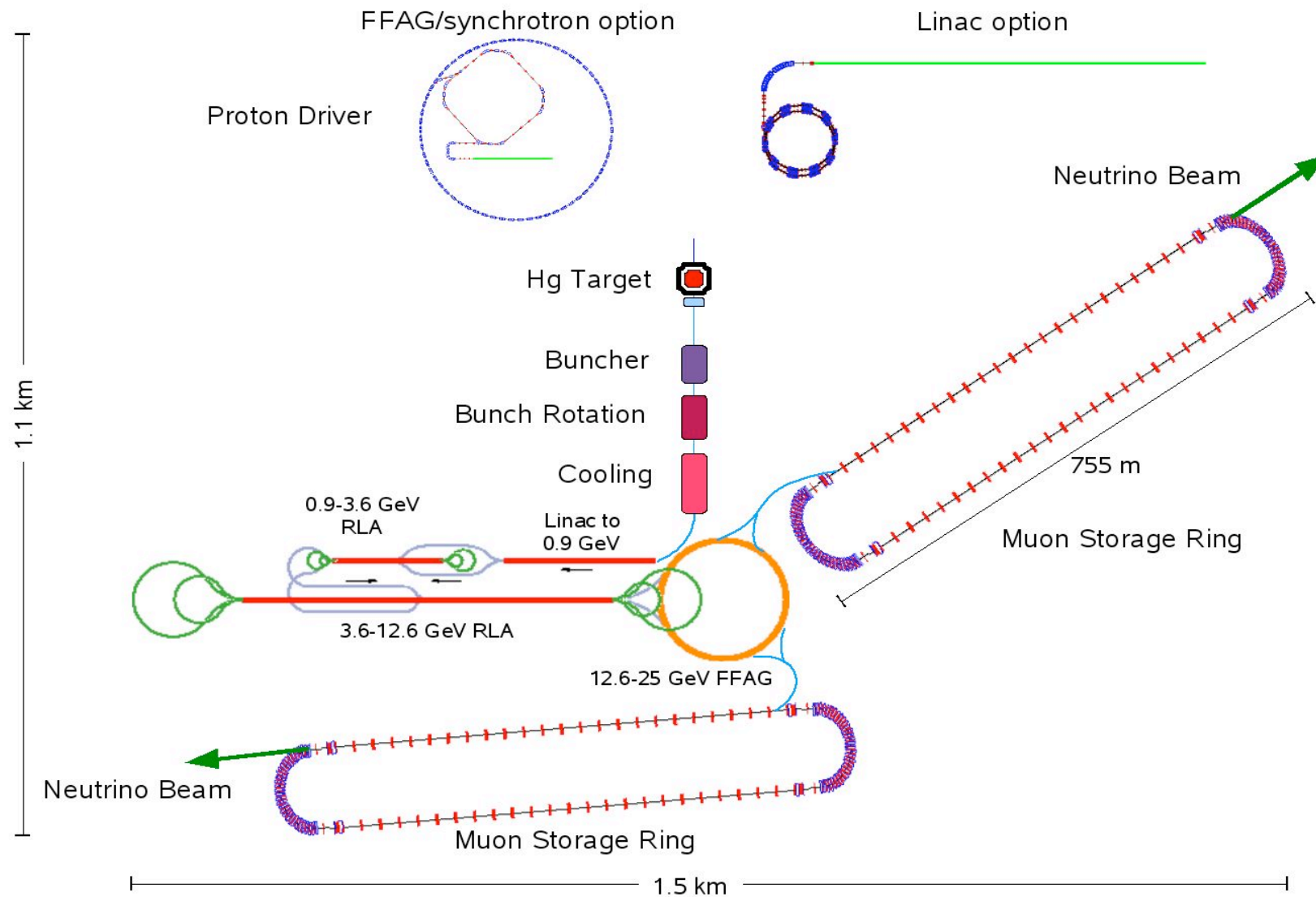
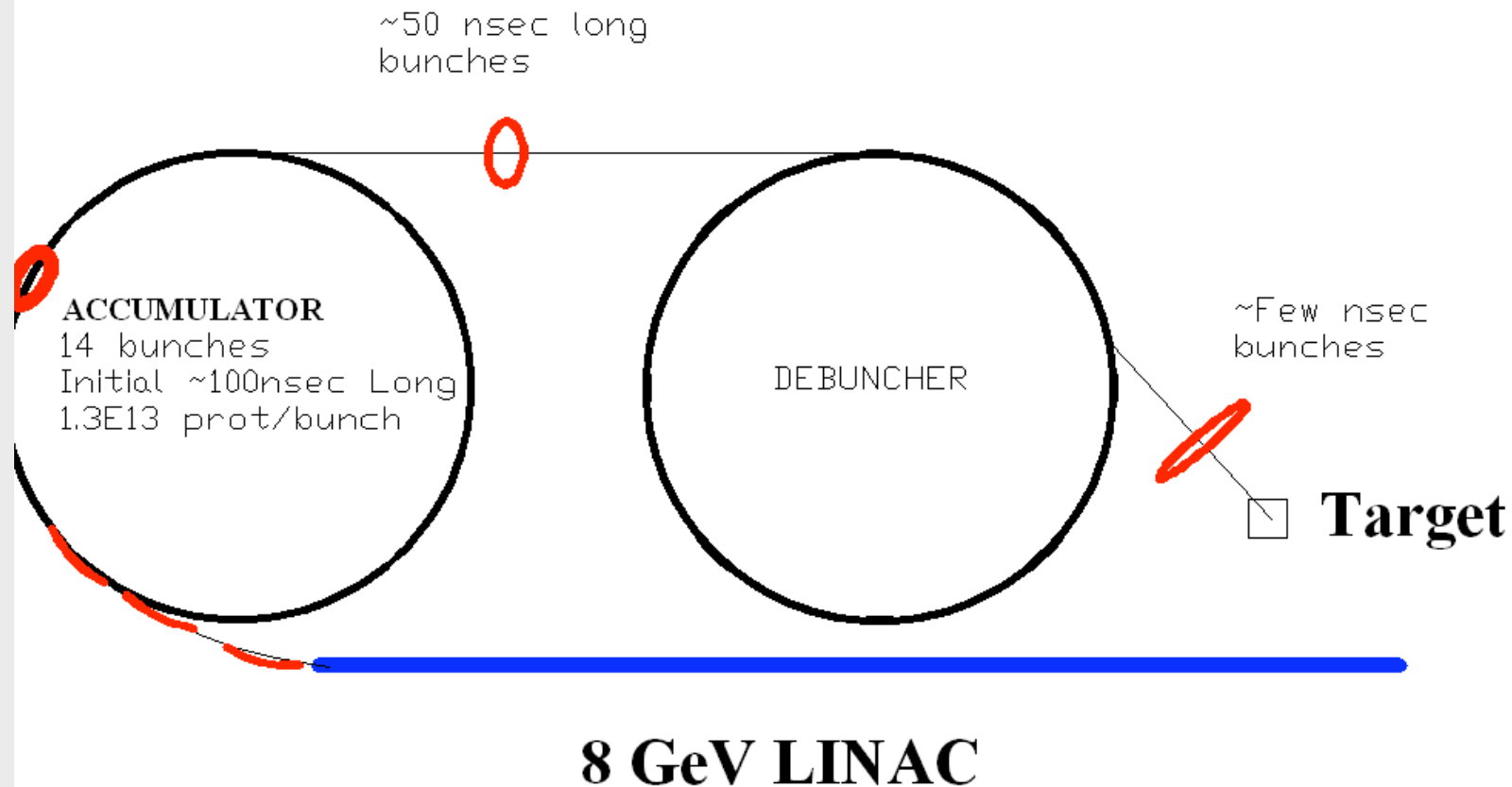
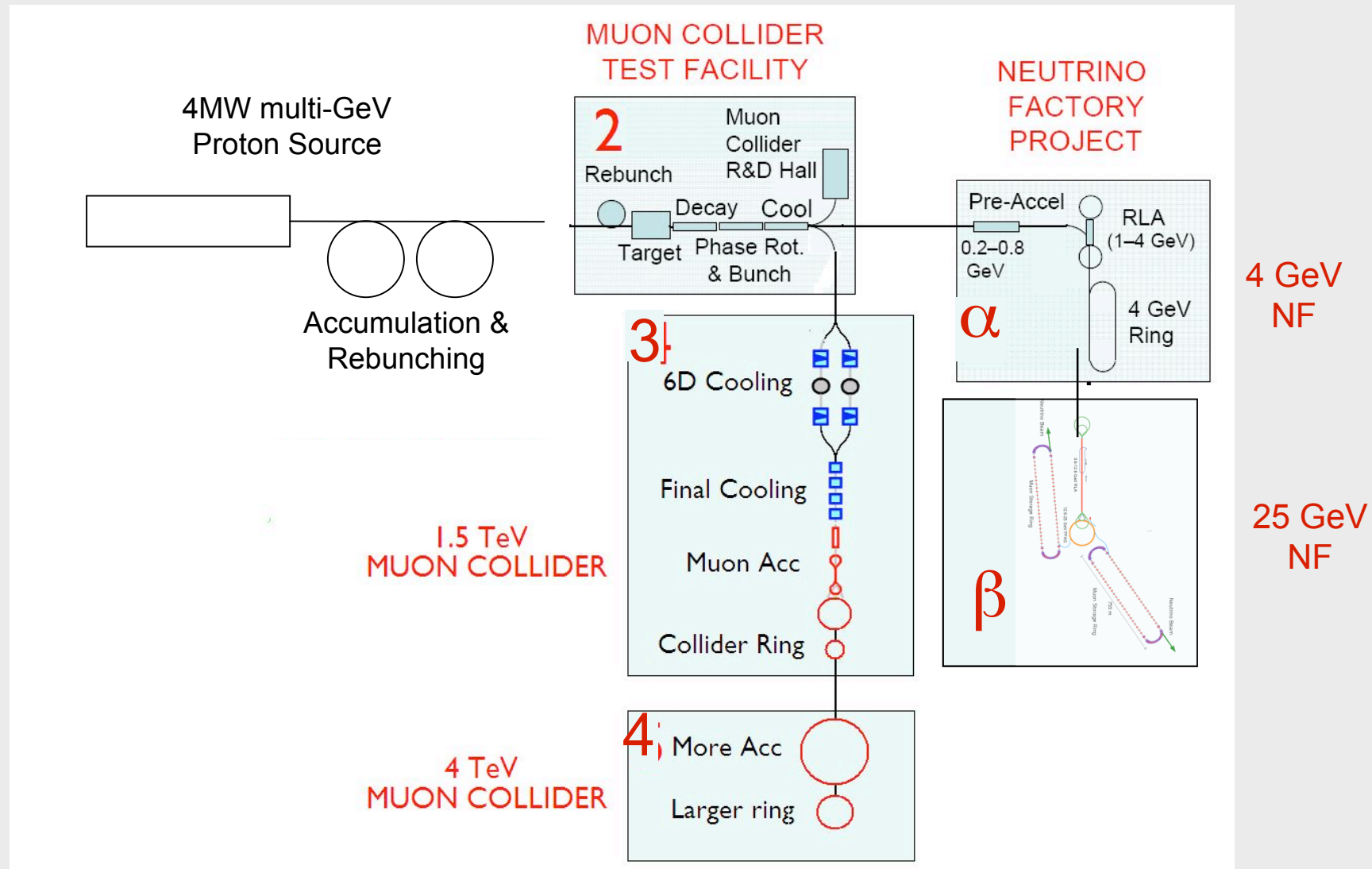


Accelerator WG summary and Preparations for the IDR



Chuck Ankenbrandt





Chuck Ankenbrandt

The most promising design approach for a proton driver consists of a full-energy H^- linac, a pair of storage rings, and an external bunch combiner.

A kinetic energy of 6 GeV is high enough to support the 4-MW proton driver capabilities required for NF and MC applications.

The linac should be hybrid (if IC-2.2 is adopted):

- the RF power should be CW;**

- the beam current should be pulsed at a high frequency.**

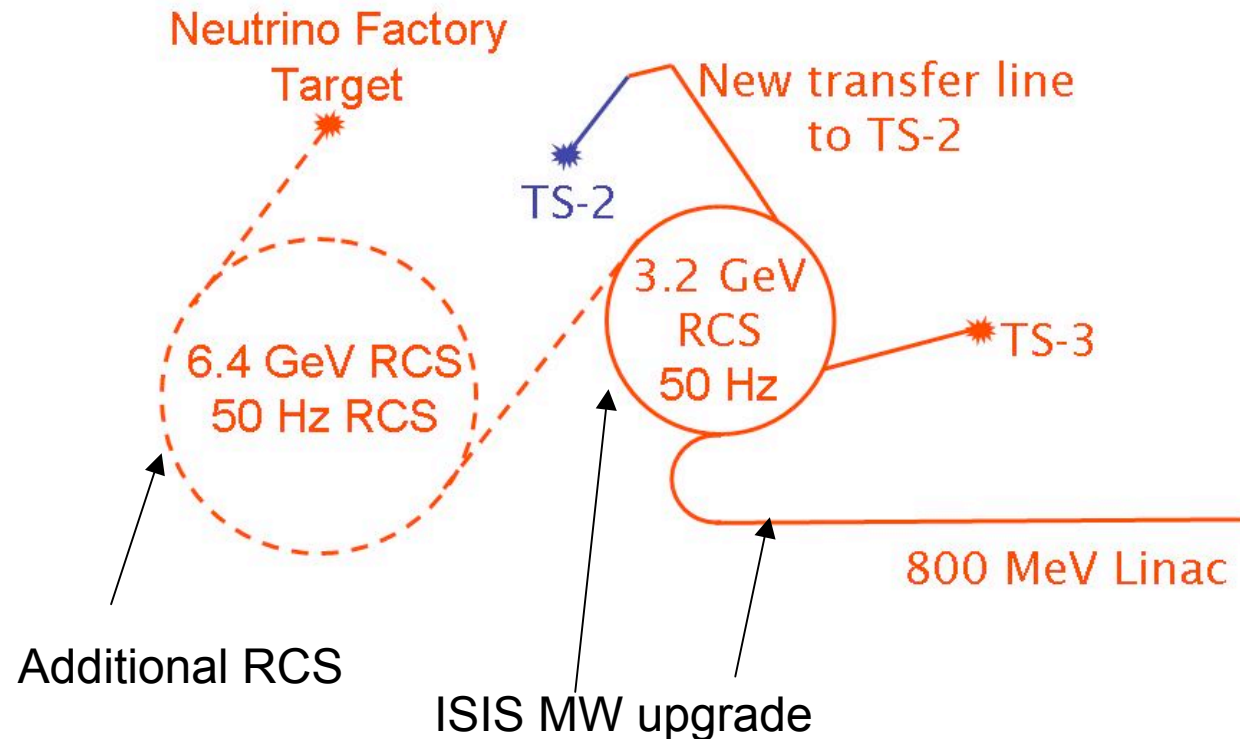
The proton driver complex can be staged:

- the first stage supporting a modest neutrino factory,**

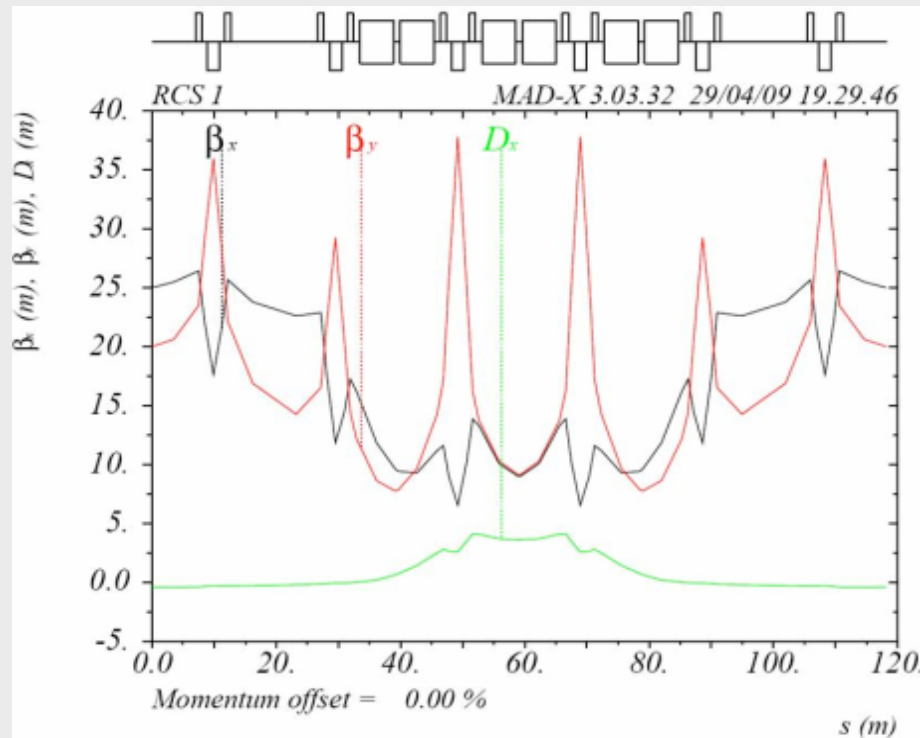
- the second stage supporting the ultimate NF,**

- the third stage supporting a high-luminosity muon collider.**

Jaroslav Pasternak



- Based on MW ISIS upgrade with 0.8 GeV linac and 3.2 GeV RCS.
- Assumes a sharing of the beam power at 3.2 GeV between the two facilities
- Requires additional RCS machine in order to meet the power and energy needs of the Neutrino Factory
- **Both facilities can have the same ion source, RFQ, chopper, linac, H⁻ injection, accumulation and acceleration to 3.2 GeV**

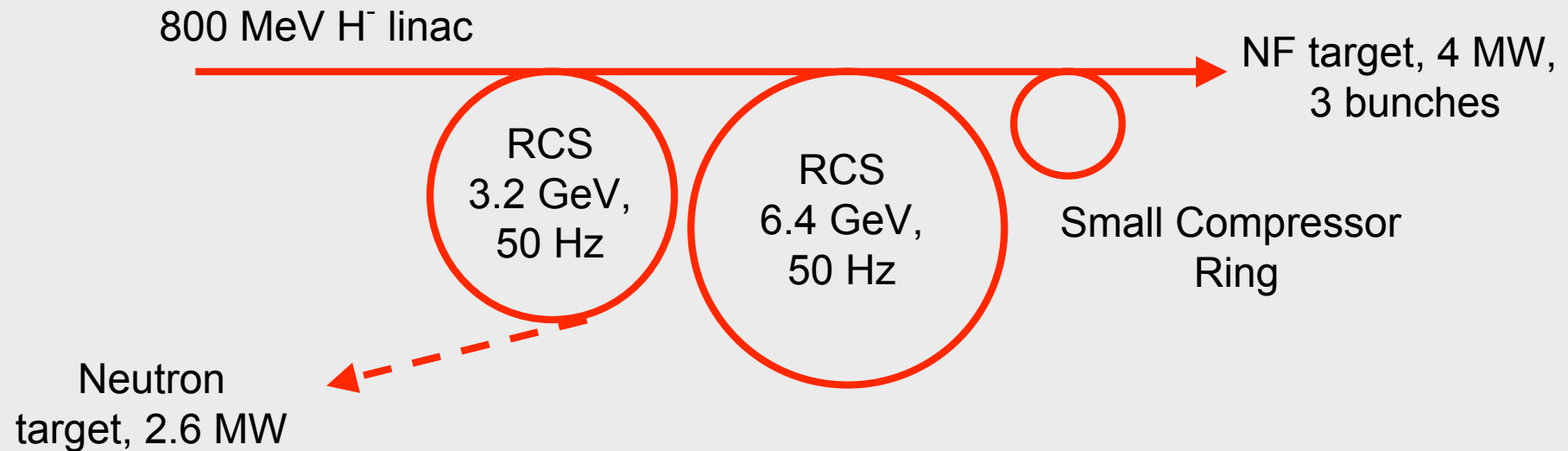


- Lattice may allow for flexibility in gamma transition choice (even with beam).
- Ring is overdesigned in order to allow for 10.3 GeV.
- Optimised solution for 6.4 GeV is in preparation!

Number of superperiods	6
Circumference	708.788 m
Harmonic number	6
RF frequency	2.4717-2.5289 MHz
Betatron tunes (Q_H , Q_V)	(7.81, 7.78)
Gamma transition	7.9056
Beam power at 6.4 GeV	4 MW for 2 bunches
Bunch area	1.8 eVs
$\Delta p/p$ at 3.2 GeV	$5.3 \cdot 10^{-3}$
Injection / extraction energy	3.2 / 6.4 [10.3] GeV
Repetition rate	50 Hz
Max B field in dipoles	1.2 T (at 10.3 GeV)
Length of long drift	12 m

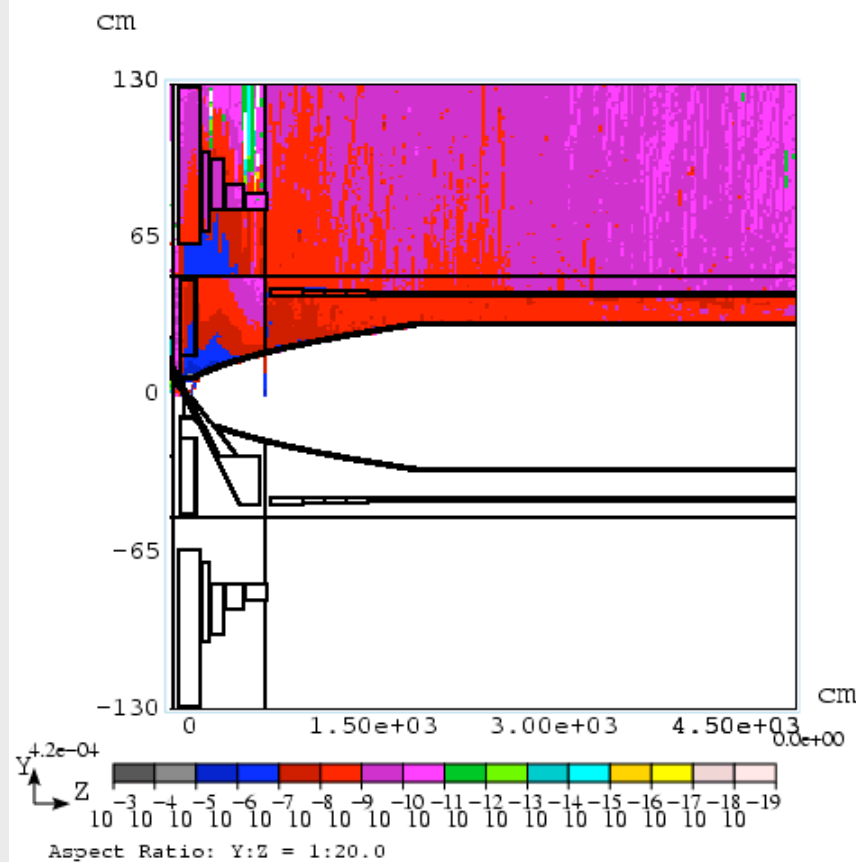
Parameters of 6.4 (10.3) GeV RCS

Jaroslav Pasternak

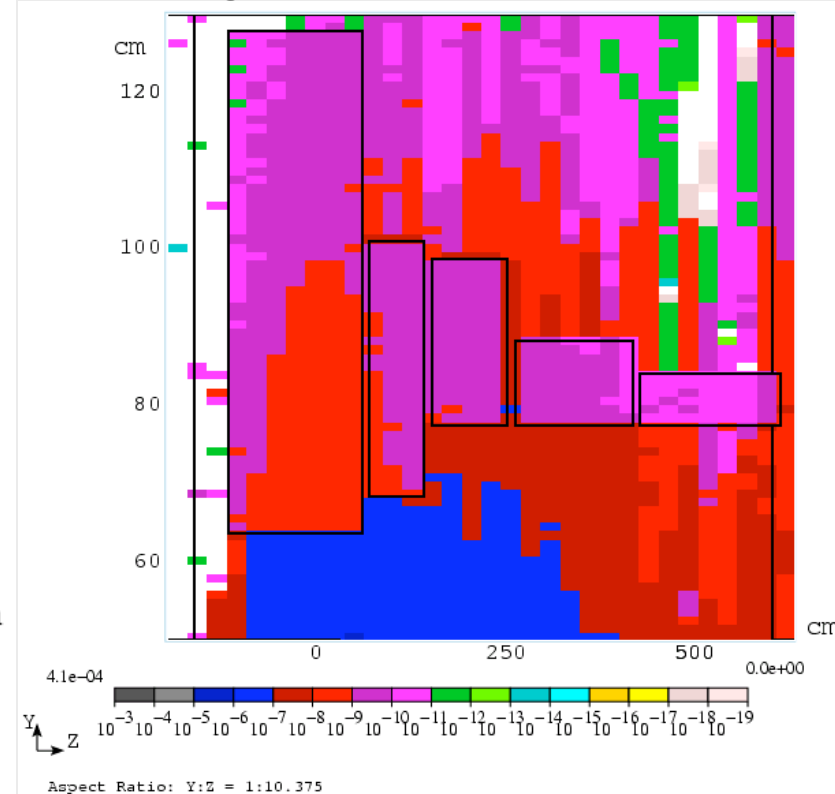


- Fast phase rotation in the dedicated compressor ring (most economic from the RF point of view, but another ring is needed).
- Bunches will be extracted one by one from the RCS.
- Compressor ring works above transition, but the rotation is very fast.
- The bunches in the RCS will wait uncompressed for 200 us, but they will come with different energies.
- We do not have a design for the compressor ring at the moment, but CERN design can be adopted.

Liquid mercury target -Energy deposition in the target area



X. Ding



Enhanced shield can decrease the power deposition in SC1 coil from 22.1kW to 4.8kW. By replacing the Res Sol by WC shield, the power deposition in SC1 coil can be decreased further to 1.3kW.



-Liquid mercury target -Hg Circulation System

V. Graves

Windows

Mercury dump and drain

Mercury vapour and condensation

Waste disposal



Previous Work -Jet Flow

- Free Jet + MHD + EDP

- Numerical Study

- R. Samulyak et al.

FronTier Code based on Front Tracking

- Surface instability
- MHD stabilizing effect
- Filament velocity in the simulations was about **25% smaller** than the experimental value

magnetic field	5T	10T	15T
experiments	54 m/s	50 m/s	35 m/s
simulations	36 m/s	27 m/s	22 m/s



0T



5T



10T



15T

Surface filaments at 150 μ s
under longitudinal magnetic field

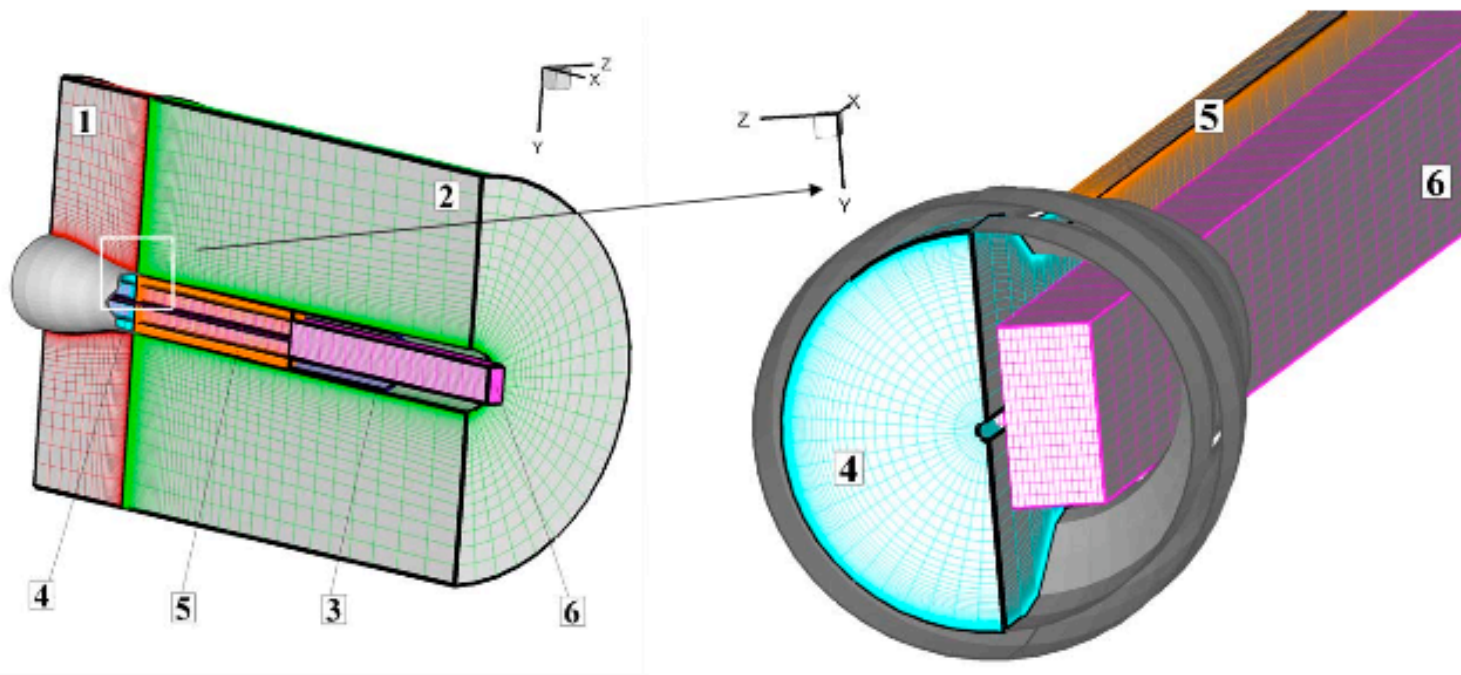
R. Ladeinde

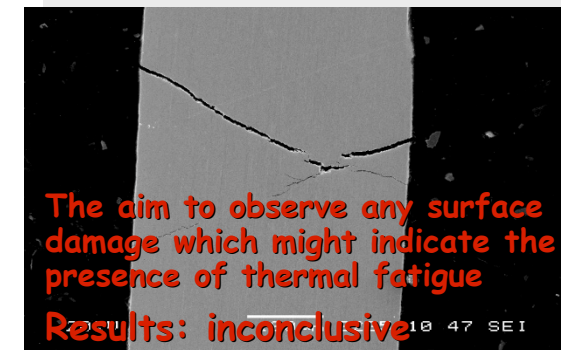
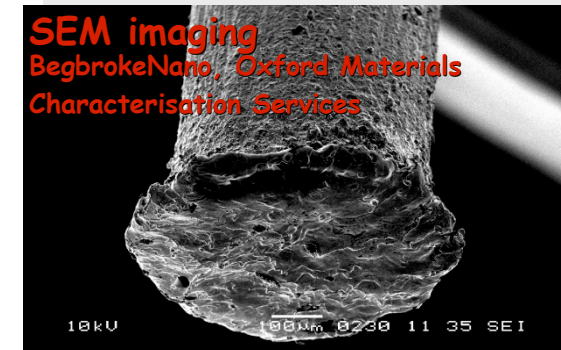
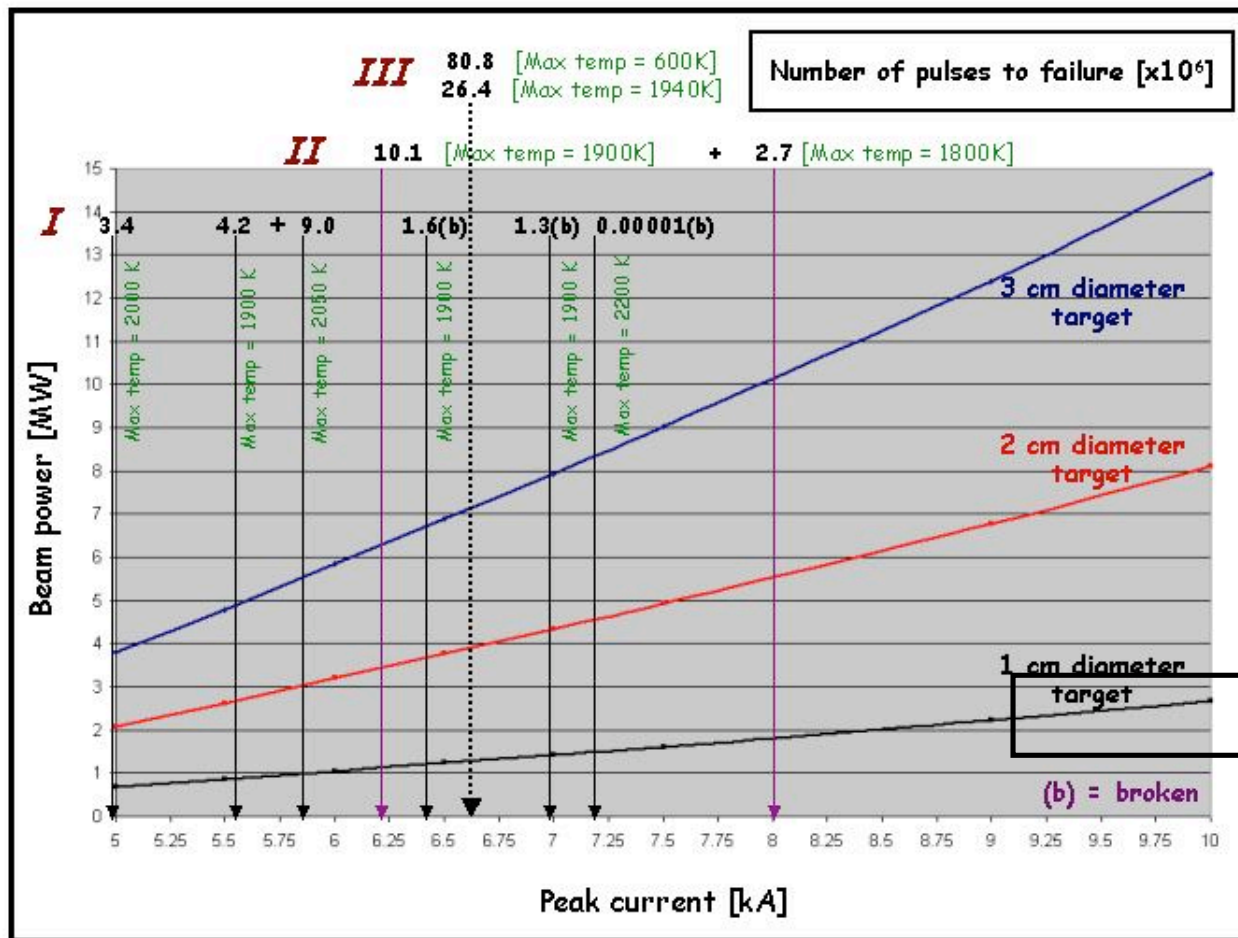
STONY
BROOK



Our Numerical Approach

- CFD-Coupled, Hybrid Nozzle/Plume Calculation
 - Viscous internal flow + plume flow
 - Avoids need to specify jet exit conditions (unknown)





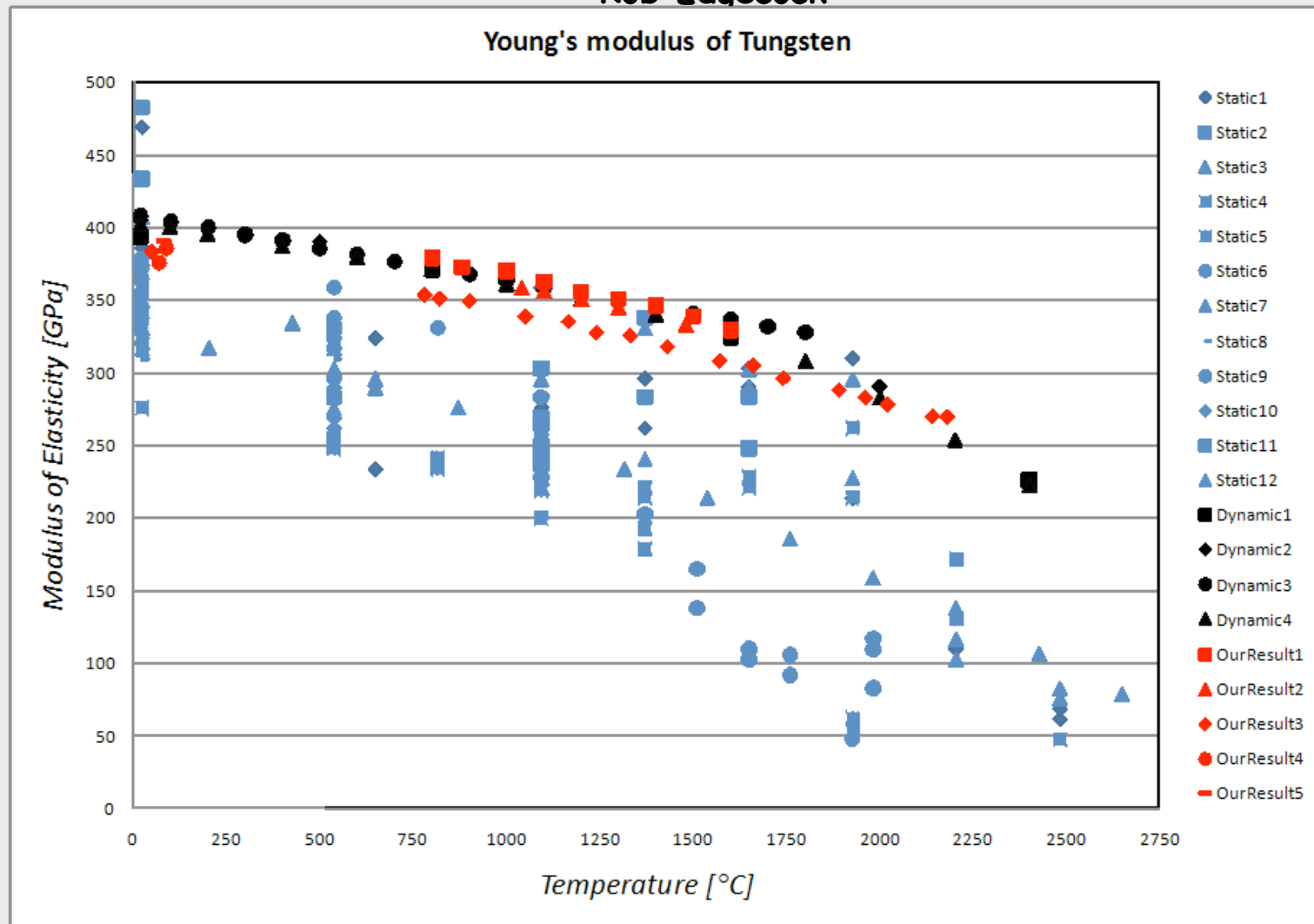
Focus now:

Measure stress;
Confirm modelling.

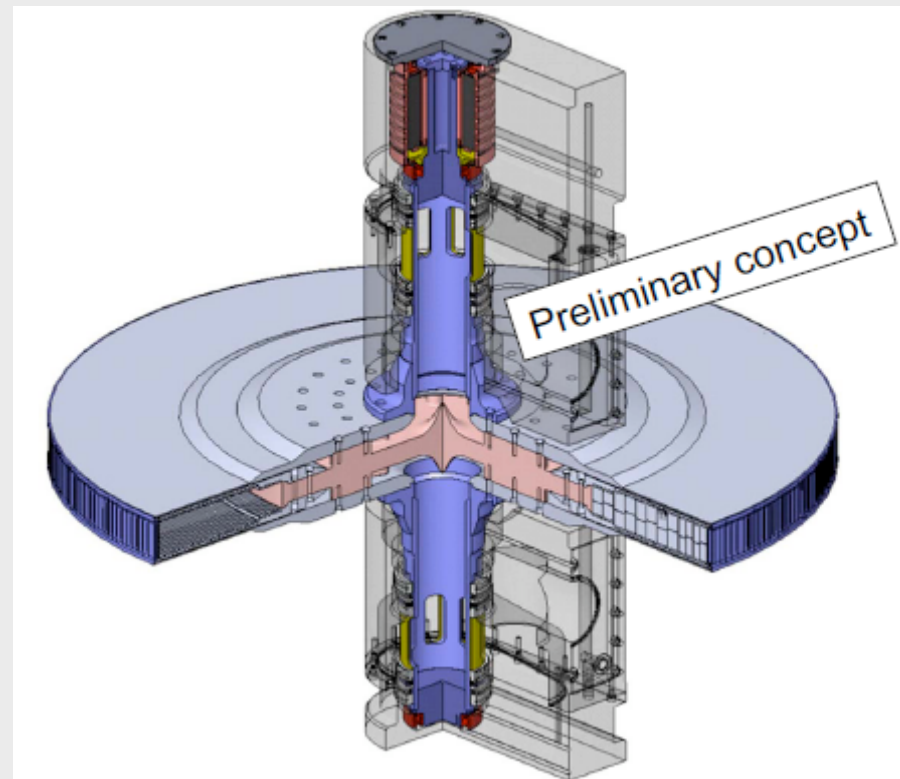
> 10 years for 2cm diameter target
> 20 years for 3cm diameter target

J. Pozimski, DSPL, 10th April 2010, P 13 / 12

Rob Edgecock



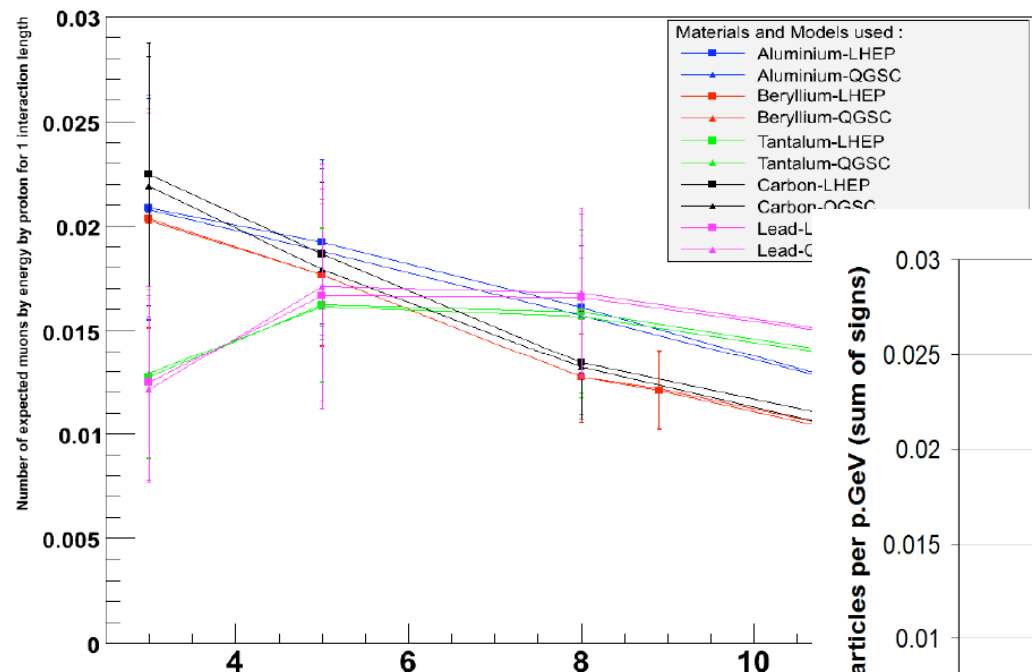
- **Current option: a wheel – being investigated now**
- **Several already used, but most relevant: design study**



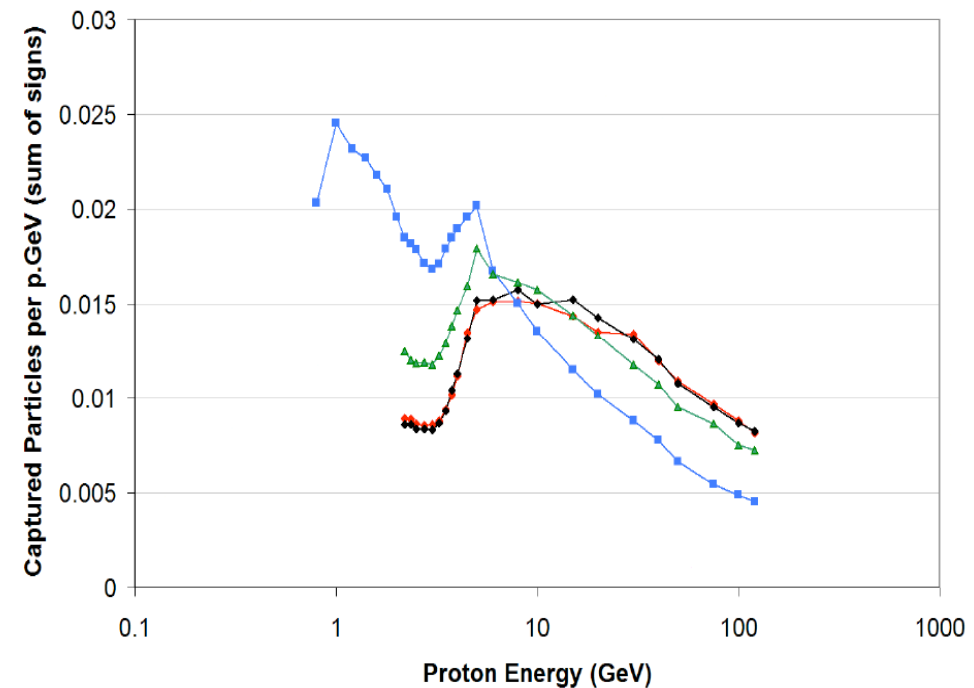
Horizontal for compatibility with baseline target station

Reweighted Harp results

**Includes NF muon
acceptance!**

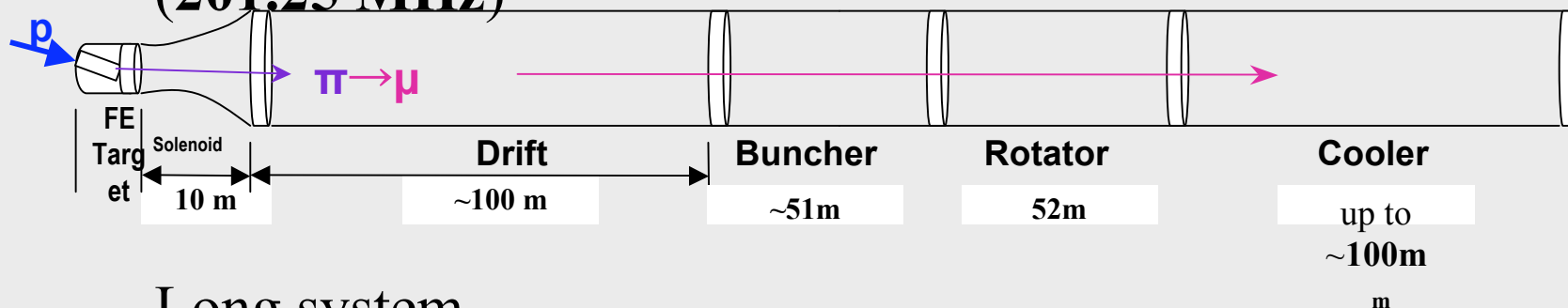


Comparison to MARS
simulations (S.Brooks)

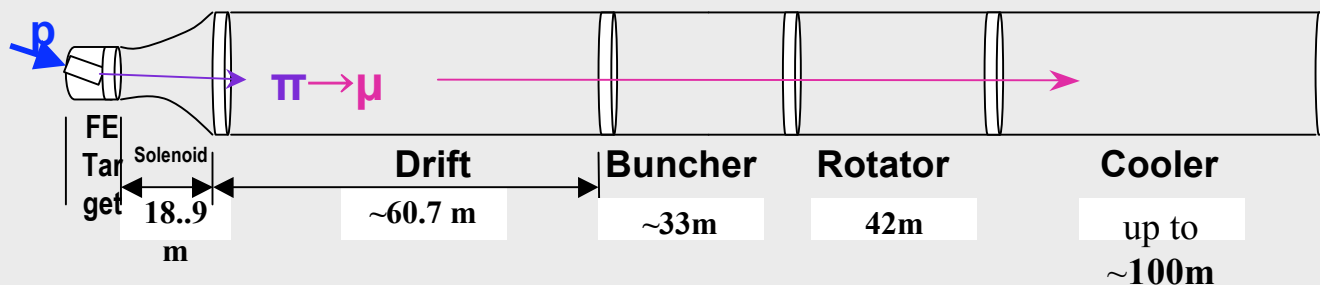


D. Neuffer

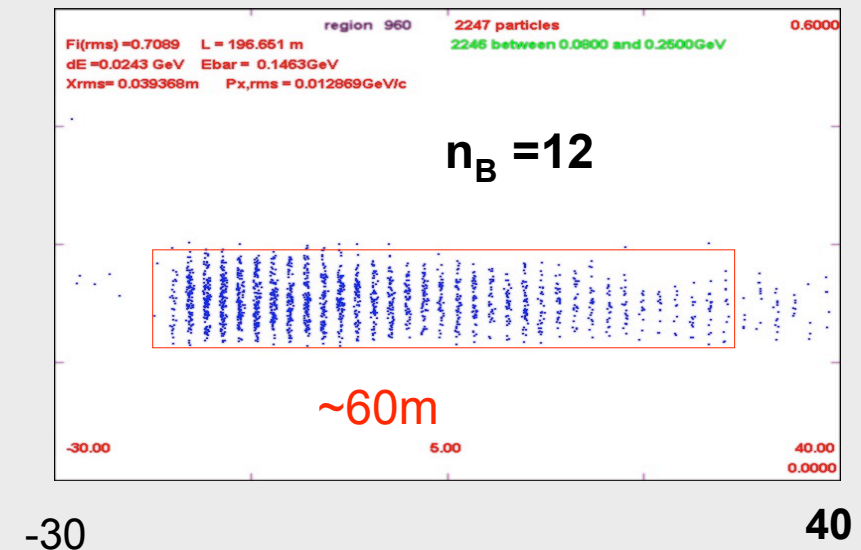
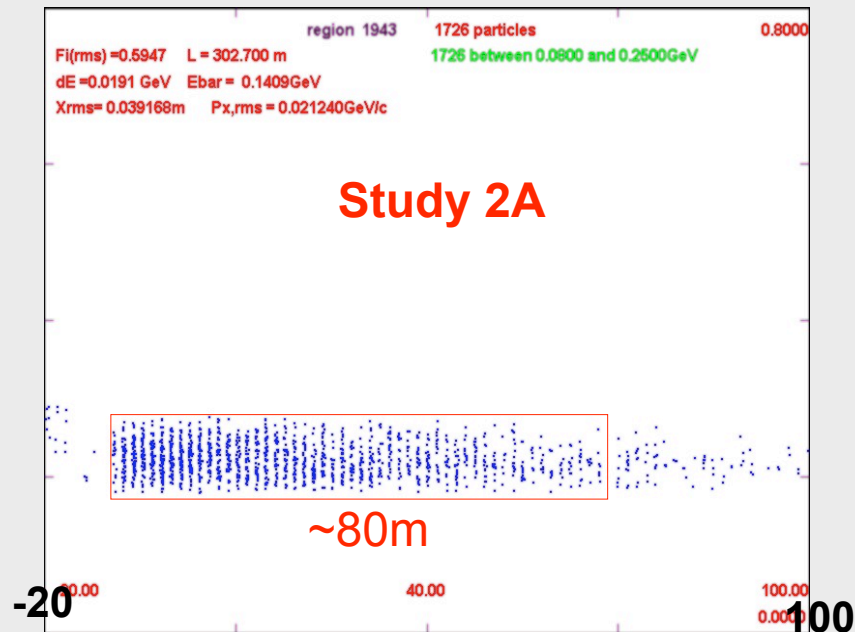
- ISS study based on $n_B = 18$ (280 MeV/c to 154 MeV/c)
 - Buncher 0 to 12MV/m; Rotator 12.5MV/m, $B=1.75T$ (201.25 MHz)



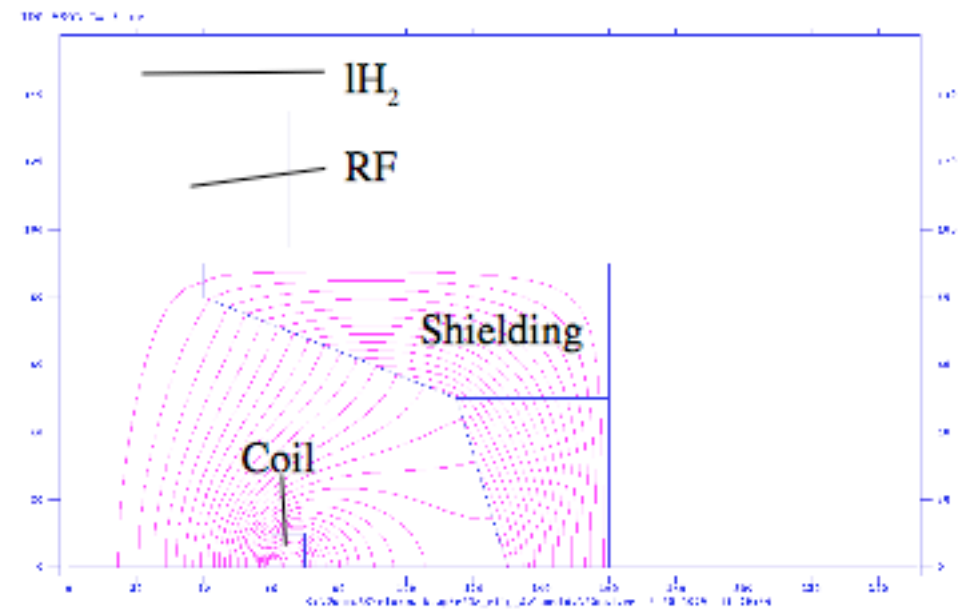
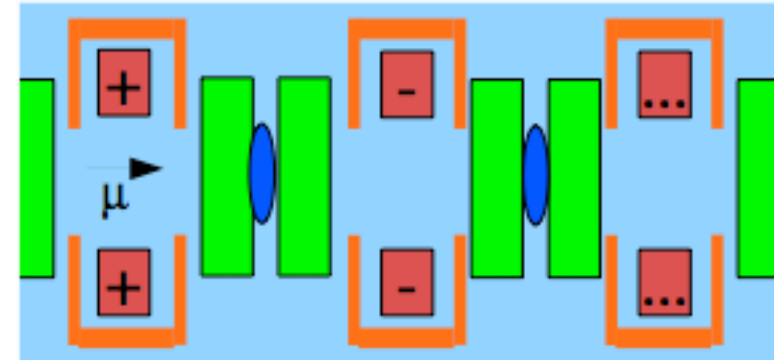
- Long system,
- Try shorter version - $n_B = 10$ (233 MeV/c to 154 MeV/c)
 - slightly lower fields (1.5T, 15MV/m)
 - Buncher 0 to 9 MV/m, Rotator 12MV/m
 - Shorter bunch train



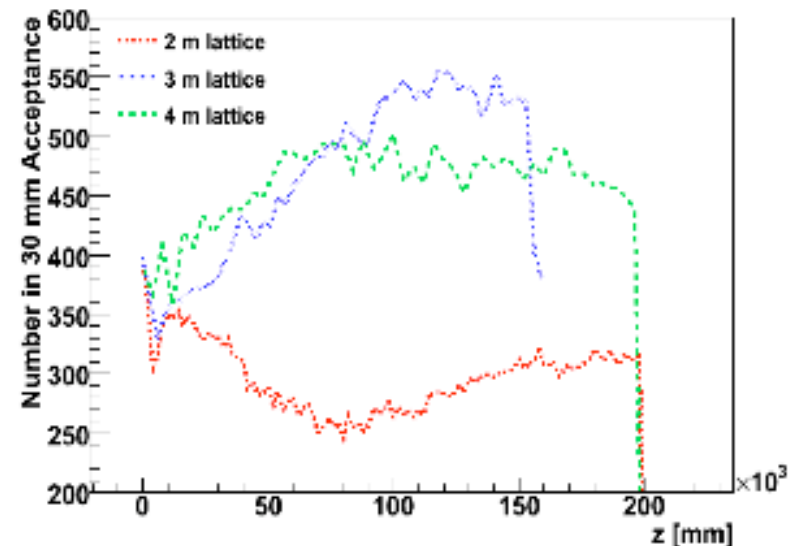
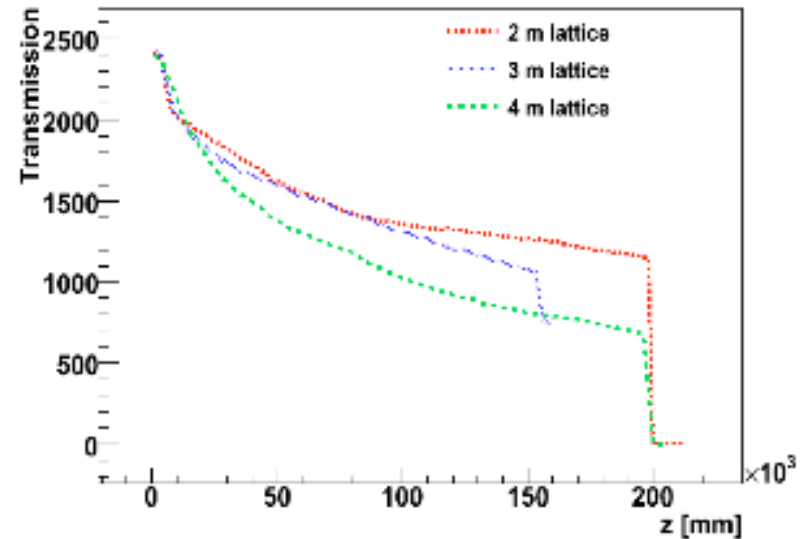
D. Neuffer

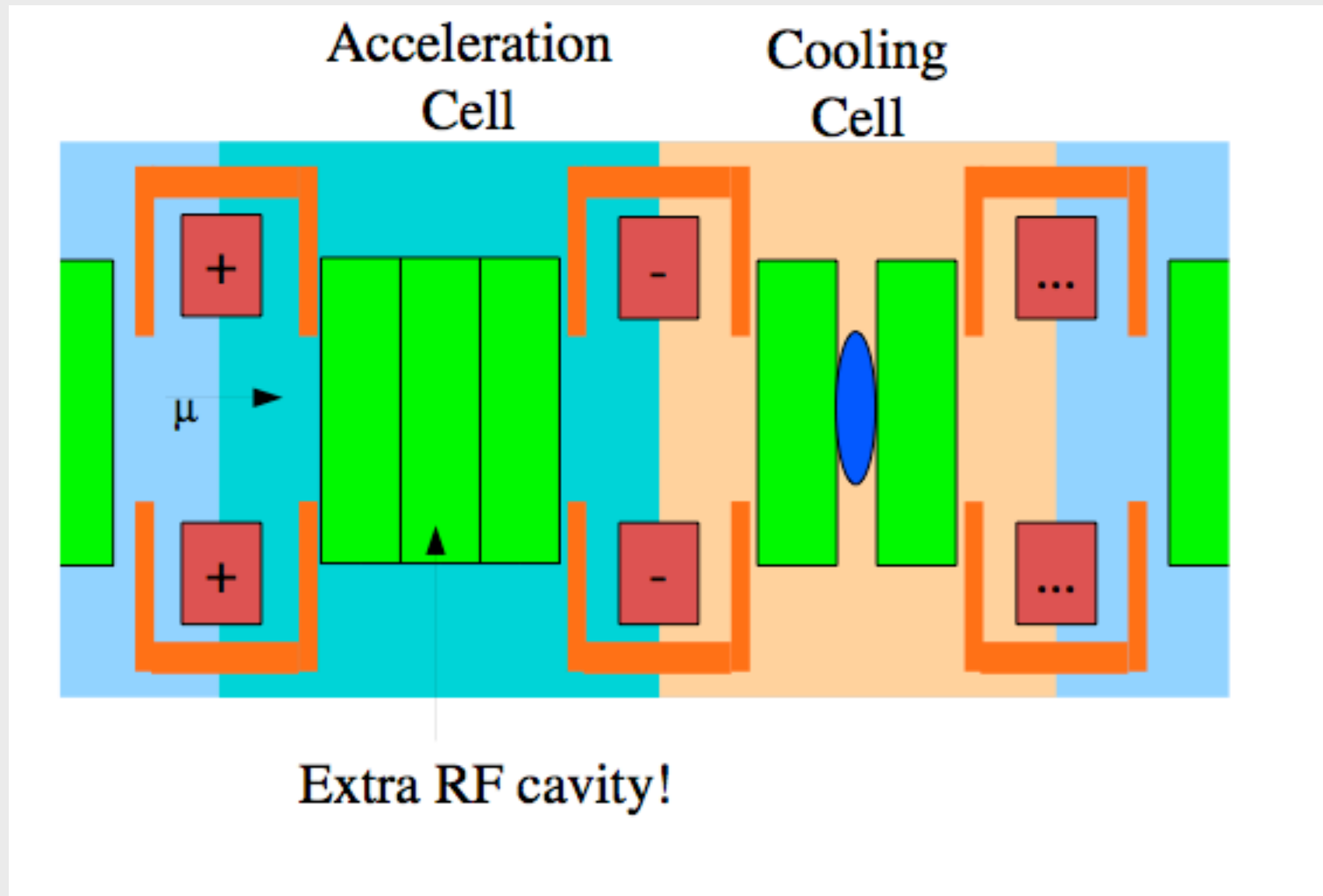


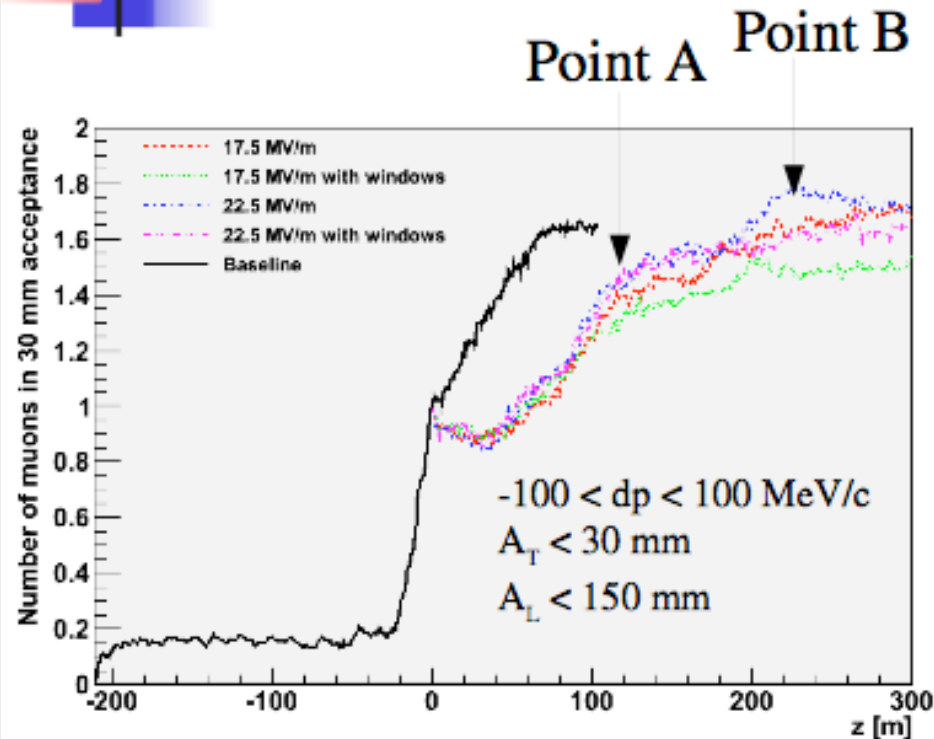
- Increase cell length to remove RF from fringe fields
 - Add shielding using iron or bucking coils
- Look at cooling section
 - This is where the RF is most limited
 - This is where optics are most demanding
- How well can we cool in this shielded scenario?
- How well can we optimise the cooling lattice?
- Try to keep RF cavities in < 0.1 T fields
- Liquid Hydrogen absorbers



- Cell length optimisation
 - Simulated using long coil option
 - Race between RF packing fraction and β function
 - Higher RF packing \Rightarrow quicker cooling
 - Shorter lattice \Rightarrow lower β function (better equilibrium emittance)
- 3m lattice is optimal
 - Worry about initial beam loss
 - Nb low statistics
 - Get ~ 40 % with long coil (a bit more optimisation is possible)
- Case for beta tapering?







- Fairly large transmission losses
 - $> \sim 50\%$
- Most of the remaining beam is inside the 30 mm acceptance
- Getting increase in rate of $\sim 70\%$
 - But with more hardware
 - Performance quite similar to baseline

- If I stop at point A - I use roughly the same amount of hardware as the baseline (RF packing fraction $\sim 1/2$ that of the baseline)
 - And lose a few muons
- I can recover baseline performance if I go to Point B
 - But those last few muons are expensive!

Pion production: 2 GeV proton beam on a 26 cm-long Hg target in 20 T field.

Drift: 30 m decay in 1.8 T field.

**Rotation: 30 1m-long cavities at 44 MHz (2 MV/m) in 1.8 T solenoid. Phasing by 1 degree shift from -121 to -4 deg.
(energy spread reduction by 2)**

Cooling I: 44 MHz RF + H₂ absorbers. (ϵ_{\perp} reduced by 40%)

Acceleration: 88 MHz (4 MV/m) RF provide acceleration to 300 MeV/c.

Cooling II: 88 MHz RF + H₂ absorbers. (ϵ_{\perp} reduced by 30%)

Pion production: 2 GeV proton beam on a 26 cm-long Hg target in 20 T field.

Drift: 30 m decay in 1.8 T field.

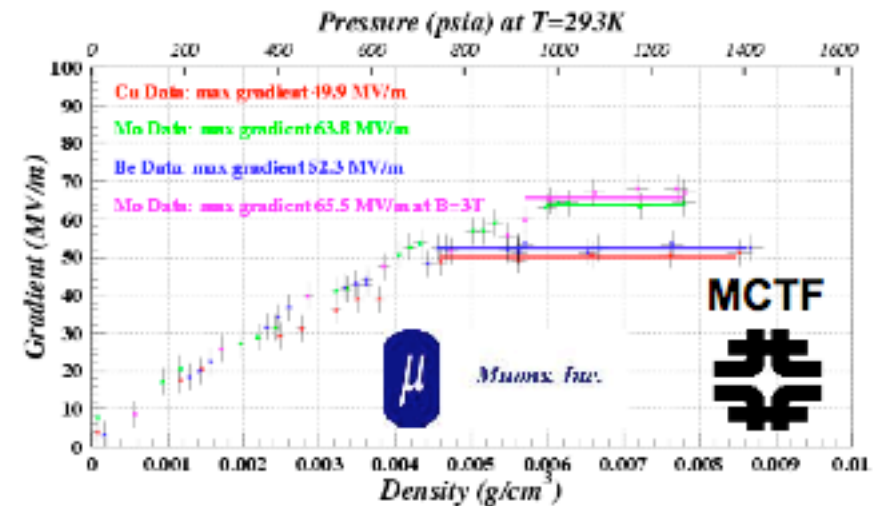
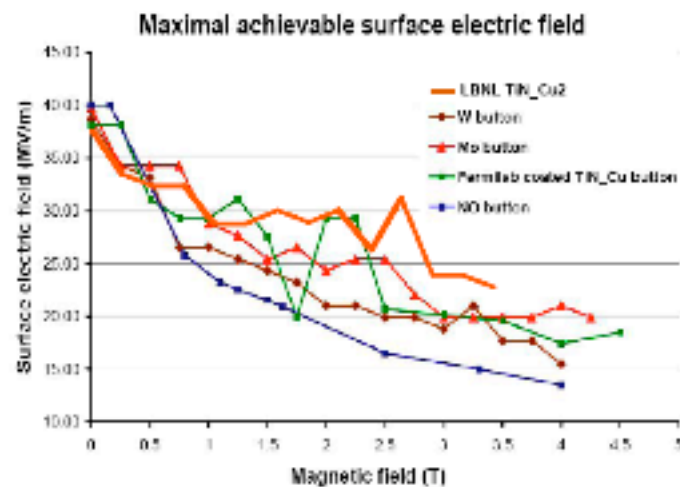
**Rotation: 30 1m-long cavities at 44 MHz (2 MV/m) in 1.8 T solenoid. Phasing by 1 degree shift from -121 to -4 deg.
(energy spread reduction by 2)**

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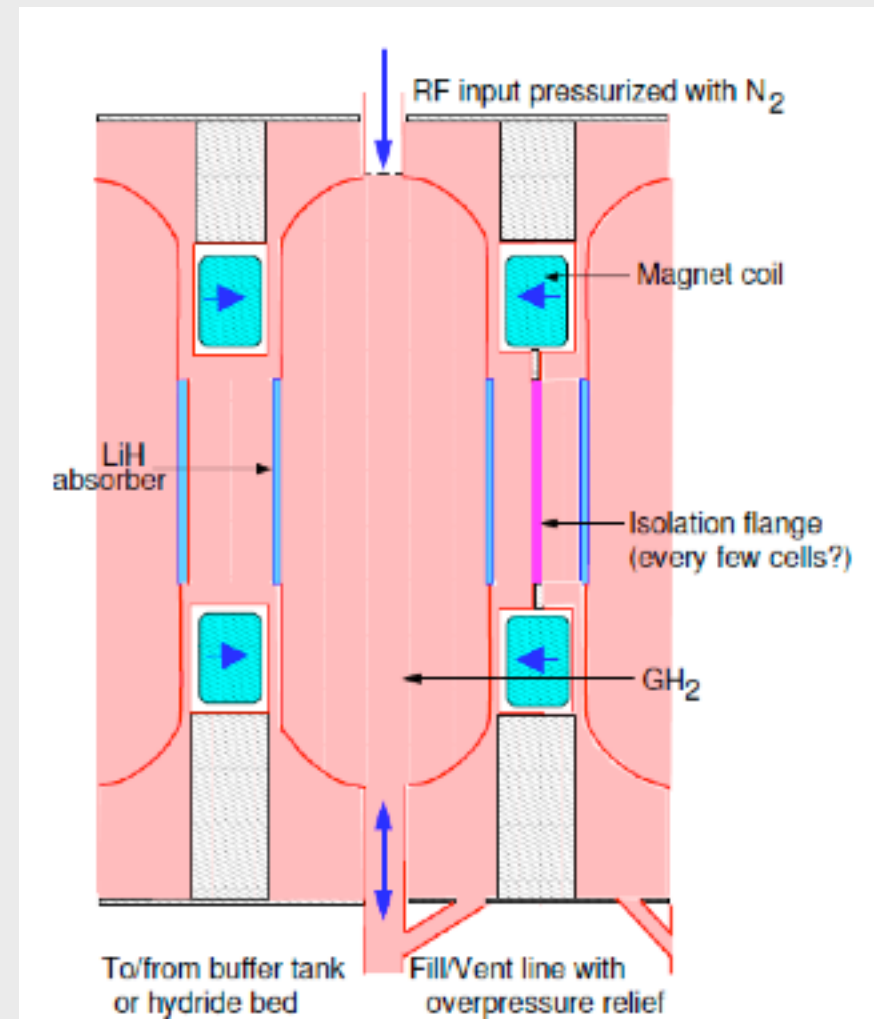
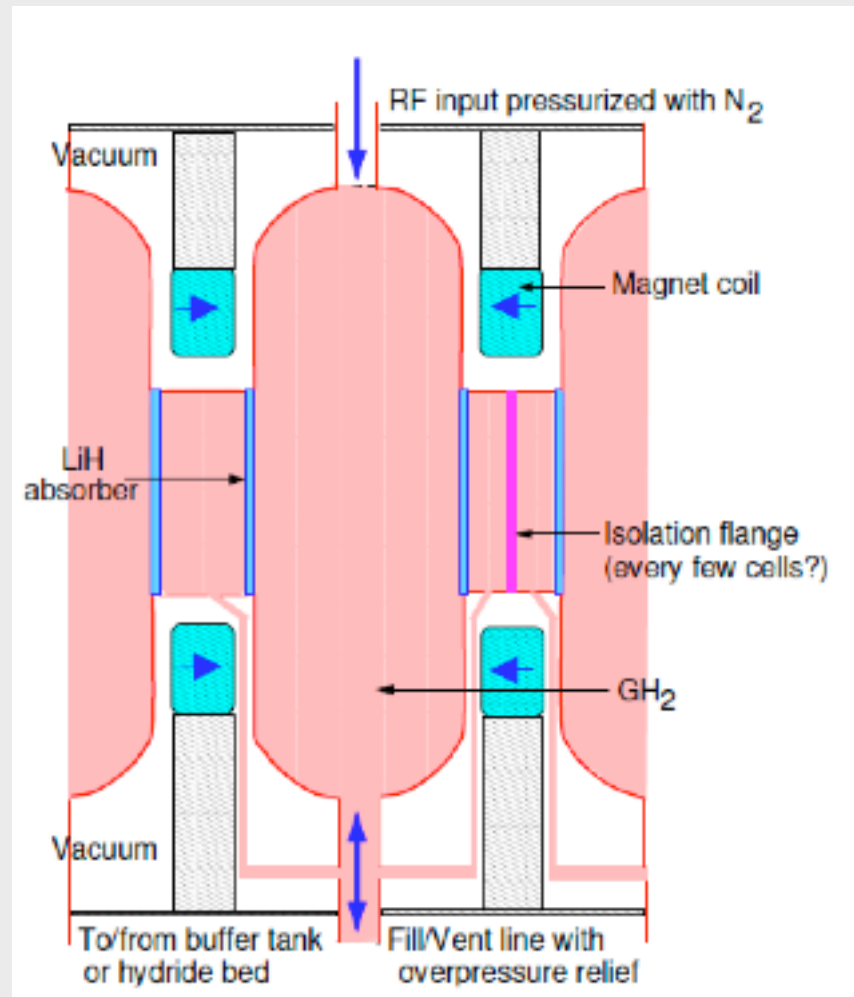
Acceleration: 88 MHz (4 MV/m) RF provide acceleration to 300 MeV/c.

Cooling II: 88 MHz RF + H₂ absorbers. (ϵ_{\perp} reduced by 30%)

- We have evidence that vacuum RF cavity gradient performance degrades in a strong magnetic field
 - alternative approach of HPRF does not
 - though it has other potential issues
- It seems prudent to begin investigating the technical aspects of implementing HPRF in a linear cooling channel
 - minimizes changes in cooling channel layout and hardware

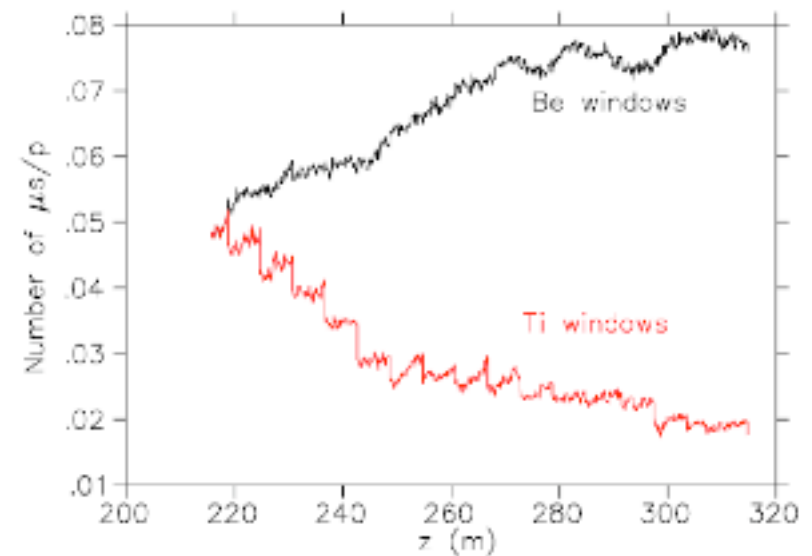
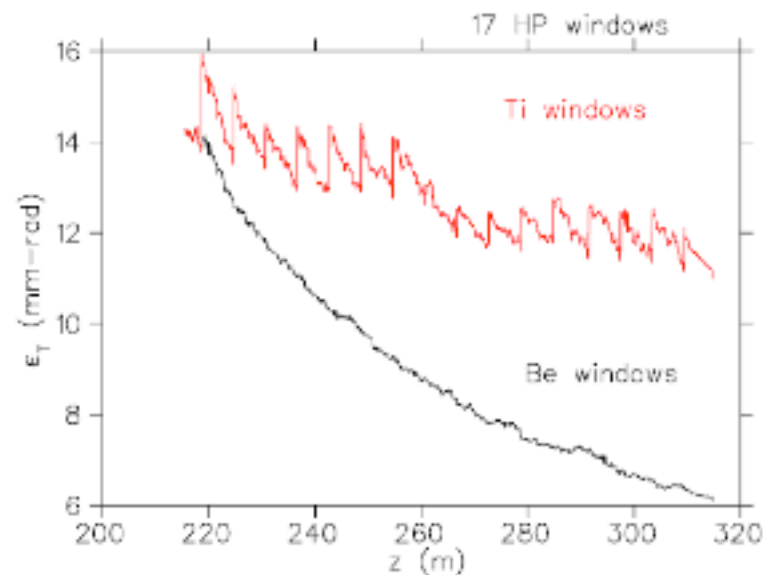


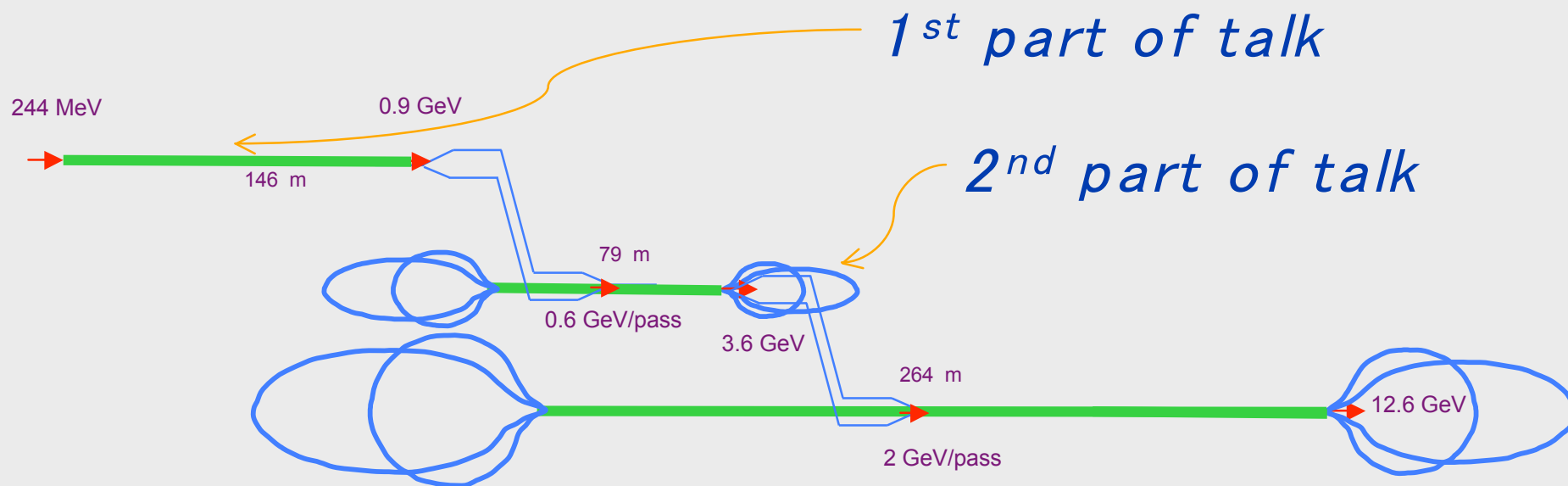
M. Zisman



M. Zisman

- To make sure we were not fooling ourselves, ran cases with both Be and Ti
 - the difference is obvious
- Will next look at Al and AlBeMet windows as time permits
 - Al is okay in terms of hydrogen embrittlement; not yet sure about other materials

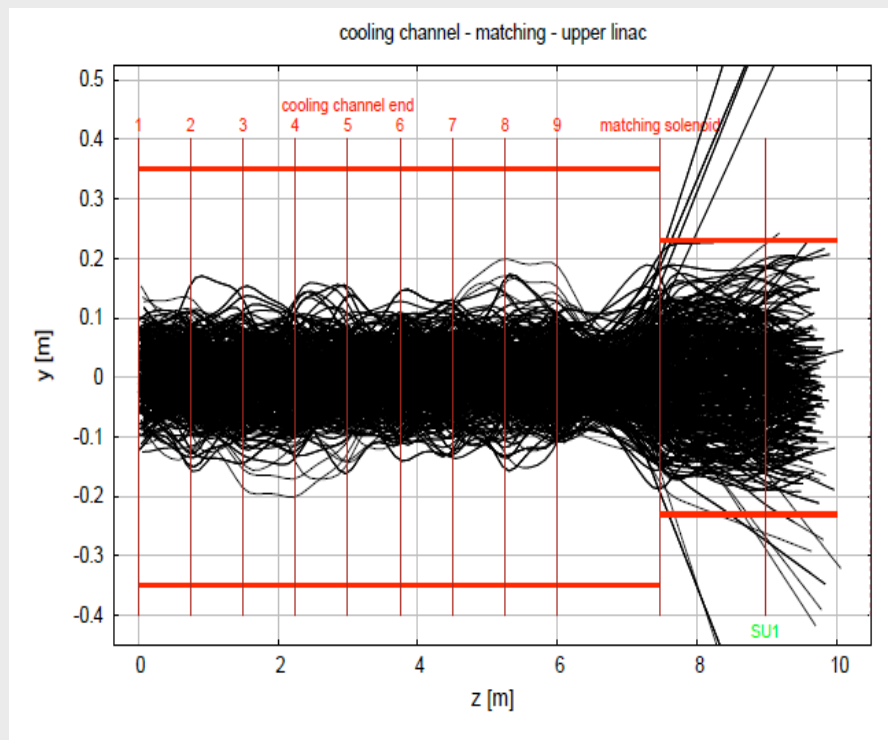




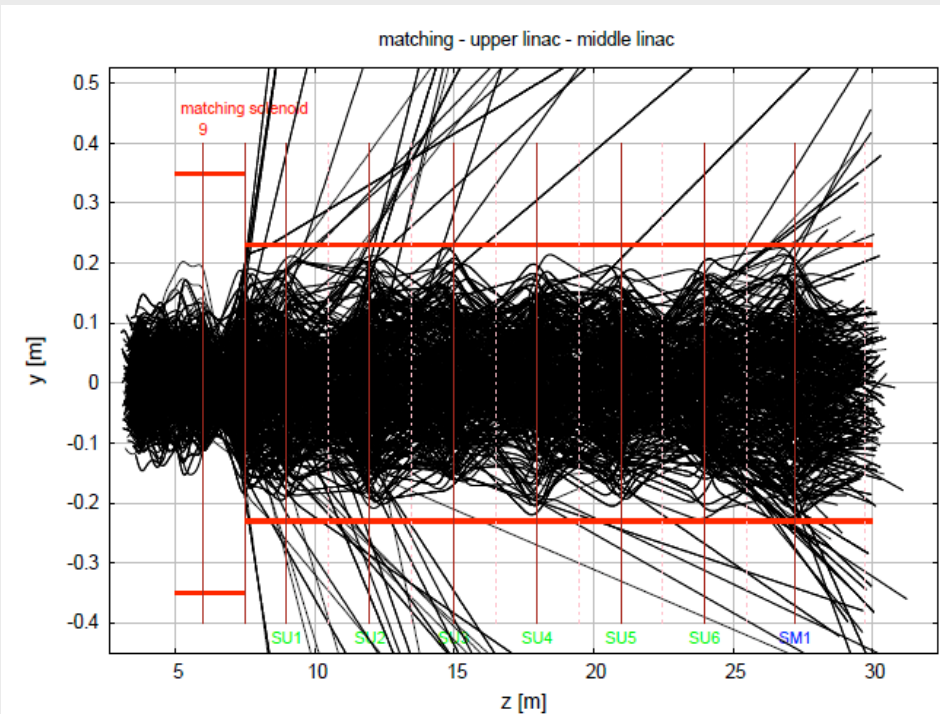
IDS Goals:

- Define beamlines/lattices for all components
- Resolve physical interferences, beamline crossings etc
- Error sensitivity analysis
- End-to-end simulation (machine acceptance)
- Component count and costing

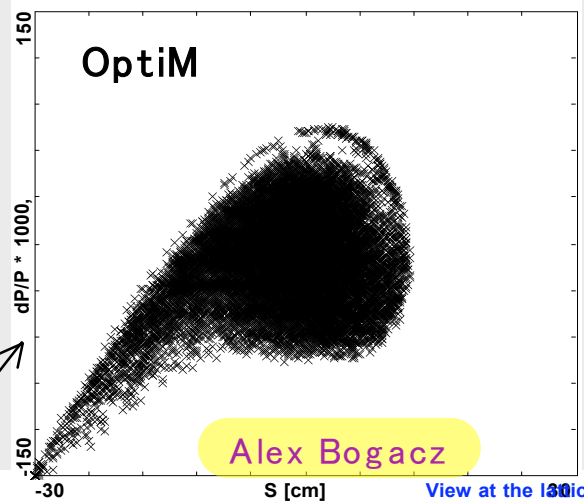
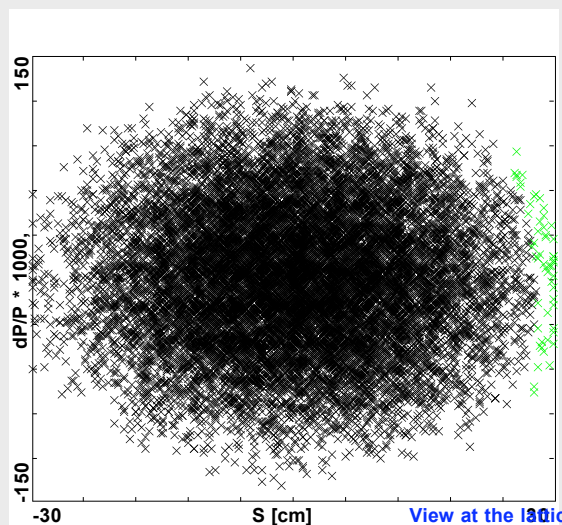
cooling → upper
linac



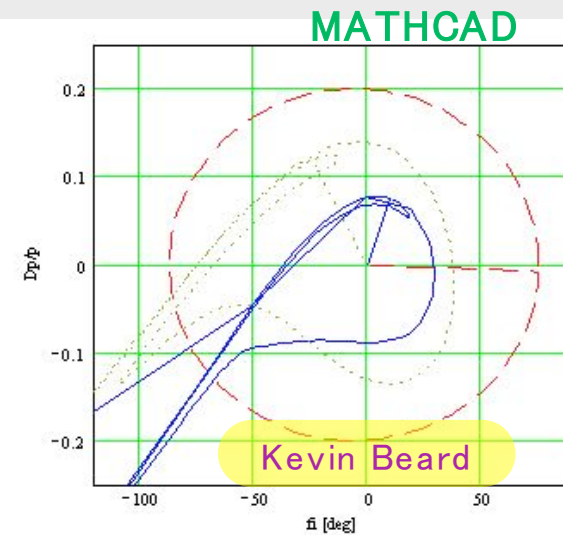
upper → middle linac



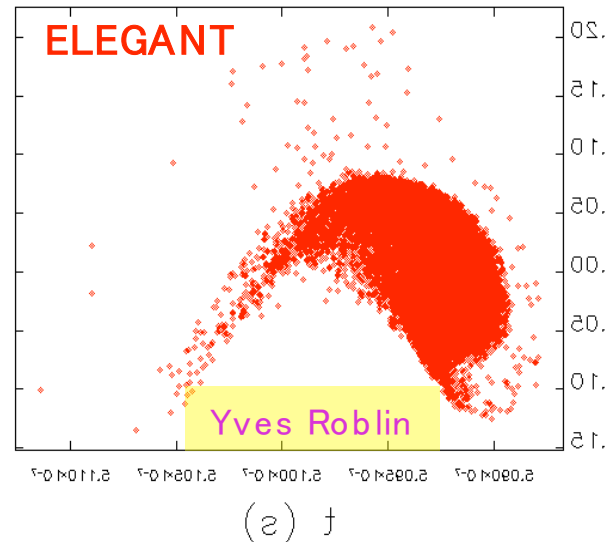
Initial distribution



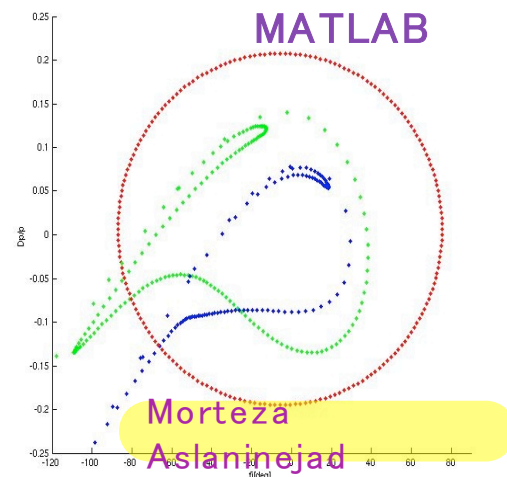
Alex Bogacz



Kevin Beard



Yves Roblin



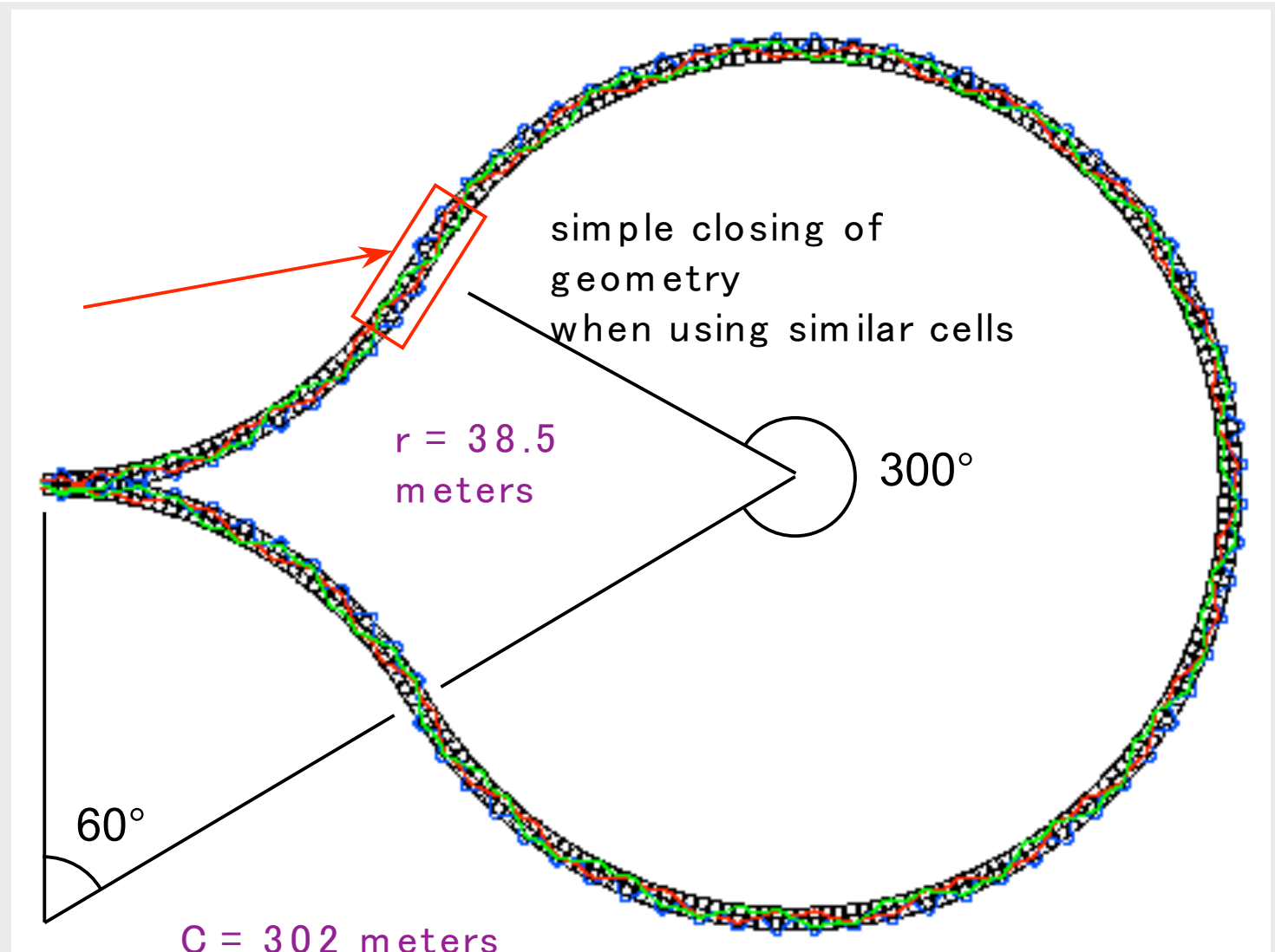
Morteza
Aslaninejad

Multipass FFAG arc

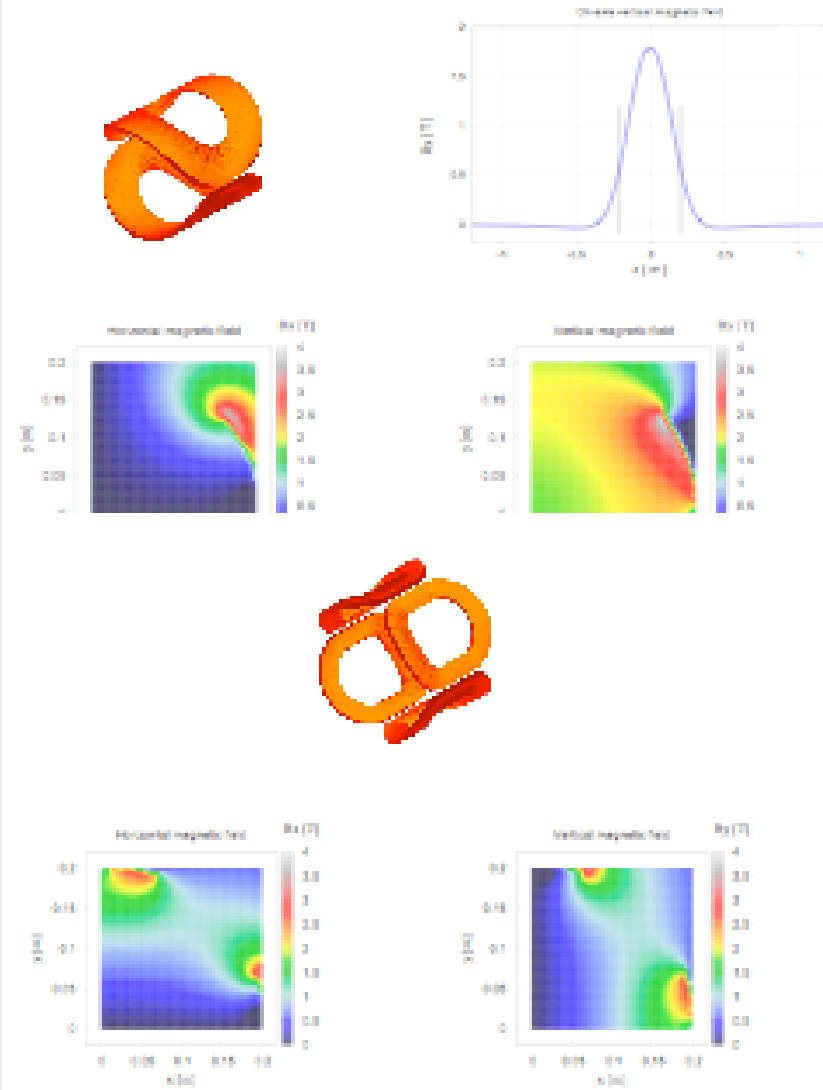
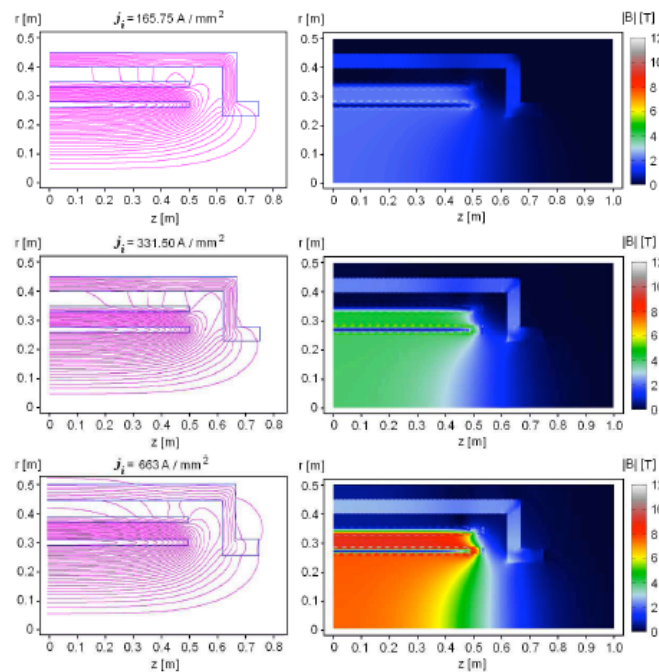
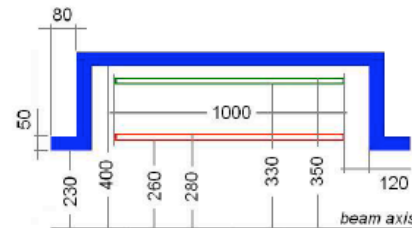
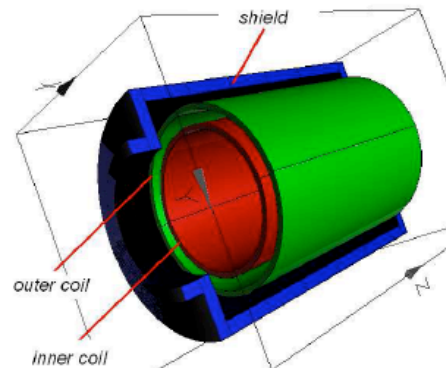
A. Bogacz



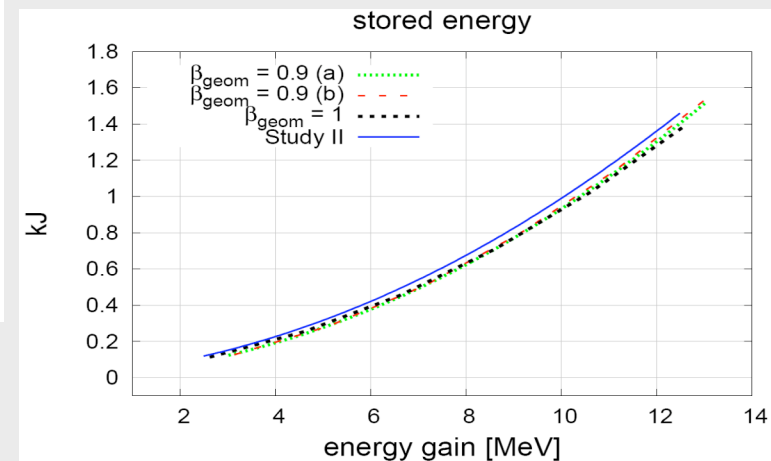
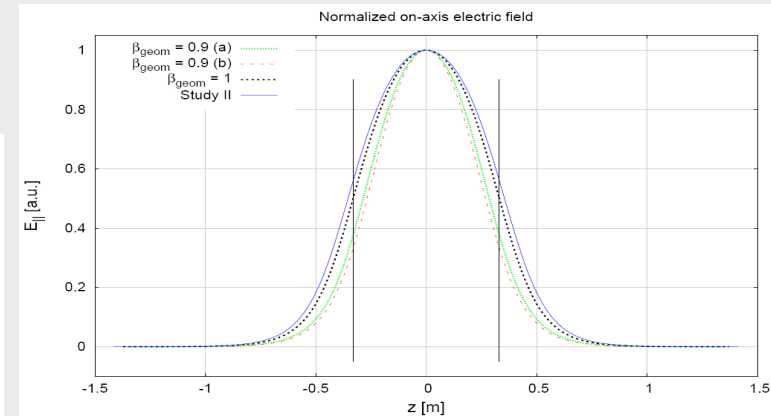
Vasiliy Morozov



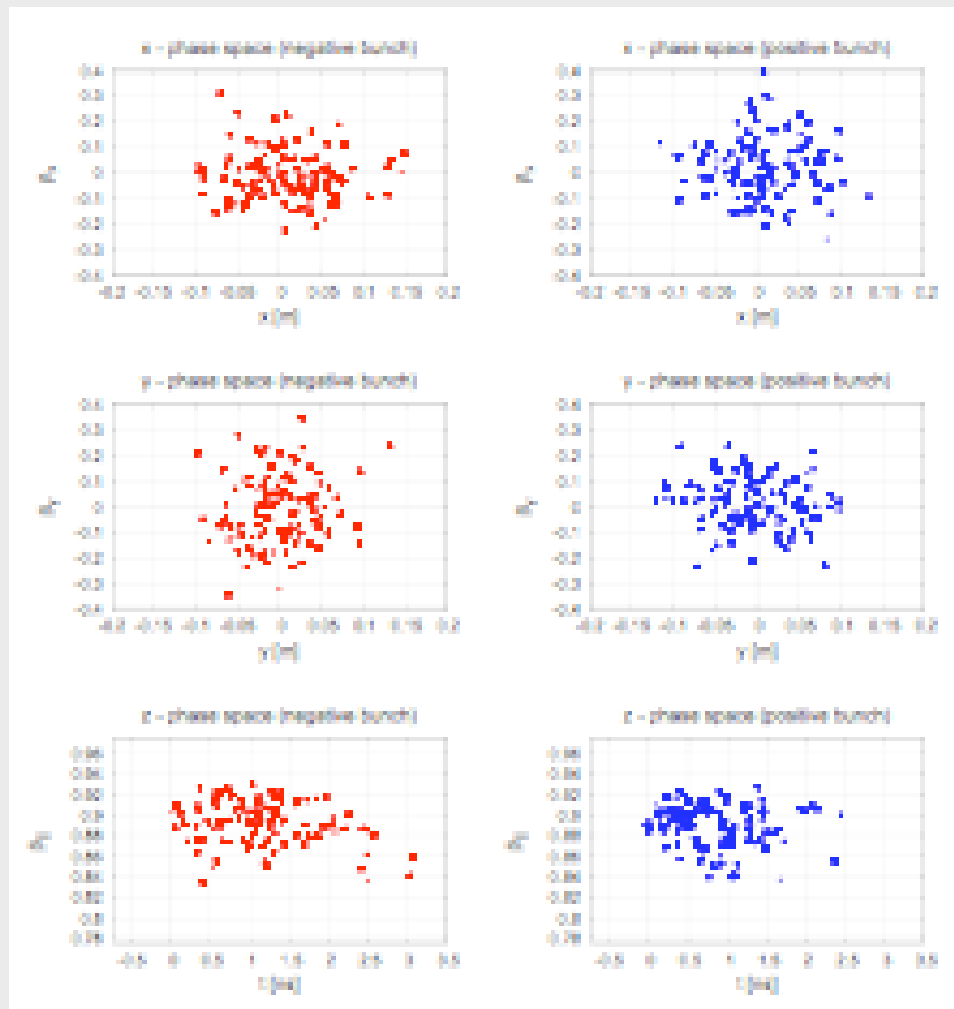
C. Bontoiu



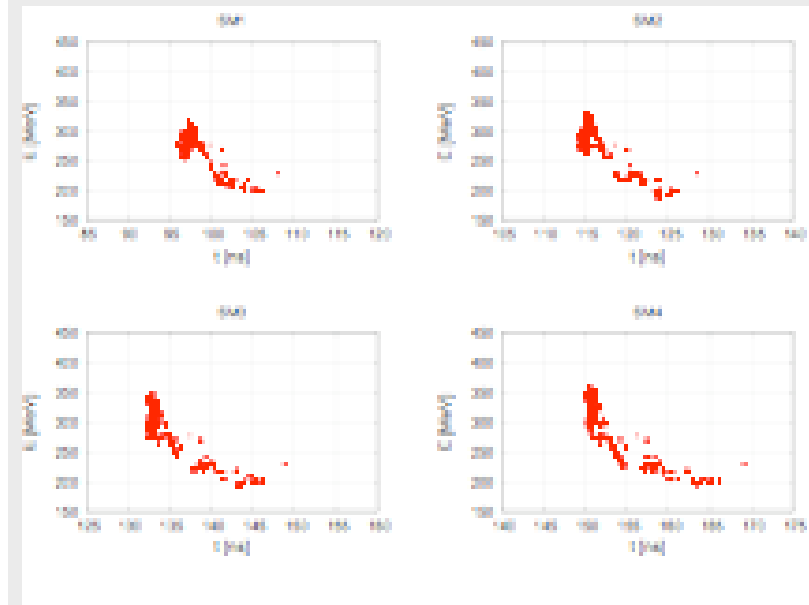
Parameter	$\beta_{geom} = 1$	$\beta_{geom} = 0.9$ (a)	$\beta_{geom} = 0.9$ (b)	Study II
l_{cav} [m]	0.7448	0.67034	0.67034	0.8282
r [m]	0.6854	0.7042	0.6804	0.6641
f_0 [MHz]	201.247	201.251	201.255	198.575
Q [10^9]	24.67	19.6	18.8	26.7
T	0.650	0.716	0.726	0.591
\hat{E} [MV/m]	26.17	27.19	27.83	26.38
\bar{E} [MV/m]	20.62	20.81	20.53	20.42
$ E _{surface}^{max}$ [MV/m]	21.70	24.87	29.45	19.75
$ H _{surface}^{max}$ [kA/m]	48.06	58.53	61.92	45.00
U [J]	712	772	797	747
$\int_{-\infty}^{+\infty} E(0,z) \cos[\omega t(z)] dz$	8.6142	9.0081	9.1336	8.8466
$\int_{-l_{cav}/2}^{+l_{cav}/2} E(0,z) \cos[\omega t(z)] dz$	10.0000	10.0000	9.9999	10.0000
$\int_{+l_{cav}/2}^{+\infty} E(0,z) \cos[\omega t(z)] dz$	-0.69204	-0.49594	-0.43320	-0.75676
correction [%]	-13.841	-9.9188	-8.6639	-15.135



Intital distribution from cooling channel



Energy spread in the linac



Parameters of a 3.6 to 12.6 GeV muon ring

Lattice type	FDF triplet
Injection/extraction energy	3.6/12.6 GeV
RF frequency	200 MHz
Number of turns	6
RF peak voltage (per turn)	1.8 GV
Synchronous energy	8.04 GeV
Mean radius	~ 160.9 m
B_{max} (@ 12.6 GeV)	3.9 T
Field index k	1390
Total orbit excursion	14.3 cm
Harmonic number h	675
Number of cells	225
Long drift length	~ 1.5 m
Horiz. phase adv. per cell	85.86 deg.
Vert. phase adv. per cell	33.81 deg.

Table 1 - Example of 3.6 to 12.6 GeV muon scaling
FFAG ring parameters.

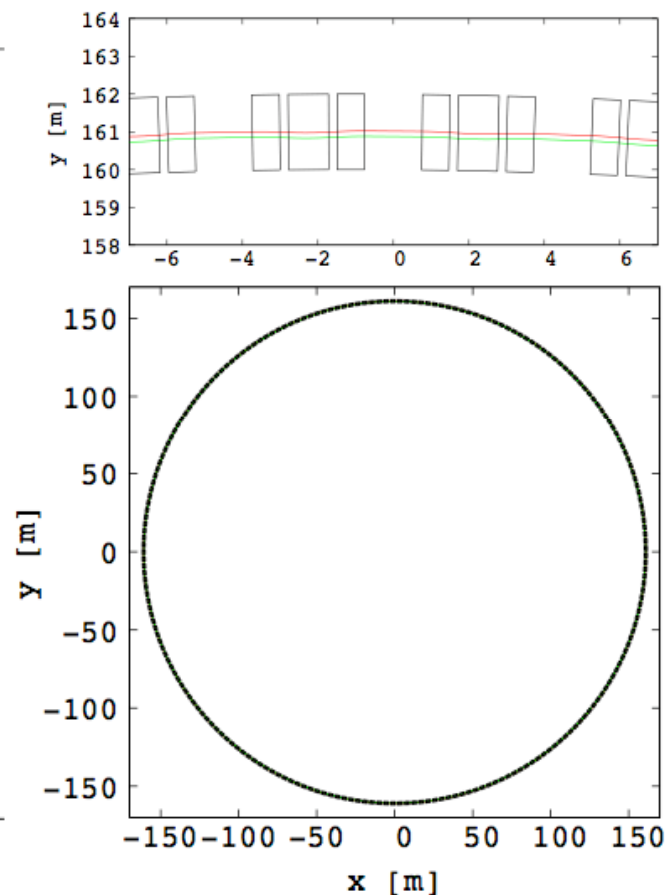
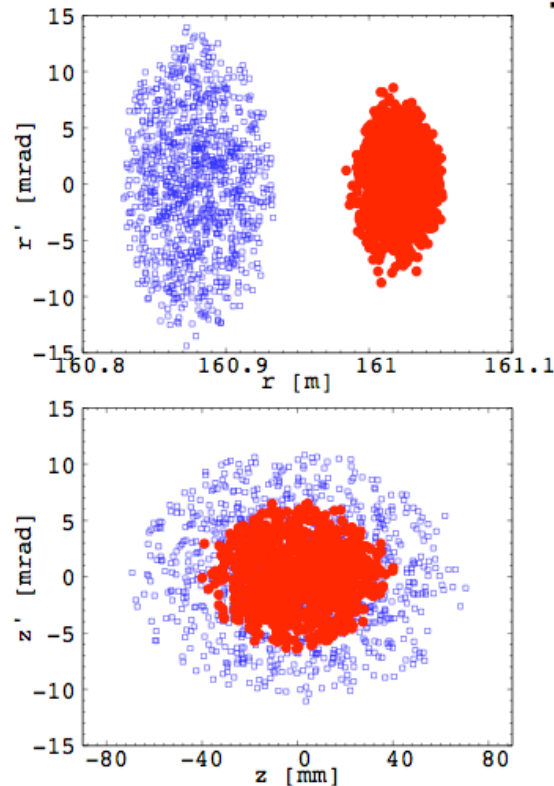


Figure 2 - Ring layout.

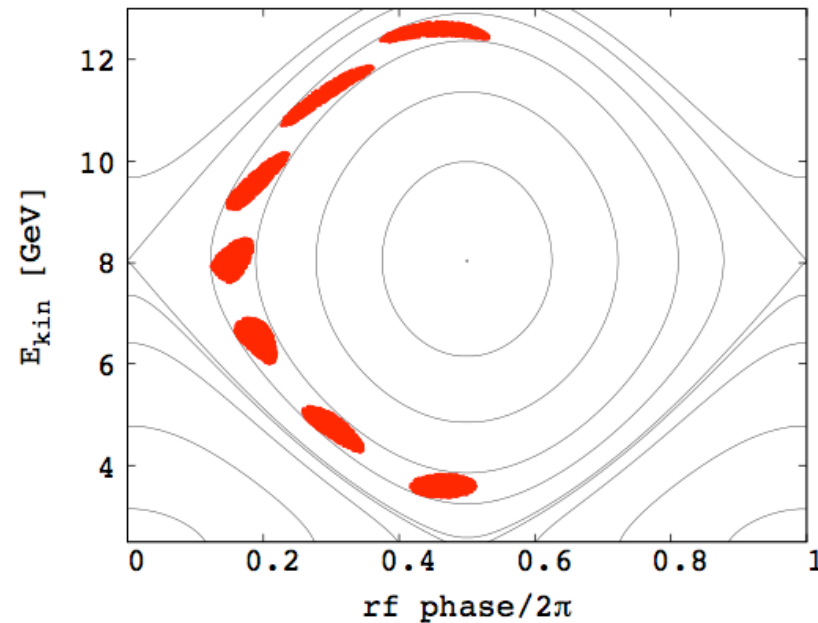
T. Planche

Full acceleration cycle - 6D tracking

- Tracking results -



Figures 8 - Initial (blue) and final (red) particles distribution in the horizontal (top), and vertical (bottom) phase space.

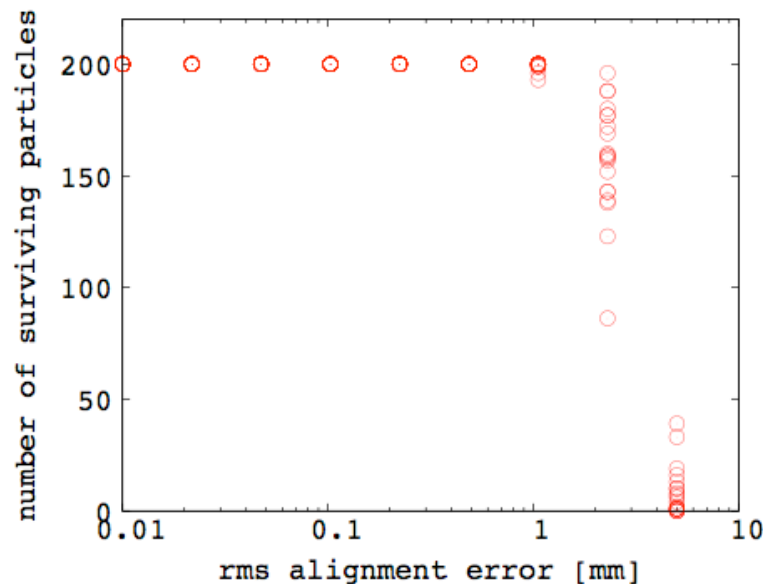


Figures 9 - longitudinal phase space plot showing a 6-turn acceleration cycle. Hamiltonian contours are superimposed.

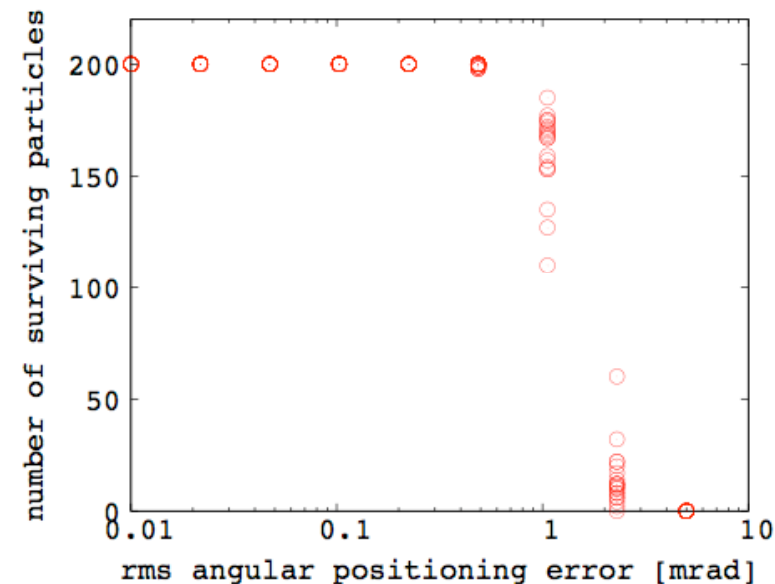
T. Planche

Study with errors

Errors in translation



Errors in rotation



Figures 10 - Number of surviving particles depending on the rms error. 20 different lattices have been generated and tested for each value of rms error.

More tolerant to errors than the NS-FFAG!

S. Berg

Most promising due to longer drifts for RF and kickers : FDFCC

Minimum E 12.6 GeV

Maximum E 25 GeV

Cells 64

Circumference 546 m

Cavity cells 88

RF Voltage 1119 MV

Decay 5.6%

Turns 11.8

D Radius 11.5 cm

F Radius 15.3 cm

D Field 6.5 T

F Field 3.6 T

D Length 1.96 m

F Length 1.29 m

Long drift 3 m

Short drift 0.5 m

S. Berg

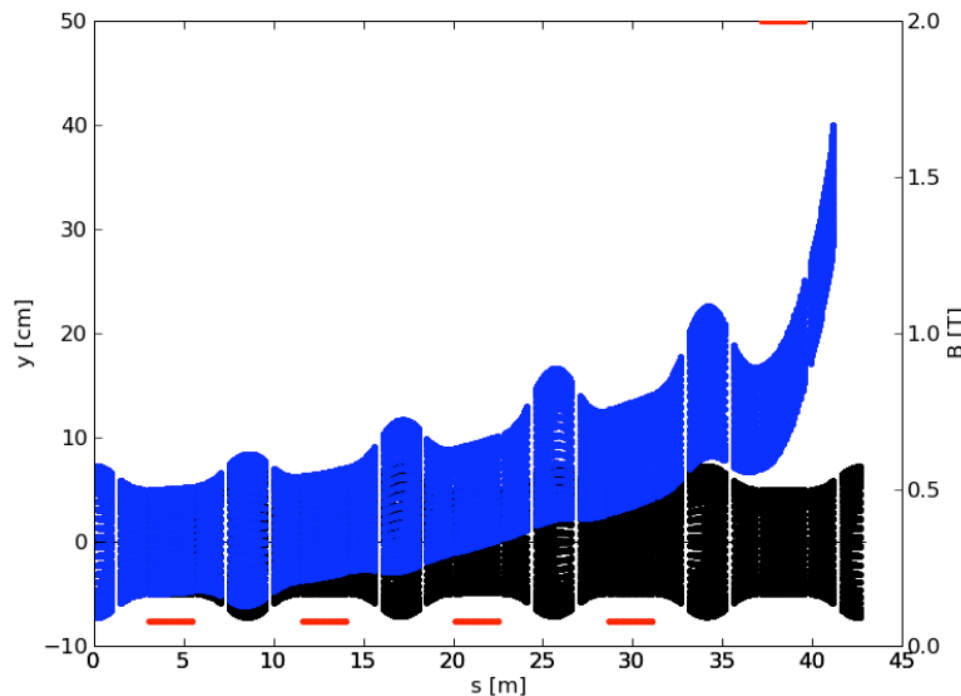
- Reduces time of flight dependence on transverse amplitude
- Downsides
 - Reduction in dynamic aperture
 - Increase in magnet apertures (cost)
 - Modification of time of flight vs. energy
- Still studying optimal choice for this
 - Some modest correction likely included

J. Pasternak

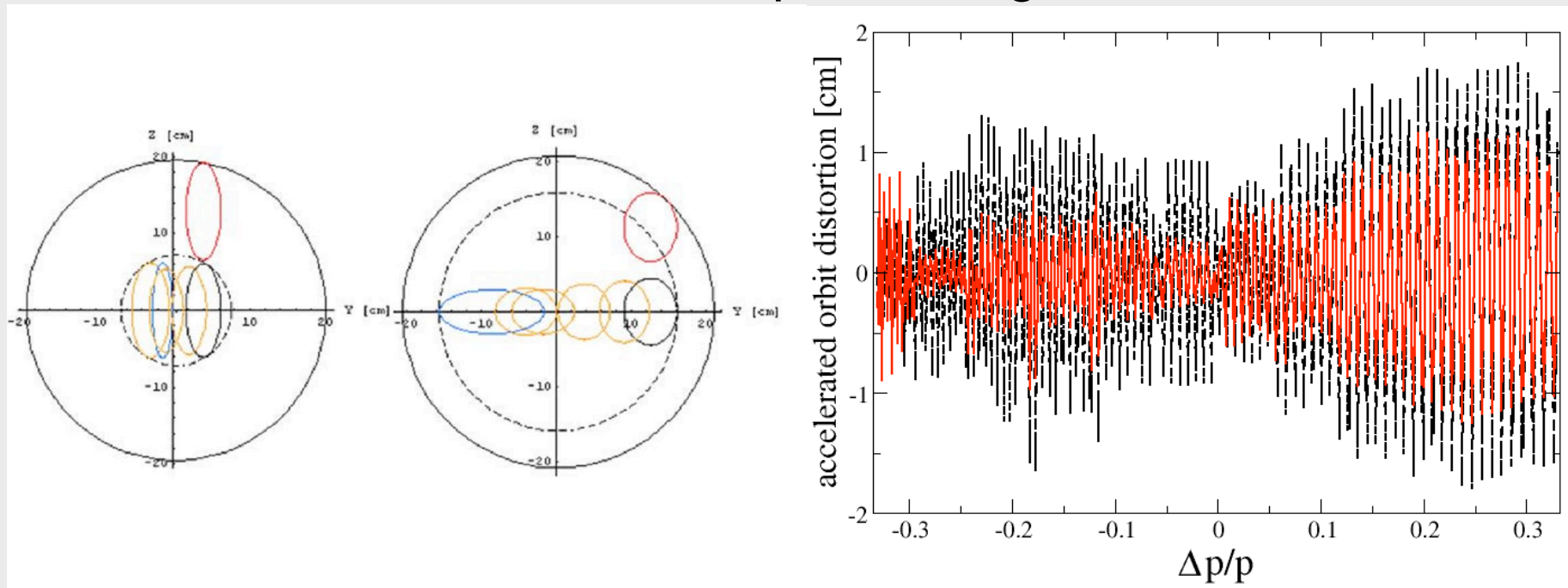
FDFCC - Vertical extraction

- 4 kickers at 0.078 T in consecutive drifts
- Septum at 2T
- Several magnets require large aperture
- Extracted beam 17cm from magnet axis of F after septum

D. Kelliher



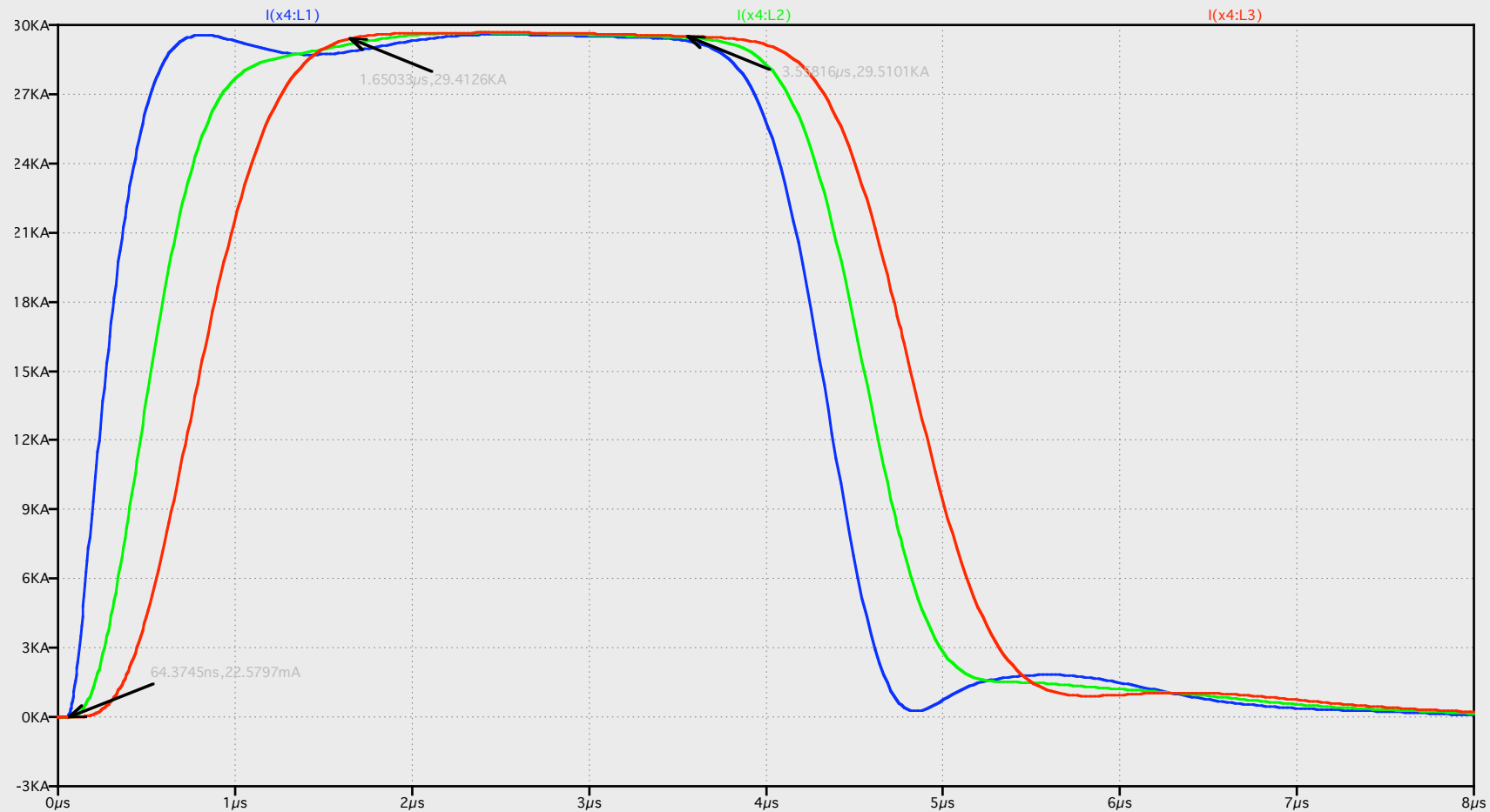
Effects of special magnets

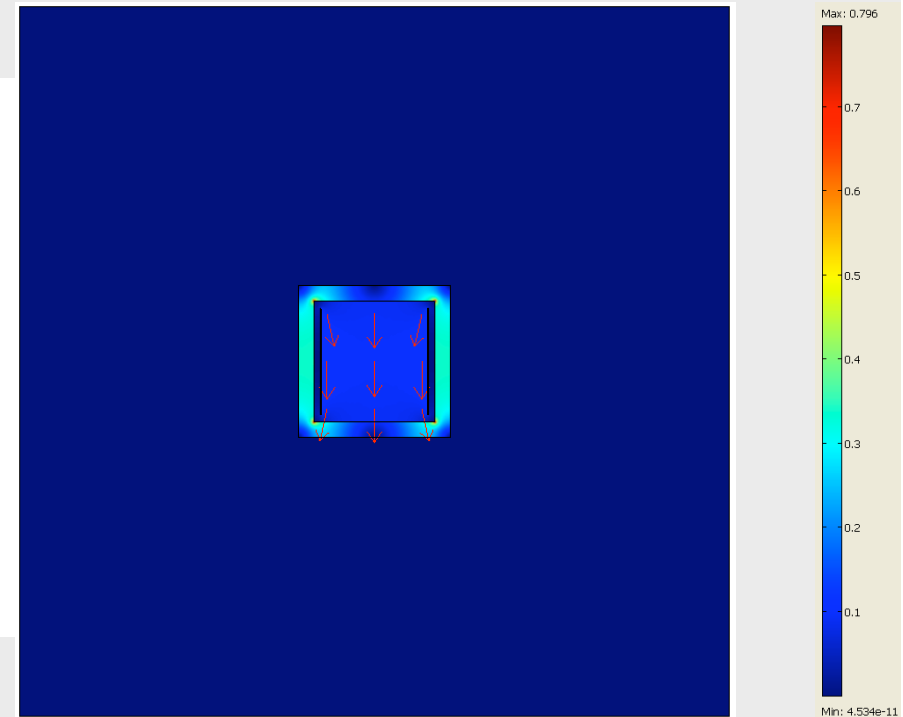
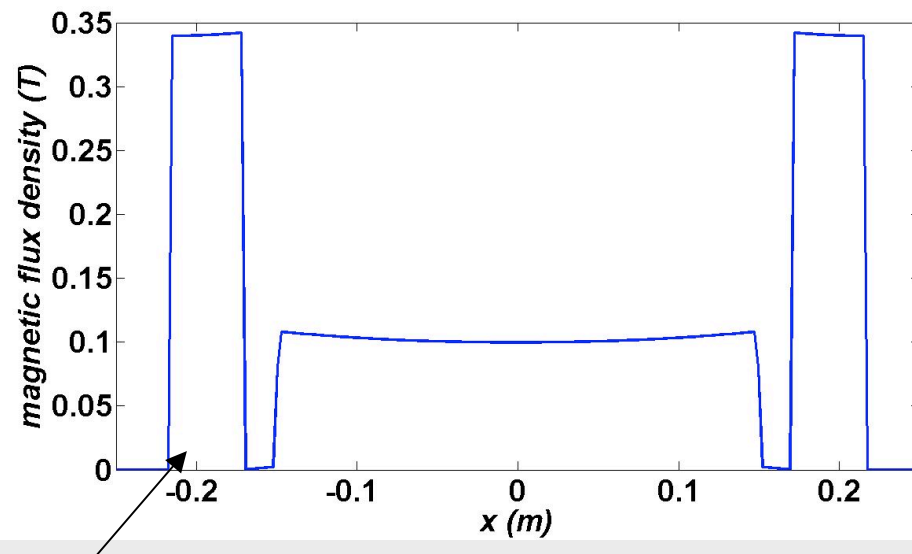


- Beams close to septum push the magnet aperture.
- Special magnets with higher aperture is needed in injection/extraction regions.
- Those magnets introduce the ring lattice symmetry breaking, which can cause accelerated orbit distortion.
- Current studies show that the effect is not dramatic, but more simulations are needed.

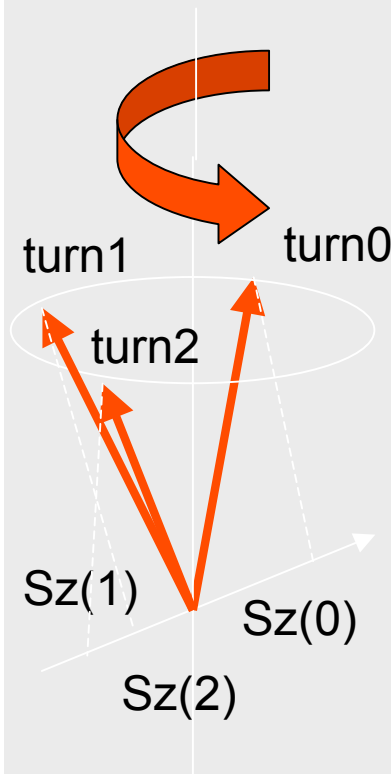
J. Pasternak

Current pulses in 3 kicker sections – „travelling wave” using PSPice





- 2D simulation to estimate the current, which is needed.
- Ferrite 3c11 is assumed for the yoke.
- Field quality is OK and it can still be improved.



- Energy can be measured using the *Polarisation of the Muon Beam* [Raja-Tollestrup – FERMILAB-Pub-97 / 402] IF some \mathcal{P} is saved after all the massage in the machines ...

- Spin precesses in a ring due to coupling with magnetic fields (bending magnets).

- At every turn spin precession is determined by the *SPIN TUNE*:

$$\omega = 2 \pi \gamma a$$

$$a = 1.16E-3$$

This determines a modulation in \mathcal{P}

- NB: if $\Delta E/E = 0 \rightarrow \gamma$ same for all muons $\rightarrow \mathcal{P}$ keeps oscillating
if $\Delta E/E \neq 0 \rightarrow \mathcal{P}$ goes to 0 after n turns

e^+ spectrum from μ -decay is a function of \mathcal{P} :

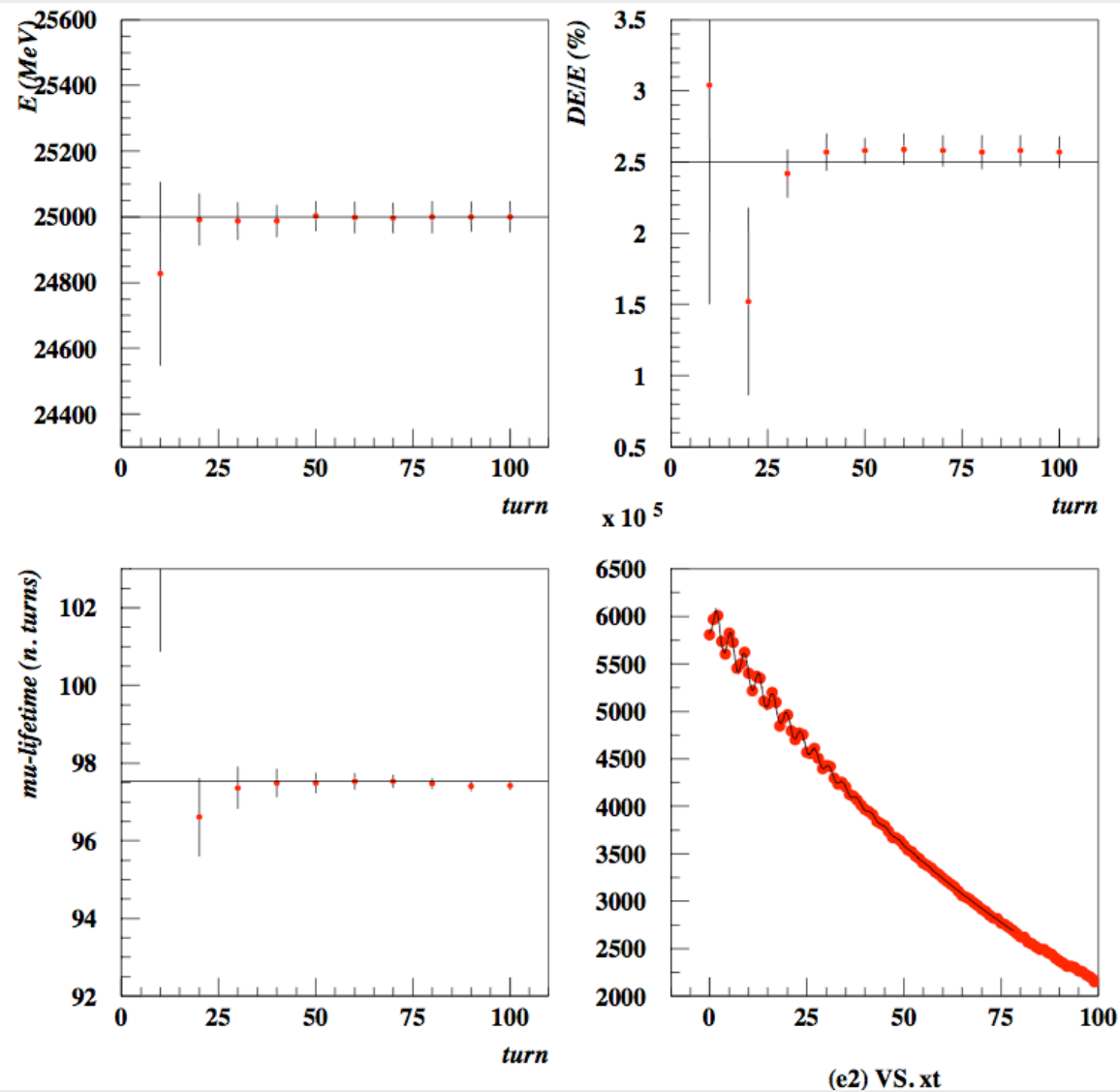
$$d^2N/dx d\cos\theta = N_0[(3-2x)x^2 - \mathcal{P}(1-x)x^2 \cos\theta] \quad (\text{CM})$$

- I have modelled the behaviour of a beam ($> 1E5$ muons) all with their spin and energy ($\Delta E/E = [0.01-0.05]$)

- Lorentz Boost

- Modulation in \mathcal{P} produces a modulation in $E(e^+)$

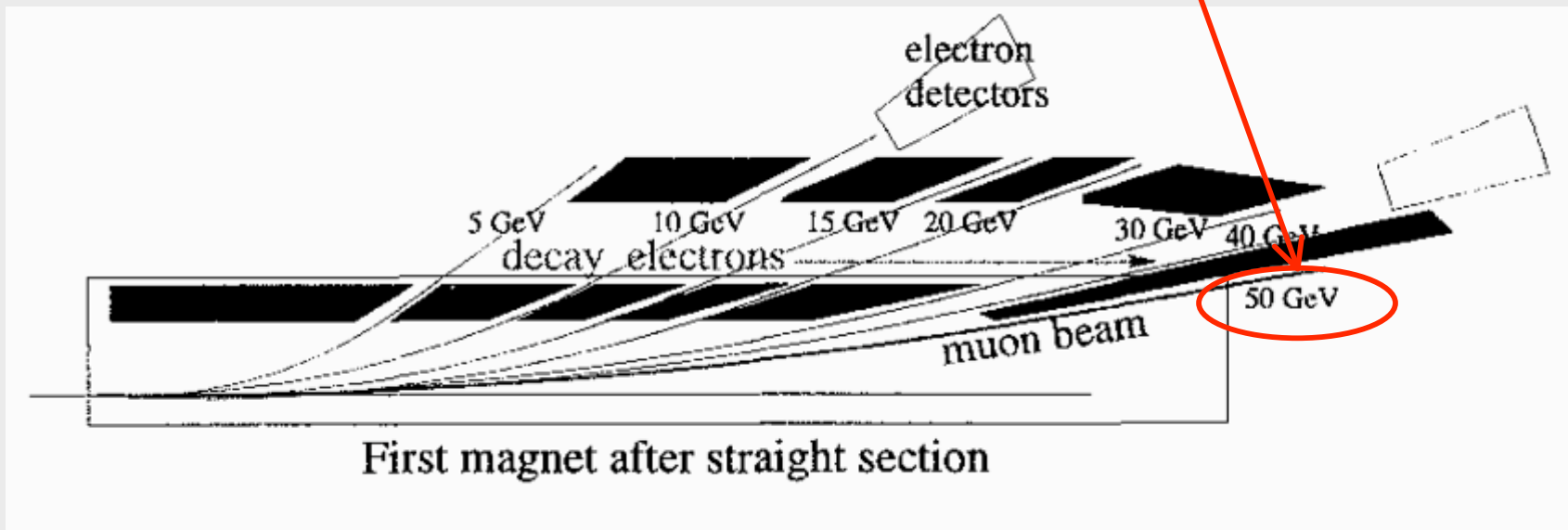
- I assume $\mathcal{P} = 27\%$ or 18% is left when filling the DK ring



M. Apollonio

It has been suggested [Blondel – ECFA 99-197(1999)] to use the *first bending magnet* after the decay straight section to SELECT electron energy bins: what does that mean today with a realistic lattice (25 GeV)?

In fact electron is emitted \sim parallel to μ (due to the high γ)



The spectral power of the 1st magnet depends on its *FIELD* and *LENGTH*

A *G4Beamline simulation* can tell us where electrons impinge after decaying somewhere along the orbit



- Define Reference Beam
- Redefine Solenoid Taper (for 1.25 T vs 1.5 T vs 1.75 T)
- Best Guess of achievable RF Voltage vs B @ 200 MHz
- Modified FS2A Front End
- Different Lithium Hydride mixtures, physics process model
- Tapering beta function
- Improved longitudinal and transverse matching
- Optimised buncher & phase rotation
- Buncher variations on a theme (shorter bunch train, etc)
- Buncher & rotator length vs gradient study
- Cooling performance degraded with different RF voltages
- Shielded cavities or Bucked coils
- High pressure RF
- Magnetically Insulated Lattice ($B \times E = 0$)
- Low Frequency Scheme
- Helical FoFo Snake
- Helical Cooling Channel



- n Description of baseline
- n ICOOL Optics Lattice
- n 2D magnet designs
- n Current densities, J vs B, tracking through field map
- n 2D RF cavity designs
- n Q-Factor for power supply costing
- n LiH absorber heat load, decision on active cooling
- n Description of one backup option for RF peak field mitigation?
- n Optics only
- n Costing
- n Hardware inventory
- n Cost scaling
- n Preliminary estimate shielding/collimation issues + requirements
- n Preliminary estimate civil engineering issues
- n How does the requirement for a CERN design study match available manpower? CERN must provide resources...

- Sub groups presented detailed plans for convergence.
- How do we organize the decision making process ?
- How do we handle alternatives and fall back options in the IDS ?
- IDR chapters should be prepared to be ready for submission to editors after RAL plenary (1st October)
- For IDS “mixed” costing according to readiness of the hardware design (global to detailed)