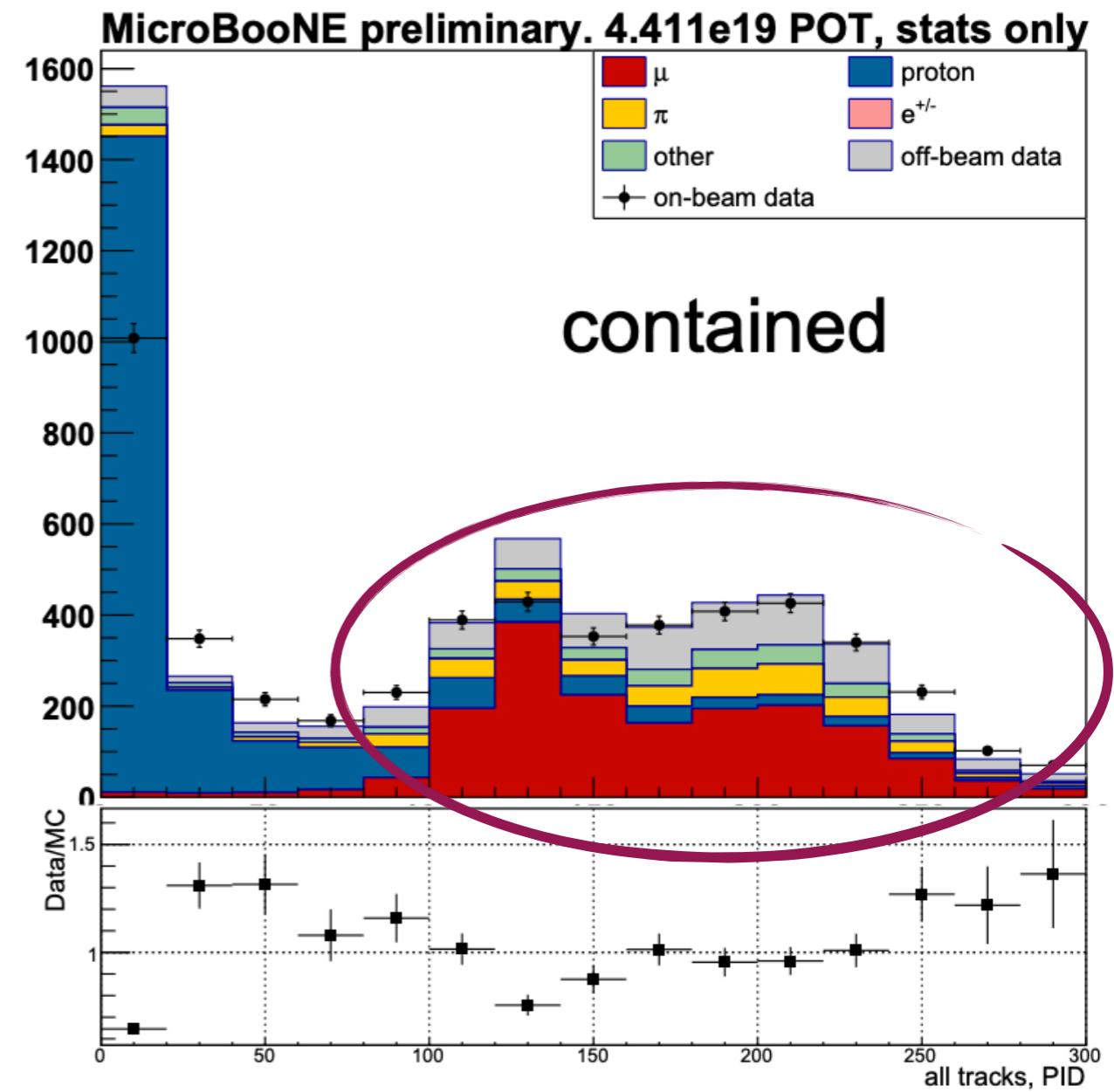
Separate charged pions from protons using deposited energy in track per unit length



# IDENTIFYING $\pi^+$

MICROBOONE-  
NOTE-1056-PUB

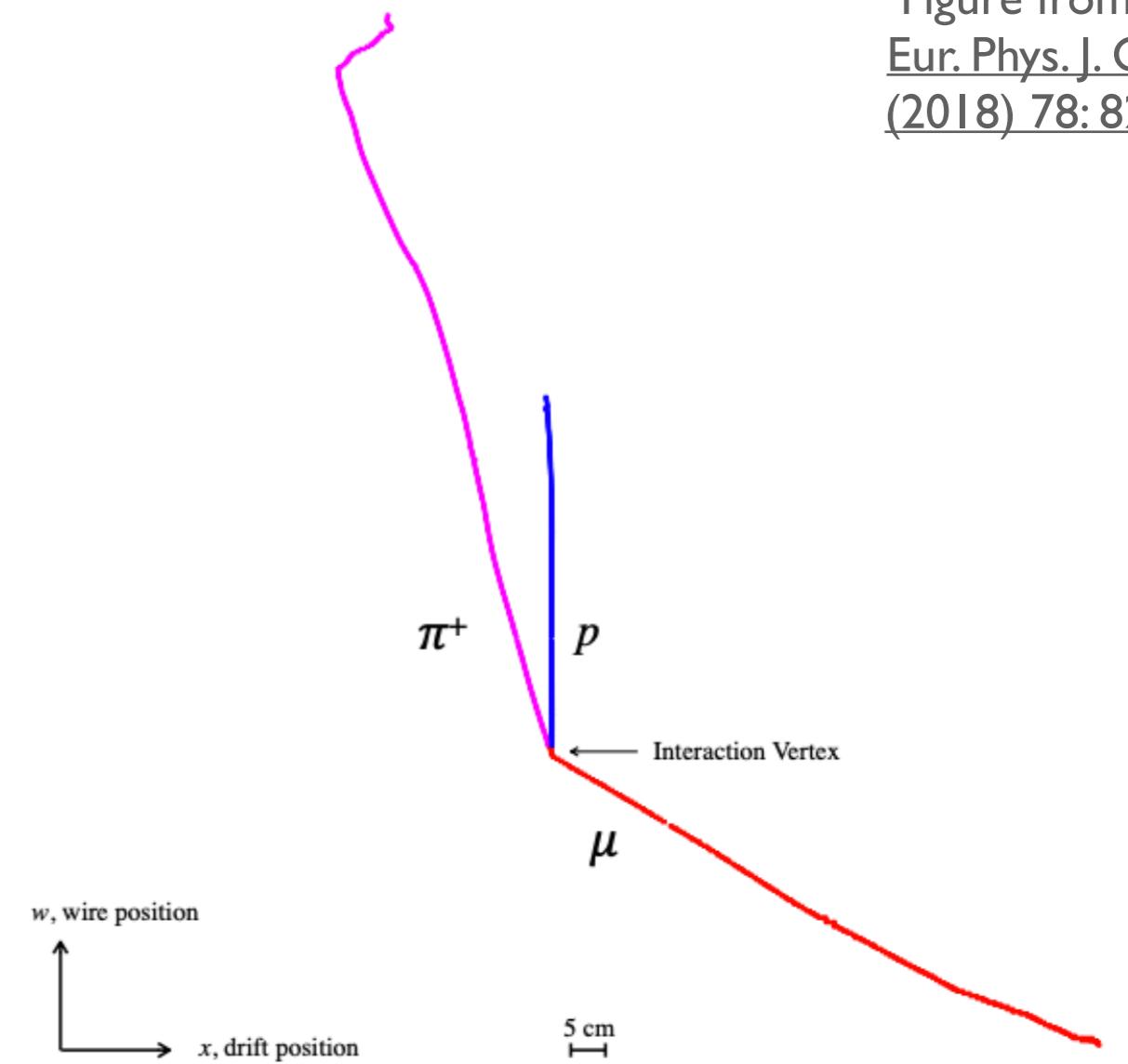
- Separating **charged pions** from **muons** is harder → at our energies,  $dE/dx$  profile is ~identical
- So how can we distinguish charged pions from muons in a CC1pi+ measurement?



# RECONSTRUCTING $\pi^+$

- Charged pions can be difficult to reconstruct
- Large amount of scattering: may be reconstructed as a shower, rather than a track
- Secondary vertices often missed: pion and daughters can be merged

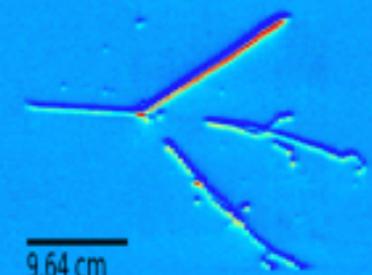
Figure from  
Eur. Phys. J. C  
(2018) 78: 82



# HADRONIC INTERACTIONS

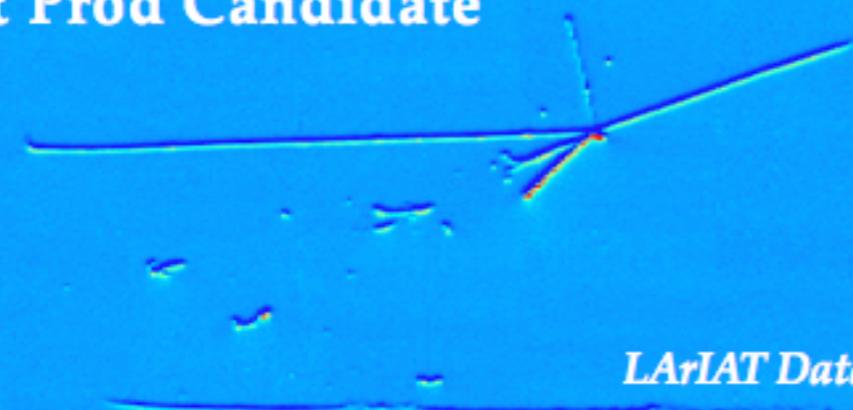
**LArIAT  
Data**

**Charge Exchange Candidate**



*LArIAT Data*

**$\pi$  Prod Candidate**



*LArIAT Data*

**Elastic Scattering Candidate**

*LArIAT Data*

**Inelastic Scattering Candidate**

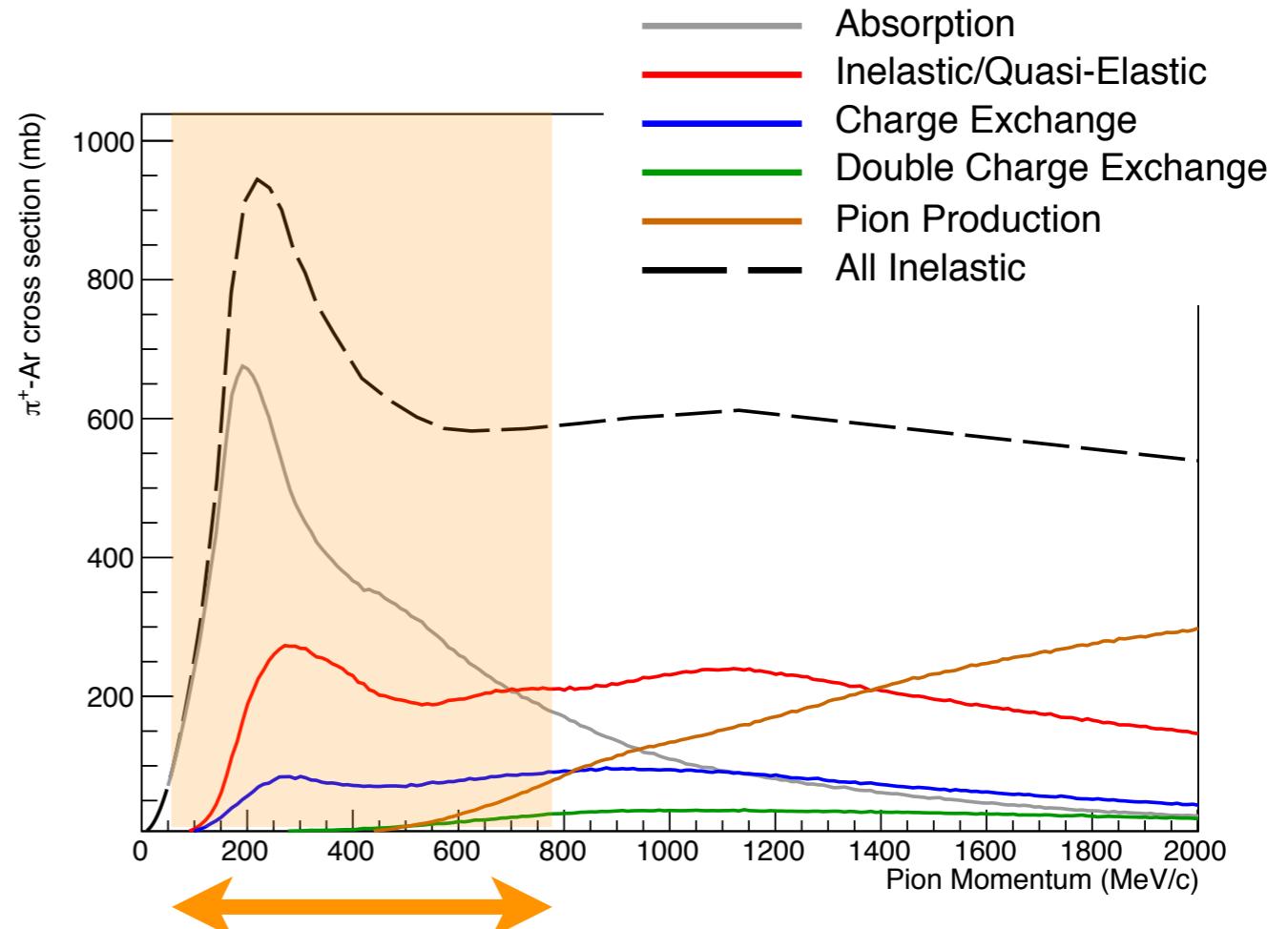
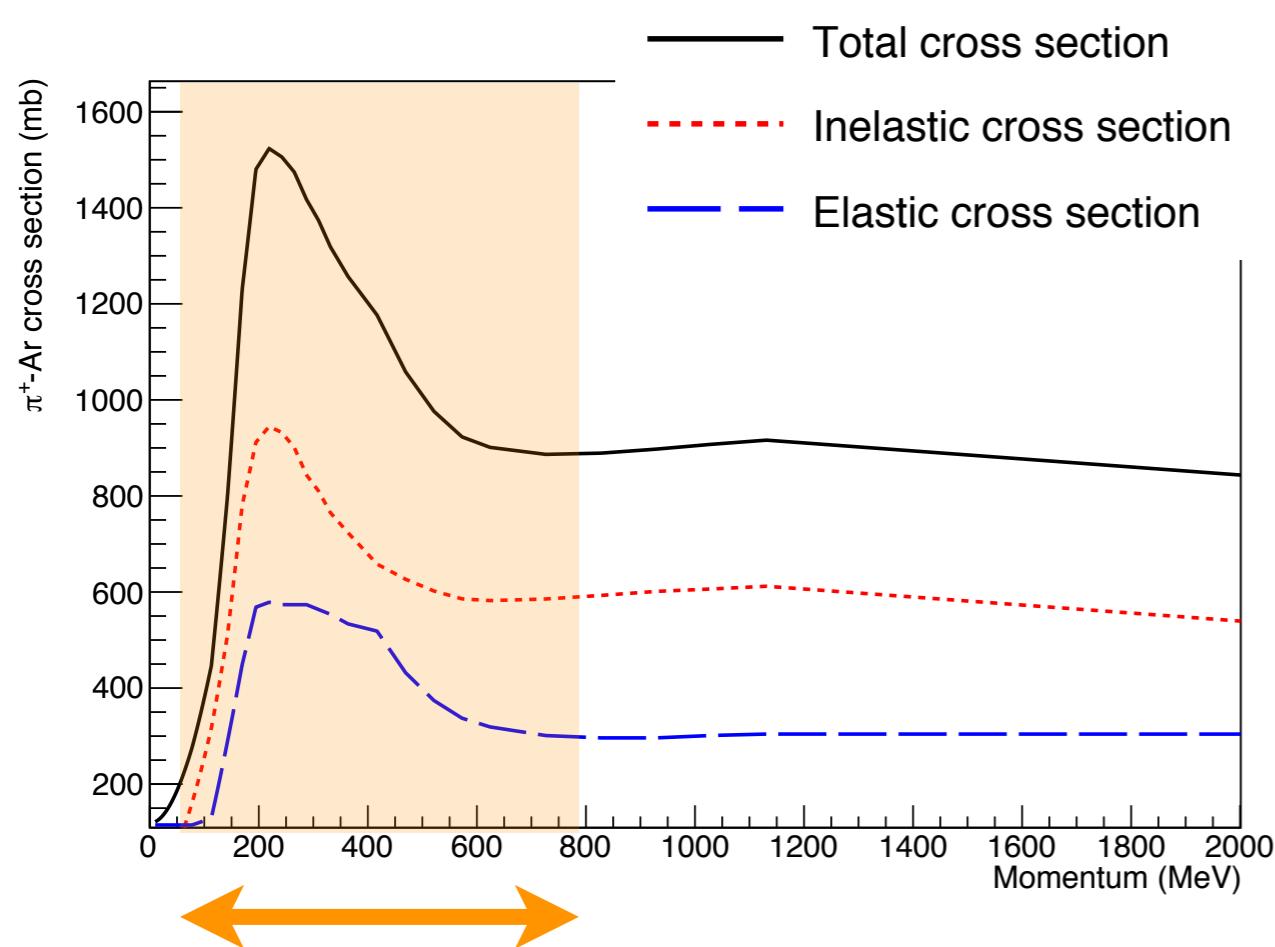
*LArIAT Data*

**Absorption Candidate ( $\pi \rightarrow 3p$ )**

*LArIAT Data*

# $\pi^+$ INTERACTIONS

Plots made using GEANTReWeight (J. Calcutt)



Roughly speaking, GENIE predicts  $\pi^+$  produced in MicroBooNE will fall in **this energy range**

# PUTTING IT ALL TOGETHER

