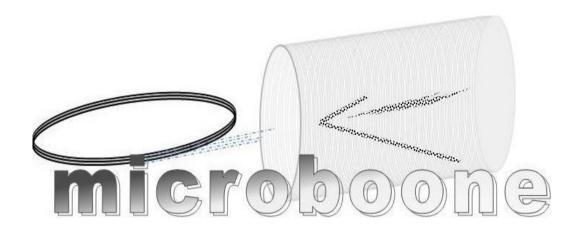
Status of the MicroBooNE LArTPC and

Prospective low-energy neutrino interaction measurements in Argon



Sowjanya Gollapinni (KSU) (on behalf of the MicroBooNE collaboration)

NuInt14, 22 May 2014 Surrey (UK)

Introduction

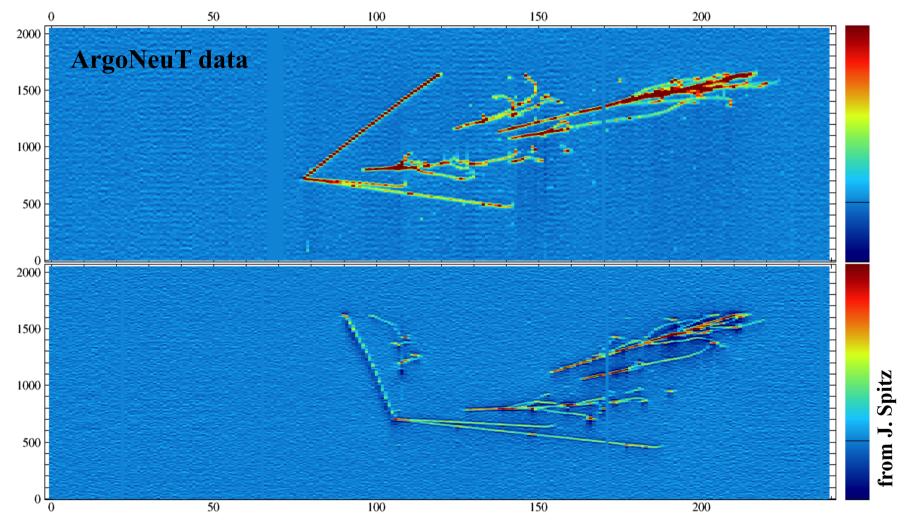
- Liquid argon time projection chambers (LArTPC) are quickly becoming the technology choice for future neutrino detectors
 - Make great precision neutrino detectors
 offer excellent spatial resolution and calorimetry
 - Can address wide array of physics goals
 low-energy neutrino cross section measurements, proton decay searches, supernova neutrinos, dark matter searches etc.
- MicroBooNE is a new 170-ton LArTPC neutrino experiment (largest so far in the U.S.!) being built on the Fermilab Booster Neutrino Beam (BNB) line.

Main physics goals

- Address the low-energy electron-like excess observed by the MiniBooNE experiment
- Make first high-statistics measurement of low-energy (~1GeV) neutrino interactions in Argon and study nuclear effects

LArTPCs are "imaging" detectors

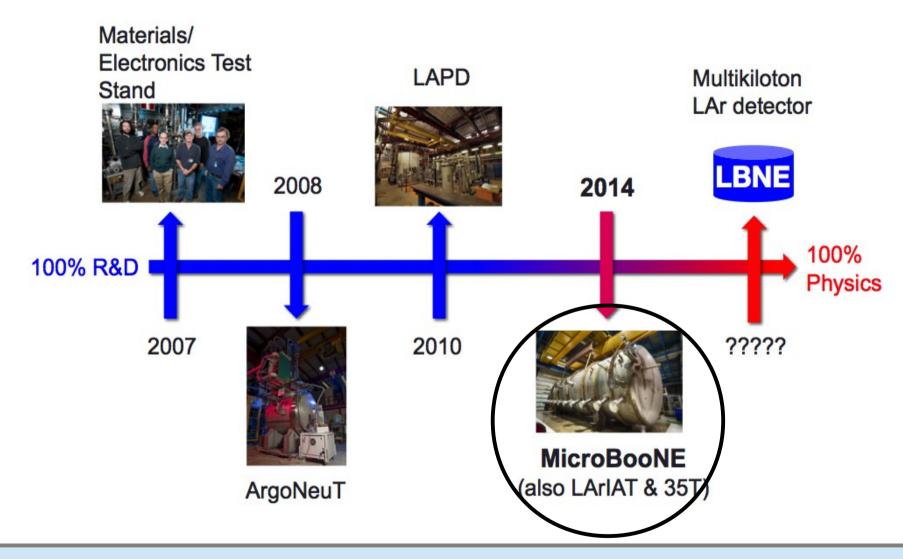
Bubble chamber quality images only in HD!



A charged current neutrino DIS event with two pi0 decays

LArTPCs provide unprecedented amount of details to study neutrino interactions in Argon!

Role of MicroBooNE in LArTPC R&D



MicroBooNE is an important step towards LArTPC R&D in establishing large-scale detectors in the U.S.

Outline

- Principle of LArTPC & why LAr?
- The MicroBooNE LArTPC
 - Design and current status
- MicroBooNE physics
 - MicroBooNE location
 - Physics motivation
 - Cross section measurements
- Summary

Why LAr as neutrino target?

Noble liquids, in general, make great neutrino targets:

- \rightarrow dense!
- → easily ionizable
- → highly scintillating (provides another detection method!)
- → If pure => high electron mobility => long drift lengths!
- → Great dielectric medium => can hold very high voltages etc.

	9	Ne	Ar	Kr	Xe	Water
Boiling Point [K] @ 1atm	4.2	27.1	87.3	120.0	165.0	373
Density [g/cm³]	0.125	1.2	1.4	2.4	3.0	1
Radiation Length [cm]	755.2	24.0	14.0	4.9	2.8	36.1
dE/dx [MeV/cm]	0.24	1.4	2.1	3.0	3.8	1.9
Scintillation [γ/MeV]	19,000	30,000	40,000	25,000	42,000	
Scintillation λ [nm]	80	78	128	150	175	

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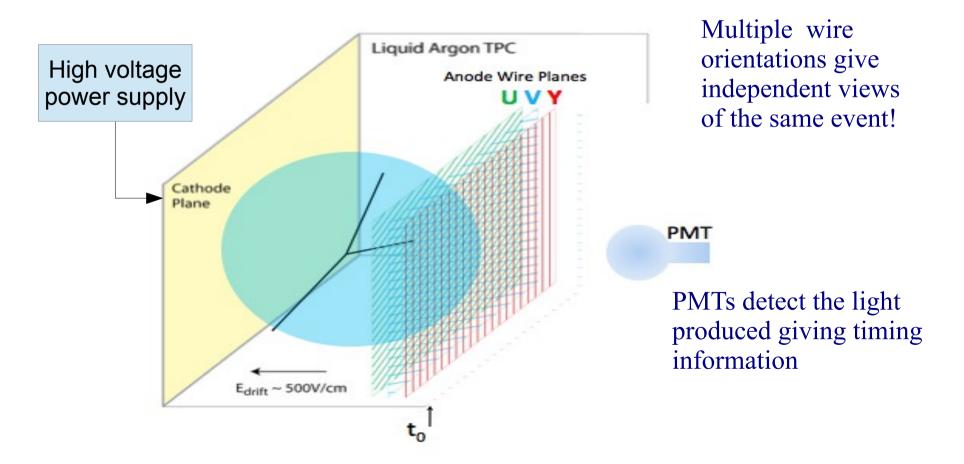


From M. Soderberg

Principal of LArTPC

Neutrino interactions with LAr in the TPC produces charged particles that cause Ionization and excitation of Argon

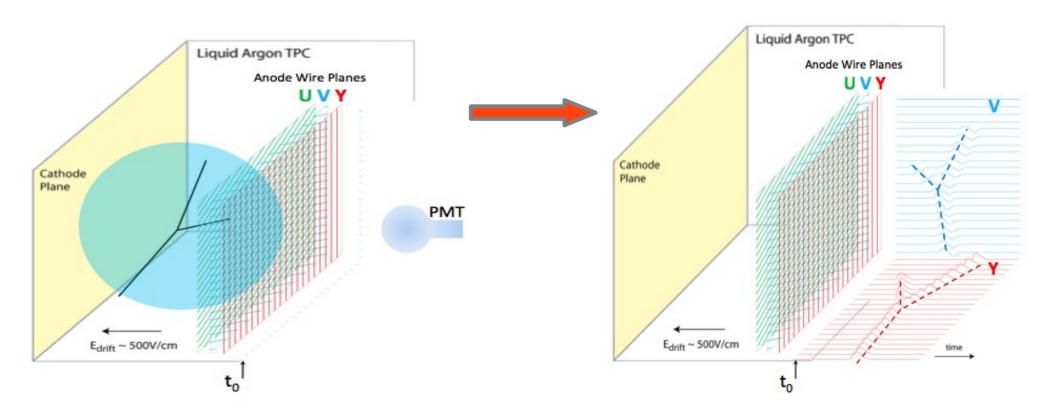
- → High E field drifts electrons towards finely segmented anode wire planes
- → Excitation of Ar produces prompt scintillation light giving "t₀" of the interaction



Principal of LArTPC

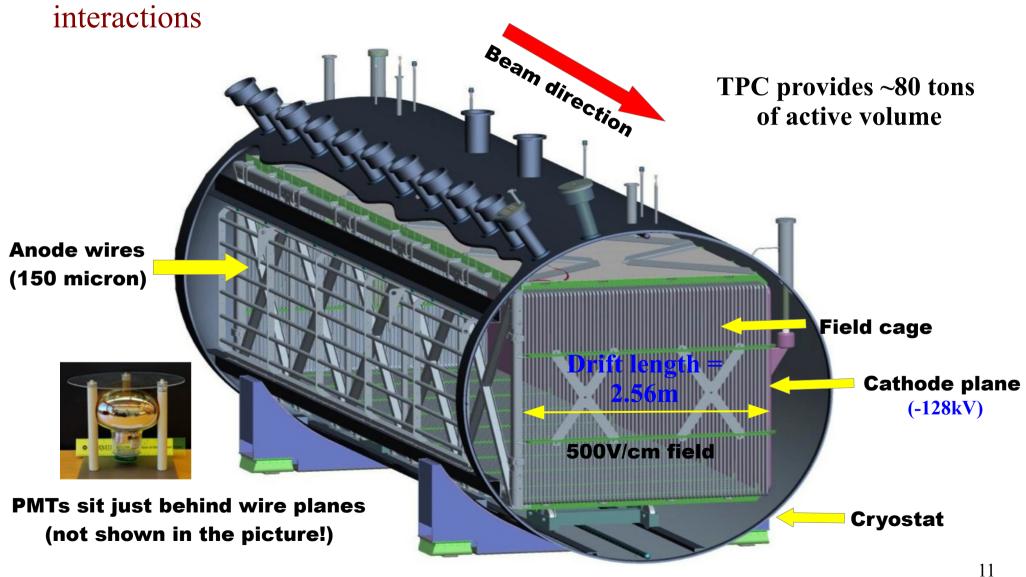
LArTPCs make 3D reconstruction possible!

- wire planes give 2D position information
- the third dimension is obtained by combining timing information with drift velocity (v_d) : $x = v_d(t-t_0) \rightarrow \text{hence}$, a "Time projection chamber"



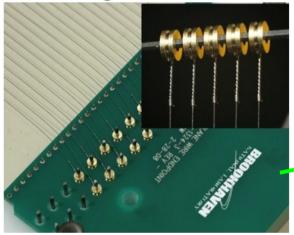
The MicroBooNE LArTPC

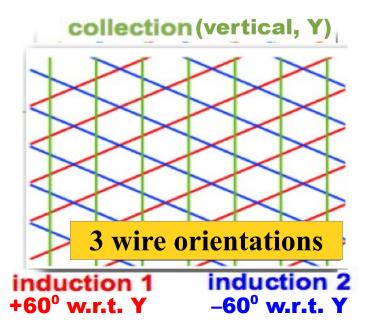
MicroBooNE Will be the largest LArTPC in U.S. to study neutrino



MicroBooNE TPC and Anode wires

Wires attach to wire carrier boards for mounting on to the TPC frame





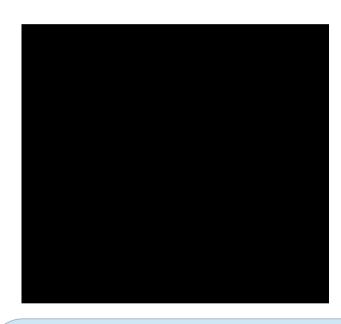
Rectangular in shape

(H = 2.33m; L = 10.37m; W = 2.56m)

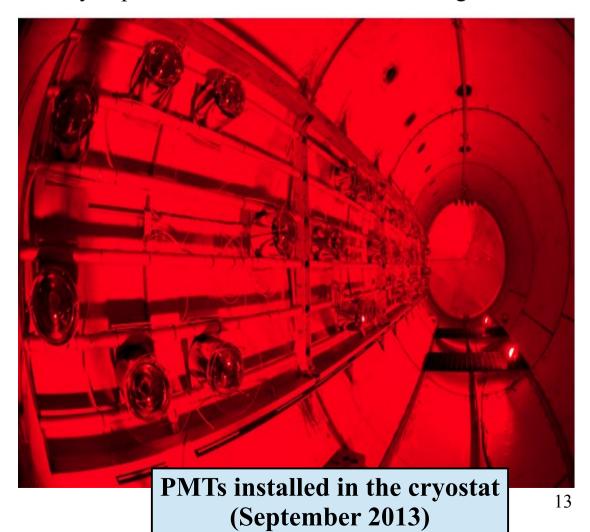


Light collection system

- MicroBooNE optical system contains 32 8-inch Photo multiplier tubes (PMT)
- De-excitation and recombination processes produce scintillation light in LAr
 - Ar emits scintillation light at $128 \text{ nm} \rightarrow \text{vacuum ultraviolet (VUV)}$ range
 - Tetraphenyl Butadiene (TPB) coated acrylic plates in front of PMTs as wavelength shifters



- Timing information very crucial for rejecting out-of-time cosmics from in-time beam interactions match "t₀" with beam gate to trigger on neutrino events
- Reduces data through-put from 200MB/s to 2MB/s



LArTPC challenges

Great technology comes with great technical challenges:

- Argon purity
 - > Argon needs to be kept ultra pure
 - Why? Impurities can capture drifting electrons affecting your signal
 → Cryogenics play a crucial role!
- High Voltage in LAr
 - > This area is not well understood (major effort from MicroBooNE)
 - Need to understand/test physics of HV in clean argon
- Cold Electronics
 - Wire signals are small, so, pre-amplifiers are immersed in the cold very close to the wire read-out → improves signal to noise ratio
- Reconstruction Software
 - No pre-existing fully automated reconstruction software for LArTPCs
 - > A major reconstruction challenge!

MicroBooNE's R&D will address all of these challenges AND serve as a "proof of principle" for future large-scale LArTPC detectors!

The Cryogenics System

Argon gas purge

- To ensure Ar purity, cryostat need to be evacuated prior to filling LAr
- Traditional vacuum evacuation of cryostats is expensive for large scale LArTPCs

MicroBooNE will purge the cryostat with Ar gas to remove air before filling

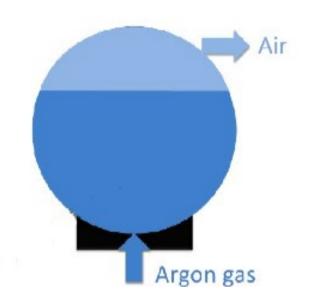
a major R&D step to guide future detectors

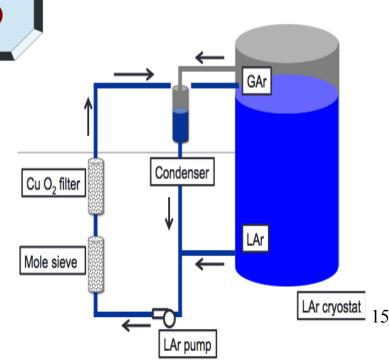
Argon gas purge has already been demonstrated to work in Liquid Argon Purity Demonstrator (LAPD) test cryostat at Fermilab

Argon Purity

- The impurities of concern are water, oxygen and nitrogen
- Nitrogen is bad for scintillation light
- Impurities need to be kept

<100 ppt (for Oxygen) <1 ppm (for Nitrogen)



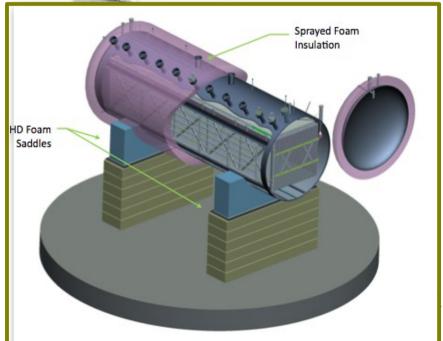


The Cryostat

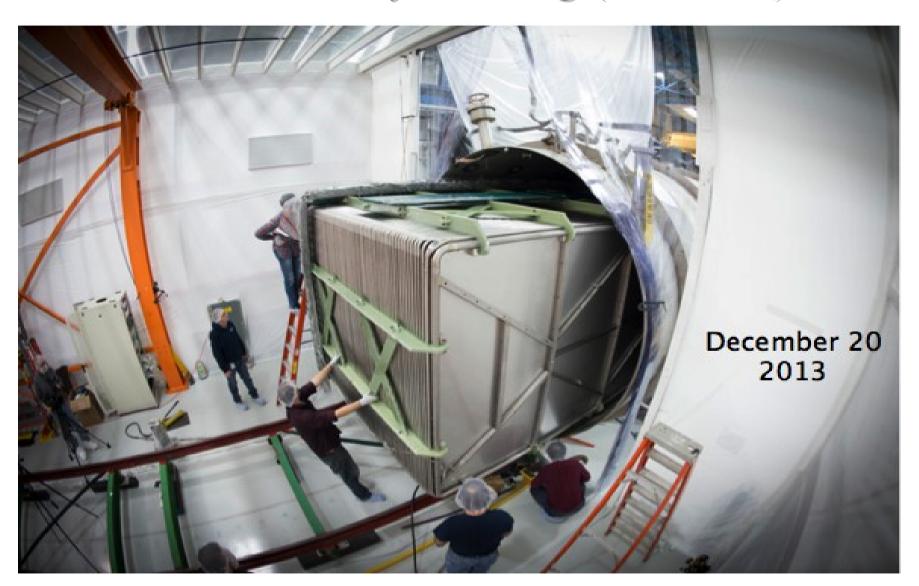
- MicroBooNE cryostat is single-walled steel cylinder with 3.5m diameter
- Can hold 170 tons of LAr
- Once inserted, cryostat will be sealed and moved to the new building and then foam insulated







Insertion of TPC into the cryostat at D0 Assembly building (Fermilab)

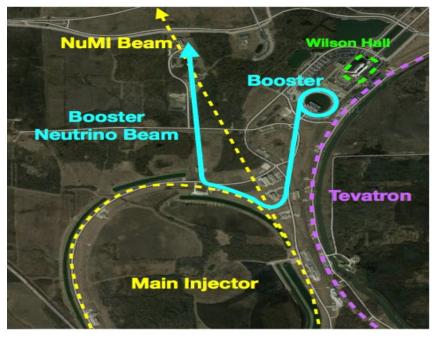


Cryostat endcap welding in progress!



- Started two days ago (May 19th)
- Going smoothly so far!

MicroBooNE location



How are neutrinos generated from BNB?

See C. Adams talk for more information

BNB flux

V

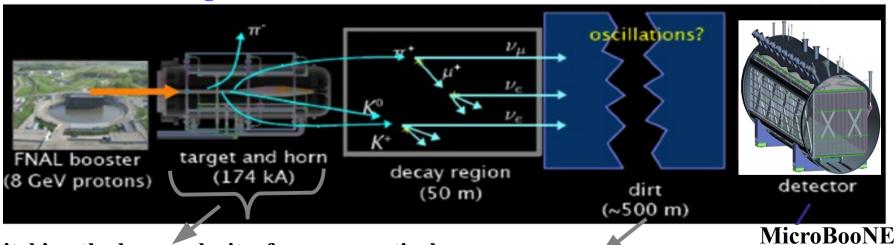
V

V

V

V

E, (GeV)



Switching the horn polarity, focuses negatively charged mesons, yielding a anti-v beam

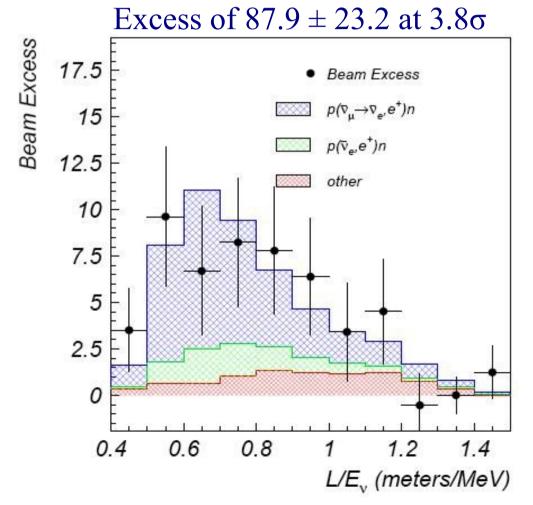
Gives similar L/E as MiniBooNE

Outline

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 - Physics motivation
 - Cross section measurements
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Physics Motivation: The LSND anomaly

• The Liquid Scintillator Neutrino Detector (LSND) experiment at Los Alamos lab saw excess of anti- v_e in an anti- v_μ beam originated by μ^+ at rest with $\langle E_y \rangle \approx 30$ MeV and L ≈ 30 m.



→ this corresponds to an oscillation probability of,

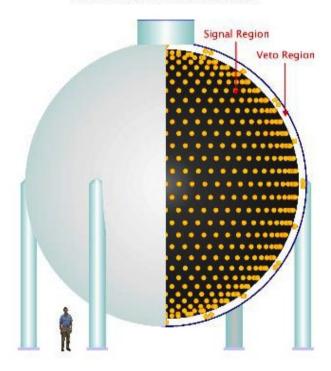
$$P(\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}) = \sin^{2}(2\theta)\sin^{2}\left(\frac{1.27L\Delta m^{2}}{E}\right)$$
$$= 0.245 \pm 0.081\%$$
(for L/E, $\approx 0.5 - 1$ m/MeV)

→ However, an appearance signal at such a small distance from the source can only correspond to non-standard oscillations (large $\Delta m^2 \sim 1 \text{ eV}^2$)!

Physics Motivation: The MiniBooNE anomaly

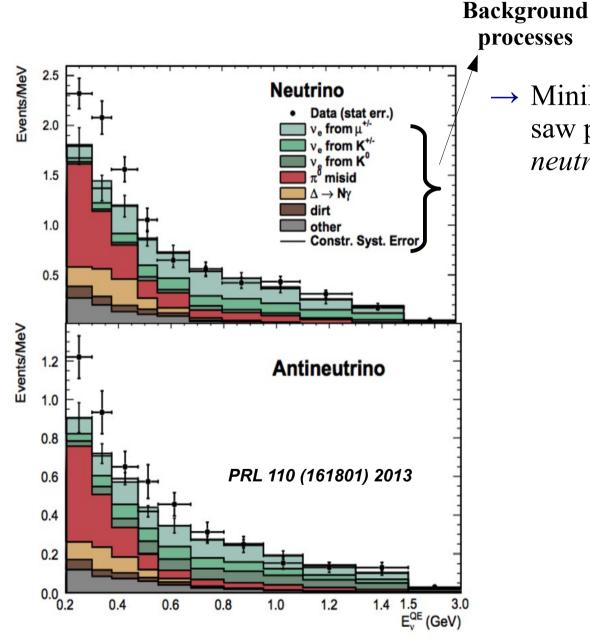
MiniBooNE, a 12m diameter cherenkov detector filled with 800 tons of mineral oil was initially conceived as an ultimate test of the LSND anomaly. → MiniBooNE sits on the Booster neutrino beam line.

MiniBooNE Detector



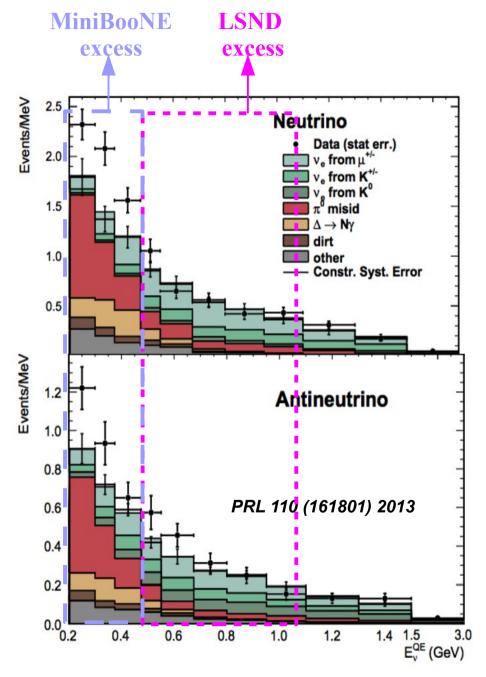
→ Similar L/E $_{\nu}$ (~ 1m/MeV) as LSND, Distance from source, L ~ 500m Typical E $_{\nu}$ ~ 500 MeV (entirely different systematics and backgrounds than those of LSND)

Physics Motivation: The MiniBooNE anomaly



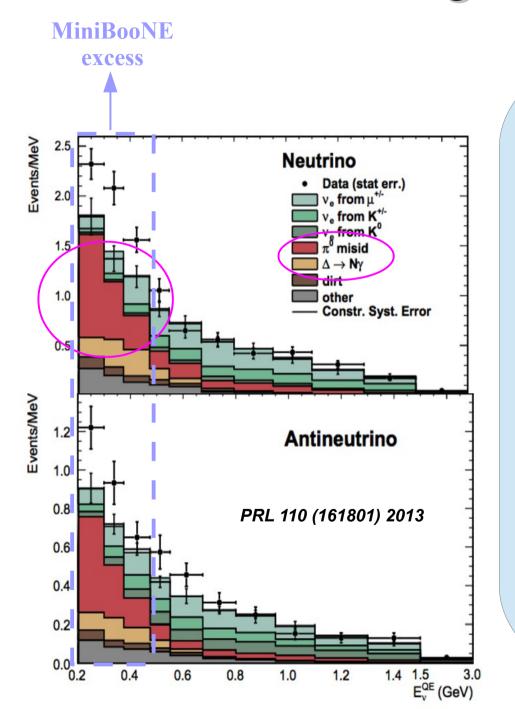
→ MiniBooNE's oscillation analysis saw pronounced excesses in both neutrino and anti-neutrino modes

Motivation: The MiniBooNE anomaly



- → MiniBooNE's oscillation analysis saw pronounced excesses in both neutrino and anti-neutrino modes But, in a different region than the LSND excess!
 - \rightarrow for energies b/n 0.2 0.475 GeV excess of 240.0 \pm 62.9 at 3.8 σ

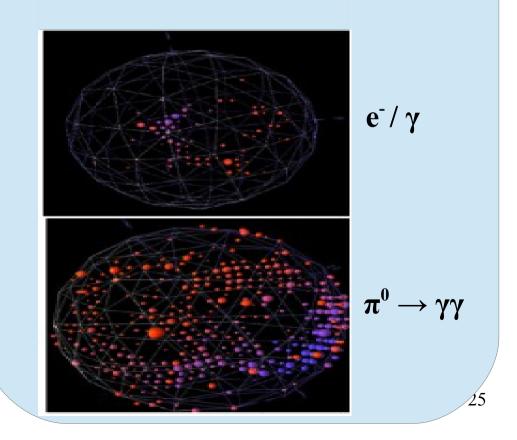
Understanding MiniBooNE excess



The real problem

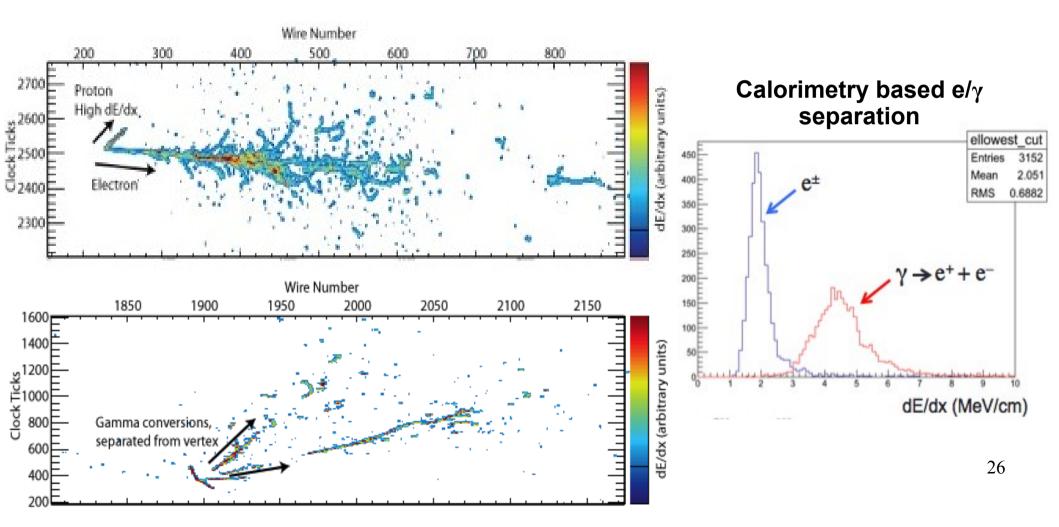
Excess could be a mis-identified background!?

A cherenkov detector problem, doesn't offer good e (signal) to γ (background) discrimination!

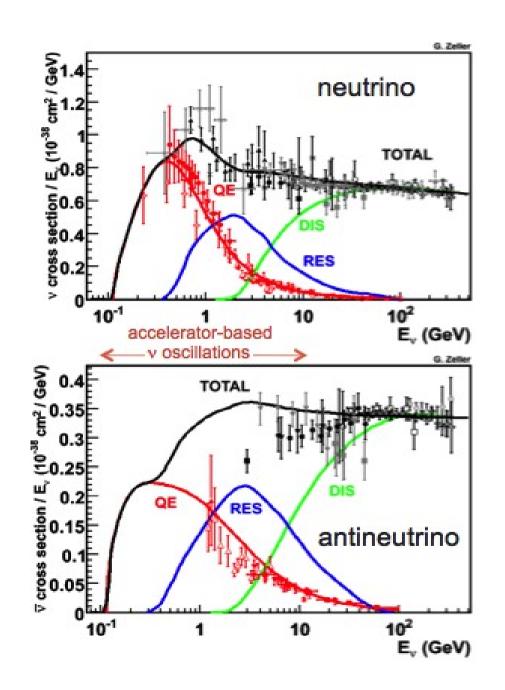


Can MicroBooNE address the MiniBooNE excess?

- LArTPCs excellent dE/dx resolution capability,
 - allows one to measure dE/dx in the first 3 cm of an EM interaction low dE/dx ~ 1 MIP => e⁻; high dE/dx ~ 2 MIPS => γ (e⁺e⁻ pair)



Neutrino cross-sections



Why measure low-energy neutrino cross sections in Argon?

Historically, low energy (up to ~ 1 GeV) σ_v are not well understood/explored Recent realization of the importance of low-energy neutrino cross-section measurements in the context of future oscillation experiments

Improved nuclear models

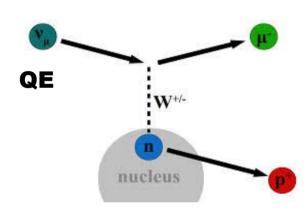
Understanding final state interactions and other nuclear effects in argon will help towards this

MicroBooNE, with its superior PID and calorimetry, will be able to make high-precision σ_{v} measurements in argon

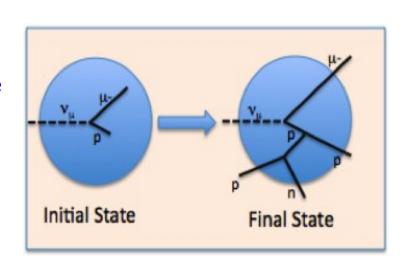
Nuclear effects in v-Ar interactions

• The Argon nucleus (18 protons and 22 neutrons) presents a complicated environment for a neutrino to interact within!

• This is particularly challenging for exclusive channels like Quasi elastic (QE), resonant pion production (RES) etc. where nuclear effects can have dramatic effects on final states



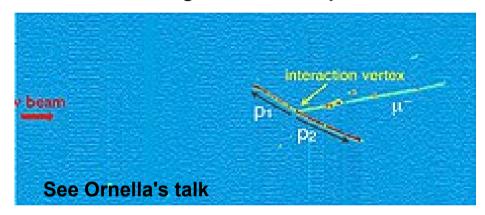
• Due to final state interactions and possible effects of correlation between target nucleons, a genuine QE interaction can often contain additional nucleons or de-excitation gammas or pions in its final state

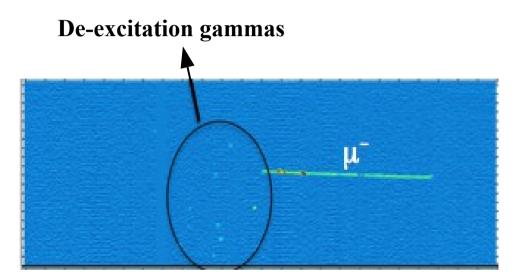


RES

Hints of Nuclear effects from ArgoNeuT experiment

One case of Multi-nucleon correlations: Neutrino interacting with a nucleon-pair emitting back-to-back protons



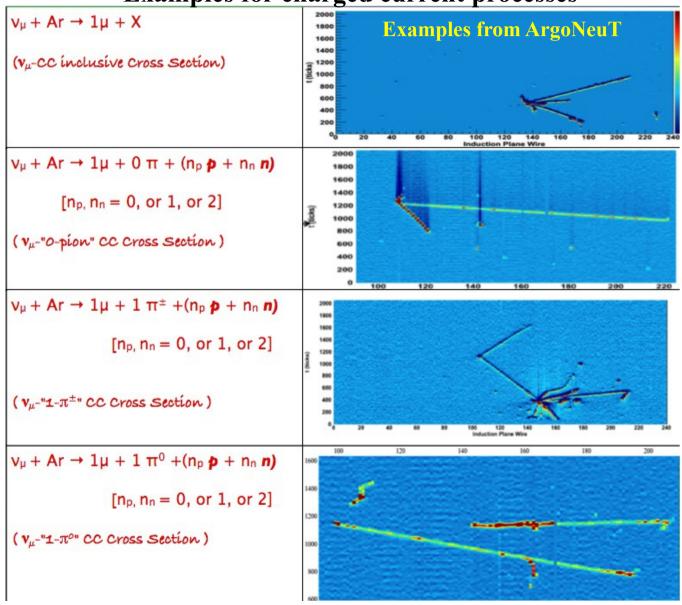


Protons and other particles emitted as part of FSI are usually low in energy (of the order of few MeV) which makes them hard to detect, this is where strength of LArTPC technique come into play:

- LArTPCs superior vertex resolution and particle identification will make it possible to identify protons and disentangle nuclear effects with extraordinary sensitivity
- ArgoNeuT already shows that they can reconstruct protons down to 21 MeV!
 - → MicroBooNE, a bigger/better version of ArgoNeuT will only do better!

Exploring Final state particle multiplicities

Examples for charged current processes



Understanding Final state (FS) particle multiplicity is crucial for measuring exclusive channels accurately

Classifying events based on FS particle multiplicity will make it most sensitive to nuclear effects

Process		No. Events
	υ _μ Events (By Final State Topology)	
CC Inclusive		88,098
$CC 0 \pi$	$ u_{\mu}N ightarrow\mu+Np$	56,580
	$\cdot \; u_{\mu} N ightarrow \mu + 0 { m p}$	12,680
	$\cdot \; u_{\mu} N ightarrow \mu + 1 \mathrm{p}$	31,670
	$\cdot u_{\mu} N ightarrow \mu + 2 { m p}$	5,803
	$\cdot \; u_{\mu}N ightarrow \mu + \geq 3p$	6,427
$CC 1 \pi^{\pm}$	$ u_{\mu}N ightarrow\mu+ ext{nucleons}+1\pi^{\pm}$	21,887
$CC \ge 2\pi^{\pm}$	$ u_{\mu}N o \mu + \text{nucleons} + \geq 2\pi^{\pm}$	1,953
$CC \ge 1\pi^0$	$ u_{\mu}N ightarrow { m nucleons} + \geq 1\pi^0$	9,678
NC Includive		33,000
NC 0 π	$ u_{\mu}N ightarrow ext{nucleons}$	21,509
NC 1 π [±]	$ u_{\mu}N \to \text{nucleons} + 1\pi^{\pm}$	4,886
NC ≥2π [±]	$ u_{\mu}N o ext{nucleons} + \geq 2\pi^{\pm}$	635
NC ≥1π ⁰	$ u_{\mu}N ightharpoonup \mathrm{nucleons} + \geq 1\pi^{0}$	6,657
	$ u_e$ Events	
CC Inclusive		567
NC Inclusive		207
Total ν_{μ} and ν_{e} l	Svente	121,099
	ν _μ Events (By Physical Process)	
OC QE	$ u_{\mu} n ightarrow \mu^- p$	48,626
OC RES	$ u_{\mu}N ightarrow \mu^{-}N$	26,852
OC DIS	$ u_{\mu}N ightarrow \mu^{-}X$	10,527
OC Coherent	$ u_{\mu}Ar ightarrow \mu Ar + \pi$	376

Estimated event rates using **GENIE**

for BNB flux of 6.6E20 POT Table from Corey Adams (Yale)

- In enumerating proton multiplicity, we assume an energy threshold on protons of 21 MeV
- The 0π topologies include any number of neutrons in the event

Summary

- TPC is fully assembled and inserted into the cryostat on December 20, 2013
- All cold and warm electronics tested successfully with the DAQ read out
- Cryostat is sealed this week and will be moved to LArTF in June
- LArTF is ready for the arrival of Cryostat, Cryogenics installation at LArTF is progressing rapidly
- MicroBooNE will start taking data soon
 - Lot to do with cosmics calibration, commissioning of DAQ etc.
 - Expect some early results on neutrino cross-sections
 - Oscillation results require at least 3 years of data to give "statistically" sensible results

Wait for MicroBooNE to uncover some rich and diverse physics all while trying to solve the MiniBooNE mystery!

Thanks very much for your attention!



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Supplementary slides

The MicroBooNE collaboration

Brookhaven Lab

Mary Bishai
Hucheng Chen
Kai Chen
Susan Duffin
Jason Farell
Francesco Lanni
Yichen Li
David Lissauer
George Mahler
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Jennet Dickinson
Georgia Karagiorgi
David Kaleko
Bill Seligman
Mike Shaevitz
Bill Sippach
Kathleen Tatum
Kazuhiro Terao
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Roberto Acciarri Bruce Baller Dixon Bogert Ben Carls Michael Cooke Herb Greenlee Cat James Eric James Hans Jostlein Mike Kirby Sarah Lockwitz Byron Lundberg Alberto Marchionni Stephen Pordes Jennifer Raaf Gina Rameika Brian Rebel Rich Schmitt Steve Wolbers

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Flavio Cavanna Ornella Palamara

Virginia Tech

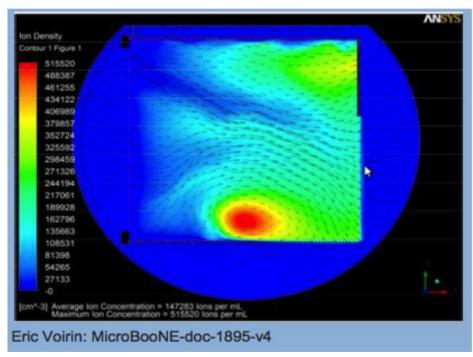
Mindy Jen Leonidas Kalousis Camillo Mariani

Yale University

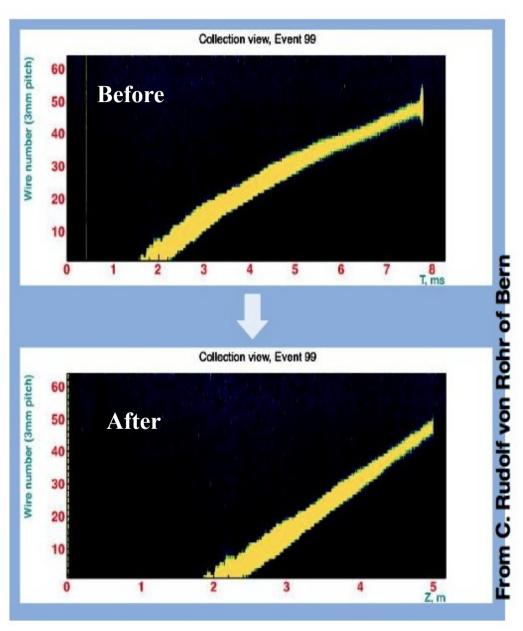
Corey Adams
Christina Brasco
Eric Church
Bonnie T. Fleming (*)
Ellen Klein
Ornella Palamara
Flavio Cavanna
Roxanne Guenette
Kinga Partyka
Andrzej Szelc

Laser Calibration

- The cosmic traffic produces significant excess of Ar+ ions which drift very slowly to the cathode (~cm/s) thus causing ion space charge accumulation
- This results in E-field distortions, that together with Ar circulation can potentially affect the tracking efficiency



• A laser delivers a straight path correcting for field distortions

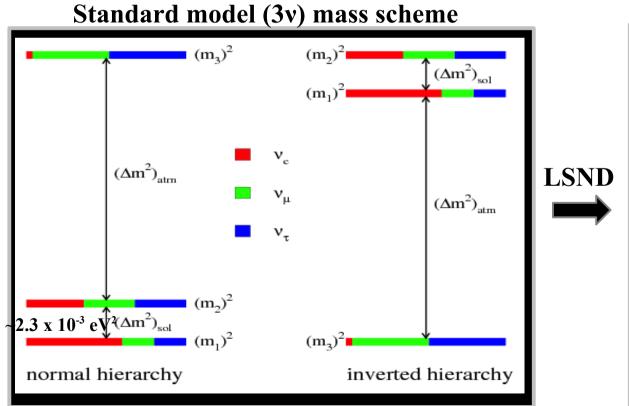


3+1 sterile neutrino model (as a possible solution to LSND)

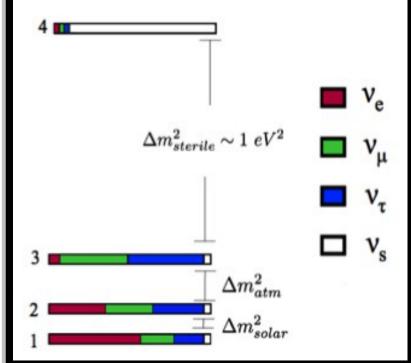
• Neutrino mass differences within Standard model:

$$(\Delta m^2)_{atm} \sim 2.3 \times 10^{-3} eV^2$$
 and $(\Delta m^2)_{sol} \sim 7.6 \times 10^{-5} eV^2$

- Now, consider a fourth neutrino that doesn't interact via weak interaction but can oscillate to any of the 3 active neutrino states
 - this is possible if the $(\Delta m^2)_{Sterile} >> (\Delta m^2)_{Atm}$ and $(\Delta m^2)_{Sterile} >> (\Delta m^2)_{Sol}$



One possible BSM (4v) mass scheme

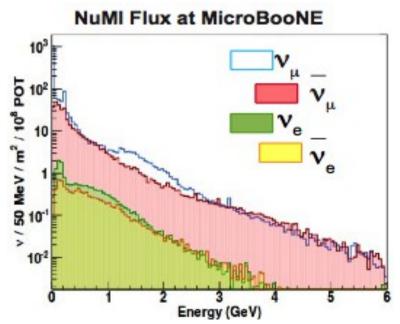


BNB and NuMI fluxes

Predicted event rates for MicroBooNE from BNB and NuMI over the next 3 years

	BNB	NuMI
Total events	145k	60k
ν _μ CCQE	68k	25k
$NC \pi^0$	8k	3k
v _e CCQE	0.4k	1.2k
POT	$6x10^{20}$	$8x10^{20}$

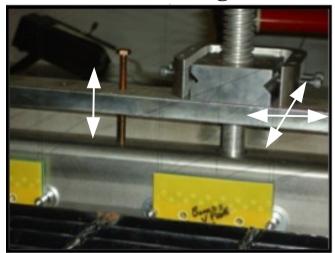
- BNB is up and running for the past 10 years
- Well understood beam systematics will result in expeditious physics measurements

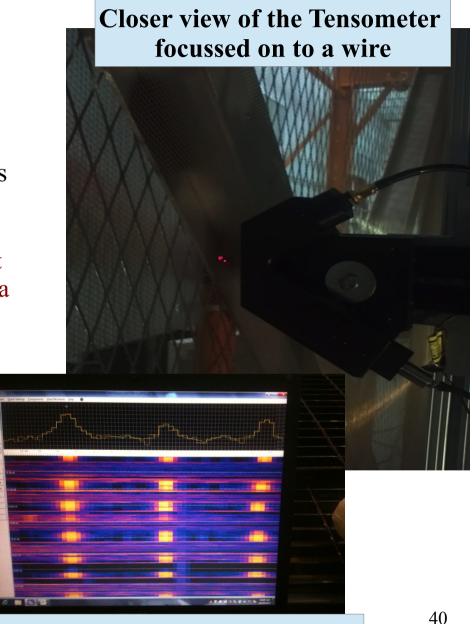


Wire tensioning

- Wires need to be tensioned to an optimum value, to prevent sagging and/or breakage!
- Tension goal was $\sim 0.7 \text{ kg}$
- A laser tensometer (from UW Madison) was used to measure and set wire tensions Laser illuminates a wire, photo diode catches the reflection. Frequency is readout when the wire vibrates and is then fed into a spectrum analyzer

TPC frame has adjustable support bars that can be used to change wire tensions

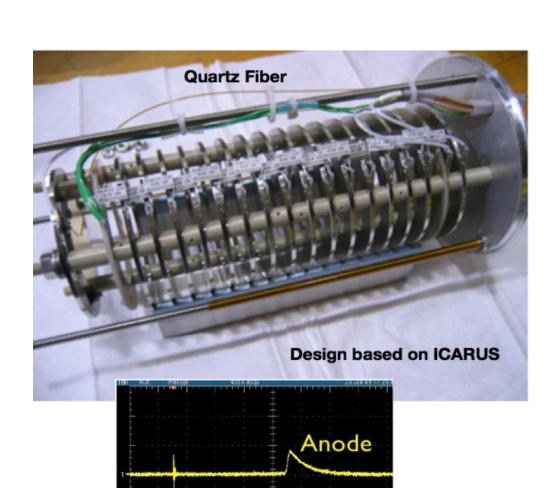




Spectrum analyzer showing uniform results on each of the wires

LAr Purity monitoring

- MicroBooNE's inline purity monitors measure argon purity Two monitors are placed in the cryostat and one is installed downstream of the filters
- Purity monitors contain a photocathode and an anode.
 A xenon flash lamp liberates electrons off of the photocathode
- Purity is measured by finding the electron lifetimes to the anode



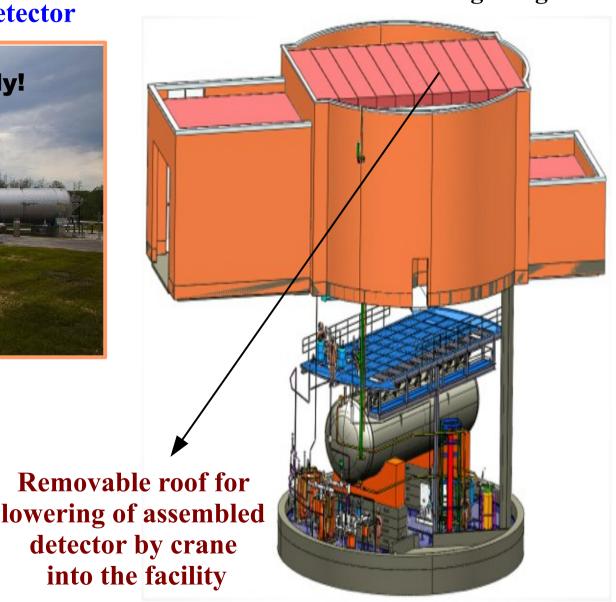
Cathode

Liquid Argon Testing Facility (LArTF)

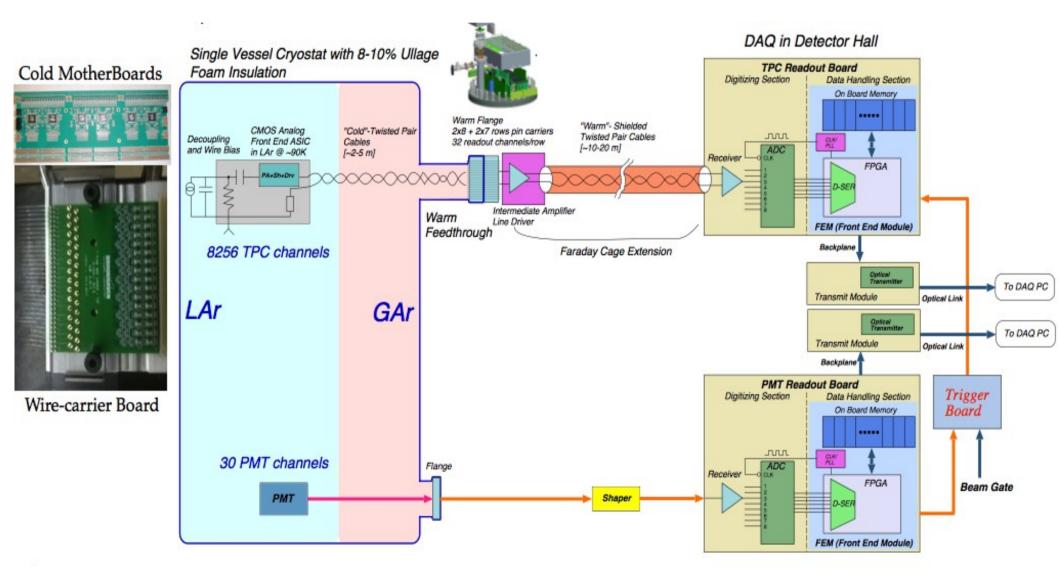
The specially designed LArTF building will house the MicroBooNE detector



LArTF actual building design

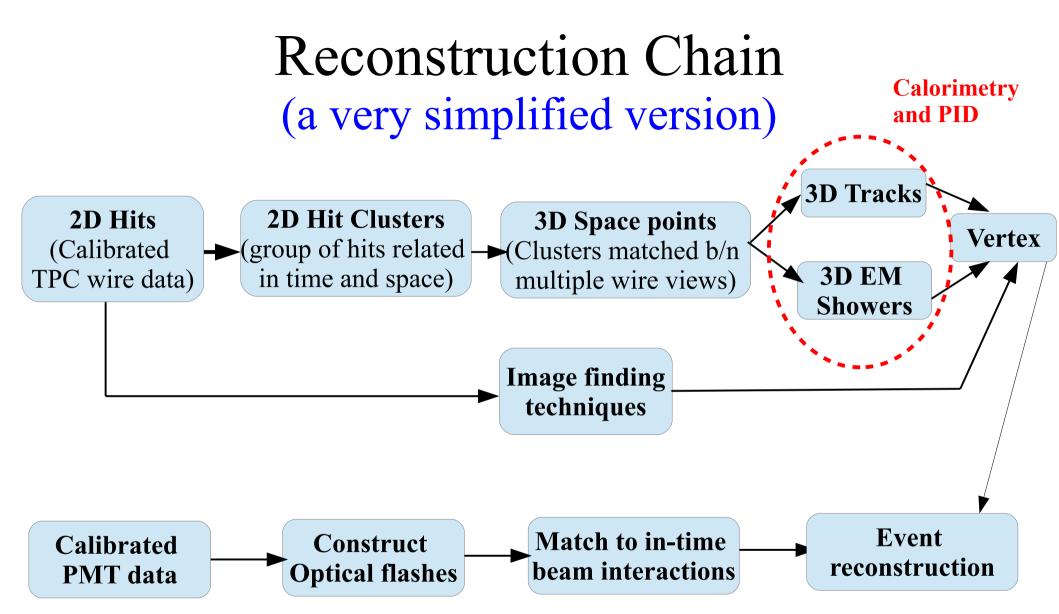


Detailed Read-out electronics chain



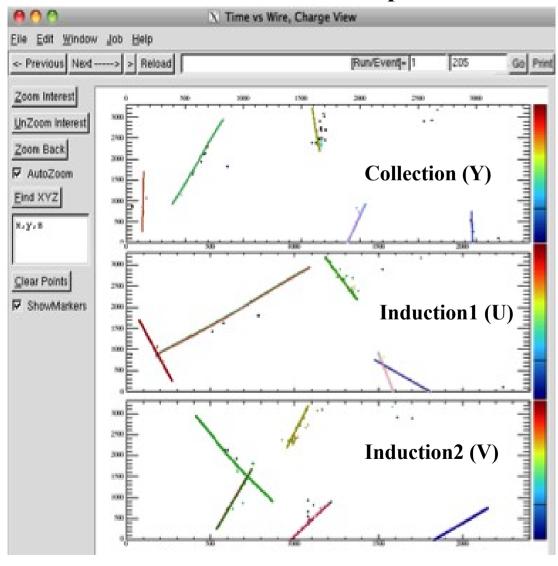
Refs:

1.) Readout Electronics Design Considerations for LAr TPC, H. Chen, ANT2013 Conference



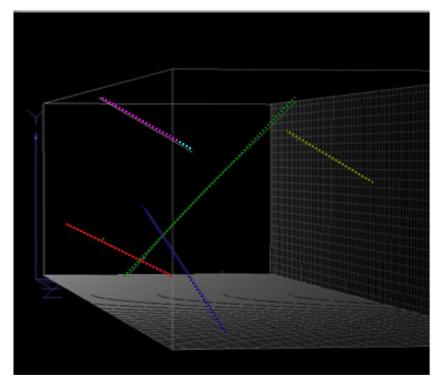
Tracking

2D views of the tracks in 3 wire plane views



Tracking algorithms maturing fast!

3D event display of cosmic tracks reconstructed with Bezier tracker



Algorithm author: Ben Jones (Bezier tracking)

Proton decay background

Some GUT models explicitly break the baryon number symmetry, thus, predicting proton decay!

MicroBooNE is not big enough to study proton decay itself

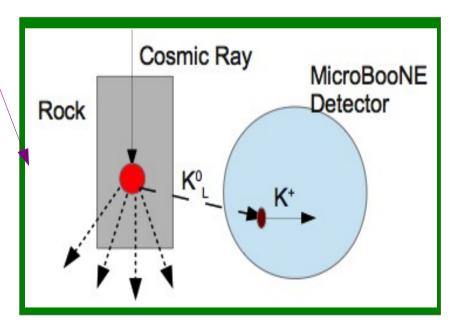
- But, MicroBooNE can study proton decay backgrounds for future experiments!

Proton decay background

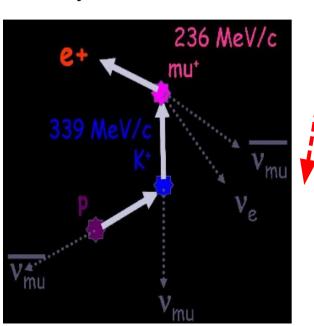
A cosmic muon interacts in a rock near the detector, produces a K_L^0 which then charge exchanges, $K_I^0 p \to K^+ n = looks like a K^+ from proton decay if right energy (339 MeV/c).$

Decay mode of interest to MicroBooNE: $p \rightarrow K^+ v$; $K^+ \rightarrow \mu^+ v_\mu$; $\mu^+ \rightarrow e^+ v_e$ (anti- v_μ)

- the distinct dE/dx pattern enables study of this 3-fold decay mode



From J. Esquivel



Supernovae neutrinos

A core-collapse supernova (SN) produces, in addition to enormous light, a burst of neutrinos

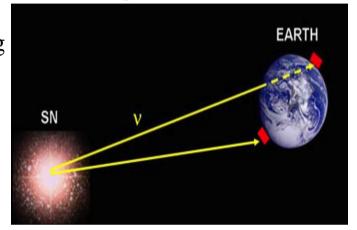
of all flavors (in few-tens-of-MeV range)

- → physics of oscillations of SN neutrinos very interesting
- → critical information on key astronomical phenomena

Water and liquid scintillator neutrino detectors,

 \rightarrow primarily sensitive to electron anti-neutrinos

 $anti-v_e + p \rightarrow n + e^+$ (inverse beta decay on free protons)

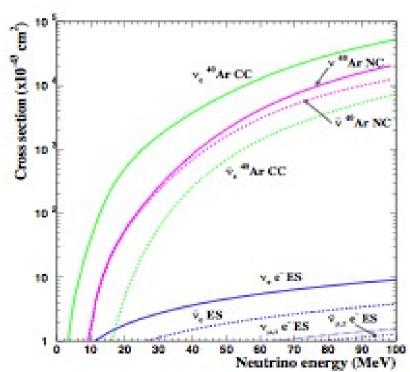


LArTPCs posses unique capability to detect SN electron neutrinos

1. CCv_e capture of SN neutrinos on Ar $v_e + Ar^{40}(18) \rightarrow K^{40}(19) + e^{-}$

Other processes:

- 2. Neutral current excitation of Ar^{40} $v_{e,\mu,\tau} + Ar^{40}(18) \rightarrow Ar^{*40}(18) + v_{e,\mu,\tau}$
- 3. Elastic scattering off electron $v_{e,\mu,\tau} + e^{-} \rightarrow v_{e,\mu,\tau} + e^{-}$



Supernovae neutrinos

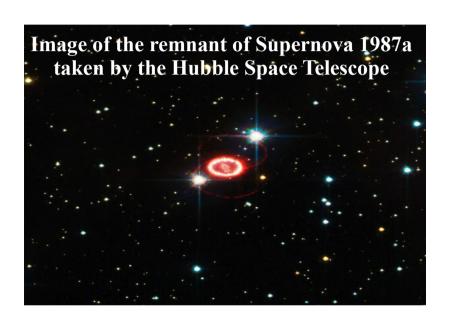
Detection requires sensitivity to low-energy gammas (<50 MeV) and electrons

• CCv_e capture on Ar can be tagged via the coincidence of emitted electron and accompanying de-excitation gamma cascade

Due to small size of MicroBooNE,

- will only see about 10-20 SN neutrinos in a duration of about 20 seconds
- A multi-kiloton detector (like LBNE) will be able to see a few hundred SN events!





Triggering on Supernovae events,

- MicroBooNE sits just below surface, too much cosmic traffic to have its own trigger!
- MicroBooNE will subscribe to SNEWS!